Compressibility Effects in Turbulence: Revisited

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STANFORD UNIVERSITY

AERONAUTICS AND ASTRONAUTICS, AND

MECHANICAL ENGINEERING

B. Olson, J. Larsson, I. Bermejo Moreno, J. Freund, P. Moin

Ack: AFOSR, ONR, NASA, SciDac

Happy 70th Sreeni!

Int. Symposium on *Turbulence from Angstroms to Light Years*, ICTS, Bangalore, Jan 21-25 2018

Outline

A Glimpse of compressibility effects on Turbulence

Fundamental Picture

More Recent Studies

Homogeneous Turbulence

Shock-Turbulence Interaction

Variable density/variable property effects

Summary and Open Issues

Physics and Modeling DNS/LES

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COMPRESSIBILITY EFFECTS ON TURBULENCE

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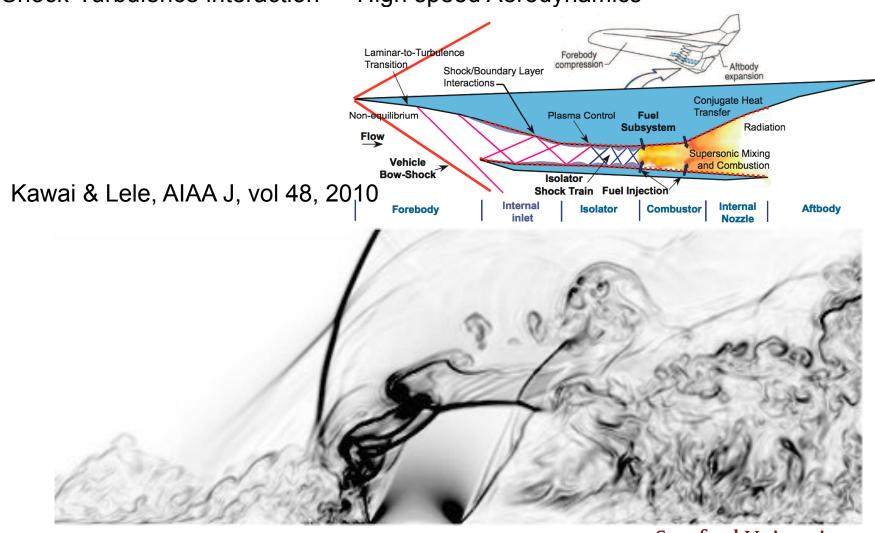
KEY WORDS: shock-turbulence interaction, supersonic mixing, turbulence modeling

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A Glimpse of compressibility effects on Turbulence

Supersonic mixing -- Scramjet
Supersonic/Hypersonic Turbulent Boundary layers - High speed Aerodynamics
Shock-Turbulence interaction -- High speed Aerodynamics

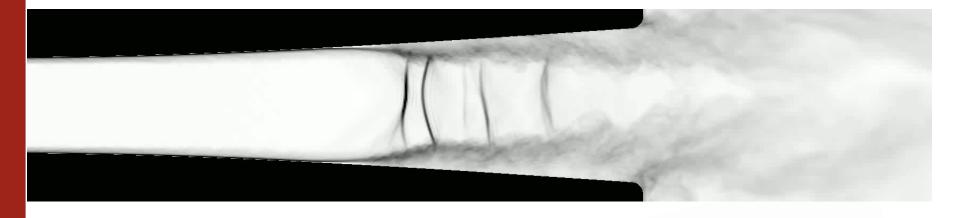


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Shock-induced Separation

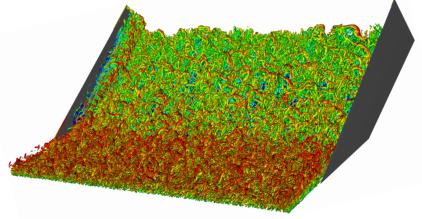
- Shock-induced separation
- Shock-induced unsteadiness

Over-expanded Planar Nozzle



3D Shock-Turbulent Boundary Layer Interaction

Compression ramp with fence

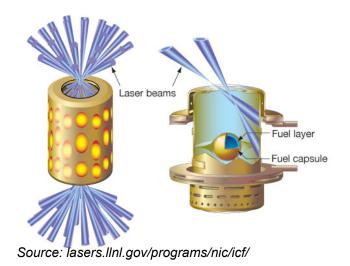


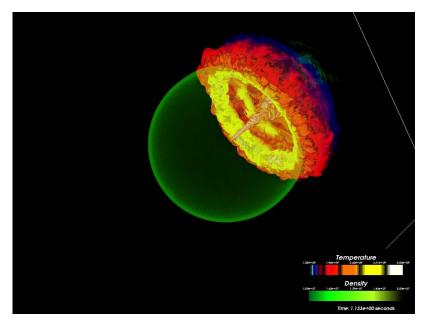
Dawson & Lele, AIAA 2015-1518

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A Glimpse of compressibility effects on Turbulence

Solar convection
Supernova collapse
Intersteller density fluctuations
Inertial confinement fusion





Source: flash.uchicago.edu

Supernova explosion

Inertial confinement fusion by shock-induced implosion of a deuterium/tritium capsule

Three Modes of fluctuations Yaglom (1948), Moyal (1952), Kovasznay (1953)

- ❖ Vortical mode or Gust Mode -- Solenoidal velocity, no pressure or density variations
- * Acoustic mode No vorticity, irrotational velocity, pressure, density fluctuation-isentropic
- **Entropic mode** No velocity or pressure, density and temperature fluctuations

Linearized modes of uniform state

Evolve independently at first order in amplitude

mode coupling: due to mean gradients and nonlinearity

Incompressible turbulence – Nonlinearly interacting vortical modes
Not just gust modes; (incompressible) Pressure effects (non-locality) important

Variable density turbulence (low M reacting flow, low M mixing) **Compressible turbulence** – density of fluid parcel changes in response to pressure

Low M_t -- Pseudo sound (Ristorcelli, 1997) Enslaved dilatational motions

Linearized modes of non-uniform state? Linear stability analysis; non-linear interactions

Primary Instability → Secondary Instability → Transition to turbulence

Useful framework but not well suited for broadband turbulence dynamics

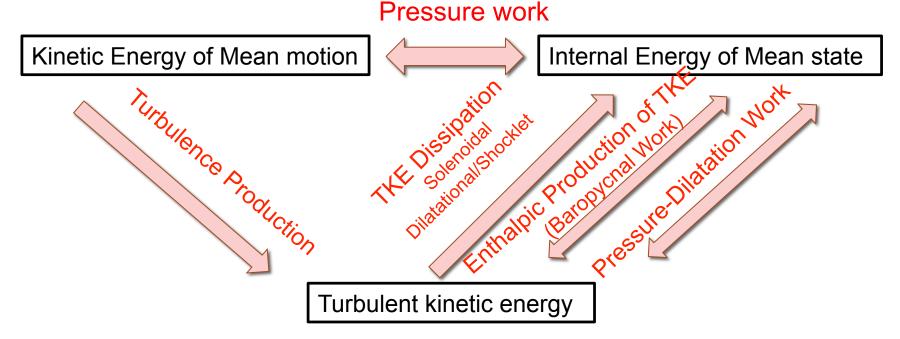
global modes/DMD/POD useful for modeling large-scale dynamics

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Decomposition of variables: **Favre decomposition** (mass weighted velocity) (define)

Energy Transfer pathways (coupling of momentum with thermodynamics)

Favre (1965), Huang et al (1995, JFM),



Compressibility Effects on Turbulence - Fundamentals **Dimensionless parameters** -- summary

Relative/Convective Mach Number $M_c M_r = (U_1 - U_2)/(C_1 + C_2)$

Turbulence Mach Number $M_t = q/C$

Gradient/Deformation Mach Number $M_g = SL/C = (L/C) / (1/S) = L / (C/S)$

Friction Mach Number $M_T = u_T / C$

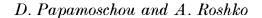
Mixing Layer $u_T / \Delta U \sim 0.2$; Boundary Layer $u_T / U_\infty \sim 1/25$ So for TBL require M_∞ which is at least 5 times larger than mixing layers for significant compressibility effects

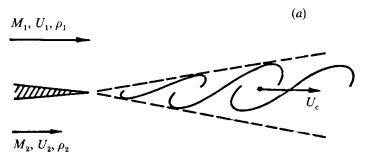
In practice, even at M_{∞} = 10 intrinsic compressibility effects seem absent! Why? Intense aerodynamic heating in supersonic/hypersonic TBL \rightarrow Very low density and high viscosity near the wall

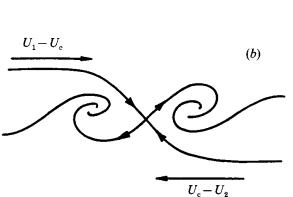
Flow Physics of Supersonic Mixing -- Observations I

Suppression of mechanical mixing

Brown & Roshko (1974), Bogdonoff (1983), Papamoschou & Roshko (1986)







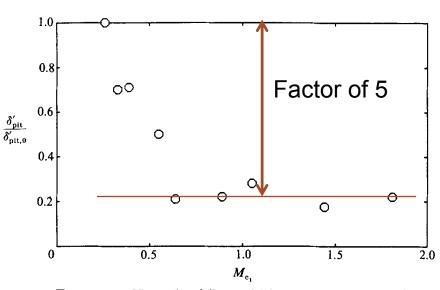


Figure 16. Normalized Pitot-thickness growth versus M_{e_i} .

$$\delta_0'(x;r,s) \simeq C_\delta \frac{(1-r)(1+s^{1/2})}{2(1+s^{1/2}r)} \left\{ 1 - \frac{(1-s^{1/2})/(1+s^{1/2})}{1+2.9(1+r)/(1-r)} \right\},\,$$

$$M_c = (U_1 - U_2)/(C_1 + C_2)$$

Compressibility Effects on Turbulence - Fundamentals Flow Physics of Supersonic Mixing -- Observations II

Explanations/Hypotheses

Dilatational/Shocklet dissipation Zeman (1990), Sarkar et al (1991)

Damping via Acoustic Radiation

Suppression of linear instability

Sandham & Reynolds (1990), Ragab & Wu (1989), Grosch & Jackson (1989), Lu & Lele (1993)

Turbulence structure change Vreman et al (1996), Pantano & Sarkar (2002)

Sonic Eddy Hypothesis Briedenthal (1992) AIAA J.

Flow Physics of Supersonic Mixing -- Observations III

Freund, Lele & Moin, JFM (2003)

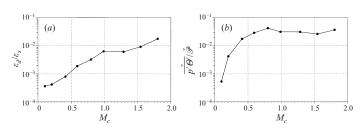


FIGURE 21. (a) Ratio of dilatational to solenoidal dissipation and (b) ratio of integrated pressure dilatation to the integrated production.

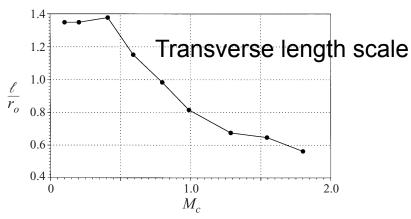
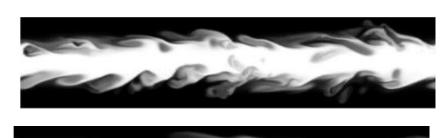
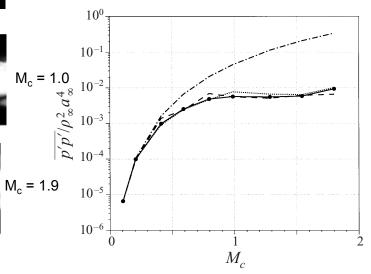




FIGURE 6. Transverse large-eddy lengthscale.







Passive tracer cuts

Absence of significant effects in Supersonic/Hypersonic TBLs Morkovin (1962), Bradshaw (1977), Smits & Dussauge (1986)

DNS Duan & Martin, JFM (2011)

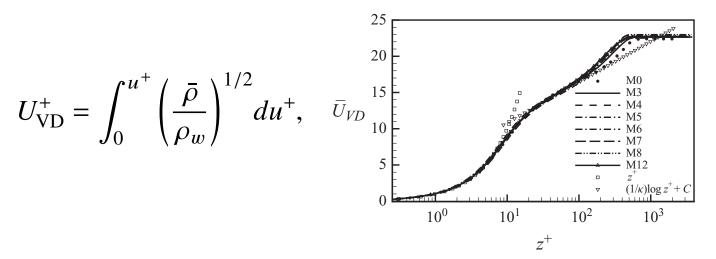
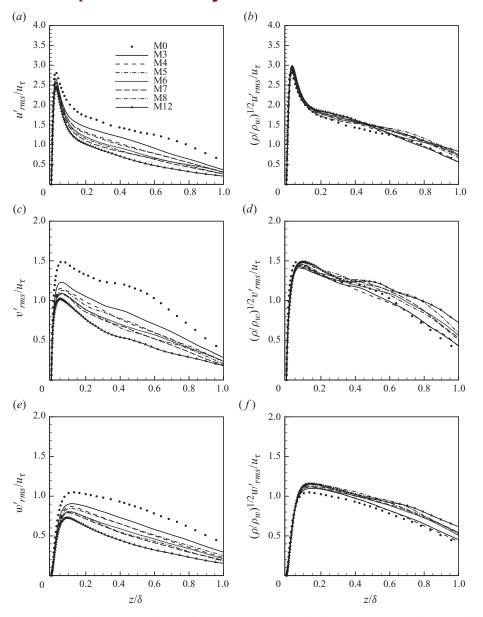


FIGURE 4. Van Driest transformed velocity for different free-stream Mach number cases.

Direct numerical simulation of hype	nic turbulent boundary layers 247
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Case	M_{δ}	$ \rho_{\delta} (\mathrm{kg} \mathrm{m}^{-3}) $	T_{δ} (K)	T_w/T_δ	$Re_{ heta}$	$Re_{ au}$	$Re_{\delta 2}$	θ (mm)	H	δ (mm)
M0	0.30	0.0900	220.0	1.00	1514.7.6	569.9	1515.2	2.76	1.41	23.0
M3	2.97	0.0910	219.9	2.51	3028.6	486.9	1586.7	0.619	5.09	8.85
M4	3.98	0.0902	219.2	3.83	4093.7	438.2	1587.4	0.658	8.01	12.0
M5	4.90	0.0962	224.4	5.31	4931.7	416.5	1578.0	0.682	11.29	15.1
M6	5.81	0.0990	230.7	7.02	5775.1	412.8	1582.2	0.730	15.4	19.7
M7	6.89	0.0929	224.2	9.49	7207.3	391.7	1586.4	0.838	20.2	28.1
M8	7.70	0.0990	232.8	11.2	7508.3	397.5	1577.5	0.861	24.2	31.8
M12	11.93	0.0906	228.0	27.6	11356.4	376.8	1577.5	1.33	46.1	84.7

TABLE 1. Dimensional boundary-layer edge and wall parameters for the DNS database.



Local density scaling

DNS Duan & Martin, JFM (2011)

FIGURE 6. Turbulence intensities and density-weighted turbulence intensities of the (a, b) streamwise, (c, d) spanwise and (e, f) wall-normal fluctuating velocity components for different free-stream Mach number cases.

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DNS/LES

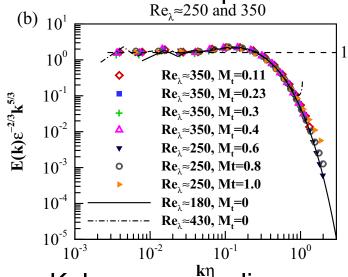
Compressibility Effects: Homogeneous Turbulence

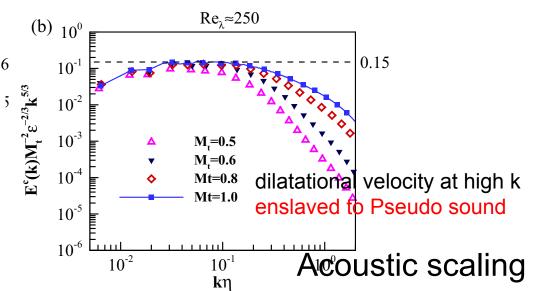
Spectral behavior

HIT Decaying and solenoidal forcing

Aluie (2011), Kritsuk et al (2013), Donzis & Jagannathan JFM (2013), Others Wang, Gotoh & Watanabe PRF (2017)



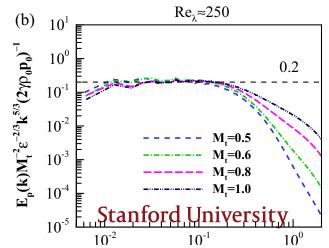




Kolmogorov scaling

TABLE III. Simulation parameters and resulting flow statistics for 1024³ grid resolution.

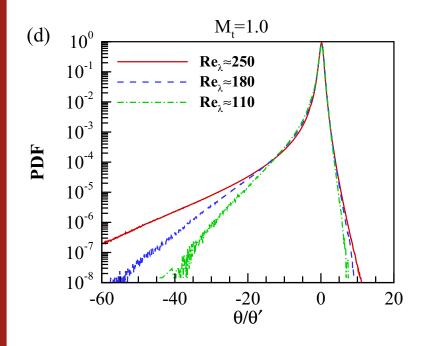
Resolutio	n Re _λ	M_t	$\eta/\Delta x$	L_I/η	λ/η	S_3	$D = \epsilon L_I / (u' / \sqrt{3})^3$
1024 ³	350	0.11	0.64	371	36.5	-0.55	0.41
1024^{3}	355	0.17	0.64	374	36.8	-0.55	0.43
1024^{3}	369	0.23	0.62	382	37.7	-0.56	0.41
1024 ³	361	0.30	0.58	395	37.4	-0.55	0.44
1024 ³	365	0.40	0.59	390	37.6	-0.54	0.43
1024 ³	253	0.51	1.00	233	31.2	-0.53	0.44
1024 ³	262	0.60	1.02	231	31.6	-0.53	0.42
1024^{3}	261	0.79	1.05	229	30.9	-0.83	0.41
1024 ³	250	1.02	1.04	226	29.1	-1.95	0.42

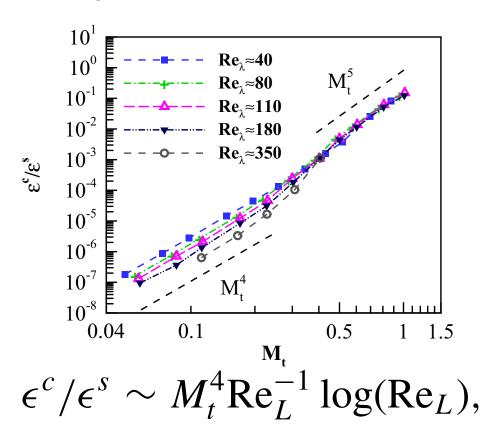


Compressibility Effects: Homogeneous Turbulence

Shocklets & Dissipation

Lee et al (1991), Samtany et al. (2001), Wang et al PRF (2017)

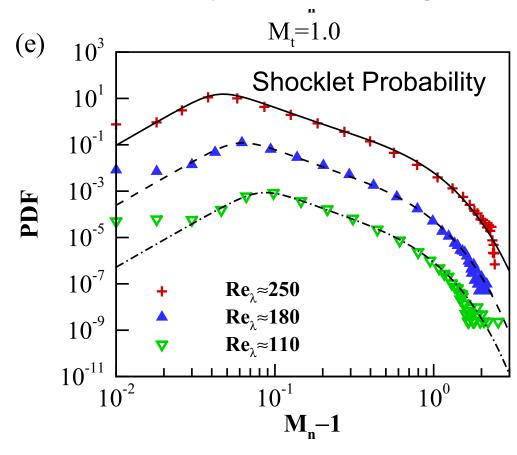




Compressibility Effects: Homogeneous Turbulence

Shocklets & Dissipation

Lee et al (1991), Samtany et al. (2001), Wang et al PRF (2017)



$$P(M_n - 1) = C_1 \exp\left[-\frac{M_n - 1}{\beta_2}\right] \left[1 + \left(\frac{M_n - 1}{M_n^{\exp} - 1}\right)^{-\beta_1}\right] \left[1 + \left(\frac{M_n - 1}{\beta_3}\right)^{-\beta_1 - 4}\right]^{-1},$$

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Canonical shock/turbulence interaction

Isotropic turbulence passing through a normal shock in a perfect gas Isolates the core interaction between turbulence and shock

Parameters:

 M_1 , M_t , Re_t

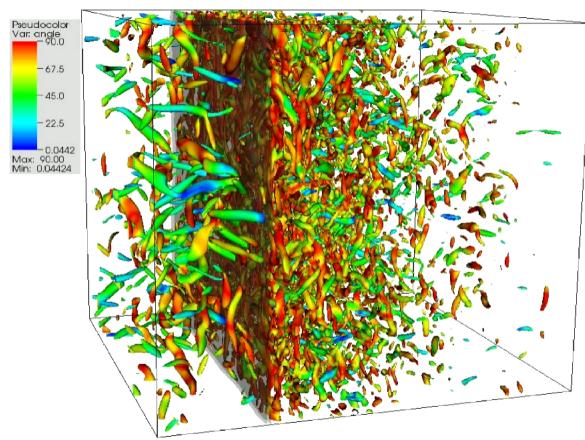
Entropy fluctuations

Acoustic fluctuations

 δ_{shock}/L

Real-gas effects

Eddies visualized by Q-criterion, colored by angle vorticity from x_1 . Shock visualized by dilatation contour.



From Larsson & Lele, Phys. Fluids, 2009

Larsson, Bermejo-Moreno & stellend FMy 2013

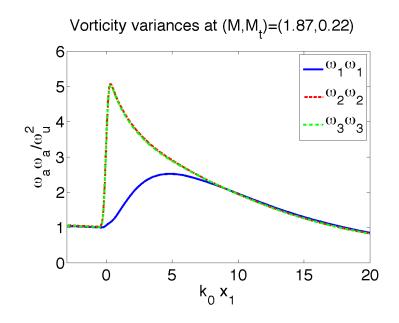
Difference compared to previous studies

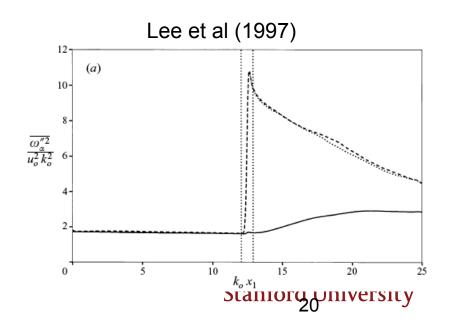
Past studies: $Re_{\lambda} = 20$ on $129*64^2$ grids (Lee et al, 1993, 1997, Mahesh et al, 1997)

Present: $Re_{\lambda} = 40$ on $1040*384^2$ grid; $Re_{\lambda} = 60$ on $1675*512^2$ grid; $Re_{\lambda} = 74$ on $2234*1024^2$ grids

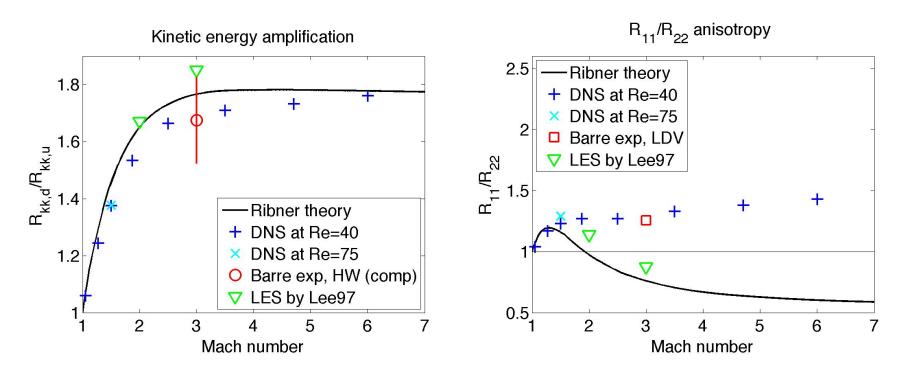
Return to local (small-scale) isotropy in present DNS, but not in past work Vorticity redistribution is a nonlinear process

Past work did not fully capture these processes due to lack of grid resolution





Reynolds stresses: amplification across shock



Linear theory predicts TKE amplification correctly, but misses the return to isotropy of small-scale turbulence and (post-shock) Reynolds stress anisotropy

Also see, Donzis (2012) Phys. Fluids and Ryu & Livescu (2014) JFM Also K. Sinha and collaborators

Shock motion and approach flow at $(M,M_t) = (1.50,0.38)$

Contours of streamwise velocity scaled by upstream speed sound (colors)

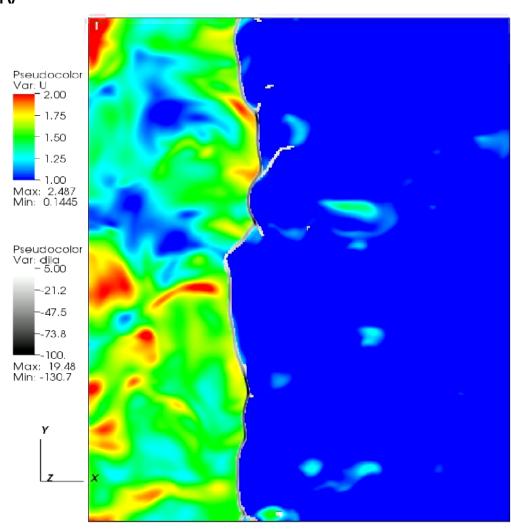
Contours of dilatation at shoc (gray scale)

u' > 0:

- "pushed back" shock
- strong shock (high compression)

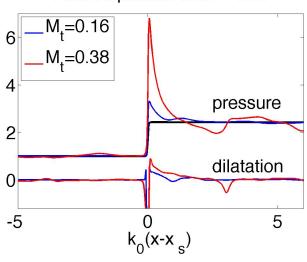
u' < 0:

- "pushed forward" shock
- weak shock compression)

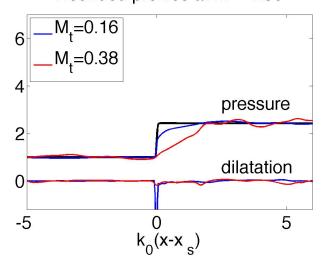


Instantaneous profiles through shock

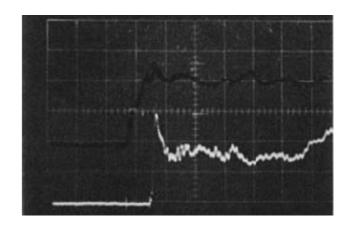
Peaked profiles at M = 1.50

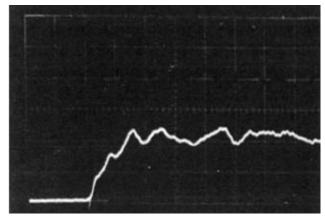


Rounded profiles at M = 1.50



Hesselink & Sturtevant (JFM 1988): "Peaked" and "rounded" pressure profiles for shock propagating through random medium at M <= 1.1





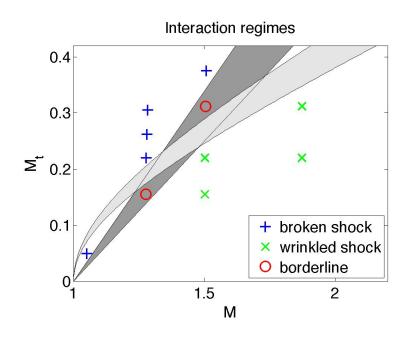
Source: Hesselink & Sturtevant, JFM 1988
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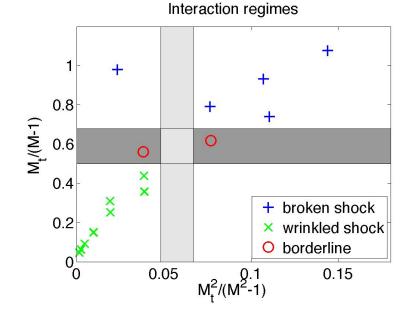
Interaction regimes

"Broken shock" regime -- existence of smooth (isentropic) profiles through shock "Wrinkled shock" regime -- every local interaction has a clear shock Under what conditions can "broken shock" regime be expected?

• Lee et al (1993): $\frac{p'_{
m turb}}{\Delta p_{
m shock}} \sim \frac{M_t^2}{M^2-1} \gtrsim {
m const}$

• Present: finite prob. of $u < c \Rightarrow \frac{M_t}{M-1} \gtrsim \text{const}$ Consistent with Donzis 2013





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Variable Property Effects: density and viscosity change

Semi-local scaling -- includes scaling of transport eqn (recall Laminar BL theory)
Coleman et al (1995), Patel et al (2016), Pecnik & Patel JFM (2017)

$$U_{\rm VD}^{+} = \int_{0}^{u^{+}} \left(\frac{\bar{\rho}}{\rho_{w}}\right)^{1/2} du^{+}, \qquad \qquad U^{+} = \int_{0}^{u^{+}} \left(\frac{\bar{\rho}}{\rho_{w}}\right)^{1/2} \left[1 + \frac{1}{2} \frac{1}{\bar{\rho}} \frac{d\bar{\rho}}{dy} y - \frac{1}{\bar{\mu}} \frac{d\bar{\mu}}{dy} y\right] du^{+},$$

Van Driest Scaling

Trettel & Larsson, PoF 2016

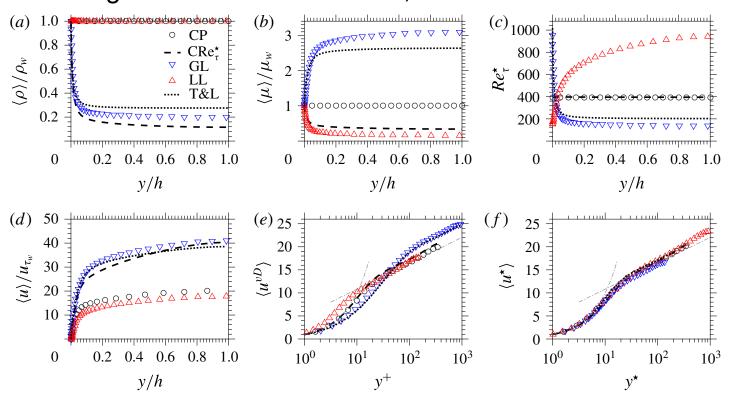


FIGURE 2. Averaged profiles for density (a), viscosity (b), semi-local Reynolds number (c), velocity (d), van Driest transformed velocity (e) and universal velocity scaling (f) for the DNS cases presented in table 1. Stanford University

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Physical understanding of compressibility effects has improved

Significant progress in some areas with DNS, LES (limitations in Re, Many non-dimensional parameters)

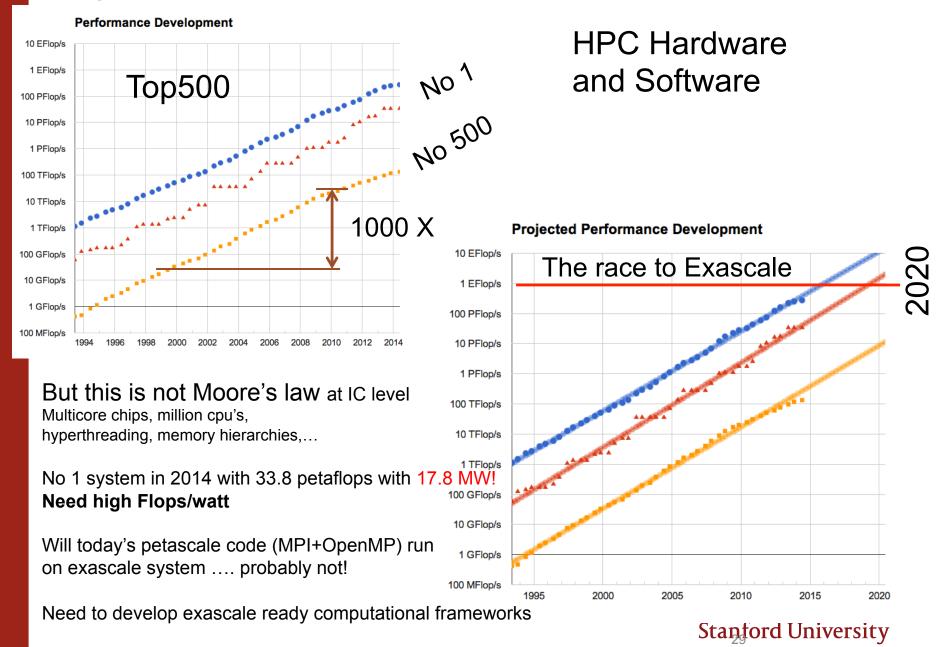
Theoretical framework still insufficient

How does acoustic communication impact vortex dynamics? Engineering models incorporating physics of variable-density and compressibility

Limited experiments with detailed turbulence measurements

Inflow conditions, boundary conditions (confinement ...)
Confirmation of phenomena and scaling laws
observed in idealized simulations

Open Issues: Shape of things to come ...?



Thank you!

Any questions?

The Challenge of High Reynolds Numbers

Spatial and temporal bandwidth increases with Re

 η/L scales as $Re^{-3/4}$

Computational cost (DNS) scales as $Re^{9/4}$ times $Re^{3/4}$ or Re^3

Computational cost Wall resolved LES scales as Re^{13/7}

Wall modeled LES reduces the cost to O(Re)

Practically the computational cost restricts the resolution which can be afforded and thus the scale bandwidth.

Need for modeling in turbulence simulations relevant to high Re continues

How to characterize?

- Effects on turbulence generation flow instability and turbulence energetics
- Effects on turbulence cascade
- Effects on turbulence dissipation
- System scale effects (multiple-scale interactions)
 - Radiative damping
 - Flow-acoustic resonance

Choice of variables?

Thermodynamic ρ ; $\log(\rho)$; $v=1/\rho$; p; $\log(p)$; s (Entropy); B (Bernoulli enthalpy)

Kinematic U_i ; ρU_i ; $\sqrt{\rho} u_i$; $\rho^{\wedge}(1/3) u_i$ (for spatial scale decomposition)

Scaling of spectra/structure functions?

Vortical turbulence -- Kolmogorov scaling

Entropy fluctuations stirred by turbulence

Random nonlinear acoustic field

Random field of shocks

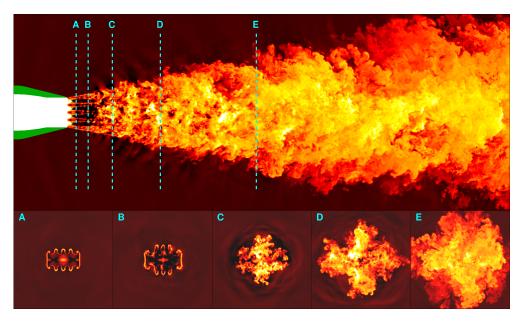
Pressure spectra -- hydrodynamic, pseudo-sound, sound

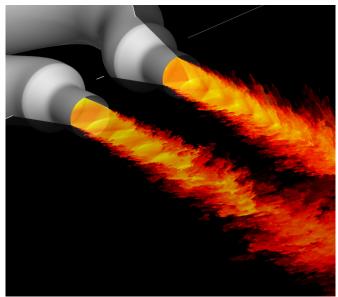
Density spectra - entropic, acoustic, pseudo-sound

Aeroacoustics

Computational methods for predicting aerodynamic noise Physics of aerodynamic noise generation Modeling noise sources and the far-field radiated noise Supersonic Jet Noise

Noise predictions using accurate LES of noise producing flow Approximate methods using source models

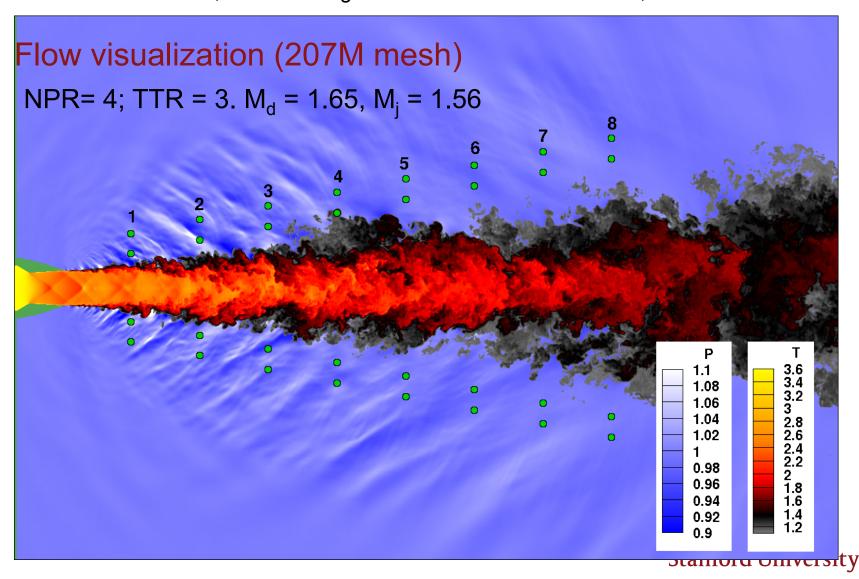




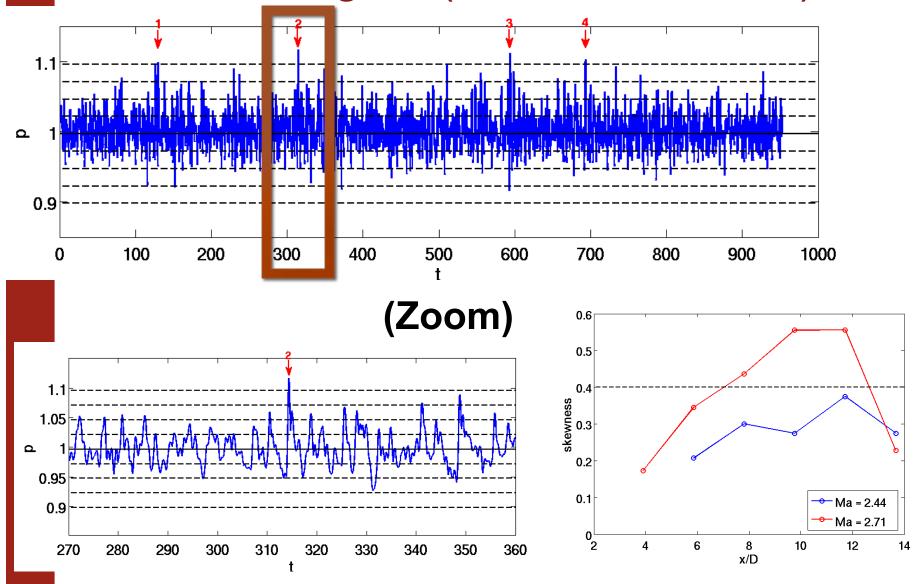
- •NASA Nozzle -- NA2C3 geometry (2:1 aspect ratio)
- CharLES solver
- Up to 500M isotropic adaptively refined mesh
- BlueGene/P with up to 160K cores (ALCC program, ANL)

Crackle: Annoying Component of Supersonic Jet Noise (Ffowcs Williams, 1975) Intermittent, Steep N-wave signature, Skewness

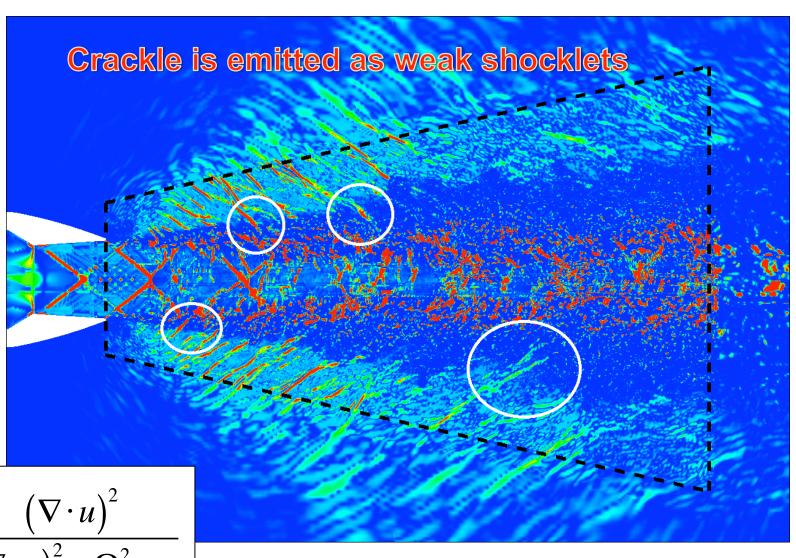
What causes crackle? Mechanism unknown –source nonlinearity vs non-lin. Propgn. **Nichols et. al. 2013**, ASME J. Eng. Gas Turbines and Power 213, Vol. 135.



Pressure signal (skewness 0.425)



Shock sensor (207M)



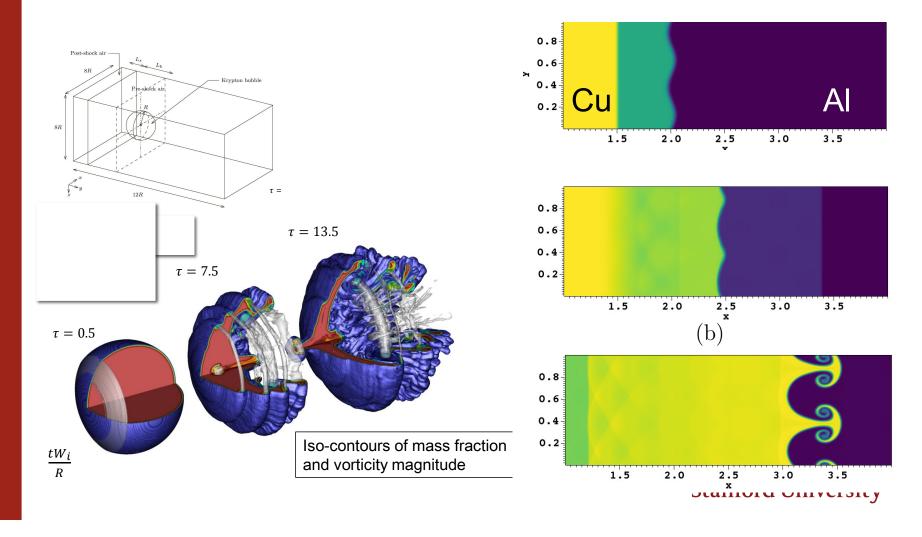
 $-\Omega^2 + \varepsilon$

(Ducros et al., 1999; Bhagatwala & Lele, 2009)

Main Research Areas: Fluid Dynamics - Turbulence

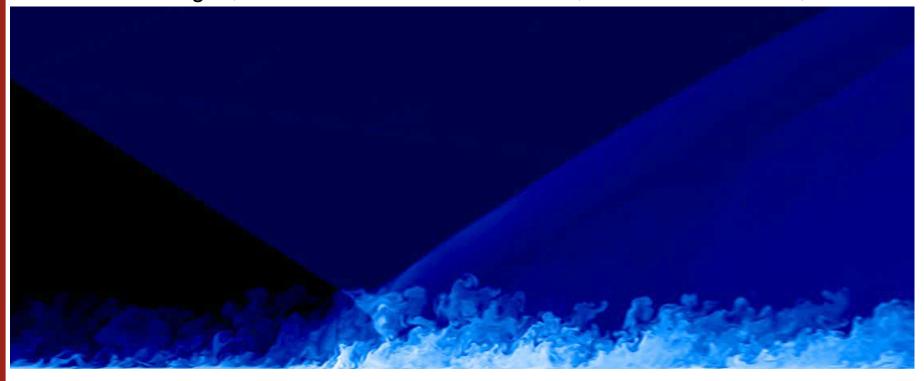
Shock-induced instability and multi-material mixing

Numerical methods for shocks/interfaces Non-linear elastic/plastic deformations Shock driven instability at solid-solid interface



Compressible turbulence, Shock-Boundary layer Interactions Capturing shock waves and turbulence

LES results Morgan, Kawai & Lele AIAA 2011-3431, J. Fluid Mech. 2013, Vol 729



M = 2.28; $Re_{\theta} = 2300$; wedge angle = 8 deg Expt: IUSTI, Marseille Temperature contours Stanford University