

Features of vigorous turbulence in experiments

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with

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1. How quickly does turbulence decay?

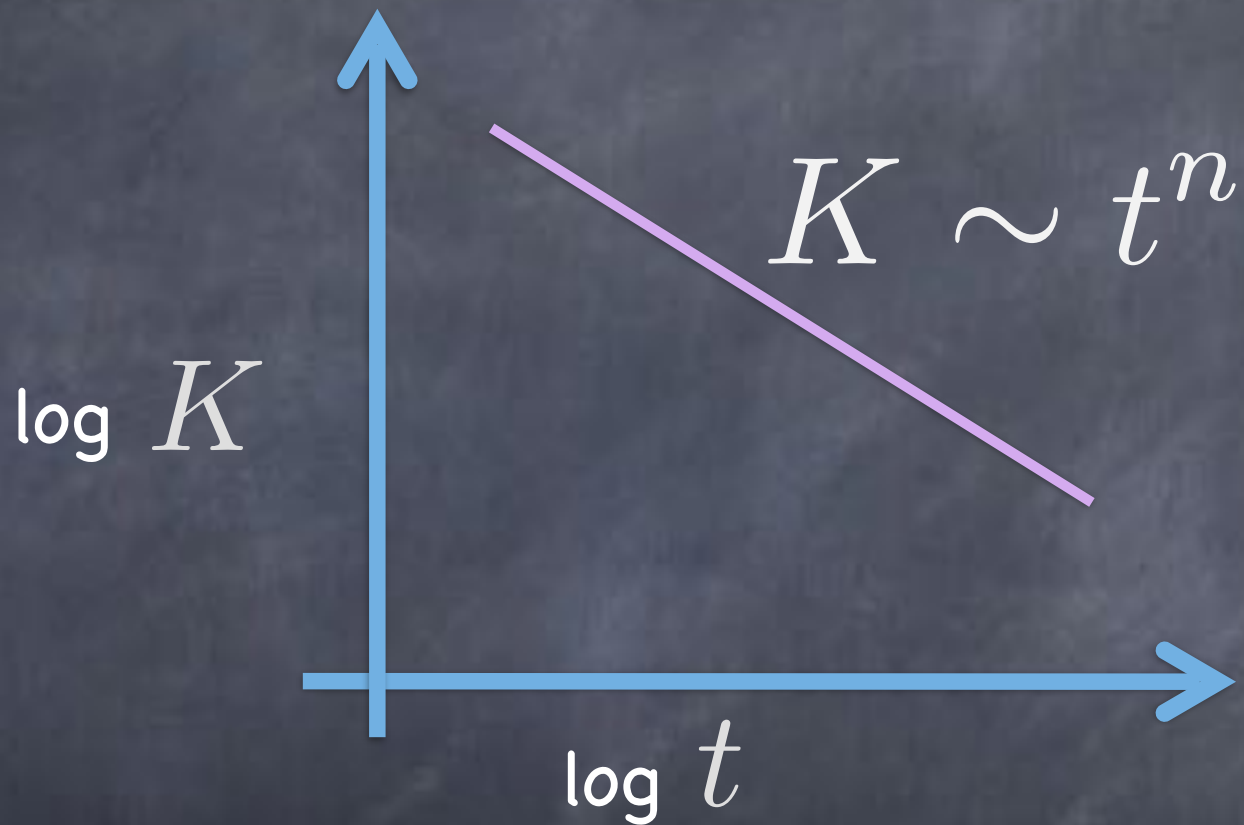


Altman and Demler (2007) *Nature*

$$\frac{dK}{dt} = -\epsilon$$

$$T \sim \frac{L}{U}$$

1. How quickly does turbulence decay?



Theoretical predictions:

— Self-similar $K \sim t^{-1}$

for $Re \rightarrow \infty$

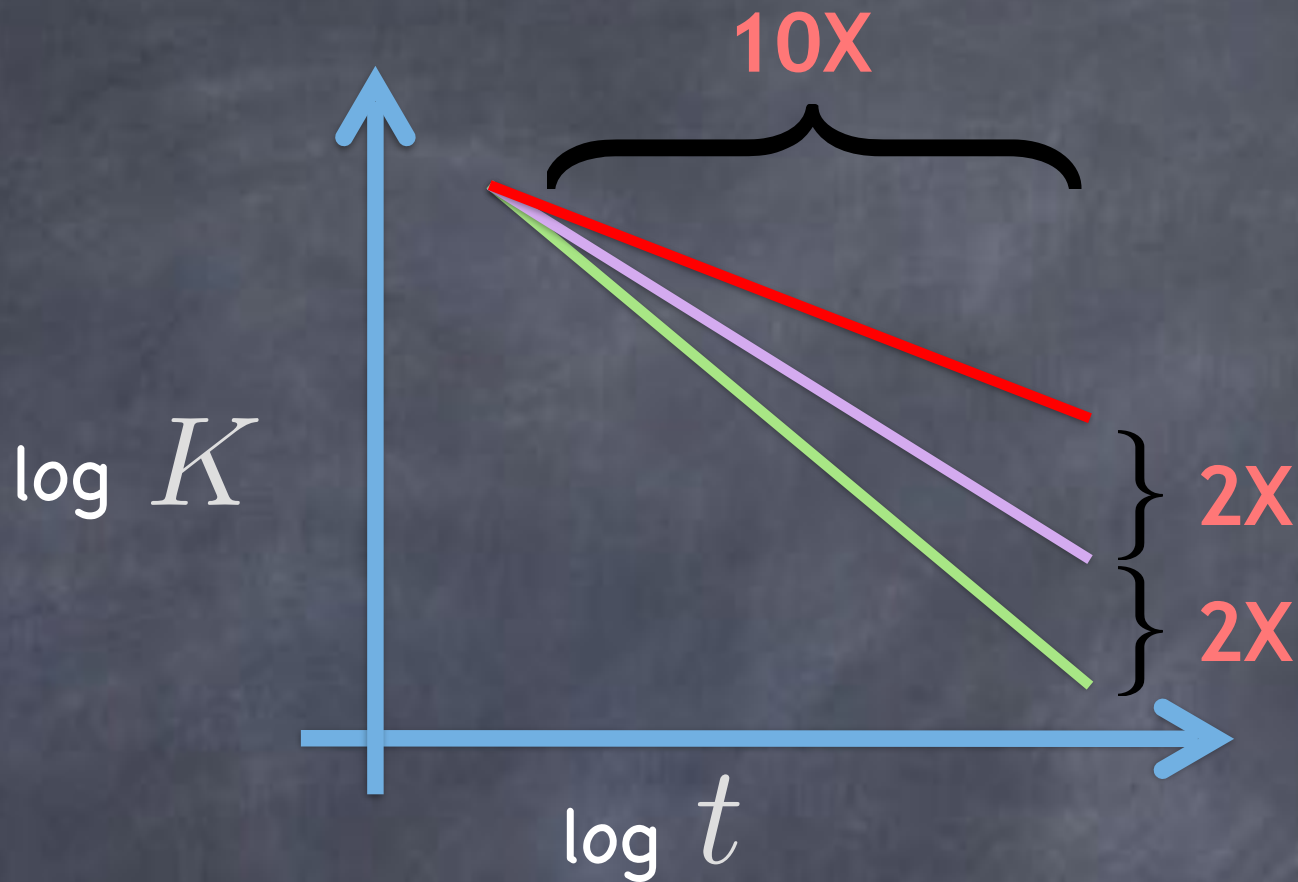
e.g. Dryden (1942) *Q. Appl. Math.*
Hinze (1975) *Turbulence*
George (1992) *Phys. Fluids*
Speziale and Bernard (1992) *JFM*

— Saffman $K \sim t^{-6/5}$

— Kolmogorov $K \sim t^{-10/7}$

depending on large-scale structure

e.g. Davidson (2011) *Phys. Fluids*

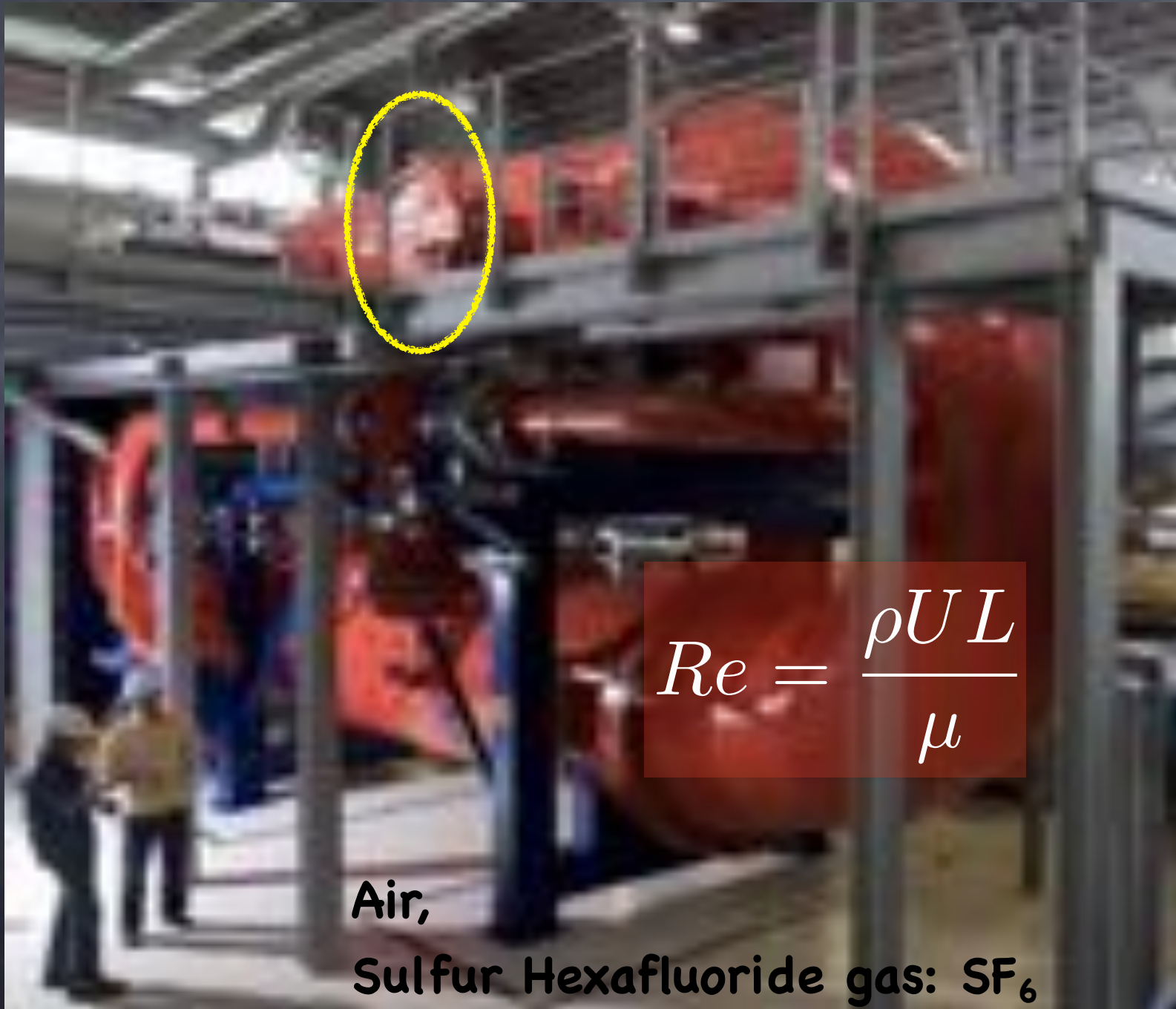


— Self-similar $K \sim t^{-1}$

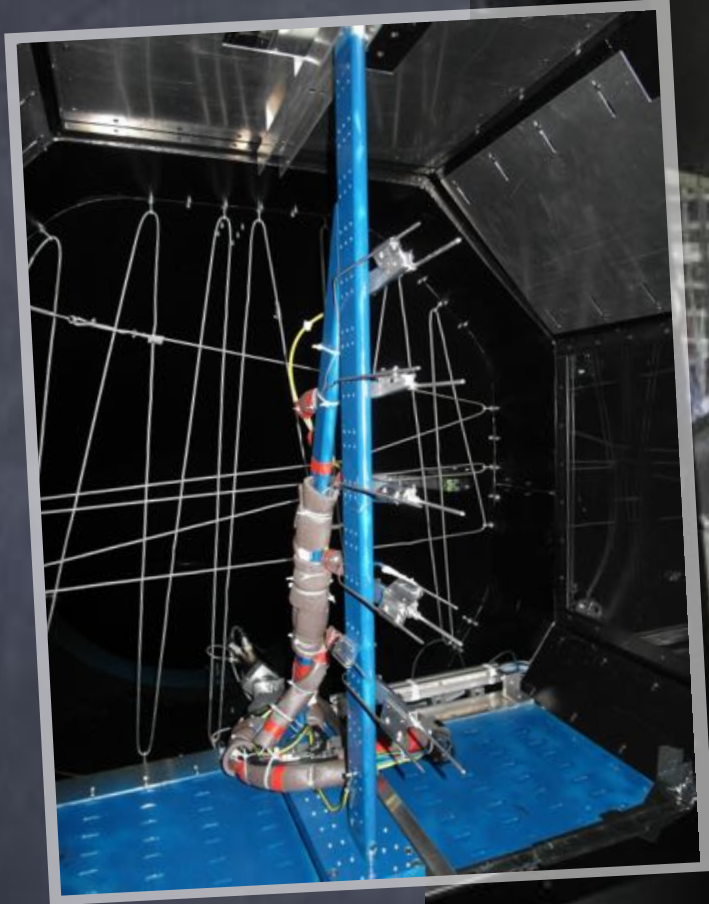
— Saffman $K \sim t^{-6/5}$

— Kolmogorov $K \sim t^{-10/7}$

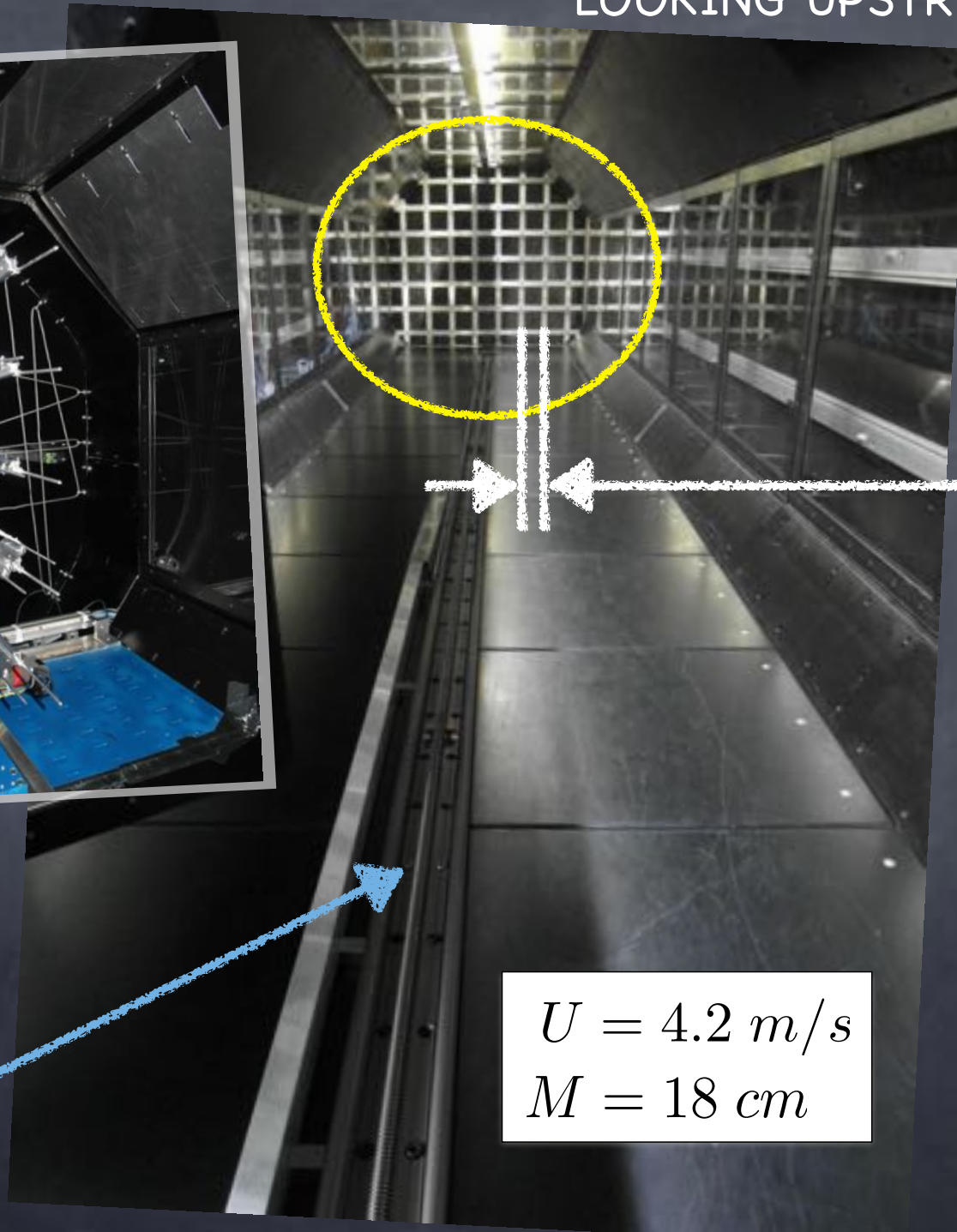
THE VARIABLE DENSITY TURBULENCE TUNNEL (VDTT)



LOOKING UPSTREAM



PROBE RAKE

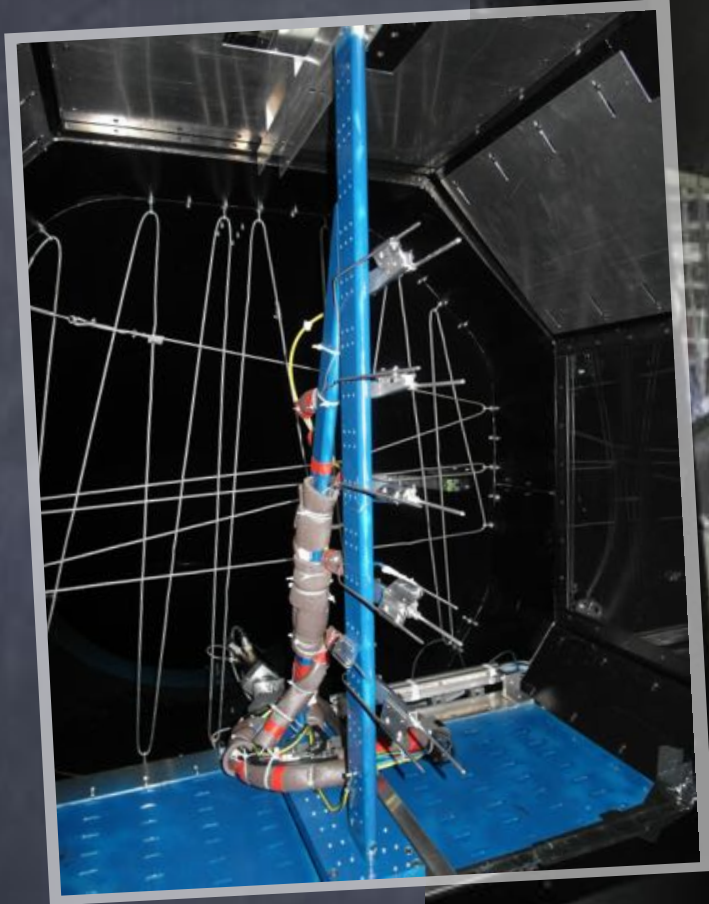


180 mm

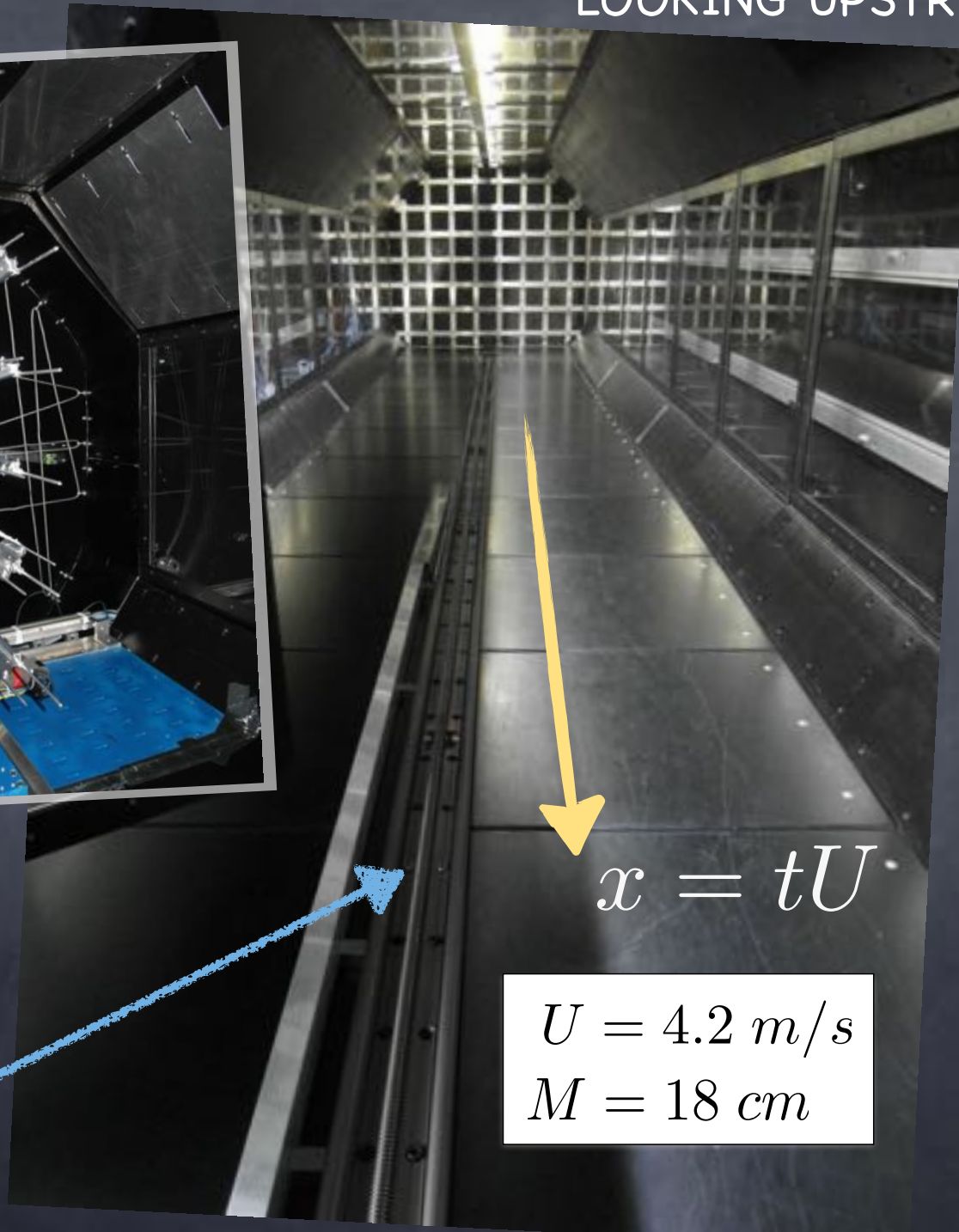
TRAVERSE

$$U = 4.2 \text{ m/s}$$
$$M = 18 \text{ cm}$$

LOOKING UPSTREAM



PROBE RAKE



$$x = tU$$

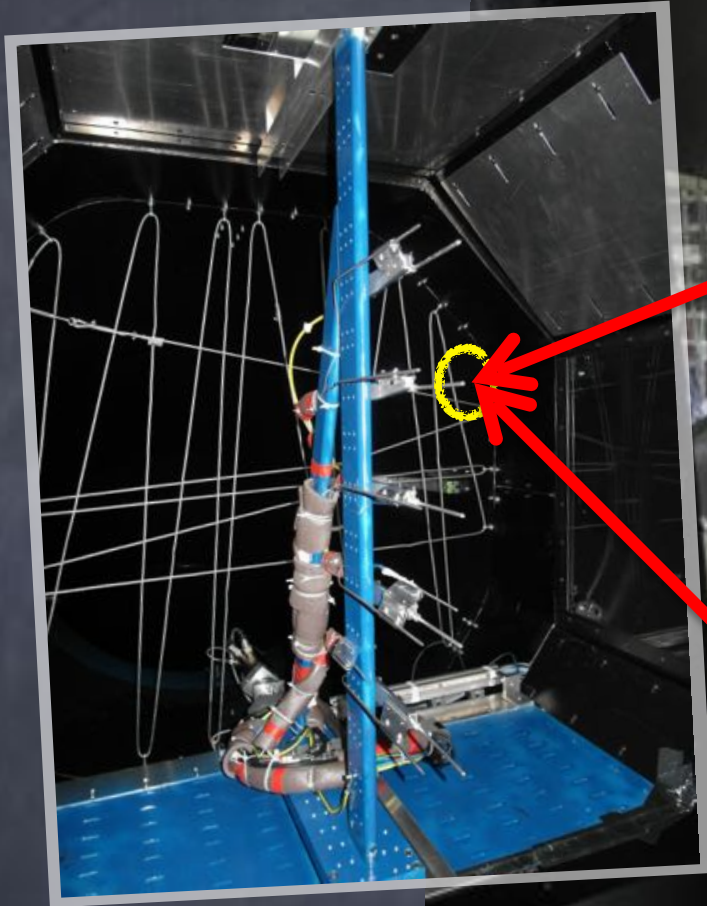
$$U = 4.2 \text{ m/s}$$

$$M = 18 \text{ cm}$$

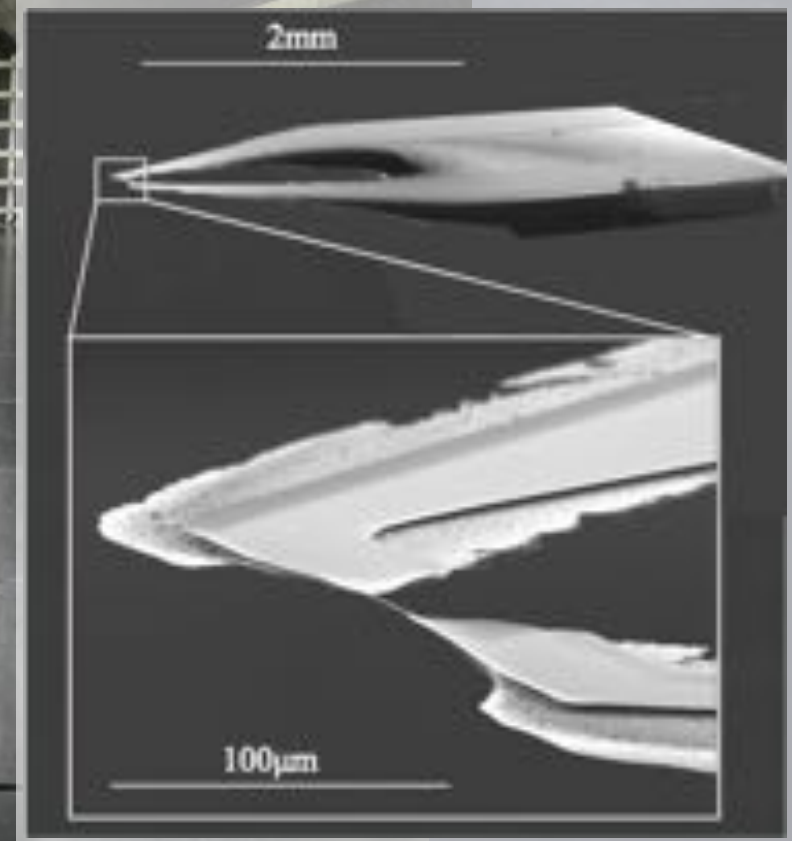
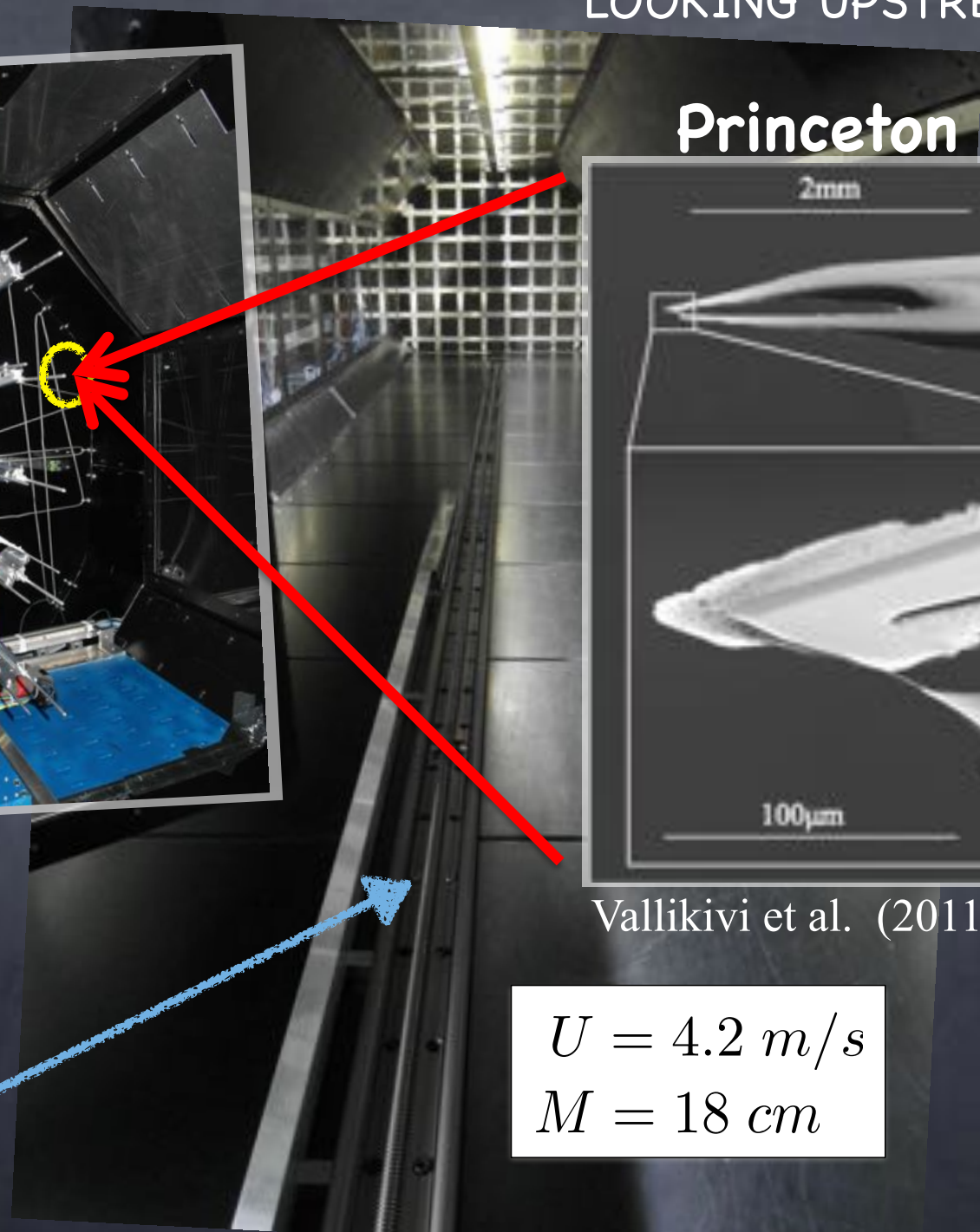
TRAVERSE

LOOKING UPSTREAM

Princeton NSTAP



PROBE RAKE

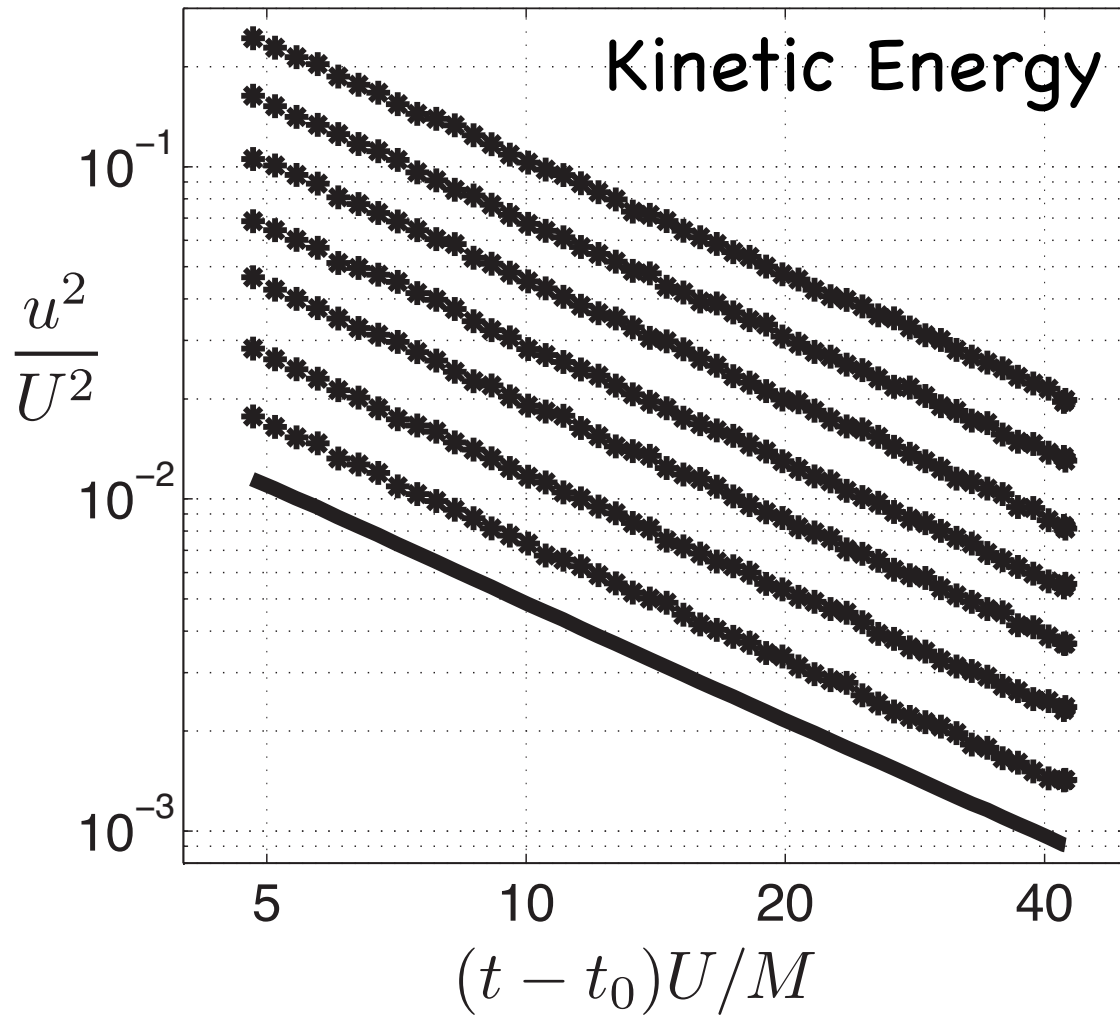


Vallikivi et al. (2011) *Expt. Fluids*

TRAVERSE

$$U = 4.2 \text{ m/s}$$
$$M = 18 \text{ cm}$$

Kinetic Energy

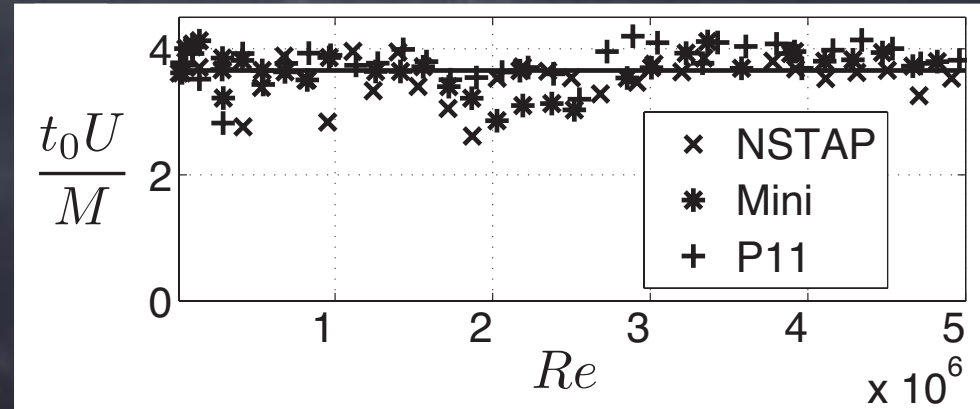


$$Re = \frac{UM}{\nu}$$

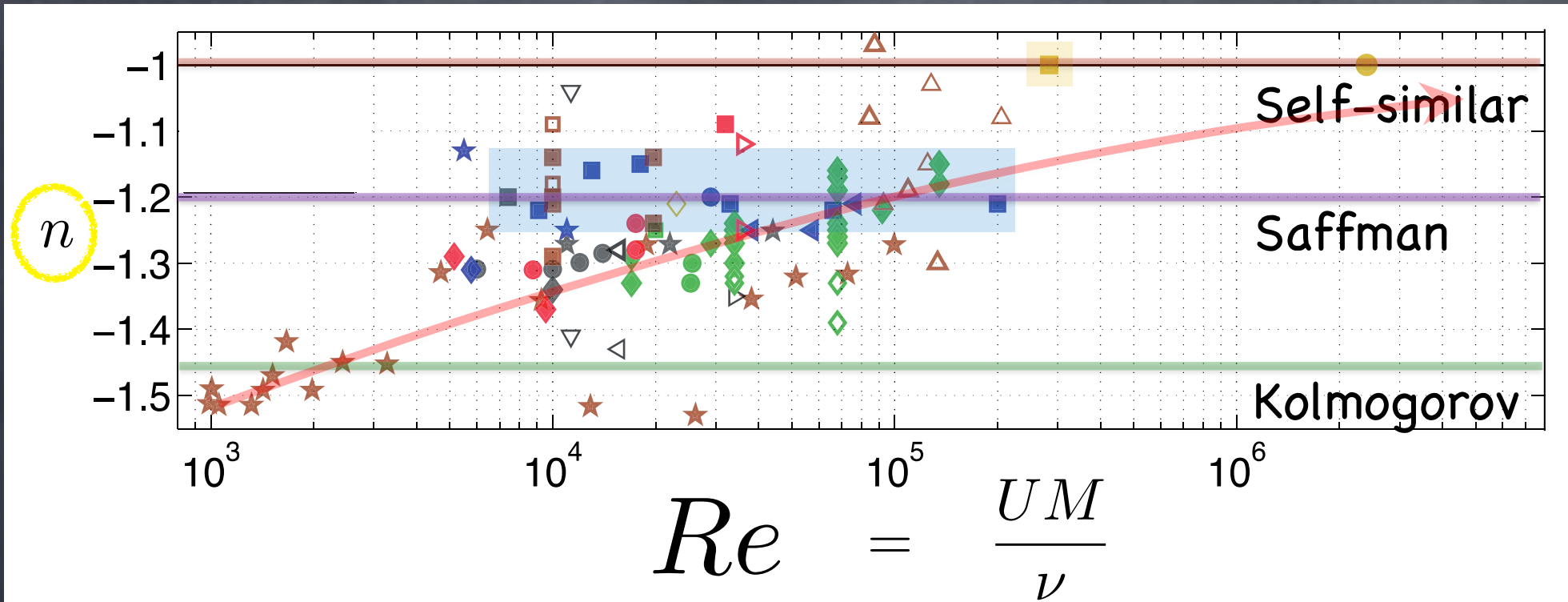
- 4.8×10^6
- 3.2×10^6
- 4.1×10^5
- 1.4×10^5
- 5.4×10^4
- 2.6×10^4

Sinhuber, Bodenschatz and Bewley (2015) *PRL*

$$\frac{u^2}{U^2} = C \left((t - t_0) \frac{U}{M} \right)^n$$



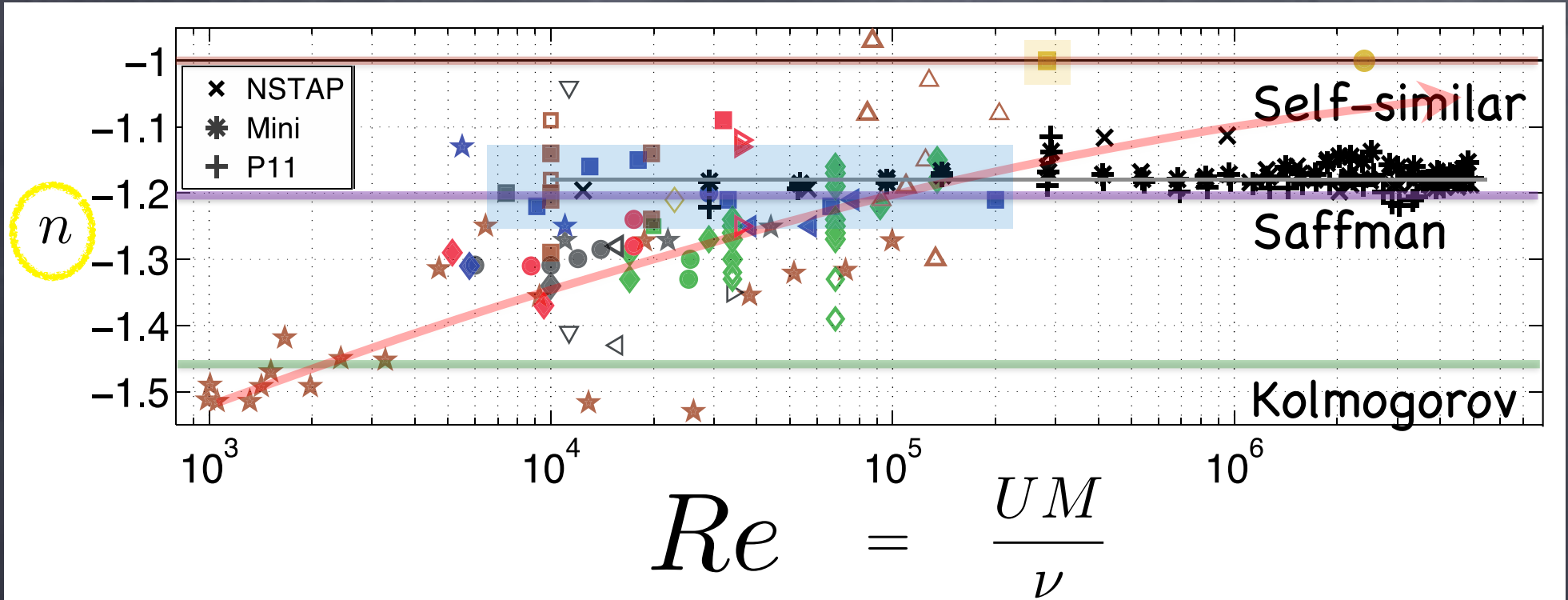
$$K \sim t^n$$



Sinhuber, Bodenschatz and Bewley (2015) *PRL*

- ★ Kurian and Fransson (2009), ★ Batchelor and Townsend (1948), ★ Wyatt (1955), ◆ Sirivat and Warhaft (1983), ◆ Yoon and Warhaft (1990), ◆ Warhaft and Lumley (1978), ◆ Comte-Bellot and Corrsin (1966), ◆ Mydlarski and Warhaft (1996), ■ Sreenivasan et al. (1980), ■ White et al. (2002), ■ Lavoie et al. (2007), ■ Antonia et al. (2003), ■ van Doorn et al. (1999), ■ Bewley et al. (2007), ● Mohamed and LaRue (1990), ● Uberoi and Wallis (1966), ● Van Atta and Chen (1968), ● Uberoi (1963), ● Kistler and Vrebalovich (1966), ▼ Poorte et al. (2002), ◀ Makita (1991), ▶ Krogstad and Davidson (2011), ◀ Valente and Vassilicos (2011), ▲ Thormann and Meneveau (2014),

$$K \sim t^n$$



Sinhuber, Bodenschatz and Bewley (2015) *PRL*

★ Kurian and Fransson (2009), ★ Batchelor and Townsend (1948), ★ Wyatt (1955), ◆ Sirivat and Warhaft (1983),
 ◆ Yoon and Warhaft (1990), ◆ Warhaft and Lumley (1978), ◆ Comte-Bellot and Corrsin (1966), ◆ Mydlarski and
 Warhaft (1996), ■ Sreenivasan et al. (1980), ■ White et al. (2002), ■ Lavoie et al. (2007), ■ Antonia et al. (2003), ■
 van Doorn et al. (1999), ■ Bewley et al. (2007), ● Mohamed and LaRue (1990), ● Uberoi and Wallis (1966), ● Van
 Atta and Chen (1968), ● Uberoi (1963), ● Kistler and Vrebalovich (1966), ▼ Poorte et al. (2002), ◀ Makita (1991),
 ▶ Krogstad and Davidson (2011), ◀ Valente and Vassilicos (2011), ▲ Thormann and Meneveau (2014),

2. What is the structure of turbulence?

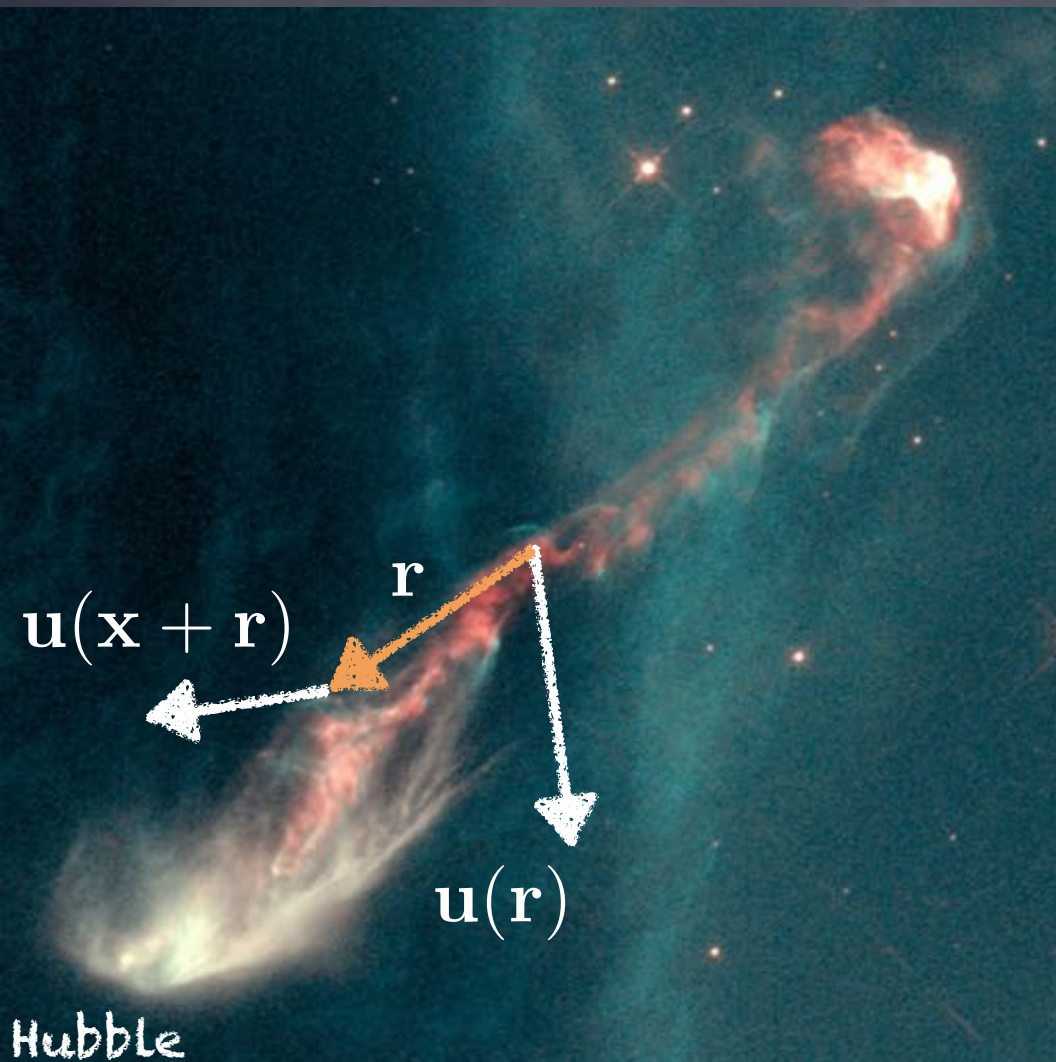
Velocity differences

$$\delta u(r) \equiv (\mathbf{u}(\mathbf{x} + \mathbf{r}) - \mathbf{u}(\mathbf{x})) \cdot \hat{\mathbf{r}}$$

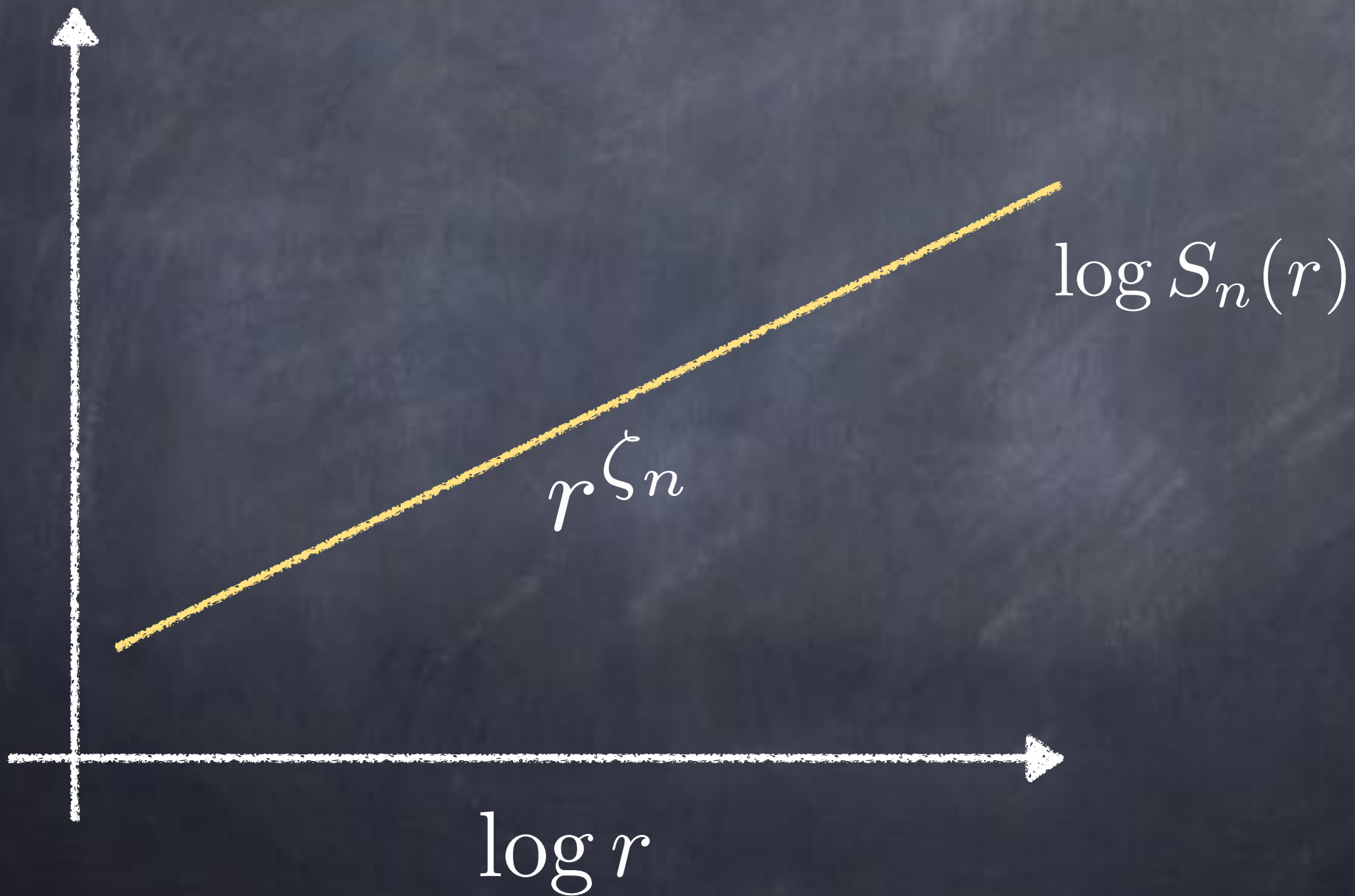
$$S_n(r) \equiv \langle \delta u^n(r) \rangle$$

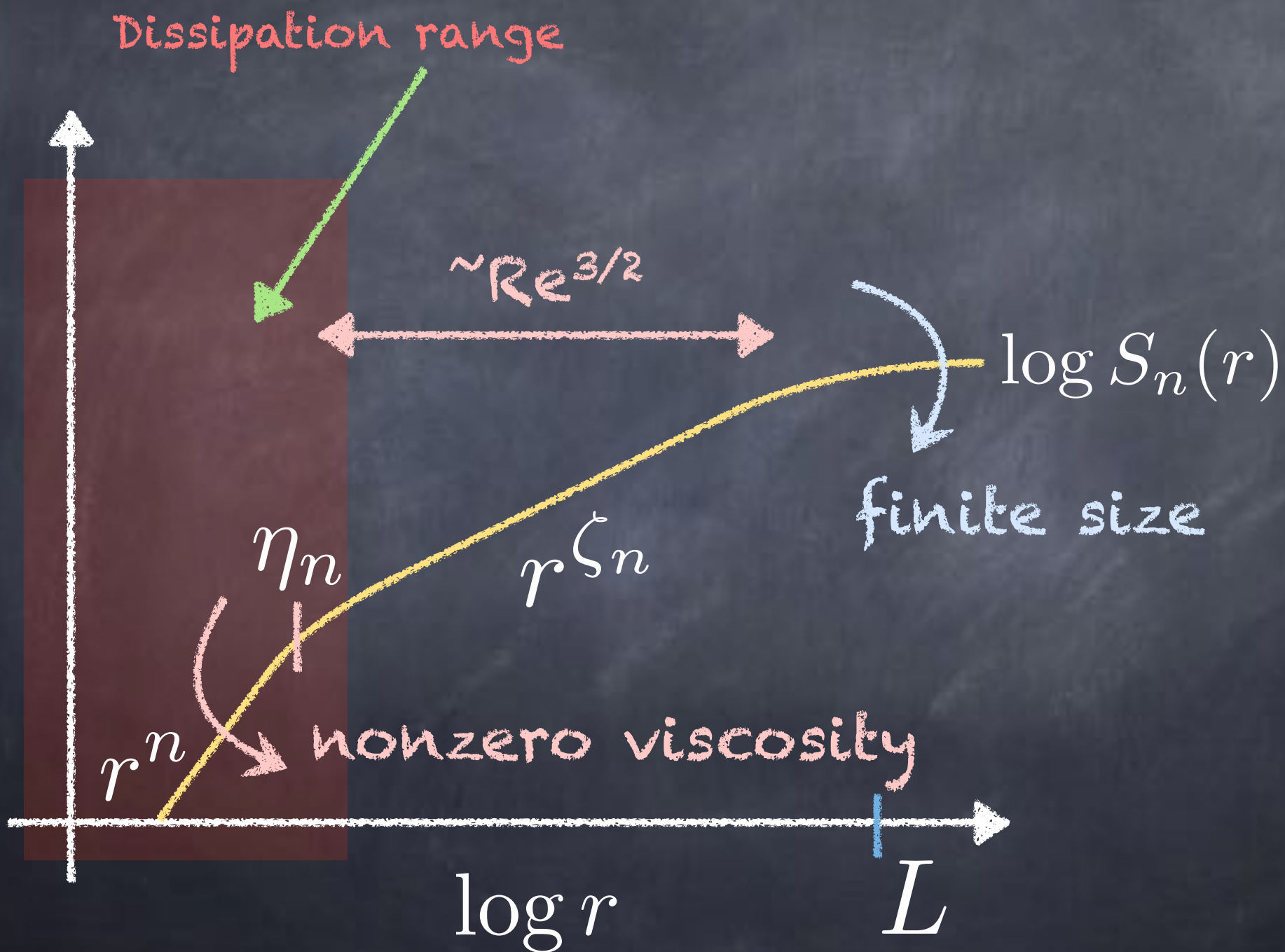
Structure functions

$$S_n \sim r^{\zeta_n}$$



Kolmogorov (1941), *etc.*



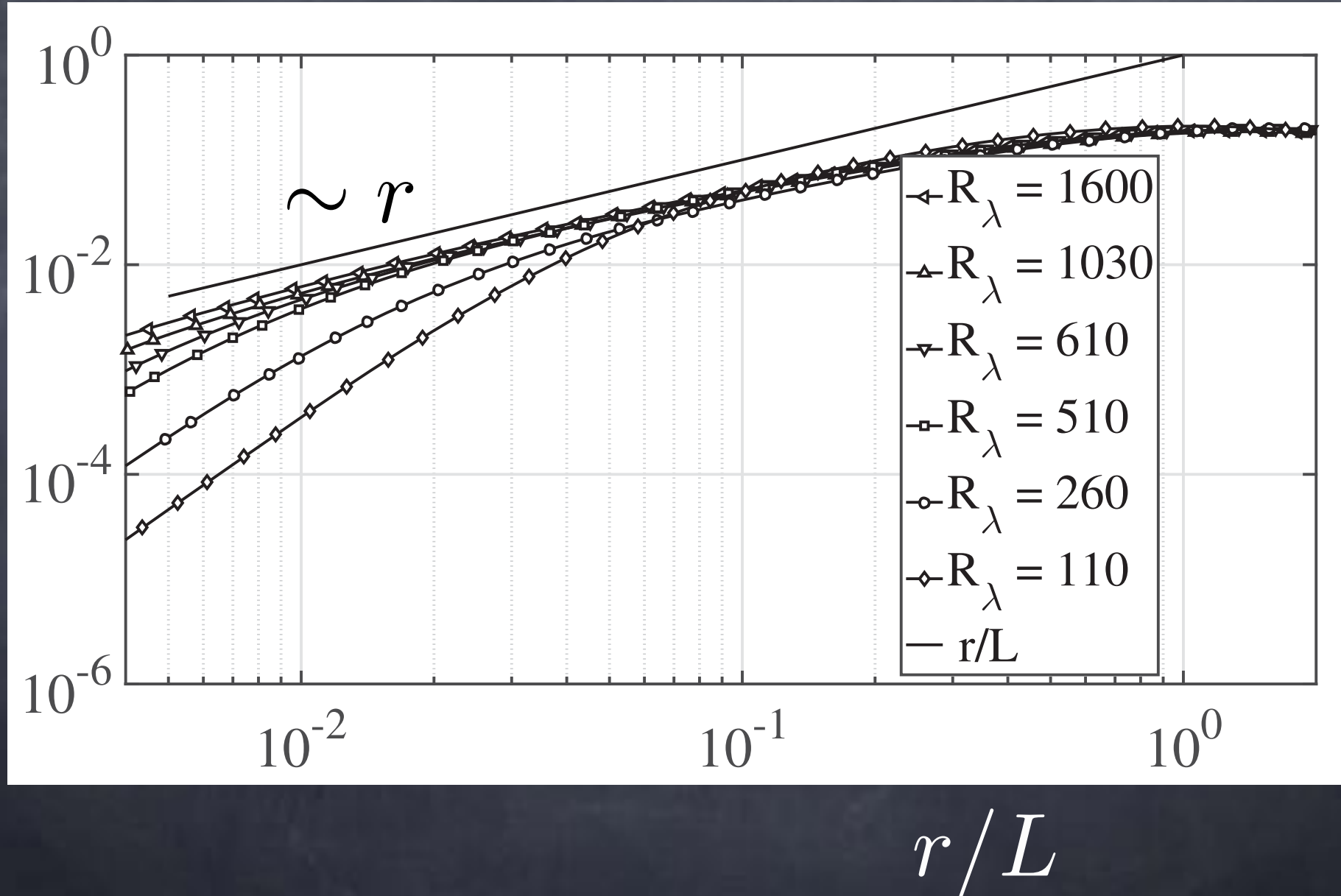


Our data

The 4/5ths Law

Kolmogorov (1941)

S_3



Our data

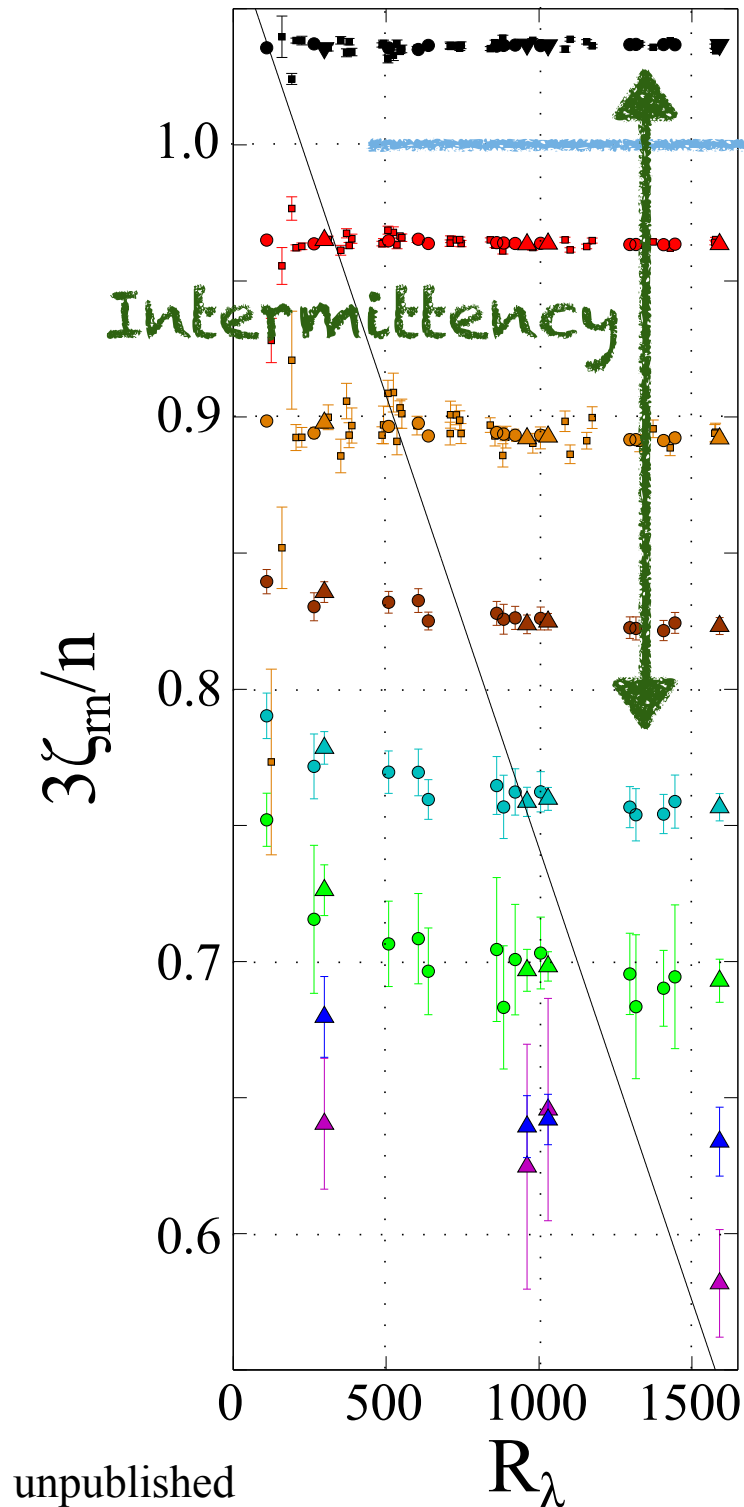
$$S_n \sim r^{\zeta_n}$$

ESS assuming

$$\zeta_3 = 1$$

for

$$\langle |\delta u^3(r)| \rangle$$



2

K41

4

6

8

10

12

14

16

n

order of moment	DNS exponents	present exponents
-0.80	-0.317	-0.313
-0.20	-0.077	-0.078
0.20	0.036	0.039
0.20	0.073	0.076
0.30	0.112	0.113
0.40	0.150	0.150
0.50	0.187	0.190
0.60	0.223	0.221
0.70	0.260	0.265
0.80	0.296	0.292
0.90	0.332	0.333
1	0.366	0.372
1.25	0.452	0.458
1.50	0.536	0.542
1.75	0.619	0.628
2	0.699	0.708
3	1	1
4	1.279	1.26
5	1.536	1.56
6	1.772	1.71
7	1.989	1.97
8	2.188	2.05
9	2.320	2.20
10	2.451	2.38

Previous results

$$S_n \sim r^{\zeta_n}$$

n ζ_n ζ_n

Sreenivasan and Dhruva (1998)

Prog. Theor. Phys. Suppl.

mass of an electron:

$$m_e = 0.000\ 548\ 579\ 909\ 067(14)(9)(2)$$

Sturm, Köhler *et al.* (2014) *Nature*

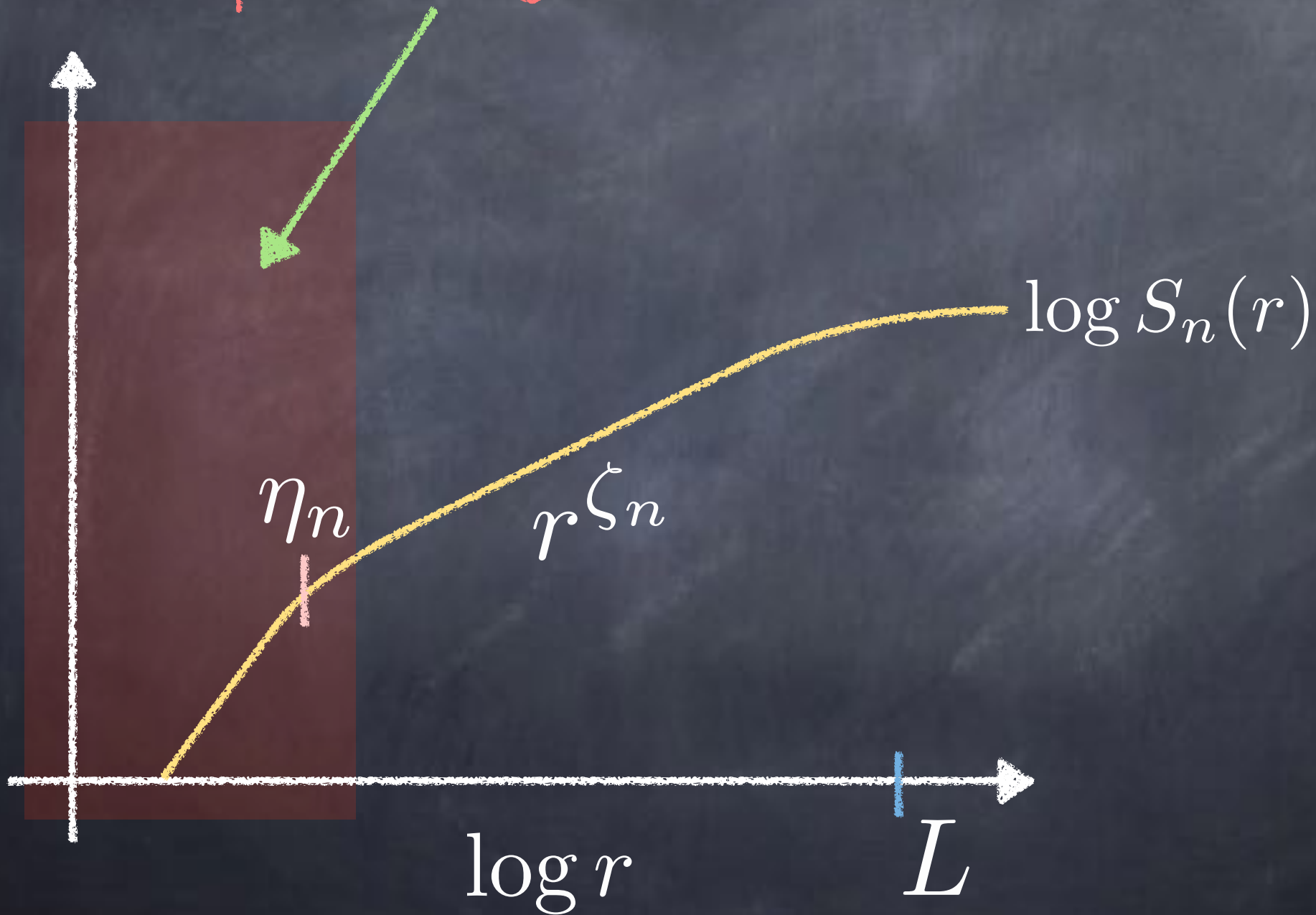
critical exponent for superfluid
phase transition in helium:

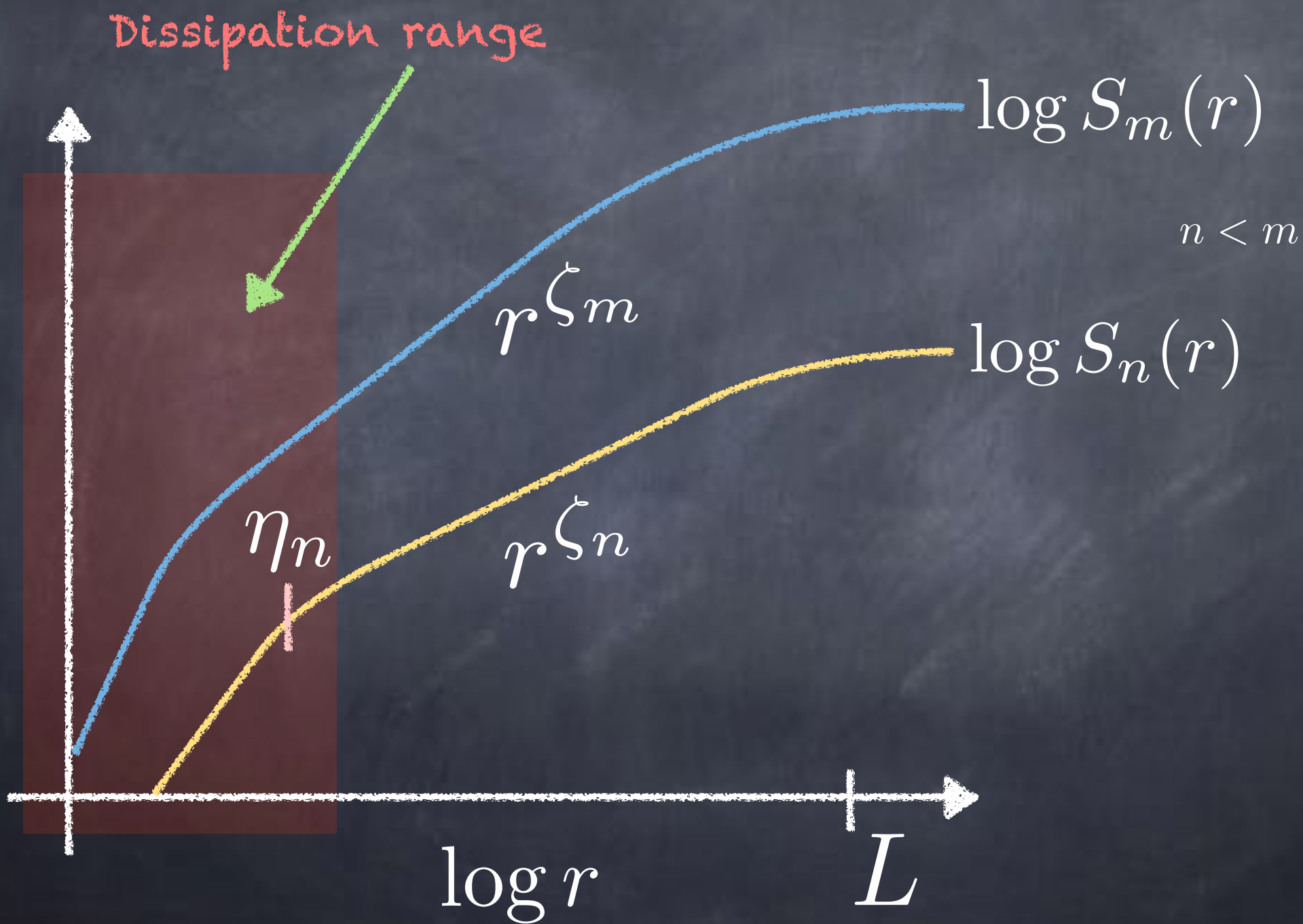
$$\alpha = -0.0127(3)$$

Lipa *et al.* (2003) *PRB*

How accurately can we determine the exponents
of turbulence?

Dissipation range





ESS:

R. Benzi, S. Ciliberto, R. Tripiccone, C. Baudet, F. Massaioli, and S. Succi, *Phys. Rev. E* **48**, R29 (1993).

C. Meneveau, *Phys. Rev. E* **54**, 3657 (1996).

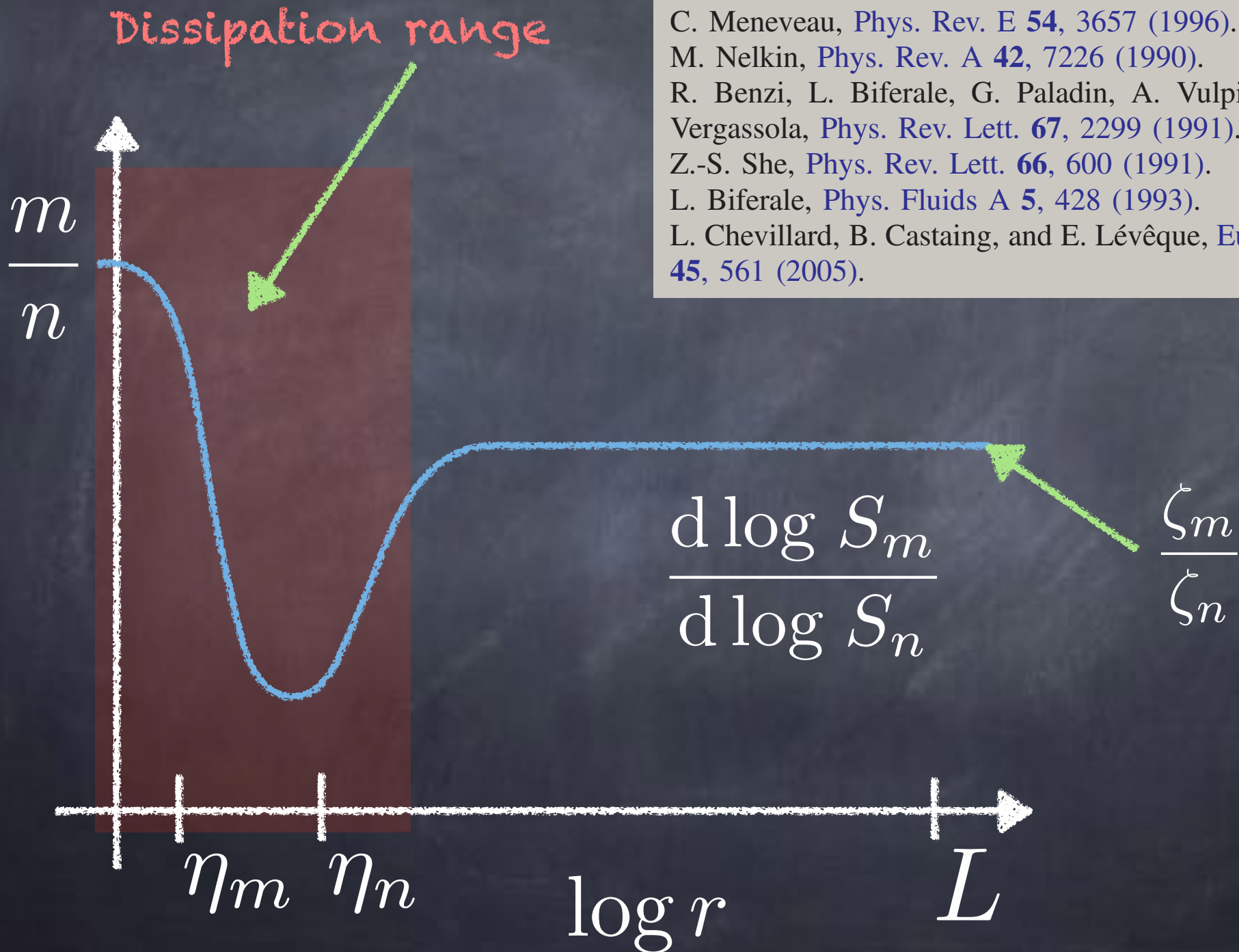
M. Nelkin, *Phys. Rev. A* **42**, 7226 (1990).

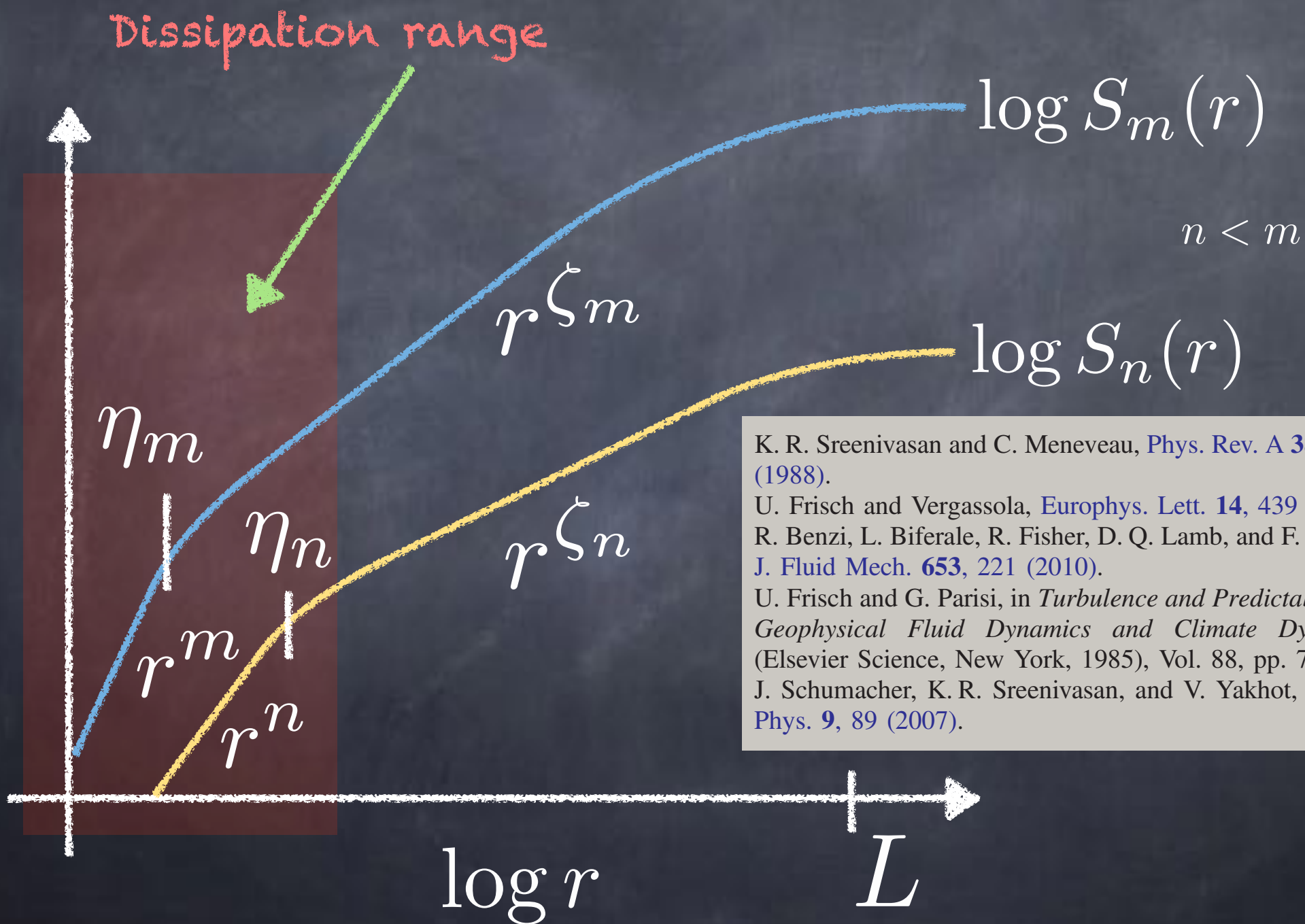
R. Benzi, L. Biferale, G. Paladin, A. Vulpiani, and M. Vergassola, *Phys. Rev. Lett.* **67**, 2299 (1991).

Z.-S. She, *Phys. Rev. Lett.* **66**, 600 (1991).

L. Biferale, *Phys. Fluids A* **5**, 428 (1993).

L. Chevillard, B. Castaing, and E. Lévêque, *Eur. Phys. J. B* **45**, 561 (2005).





K. R. Sreenivasan and C. Meneveau, *Phys. Rev. A* **38**, 6287 (1988).

U. Frisch and Vergassola, *Europhys. Lett.* **14**, 439 (1991).

R. Benzi, L. Biferale, R. Fisher, D. Q. Lamb, and F. Toschi, *J. Fluid Mech.* **653**, 221 (2010).

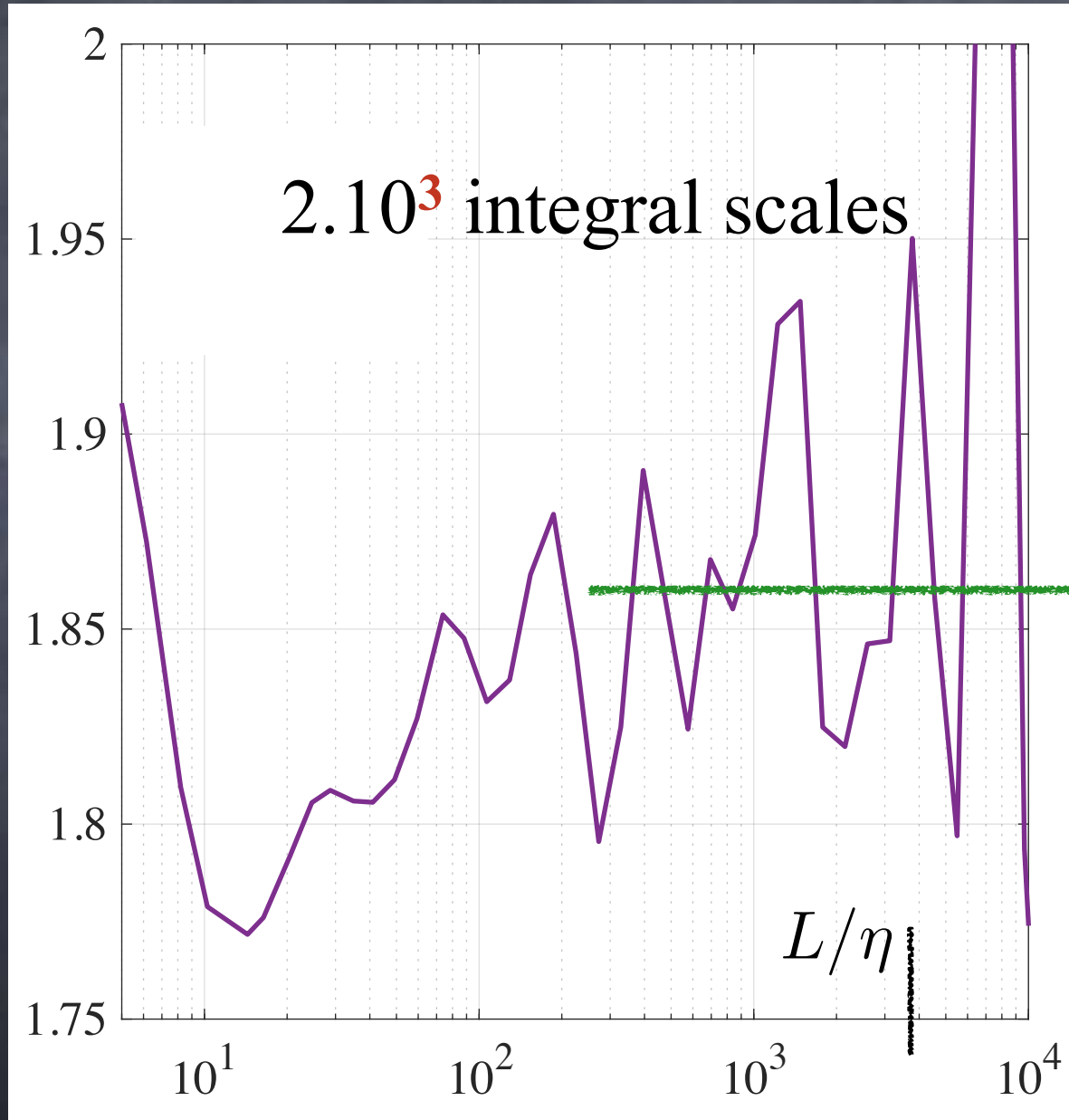
U. Frisch and G. Parisi, in *Turbulence and Predictability in Geophysical Fluid Dynamics and Climate Dynamics* (Elsevier Science, New York, 1985), Vol. 88, pp. 71–88.

J. Schumacher, K. R. Sreenivasan, and V. Yakhot, *New J. Phys.* **9**, 89 (2007).

How much data?

$$R_\lambda = 1030$$

$$\frac{d \log S_4}{d \log S_2}$$



0.2 km
of data

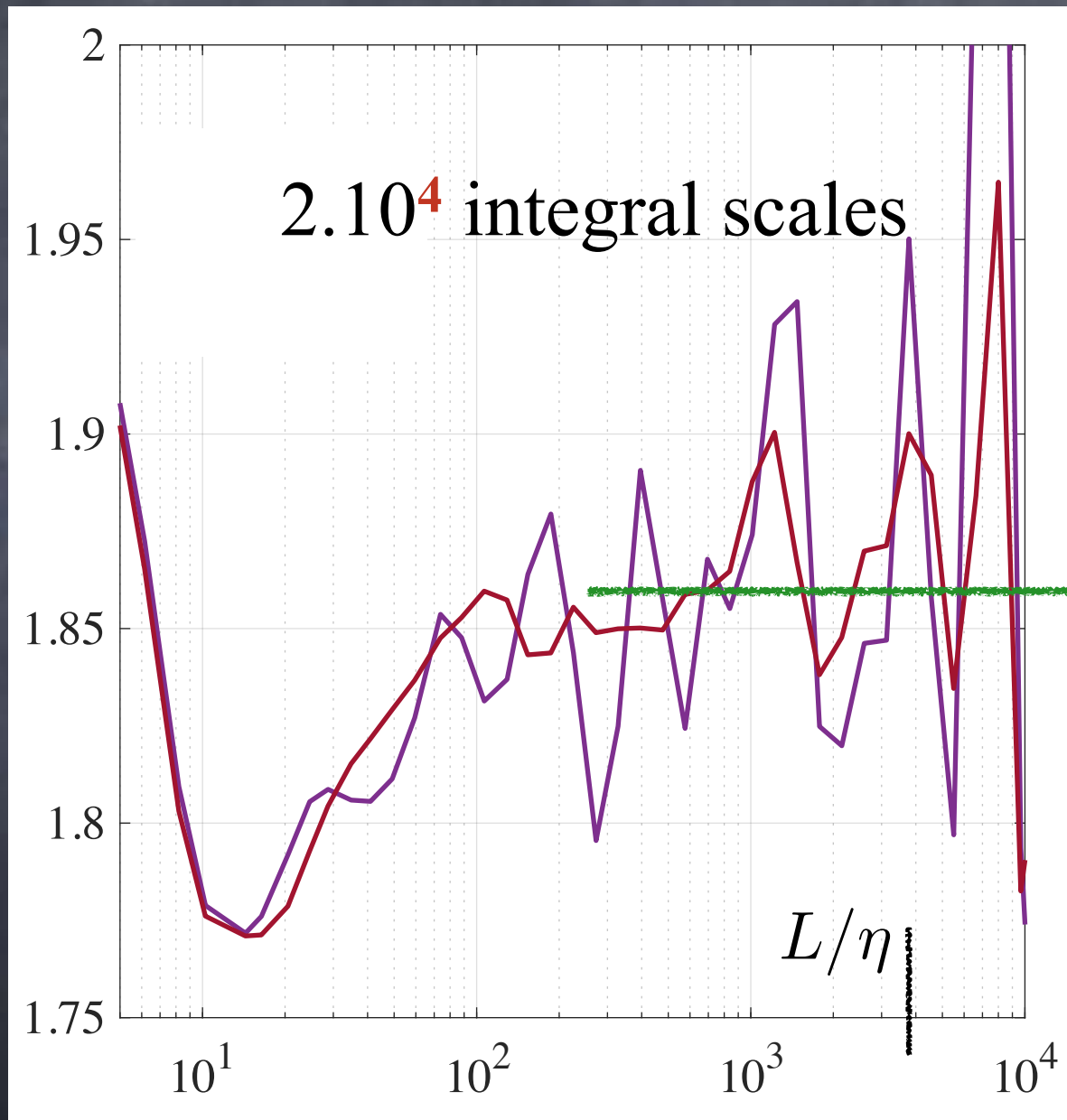
$$\frac{\zeta_4}{\zeta_2}$$

$$r/\eta$$

How much data?

$$R_\lambda = 1030$$

$$\frac{d \log S_4}{d \log S_2}$$



2 km
of data

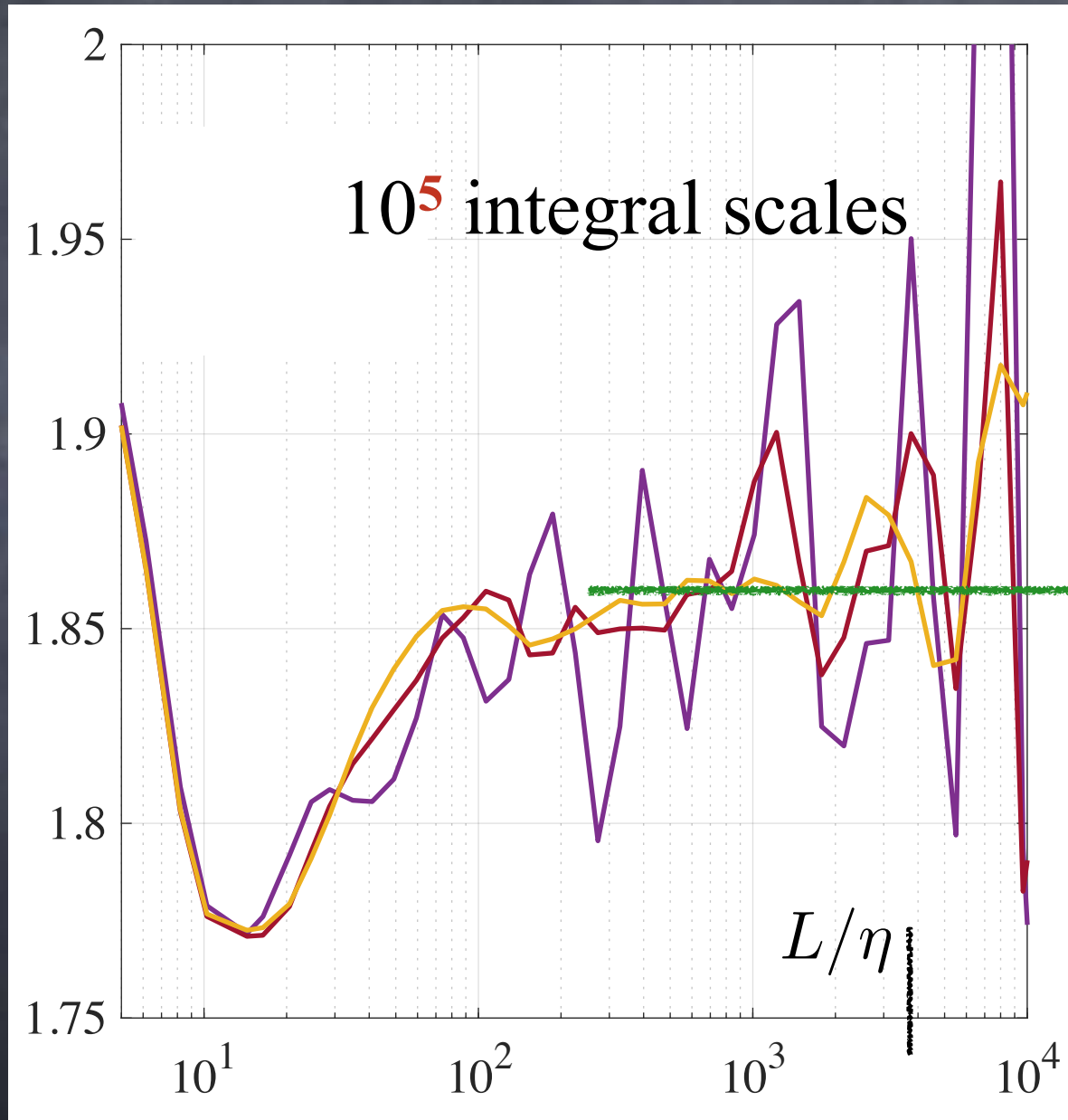
$$\frac{\zeta_4}{\zeta_2}$$

$$r/\eta$$

How much data?

$$R_\lambda = 1030$$

$$\frac{d \log S_4}{d \log S_2}$$



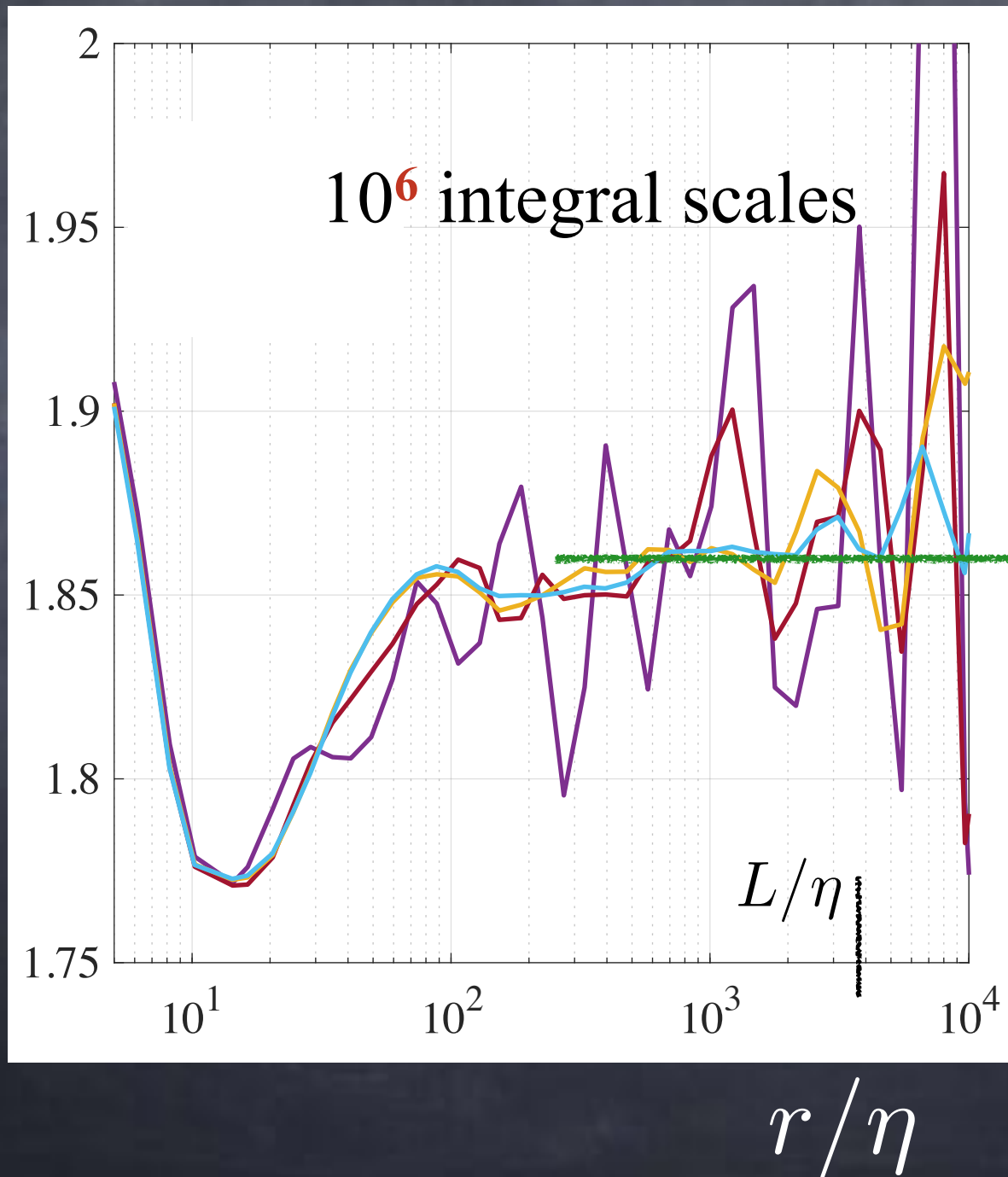
10 km
of data

$$\frac{\zeta_4}{\zeta_2}$$

How much data?

$$R_\lambda = 1030$$

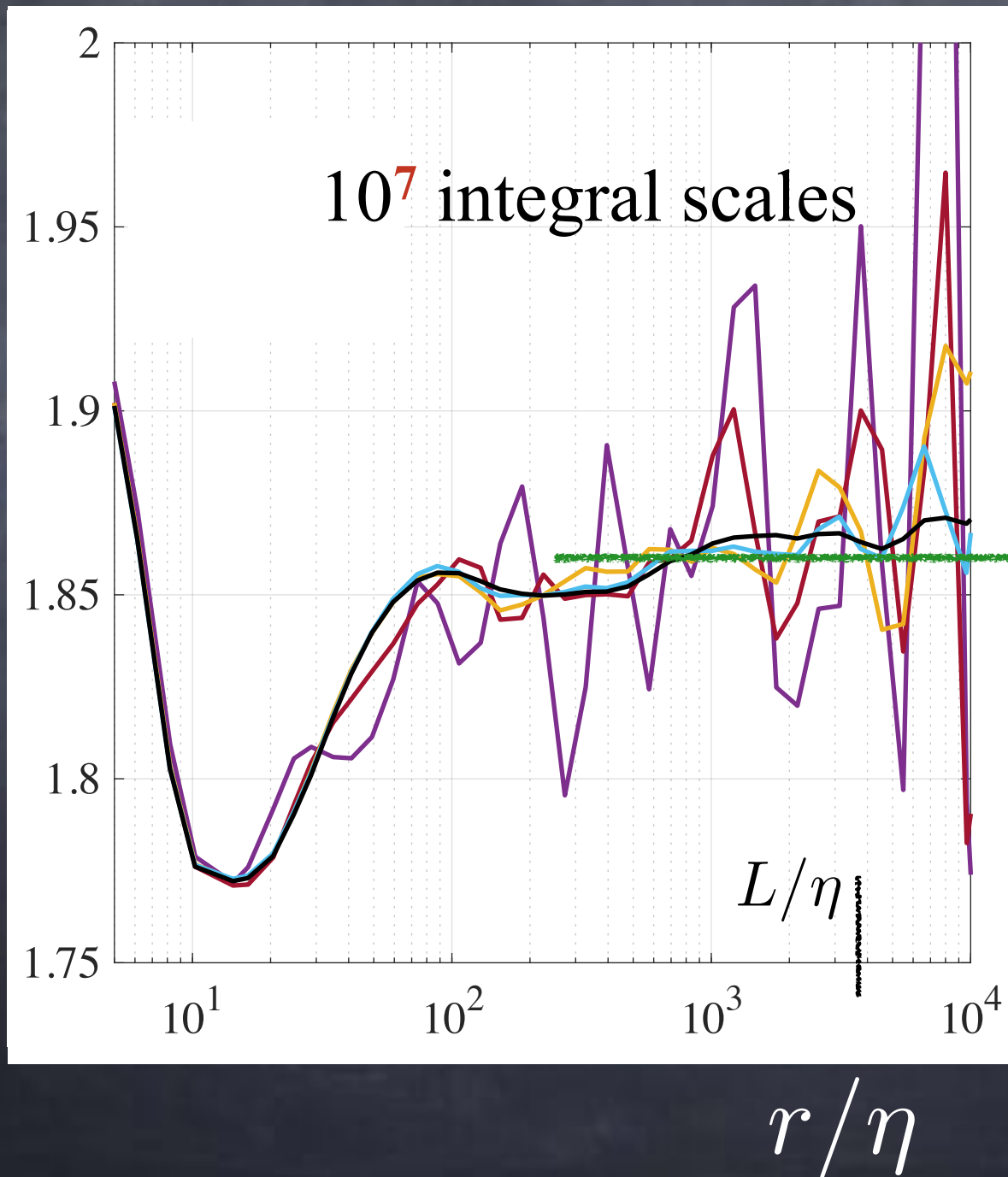
$$\frac{d \log S_4}{d \log S_2}$$



How much data?

$$R_\lambda = 1030$$

$$\frac{d \log S_4}{d \log S_2}$$

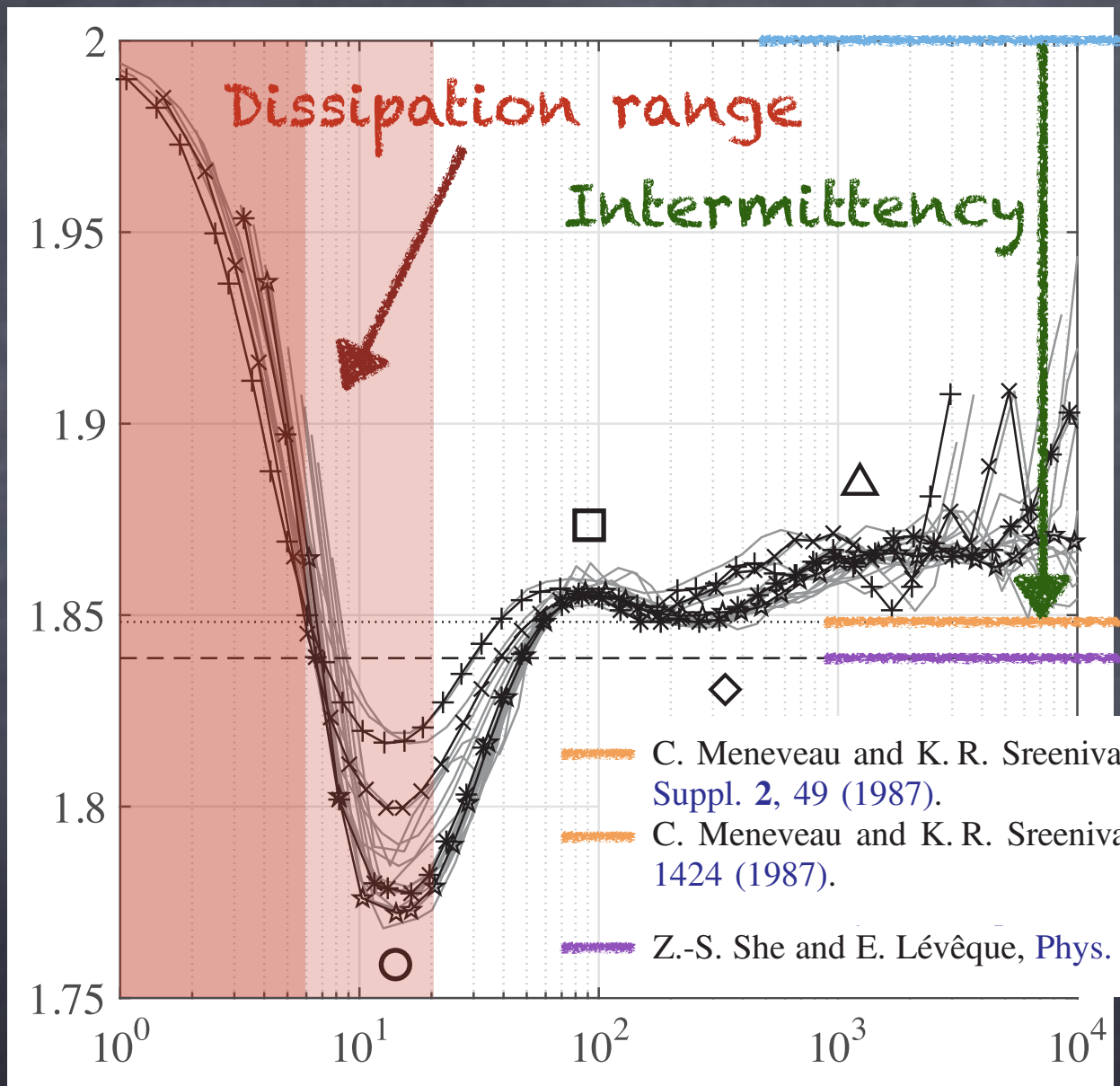


1000 km
of data

$$\frac{\zeta_4}{\zeta_2}$$

For various Reynolds numbers:

$$\frac{d \log S_4}{d \log S_2}$$



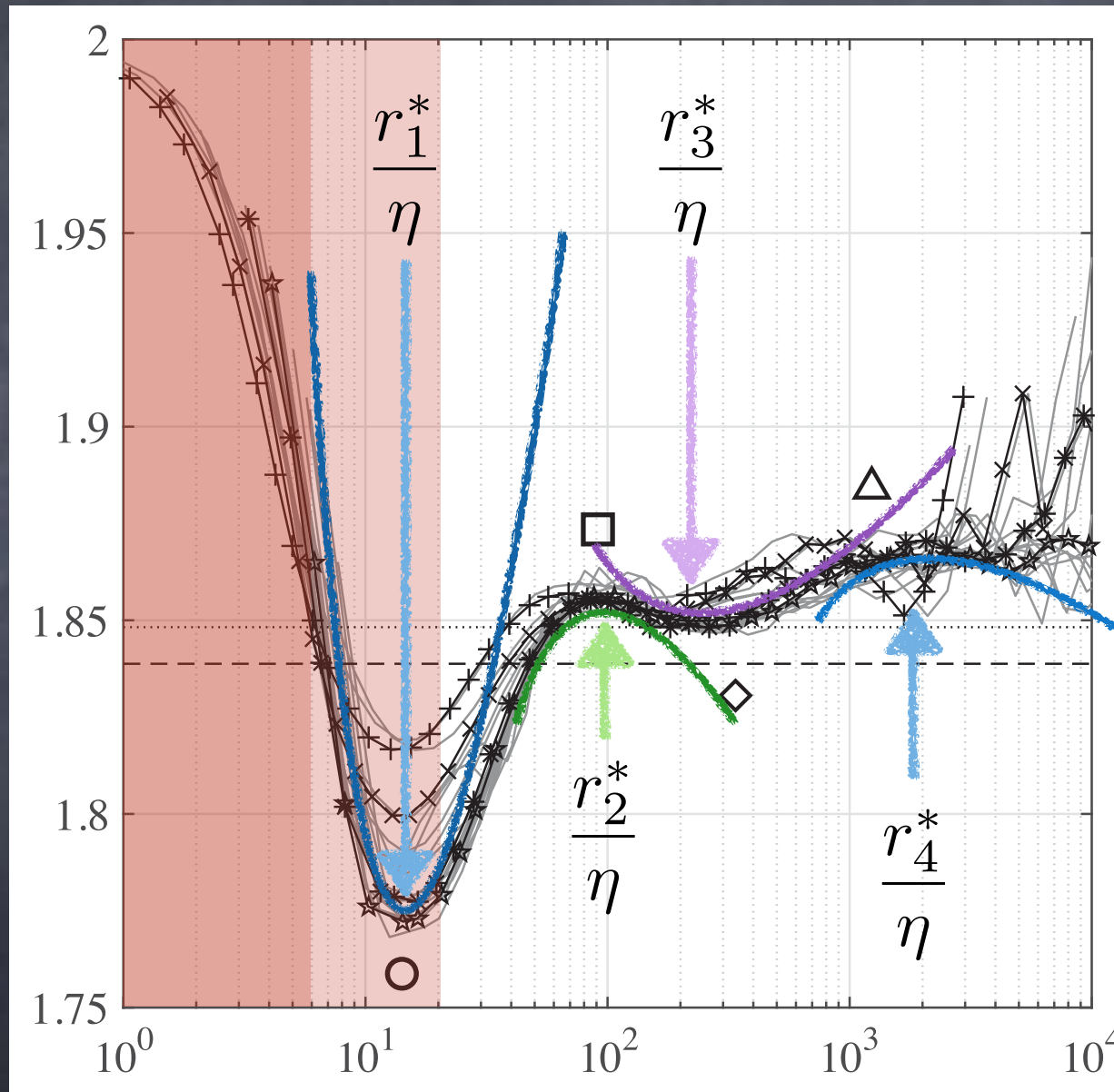
K41

- R_λ
- 260 (+)
 - 510 (x)
 - 880 (*)
 - 1030 (☆)

$$r/\eta$$

For various Reynolds numbers:

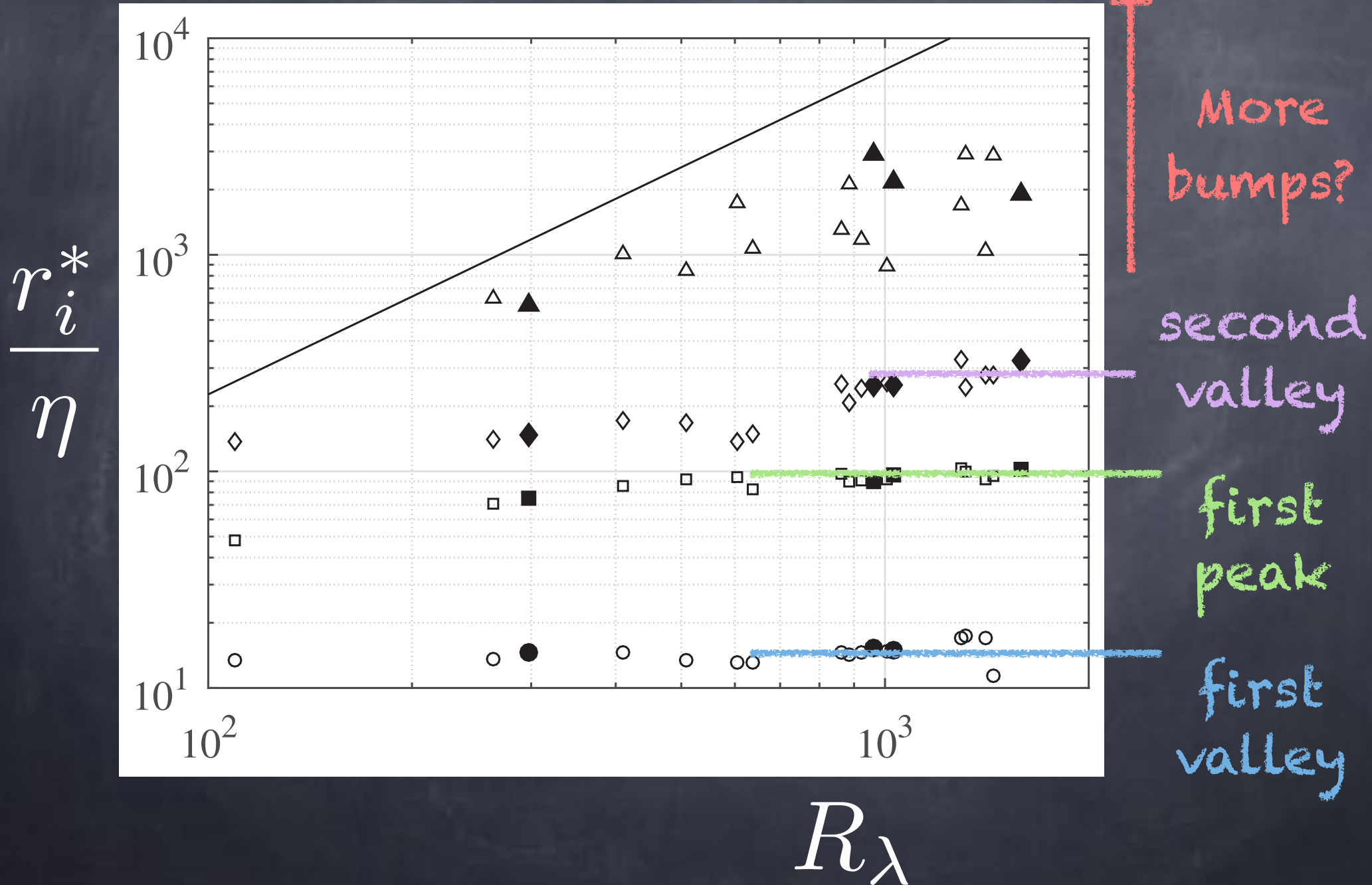
$$\frac{d \log S_4}{d \log S_2}$$

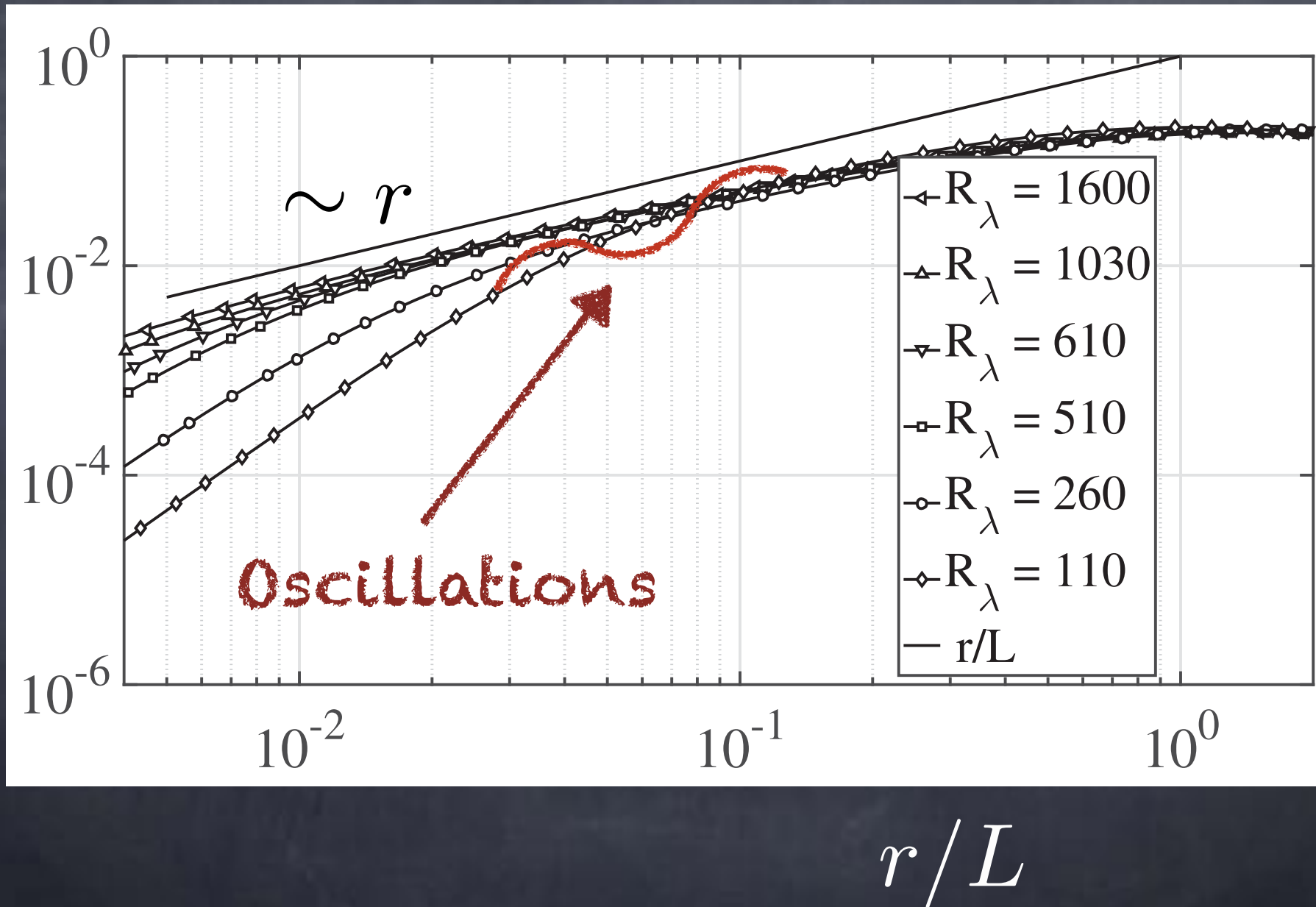


R_λ
 260 (+)
 510 (x)
 880 (*)
 1030 (☆)

$$r/\eta$$

For various Reynolds numbers:



S_3 

Implications

- Structure functions oscillate.
- To measure scaling exponents accurately, these oscillations need to be understood theoretically.



The Bewley applied turbulence lab



Thanks!



Cornell University

Come to Cornell!

THANK YOU!

Active grid

E. Cekli
F. Köhler
J. Kassel
F. Lachaussée
H. Grajewski
H. Zhang
J. Liu
K. Griffin
N. Wei
W. v.d.Water
Z. Warhaft

Wind tunnel

A. Costanzo
C. Küchler
H. Nobach
H. Eckelmann
G. Schewe

NSTAP

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M. Hultmark
A. Smits

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