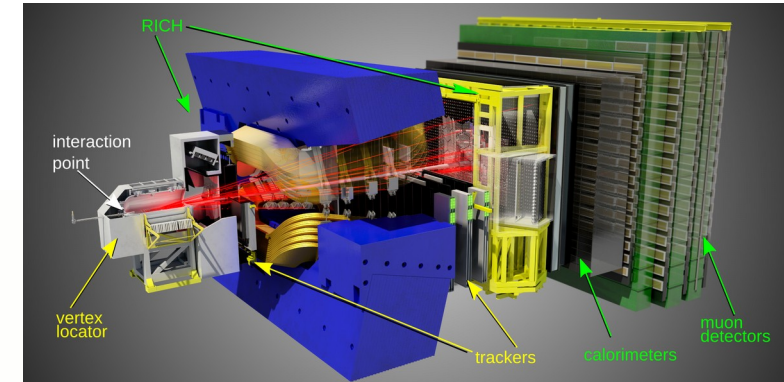
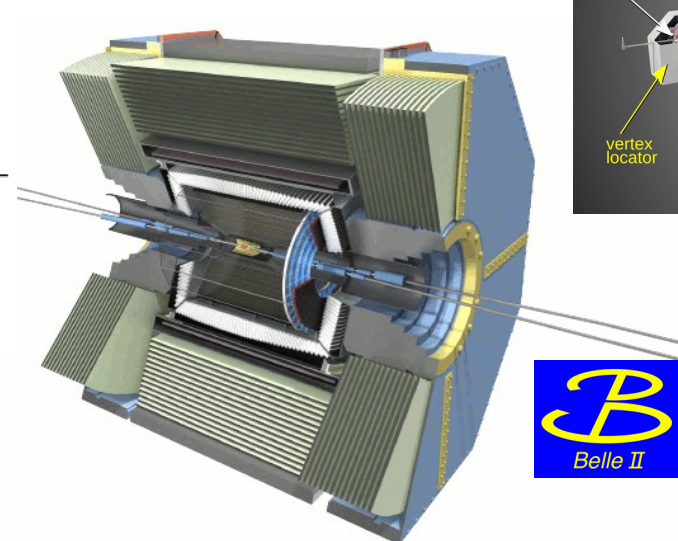
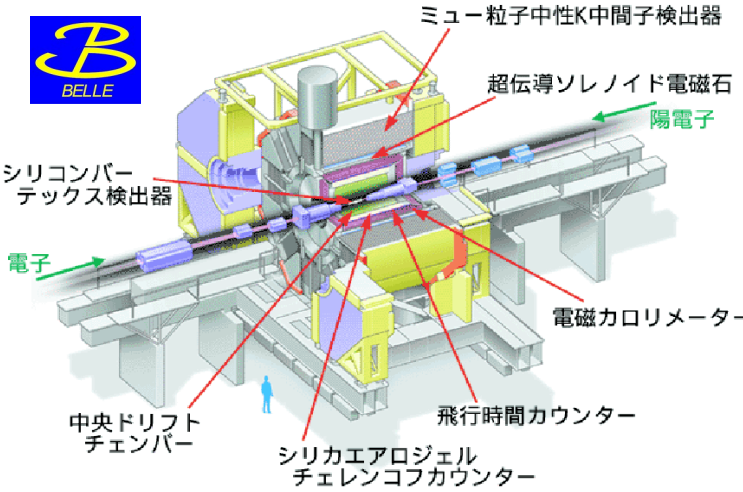
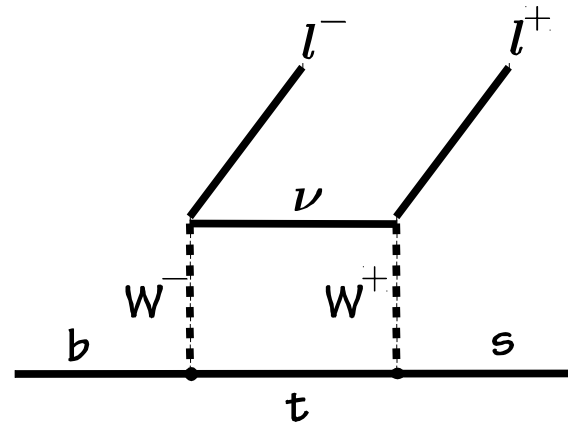
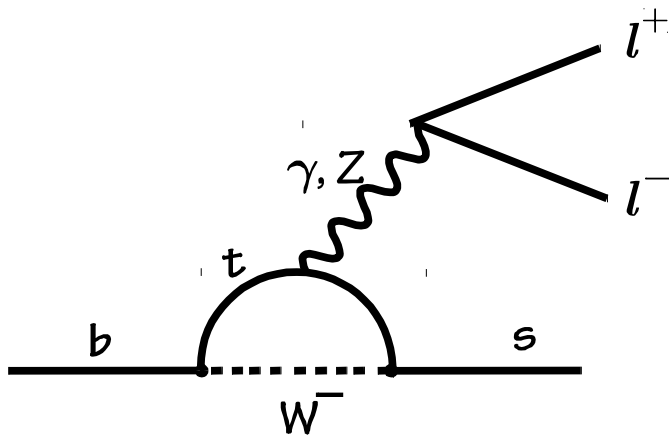
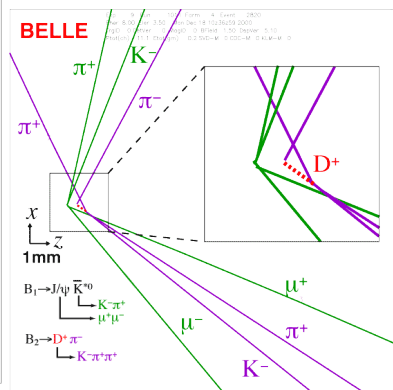
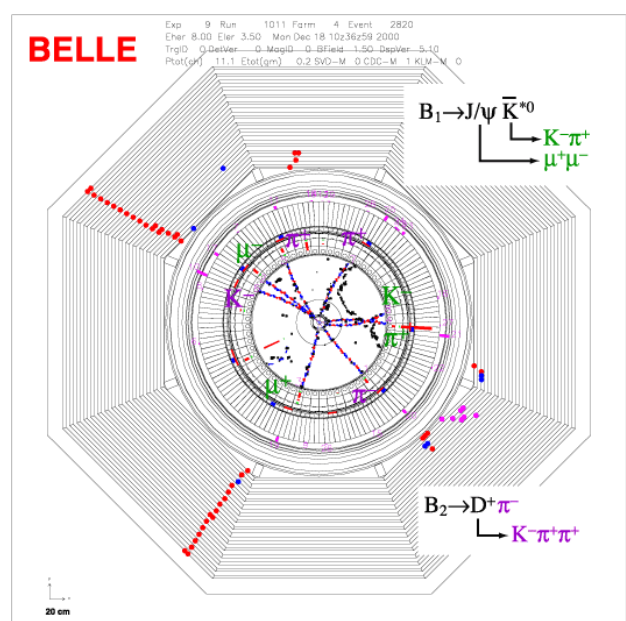
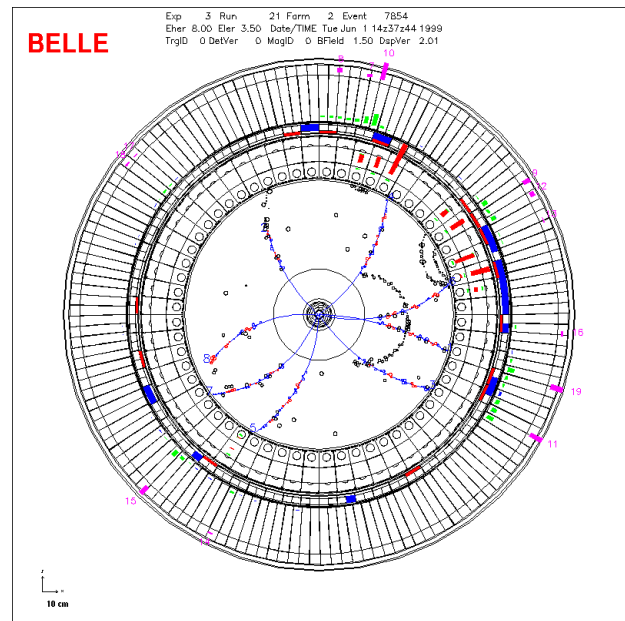
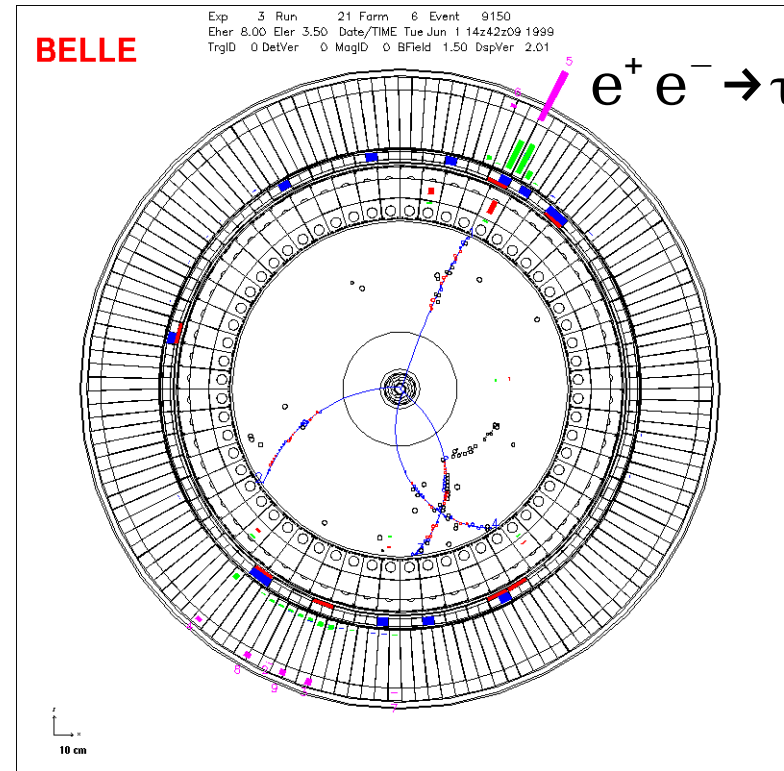
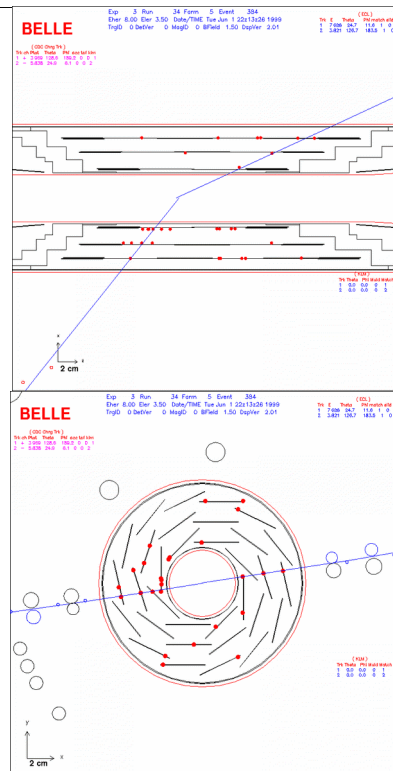
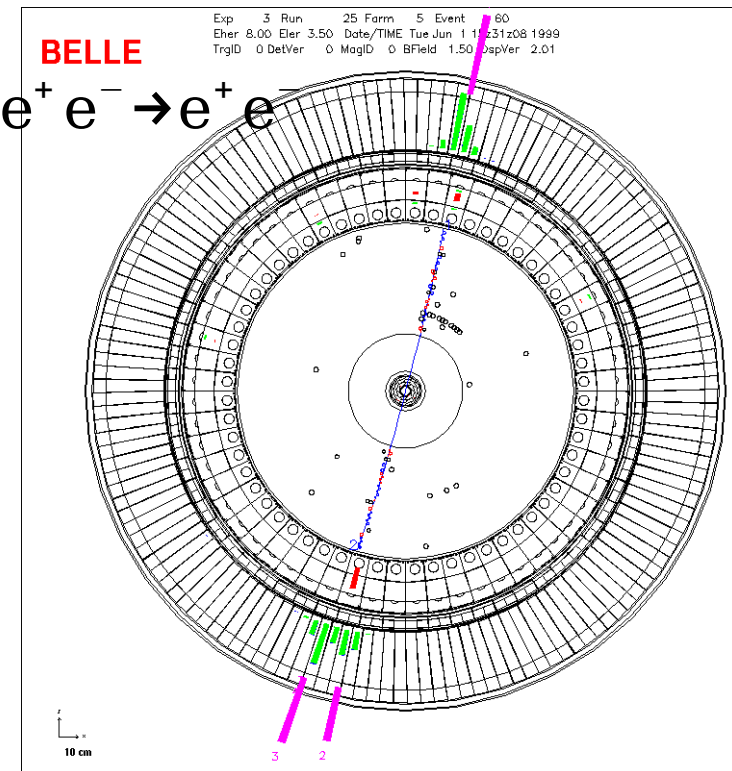


Beautiful paths to probe physics beyond the standard model of particles

K. Trabelsi
karim.trabelsi@in2p3.fr

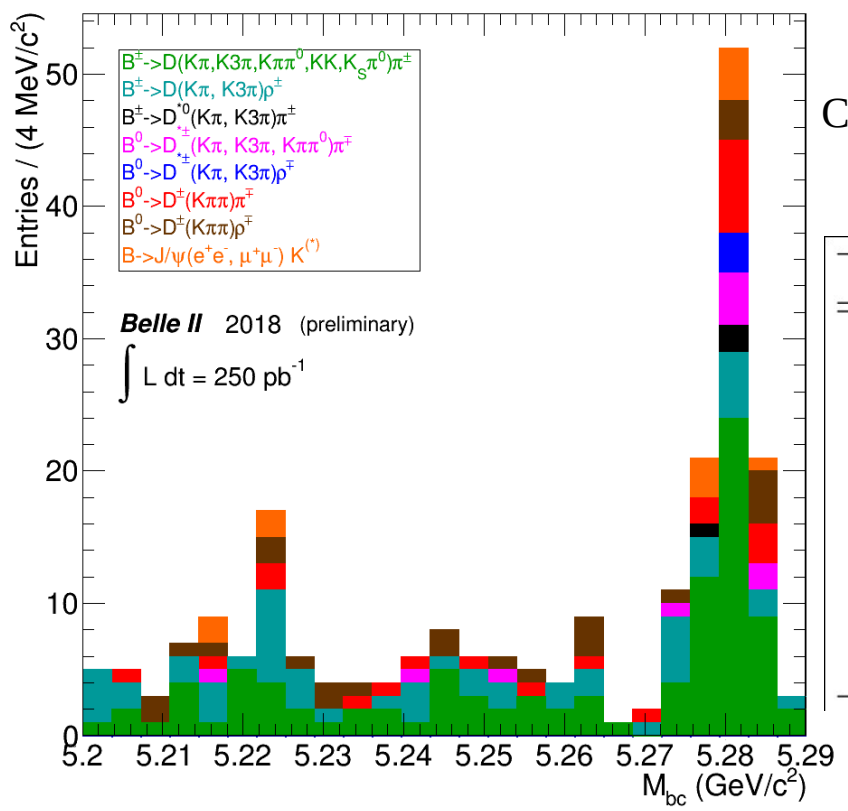
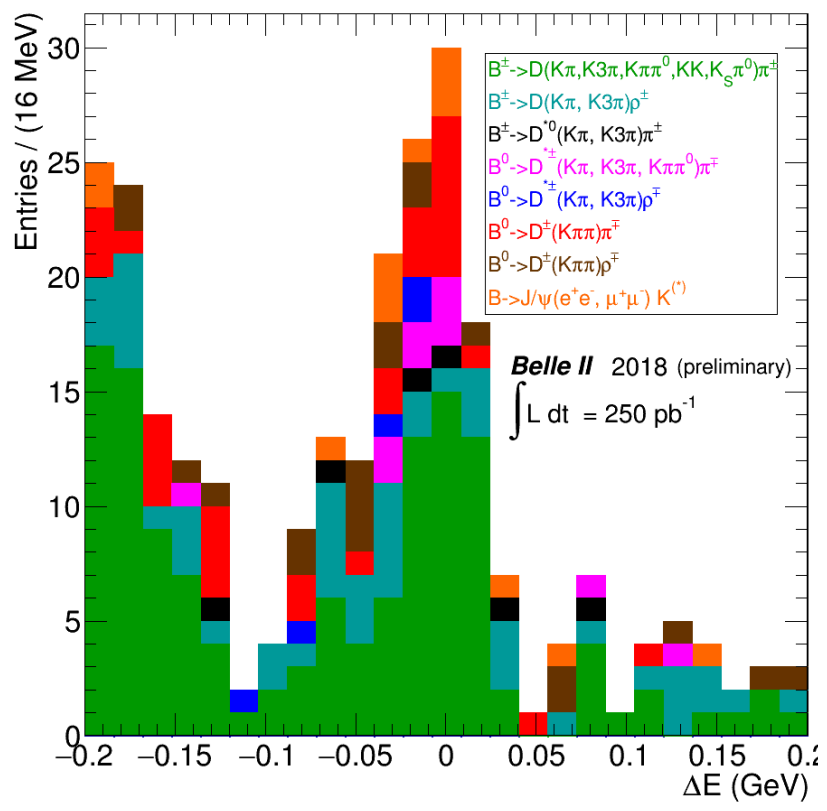


How to detect particles in Belle



Rediscovering beauty: $B \rightarrow D^{(*)} h + B \rightarrow J/\psi K^{(*)}$

with very limited statistics ($< 1 \text{ fb}^{-1}$), Belle II can rediscover the B meson



Candidates in signal box
 $(M_{bc} > 5.27 \text{ GeV}/c^2,$
 $|\Delta E| < 0.050 \text{ GeV})$

Mode	yield
$B^\pm \rightarrow D\pi^\pm$	51
$B^\pm \rightarrow D\rho^\pm$	16
$B^\pm \rightarrow D^*\pi^\pm$	3
$B^0 \rightarrow D^{*\pm}\pi^\mp$	7
$B^0 \rightarrow D^{*\pm}\rho^\mp$	3
$B^0 \rightarrow D^\pm\pi^\mp$	13
$B^0 \rightarrow D^\pm\rho^\mp$	8
$B \rightarrow J/\psi K^{(*)}$	8

~ 100 evts

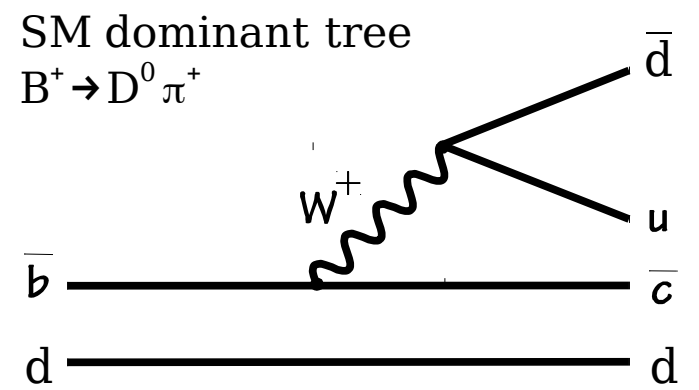
Show capacity for charm physics in $e^+ e^- \rightarrow c \bar{c}$

- D^0, D^+, D^*
- Cabibbo favoured and suppressed modes

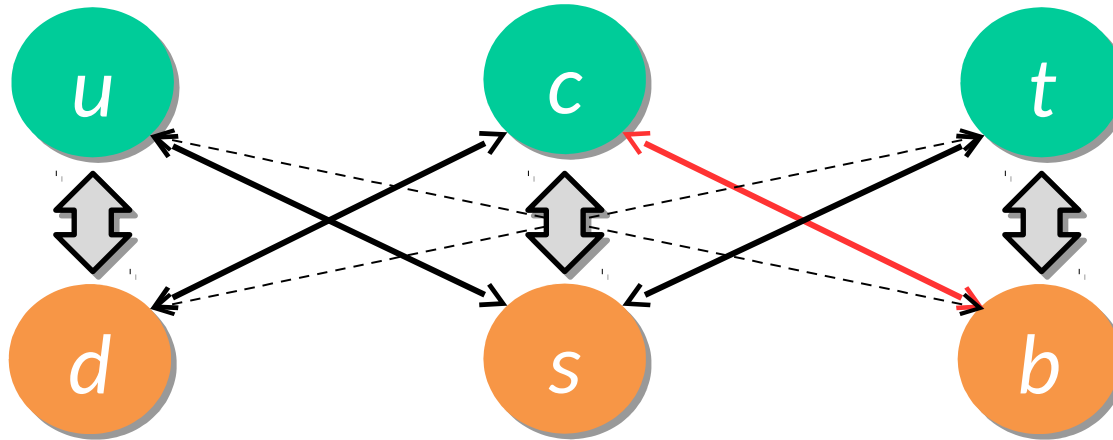
...for B-physics

- hadronic modes from $b \rightarrow c$
- semileptonic decay modes from $b \rightarrow c$

that is for dominant decays ...
... we are looking for rare decays



But you said "rare B decays"...



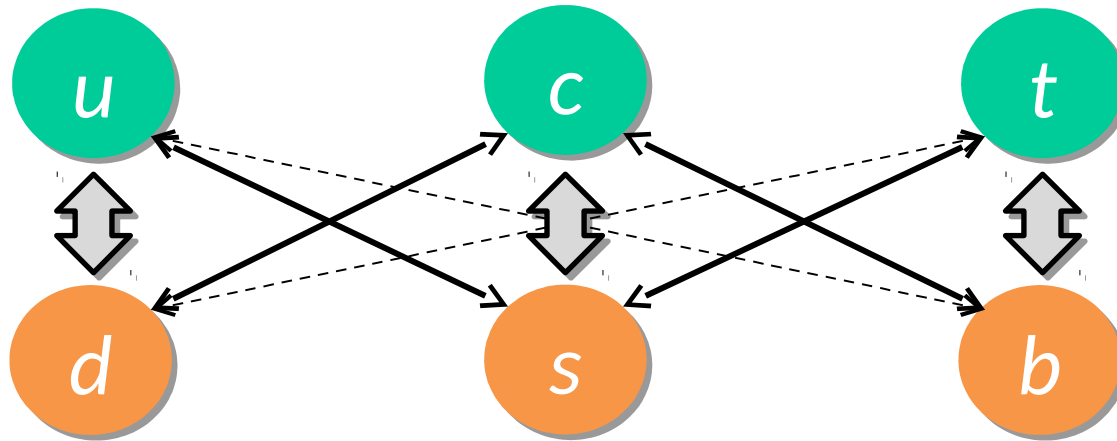
What are rare decays ?

Dominant decays: Not rare

Phase space suppressed decays: Not that rare

Cabibbo-suppressed decays: Some call them rare

$B(D^0 \rightarrow K^- \pi^+)$ versus $B(D^0 \rightarrow \pi^- \pi^+)$ $B(b \rightarrow c l^+ \nu)$ versus $B(b \rightarrow u l^+ \nu)$



What are rare decays ?

Dominant decays: Not rare

Phase space suppressed decays: Not that rare

Cabibbo-suppressed decays: Some call them rare

$$\frac{B(D^0 \rightarrow K^- \pi^+)}{B(D^0 \rightarrow \pi^- \pi^+)} = 28 \quad \frac{B(b \rightarrow c l^+ \nu)}{B(b \rightarrow u l^+ \nu)} = 135$$

What are rare decays ?

Dominant decays: Not rare

Phase space suppressed decays: Not that rare

Cabibbo-suppressed decays: Some call them rare

Colour-suppressed decays: Not really rare

$$B(B^0 \rightarrow D^- \pi^+) = (3.5 \pm 0.9) 10^{-3},$$

$$B(B^0 \rightarrow \bar{D}^0 \pi^0) = (2.9 \pm 0.3) 10^{-4},$$

while they are both $b \rightarrow c W$ and $W \rightarrow u \bar{d}$ transitions.

What are rare decays ?

Dominant decays: Not rare

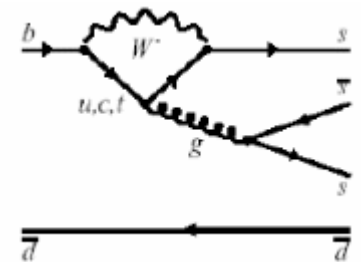
Phase space suppressed decays: Not that rare

Cabibbo-suppressed decays: Some call them rare

Colour-suppressed decays: Not really rare

Hadronic FCNC decays: Not the topic of our lecture

- For instance $B \rightarrow \phi K_S^0$, or $B \rightarrow K_S^0 K \pi \dots$



What are rare decays ?

Dominant decays: Not rare

Phase space suppressed decays: Not that rare

Cabibbo-suppressed decays: Some call them rare

Colour-suppressed decays: Not really rare

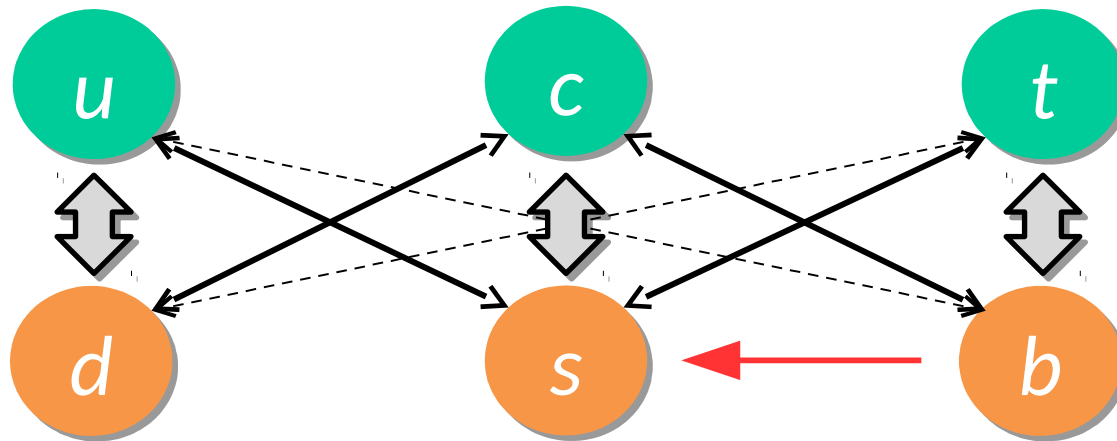
Hadronic FCNC decays: Not the topic of our lecture

- For instance $B \rightarrow \phi K_S^0$, or $B \rightarrow K_S^0 K \pi \dots$
- Or $B^0 \rightarrow \phi K_S^0$, or the penguin contribution to $B \rightarrow J/\psi K_S^0$

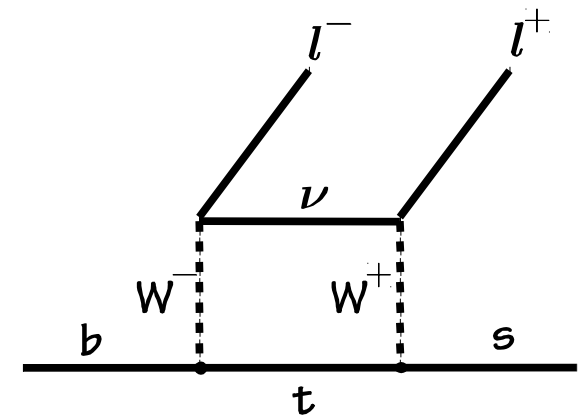
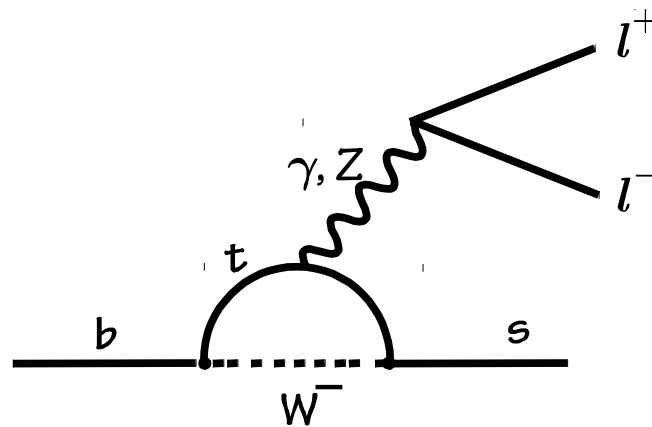
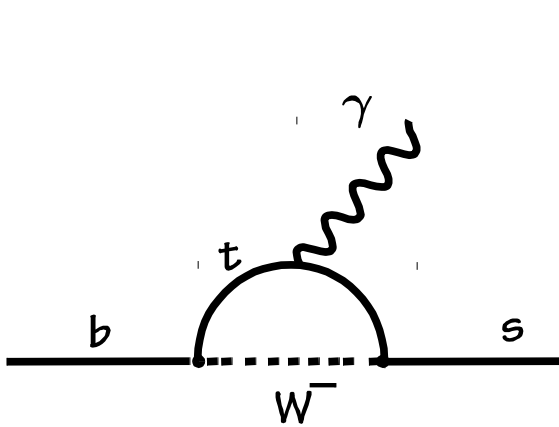
Electroweak FCNC penguins: That's rare !

- $b \rightarrow s \gamma$
- $b \rightarrow s l l$
- And friends...

Rare B decays



- FCNC: Flavour Changing Neutral Current
- FCNC are strongly suppressed in the SM: only loops + GIM mechanism

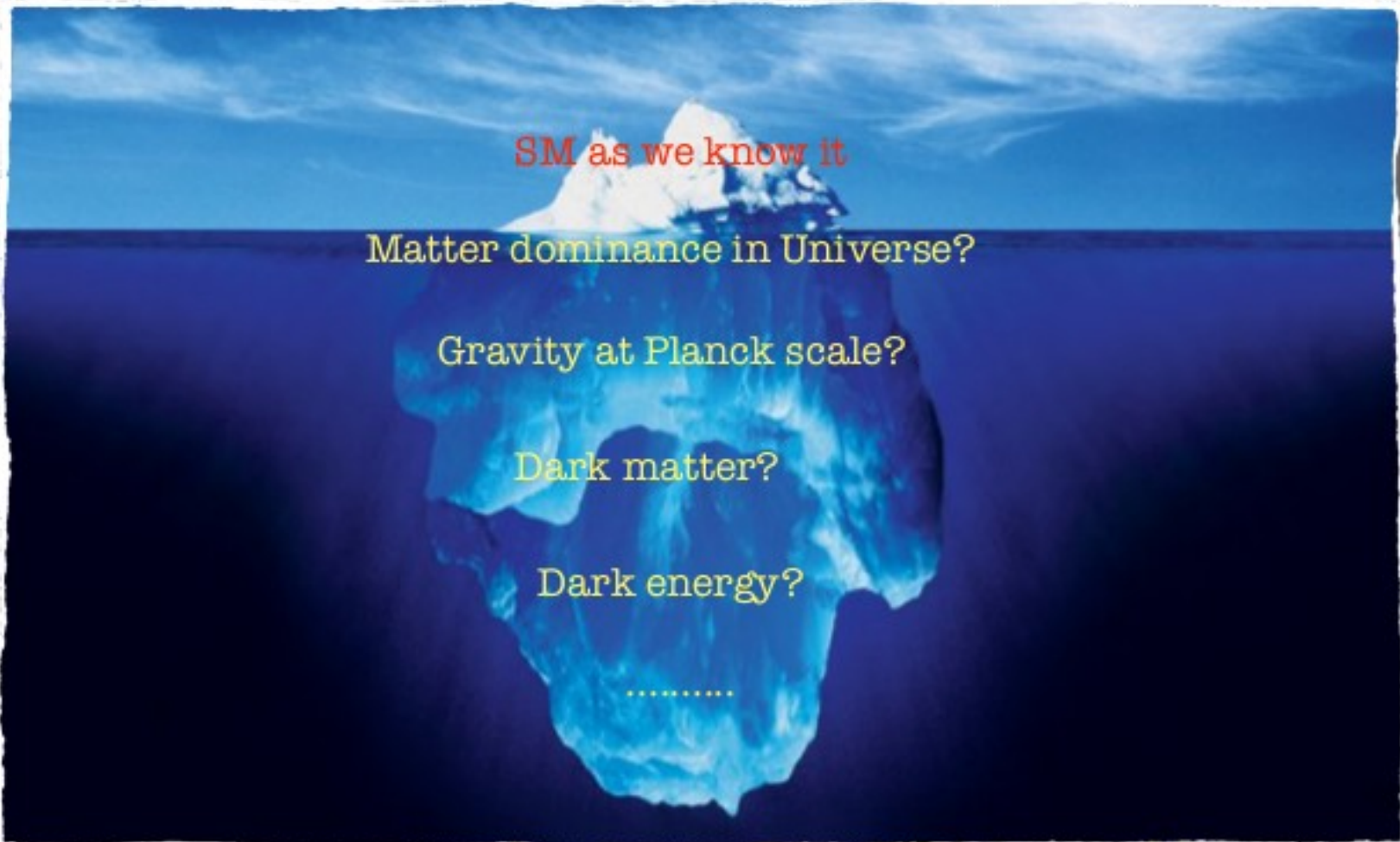


Motivations for NP

SM, are we done ?

No. Open questions

D. Tonelli



These and many other questions fuel the strong and wide-spread prejudice that the SM is **completed at high-energy by new particles and interactions**

How do you we search for new particles ?

Direct vs Indirect Searches

complementarity with LHC

Why flavor physics ?

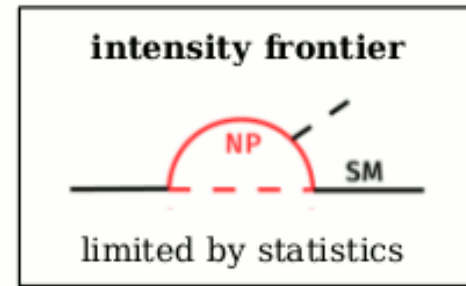
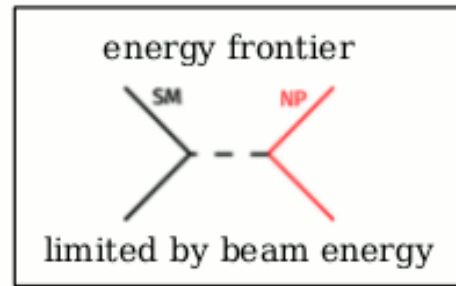
1 TeV

ATLAS SUSY Searches* - 95% CL Lower Limits
December 2017

ATLAS Preliminary
 $\sqrt{s}=7, 8, 13 \text{ TeV}$

Model	$k_{\mu\nu}, T, \tilde{g}$	Jets	$A_{eff}^{(4\ell)}$ (40%)	Mass limit	$\sqrt{s}=7, 8 \text{ TeV}$	$\sqrt{s}=13 \text{ TeV}$	Reference
Relaxed Constraints	$\tilde{g}, \tilde{u}, \tilde{d}$	0	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
1-2 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	1-2	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
3-4 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	3-4	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
5-6 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	5-6	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
7-8 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	7-8	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
9-10 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	9-10	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
11-12 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	11-12	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
13-14 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	13-14	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
15-16 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	15-16	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
17-18 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	17-18	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
19-20 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	19-20	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
21-22 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	21-22	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
23-24 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	23-24	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
25-26 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	25-26	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
27-28 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	27-28	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
29-30 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	29-30	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
31-32 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	31-32	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
33-34 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	33-34	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
35-36 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	35-36	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
37-38 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	37-38	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
39-40 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	39-40	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
41-42 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	41-42	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
43-44 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	43-44	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
45-46 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	45-46	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
47-48 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	47-48	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
49-50 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	49-50	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
51-52 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	51-52	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
53-54 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	53-54	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
55-56 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	55-56	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
57-58 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	57-58	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
59-60 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	59-60	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
61-62 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	61-62	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
63-64 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	63-64	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
65-66 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	65-66	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
67-68 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	67-68	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
69-70 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	69-70	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
71-72 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	71-72	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
73-74 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	73-74	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
75-76 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	75-76	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
77-78 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	77-78	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
79-80 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	79-80	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
81-82 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	81-82	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
83-84 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	83-84	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
85-86 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	85-86	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
87-88 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	87-88	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
89-90 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	89-90	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
91-92 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	91-92	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
93-94 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	93-94	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
95-96 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	95-96	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
97-98 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	97-98	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026
99-100 jets	$\tilde{g}, \tilde{u}, \tilde{d}$	99-100	2.0	300 GeV	1.0 TeV	1.3 TeV	ATLAS-CONF-2017-026

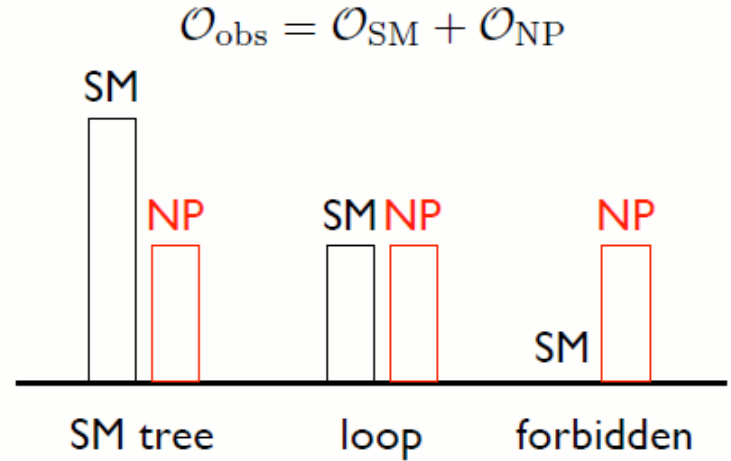
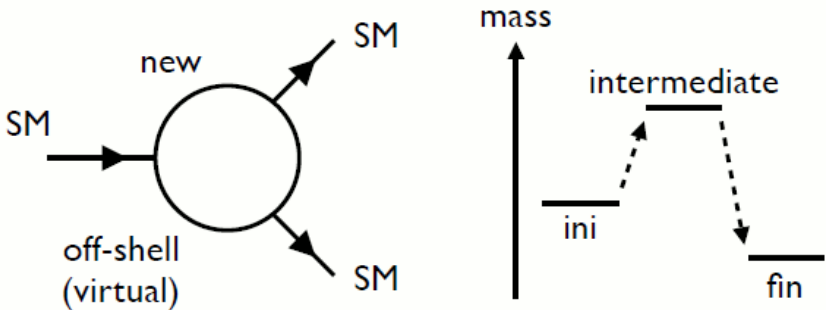
*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, cf. ref. for the search regions.



→ NP beyond the direct reach of the LHC

Three classes of SM processes (M. Endo)

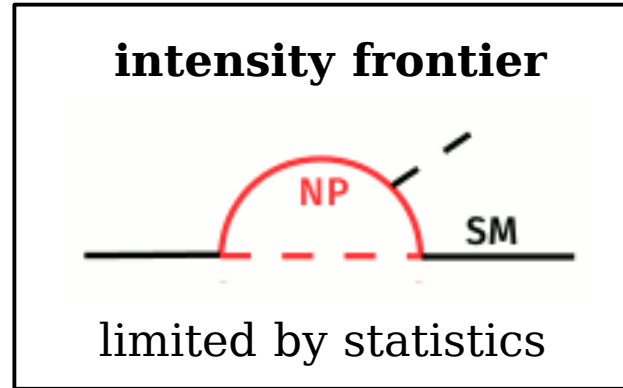
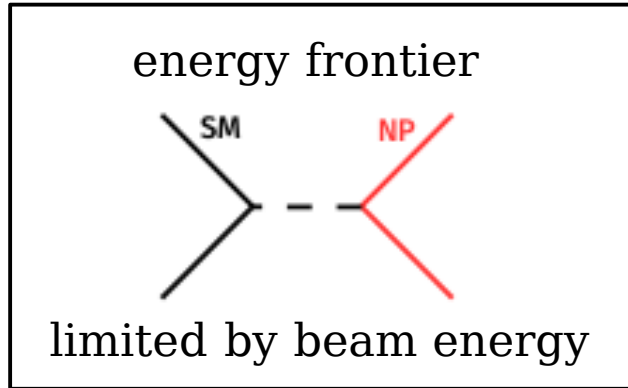
> ~100GeV (1TeV), if interaction is weak (strong)
New particle via quantum effects



No sharp cutoff for energy scale (cf. LHC search)
— suppressed by $(E/\Lambda)^n$

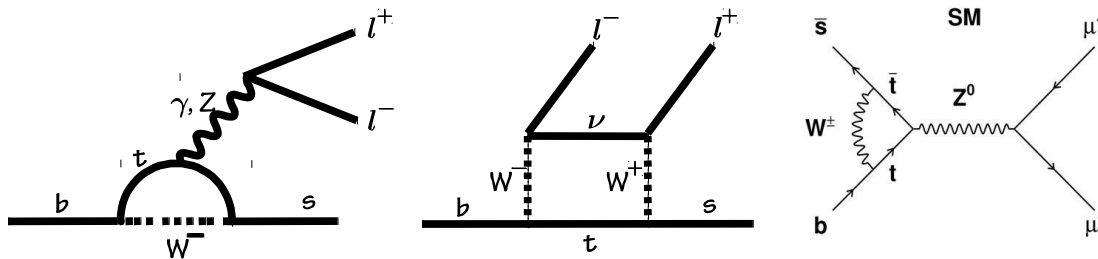
Rare B decays

- FCNC are strongly suppressed in the SM: only loops + GIM mechanism
- Any new particle generating new diagrams can change the amplitudes

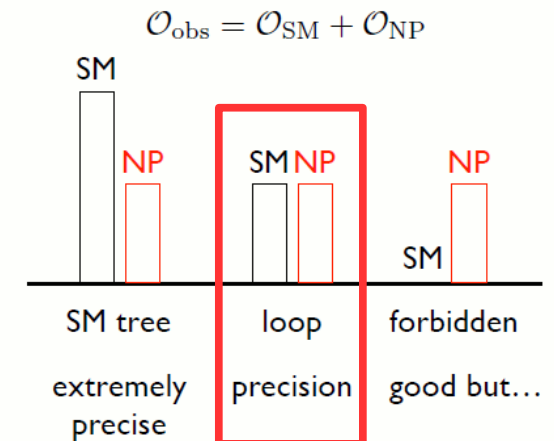


→ NP beyond the direct reach of the LHC

New particles can for example contribute to loop or tree level diagrams **by enhancing/suppressing decay rates, introducing new sources of CP violation or modifying the angular distribution of the final-state particles**



Three classes of SM processes



Why rare decays ?

We want to find new physics indirectly !

No new physics at tree level: we would have noticed ?

Interference of tree interactions and new physics: this is what CP violation does

Interference of loop induced decays and new physics:

- Only allowed in loops
- Could be SM Z and W , or anything else that is heavy

Experimental aspects:

- You want to measure a 50% effect on a rare decay , not a 1% effect on the neutron lifetime . That's very hard .

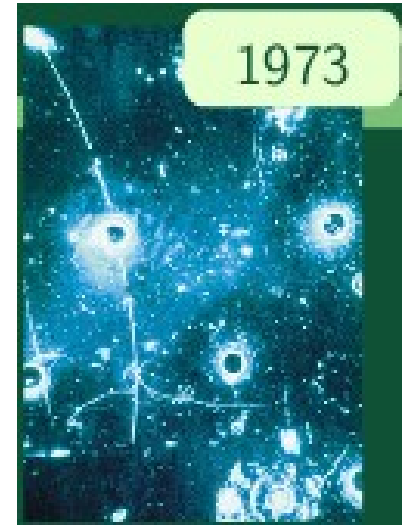
⇒ Statistic versus systematic error

Theoretical **clean**: There are many rare decays that are theoretically clean. This is needed as in the end you will compare a measured effect to an SM prediction.

Indirect searches

Sensitive to New Physics effects

- When was the Z discovered ?
 - 1973 from $N \nu \rightarrow N \nu$?
 - 1983 at SpS ?
- c quark postulated by GIM, third family by KM



Estimate masses

- t quark from $B\bar{B}$ mixing

Get phases of couplings

- Half of new parameters
- Needed for a full understanding

Look in lepton and **flavour** sectors

→ CP asymmetry in the Universe

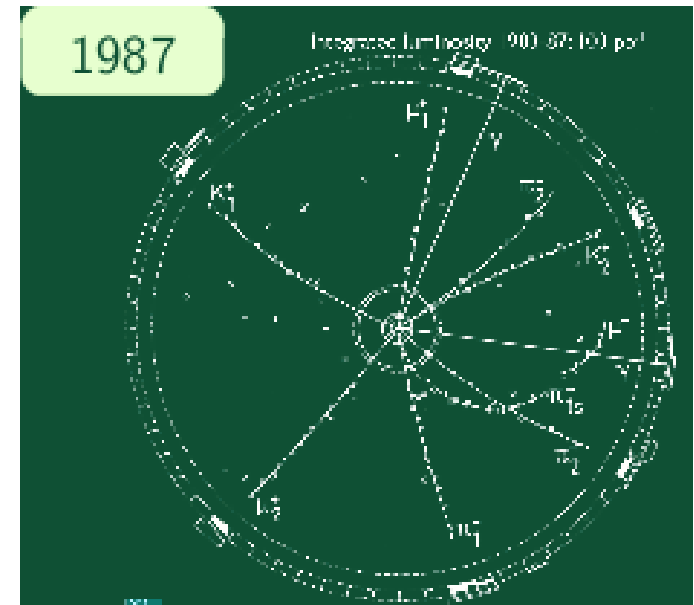
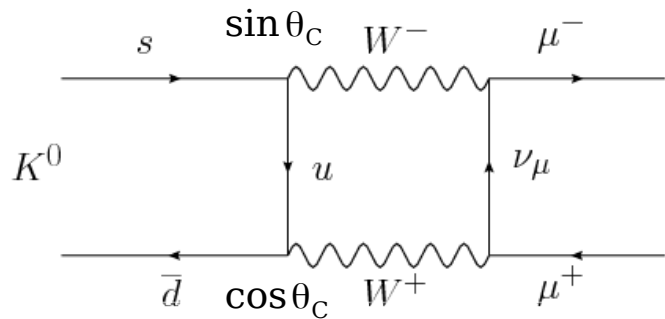


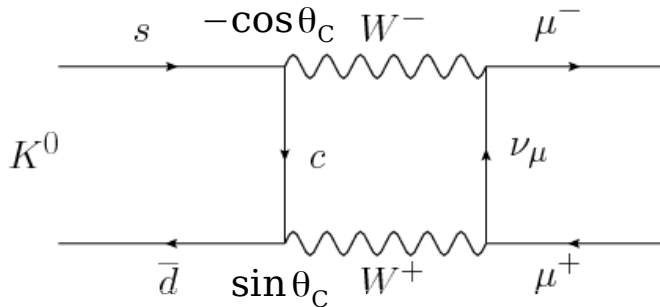
Illustration of indirect search: $K_L^0 \rightarrow \mu\mu$

$K_L \rightarrow \mu^+ \mu^-$ decay can be generated by the box diagram:



$K_L^0 \rightarrow \mu\mu$ was not observed though expected
 Now BF is measured to be $(6.84 \pm 0.11)10^{-9}$
 [Ambrose et al, 2000]

in a renormalisable gauge theory, is expected to give a branching ratio of $g^4 \sim \alpha^2 \sim 10^{-4}$, with α the fine structure constant.



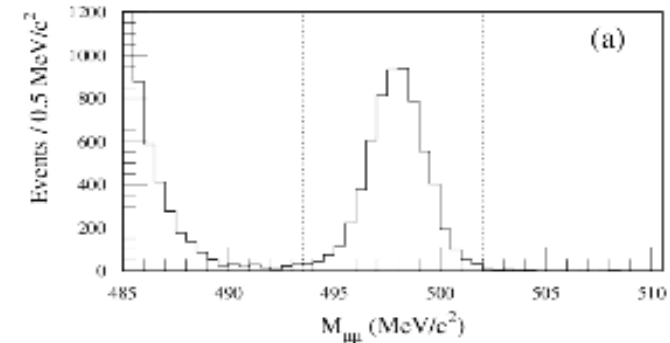
GIM observed that, with a fourth quark, there is a second diagram, with c replacing u. In the limit of exact flavour symmetry, the two diagrams cancel.

[Glashow, Iliopoulos and Maiani, 1970]

The breaking of flavour symmetry induces a mass difference between the quarks, so the sum of the two diagrams is of order $g^4(m_c^2 - m_u^2)/m_W^2 \sim \alpha^2 m_c^2/m_W^2$.

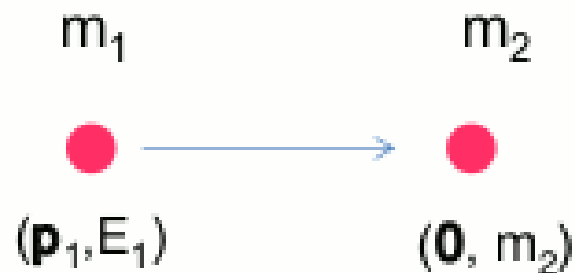
With the measured charm quark mass $m_c \sim 1.27$ GeV, the predicted rates are in agreement with observation.

⇒ but no experimental evidence of a fourth quark...



Proton beam

Fixed target experiment

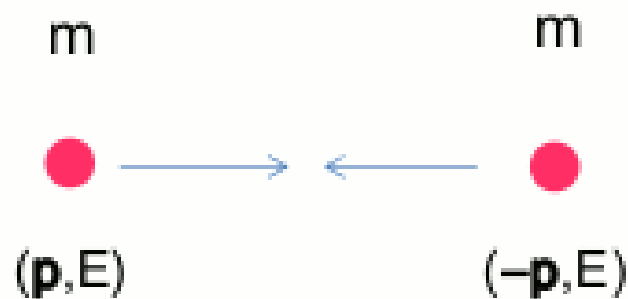


$$\sqrt{s} = 25 \text{ GeV} \gg 3 \text{ GeV}$$

- In proton collision, other particles (than J/ψ) are also produced.

Electron beam

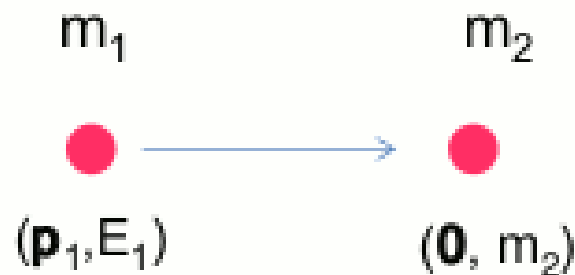
Collider experiment



$$\sqrt{s} = 3 \text{ GeV}$$

- CM energy is efficiently used to produce J/ψ .

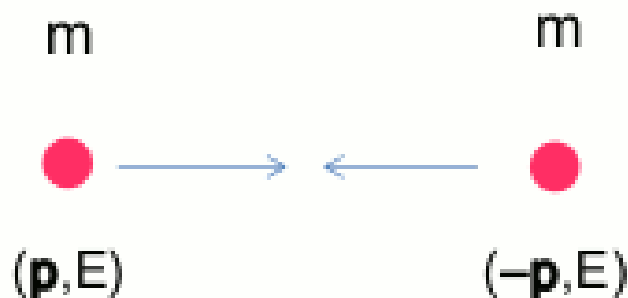
Proton beam
Fixed target experiment



$m_1 = 0.938 \text{ GeV}$
 $\sim 1 \text{ GeV}$: proton
 $m_2 \sim 9 \text{ GeV}$: Beryllium
 $E_1 = 30 \text{ GeV}$

CM energy (center-of-mass) $\sqrt{s} = \sqrt{(E_1 + m_2)^2 - |\vec{p}_1|^2}$
 $= \sqrt{2E_1m_2 + m_1^2 + m_2^2}$
 $= 25 \text{ GeV} \gg 3 \text{ GeV}$

Electron beam
Collider experiment



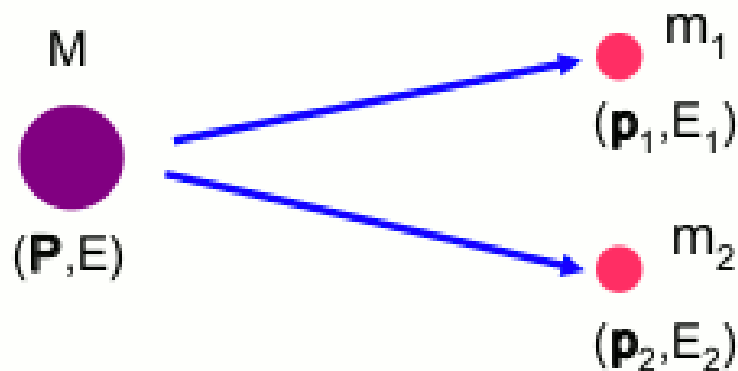
$m = 0.511 \text{ keV}$
 $E = 1.5 \text{ GeV}$

CM energy $\sqrt{s} = 2E = 3 \text{ GeV}$

How to detect particles

Most short-lived particles generated by the collision (, in which we are interested), decay inside the detector, but we can reconstruct them if we know the 4-momentum of decay products.

Simple case: 2-body decay.

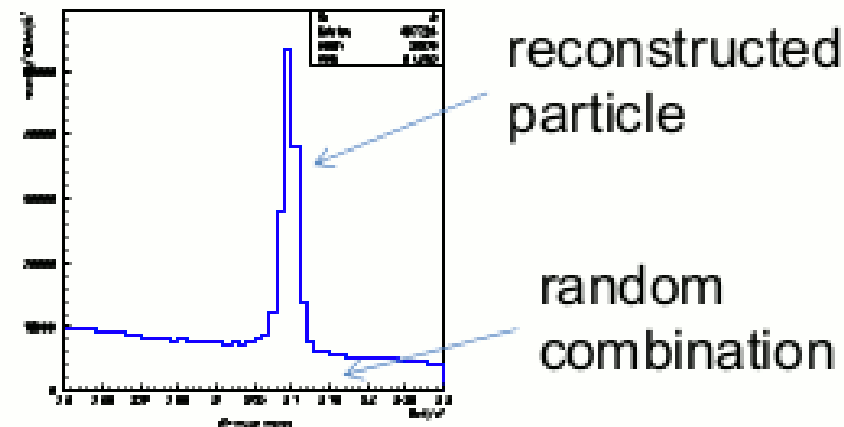


energy and momentum conservation

$$E = E_1 + E_2$$

$$\mathbf{P} = \mathbf{p}_1 + \mathbf{p}_2$$

$$M^2 = E^2 - |\mathbf{P}|^2 = (E_1 + E_2)^2 - |\mathbf{p}_1 + \mathbf{p}_2|^2$$

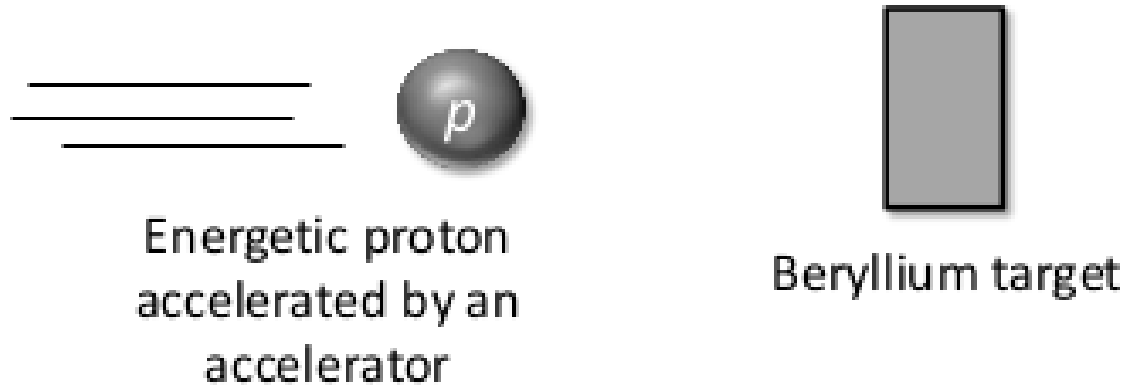


In reality, there are many particles in a final state; we don't know which is the correct combination.

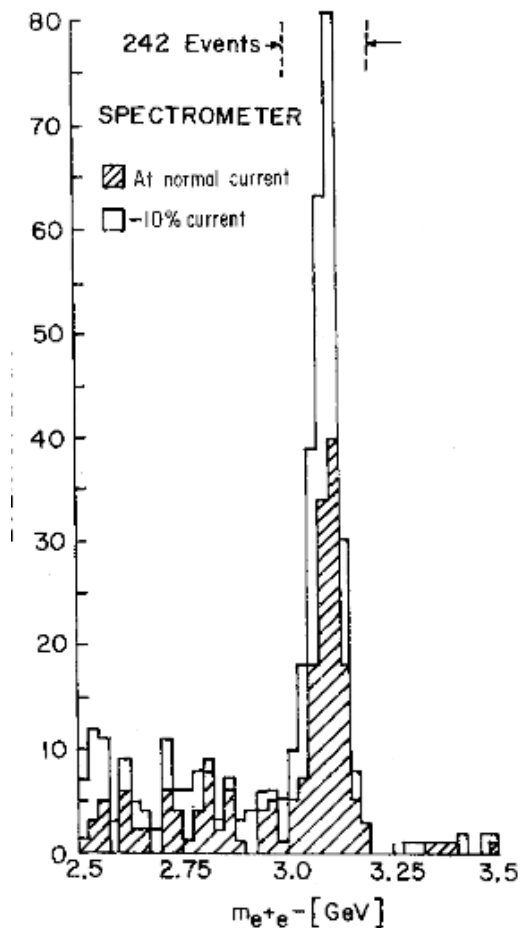
J/ψ (1974)

Experiment carried out by S. Ting group at Brookhaven National Laboratory

– fixed target experiment:



- particle found at $3.0 \text{ GeV}/c^2$
- They coined name of "J" to the particle

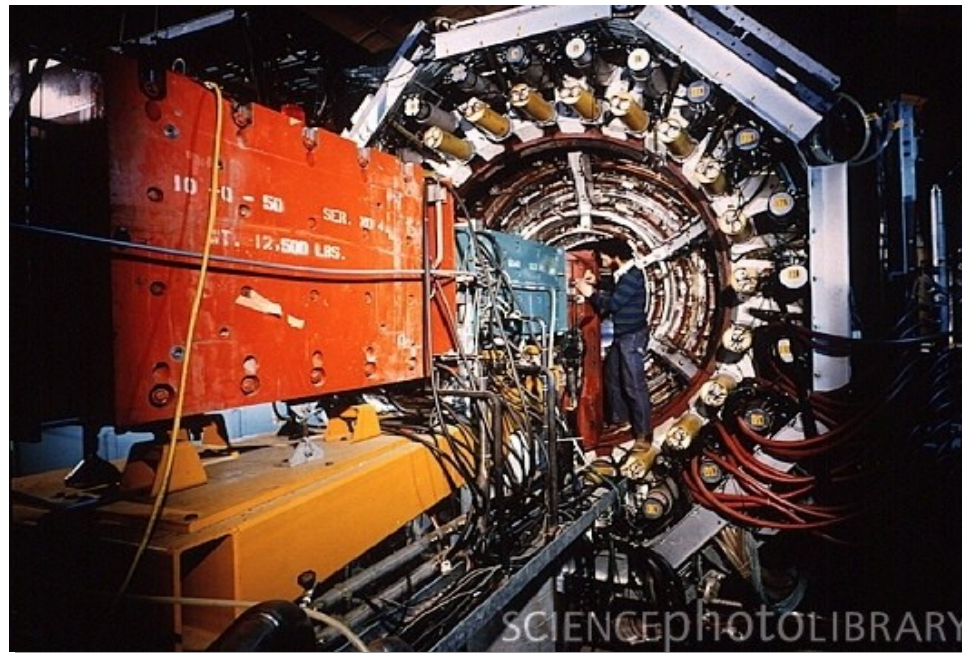


J/ψ (1974)

Experiment carried out by B. Richter group

– $e^+ e^-$ collider experiment

Contrarily to the S. Ting's group, B. Richter's group tried to find out a new particle by scanning the $e^+ e^-$ collision energy from 2.4 GeV by 0.2 GeV step



MARK-I detector at SLAC

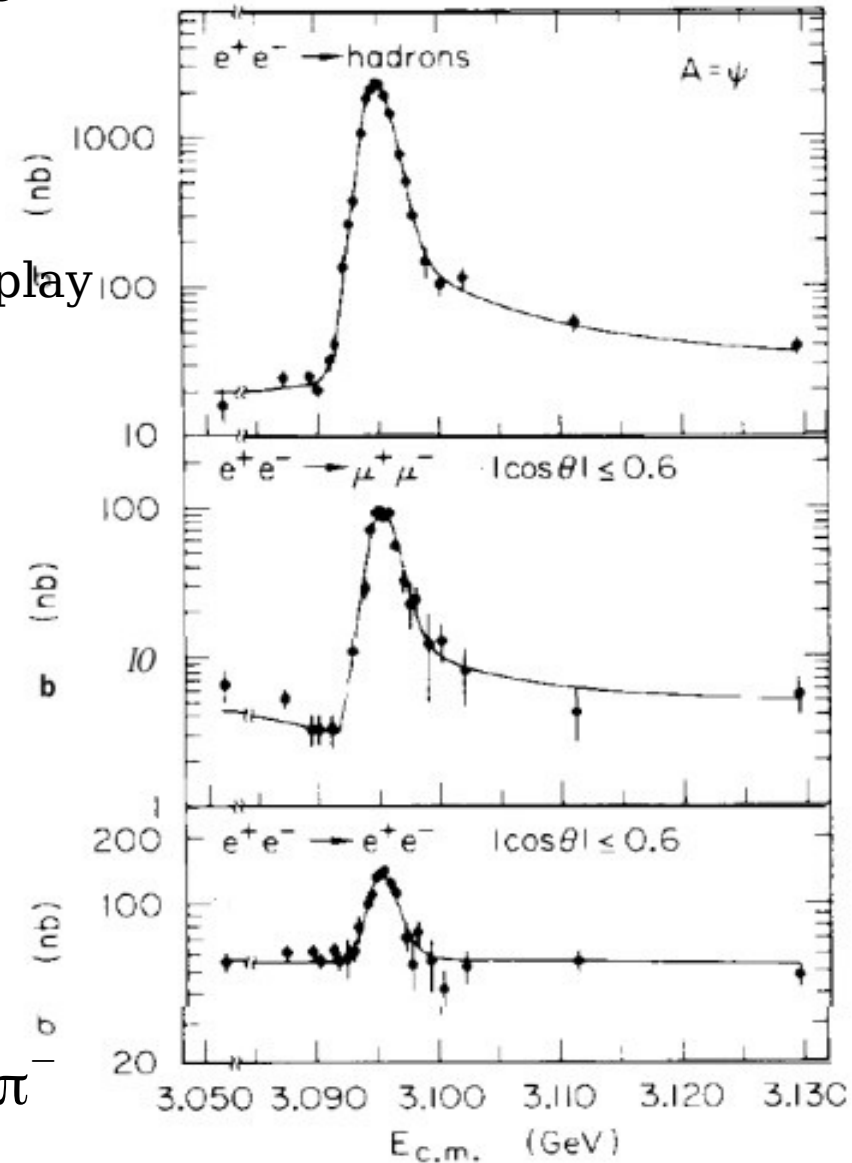
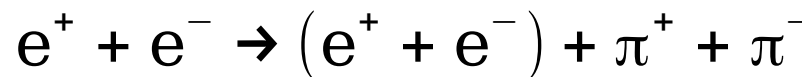
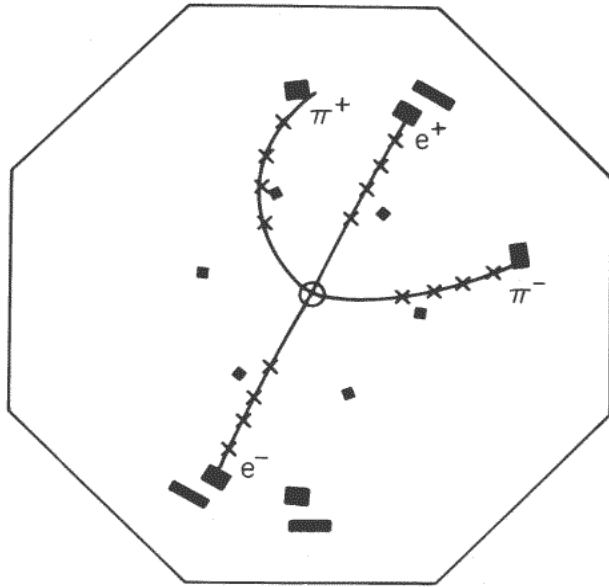
J/ψ (1974)

Experiment carried out by B. Richter group

– They observed a bump at 3 GeV/c²

– Event display

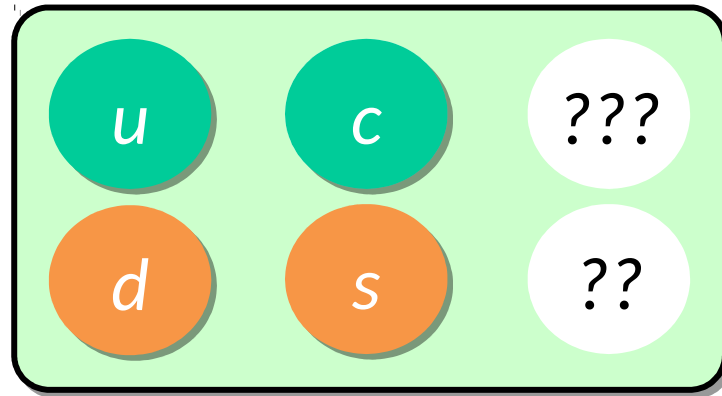
The particle name was taken from its event display



J/ψ (1974)

Discovery of the 4th quark

Finally, the J/ψ particles were identified as c c̄ mesons



Two names for the same particle

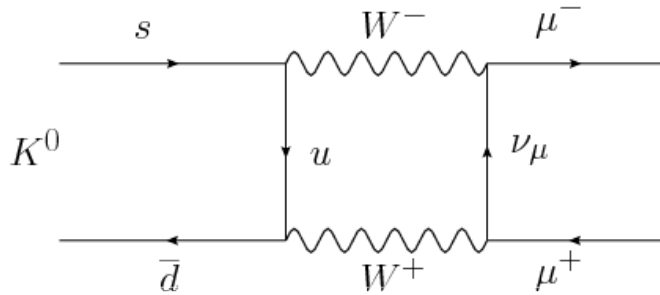
As both groups published the discoveries of J and ψ on the same day (11th Nov 1974), the particle was given 2 names: J/ψ

November revolution

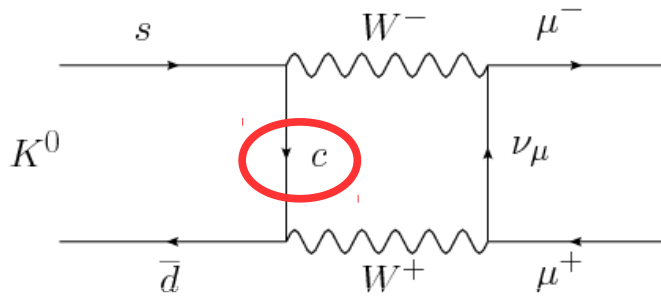
⇒ Nobel Prize 1976 rewarded Richter and Ting

indirect search: $K_L^0 \rightarrow \mu\mu$

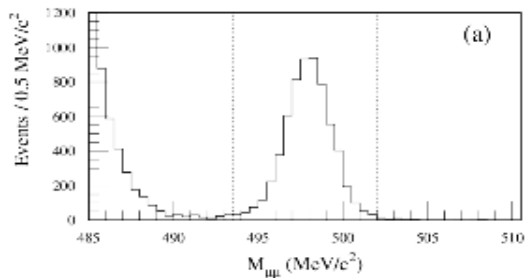
$K_L^0 \rightarrow \mu^+ \mu^-$ decay can be generated by the box diagram:



in a renormalisable gauge theory, is expected to give a branching ratio of $g^4 \sim \alpha^2 \sim 10^{-4}$, with α the fine structure constant.



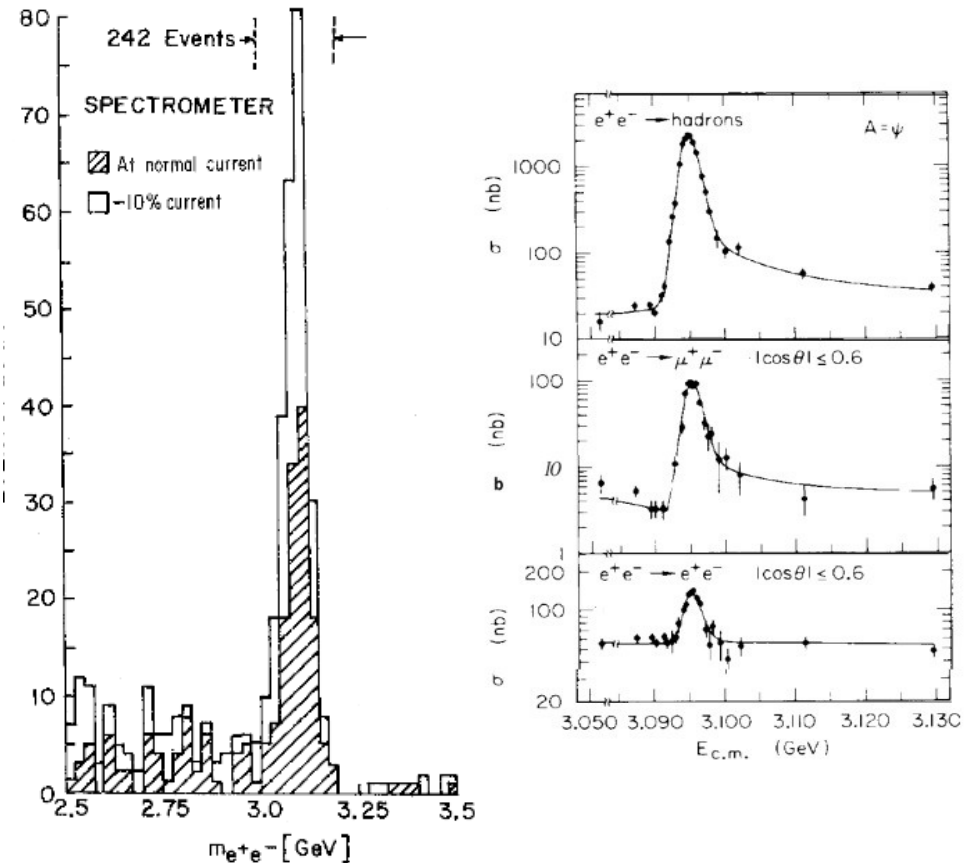
$K_L^0 \rightarrow \mu\mu$ was not observed though expected
 Now BF is measured to be $(6.84 \pm 0.11) 10^{-9}$ [Ambrose et al, 2000]



direct search: $J/\psi \rightarrow ee$

\rightarrow c quark eventually observed in 1974 [Ting], [Richter]

J/ ψ



With the measured charm quark mass $m_c \sim 1.27 \text{ GeV}$, the predicted rates are in agreement with observation.

W and Z bosons discoveries

The weak force is essentially as strong as the electromagnetic force, but it appears weak because its influence is limited by the large mass of the Z and W bosons. Their mass limits the range of the weak force to about 10^{-18} meters, and it vanishes altogether beyond the radius of a single proton.

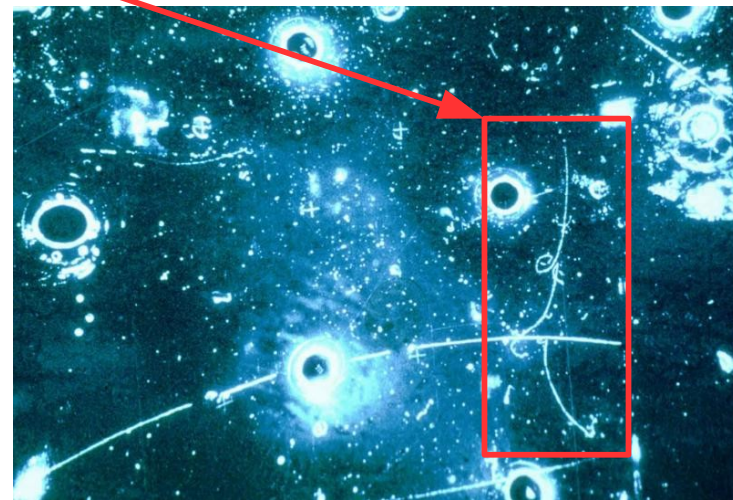
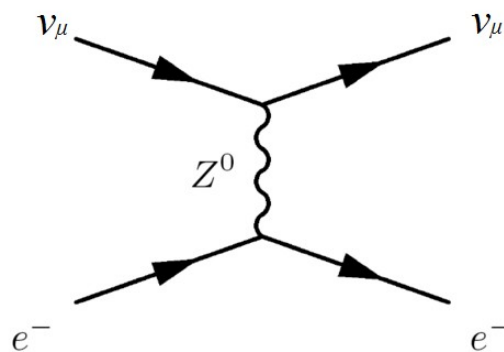
Sheldon Glashow, Abdus Salam and Steven Weinberg developed in the 1960s the theory in its present form, when they proposed that the weak and electromagnetic forces are actually different manifestations of one electroweak force.

First, in 1973, came the observation of neutral current interactions as predicted by electroweak theory at Gargamelle bubble chamber (Andre Lagarrigue et al)

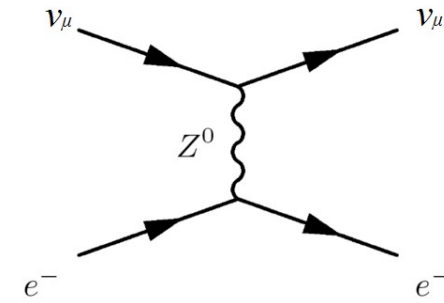
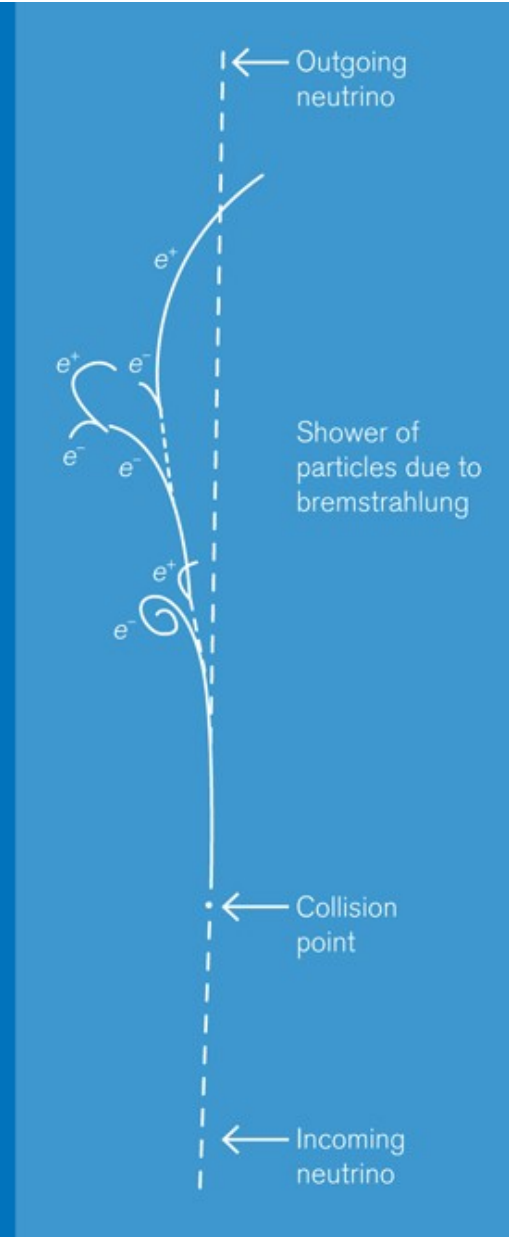
Neutrinos are particles that interact only via the weak interaction, and when the physicists shot neutrinos through the bubble chamber they were able to **detect evidence of the weak neutral current, and hence indirect evidence for the Z boson.**



neutrino beam



W and Z bosons discoveries



W and Z bosons discoveries

Super Proton Synchrotron, proton-antiproton collider, where unambiguous signals of W bosons were seen in January 1983 during a series of experiments made possible by Carlo Rubbia and Simon van der Meer. Experiments are UA1 and UA2.

270 GeV per beam, enough energy to produce W and Z particles
first general purpose 4π experiment in high energy physics



the **Central Detector**, a big drift chamber to track charged particles in the 0.7 T field of the dipole magnet

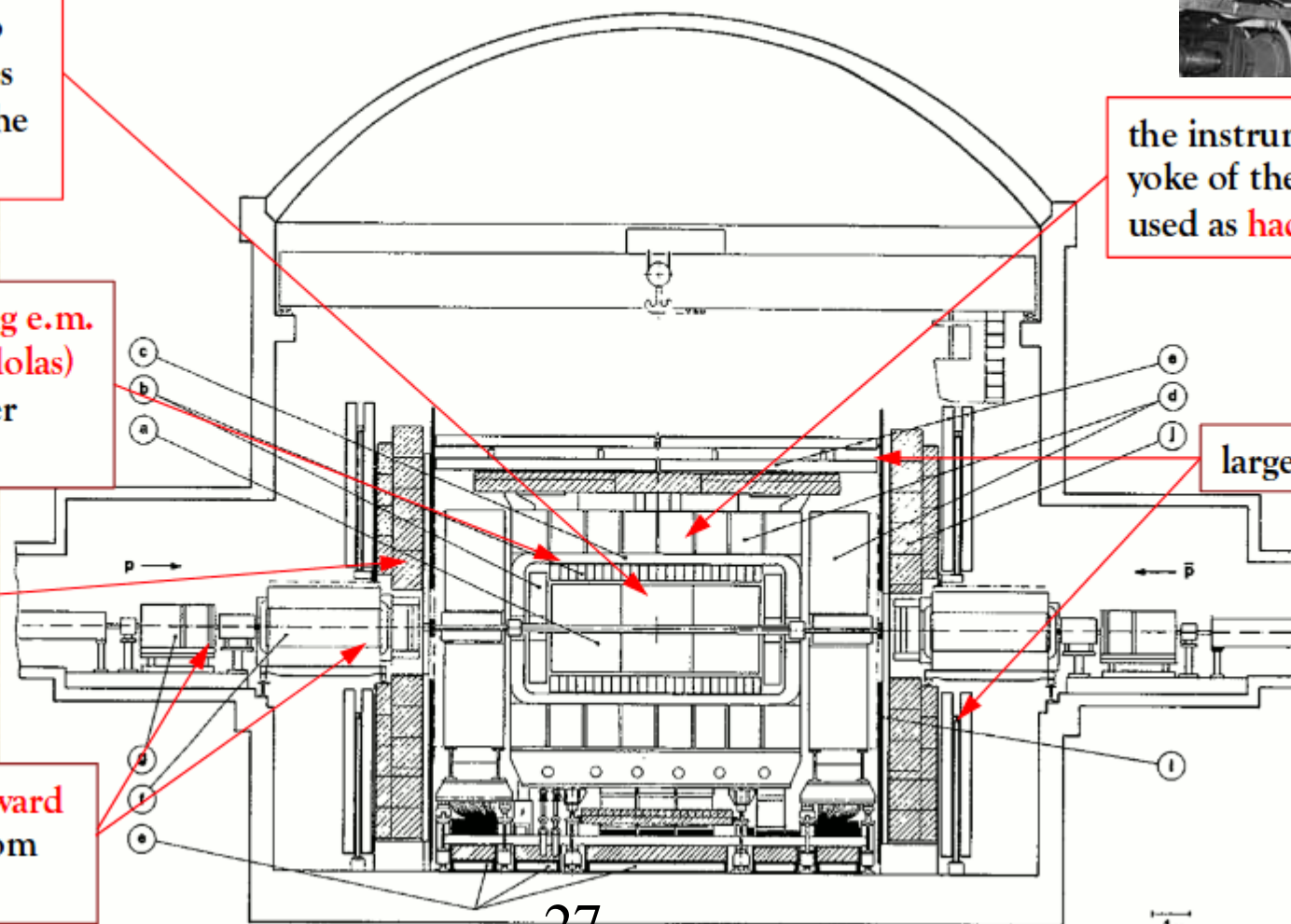
the instrumented return yoke of the dipole magnet used as **hadron calorimeter**

the fine grain **sampling e.m. calorimeter (the Gondolas)** readout by wave shifter bars (BBQ)

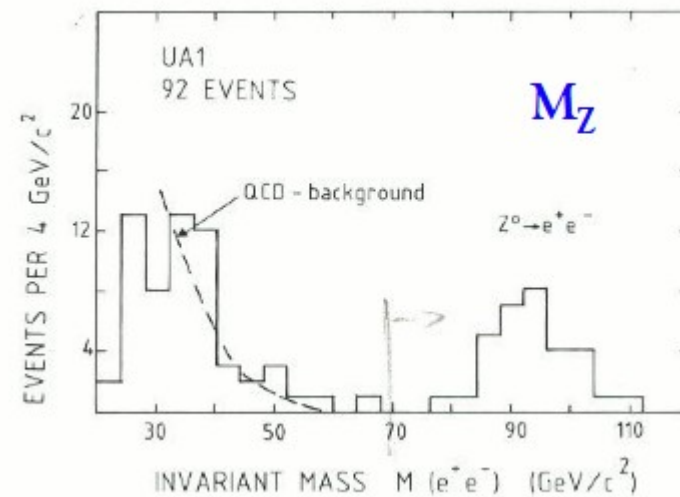
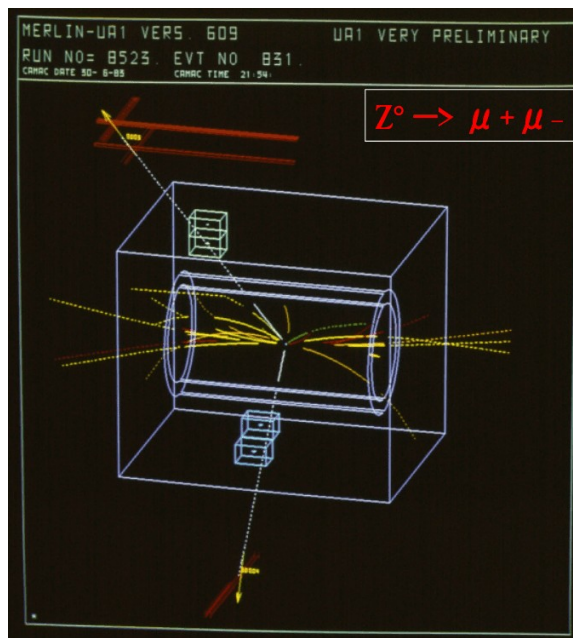
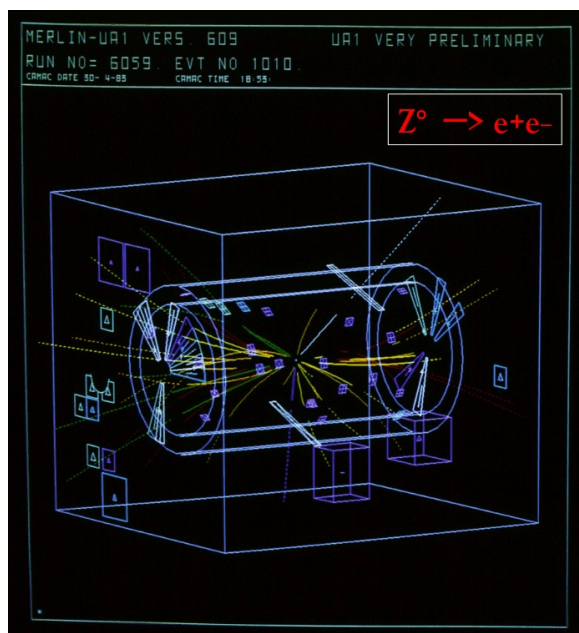
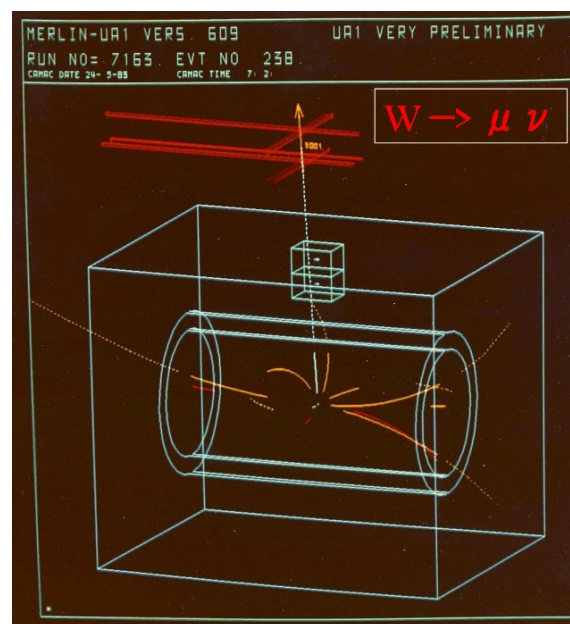
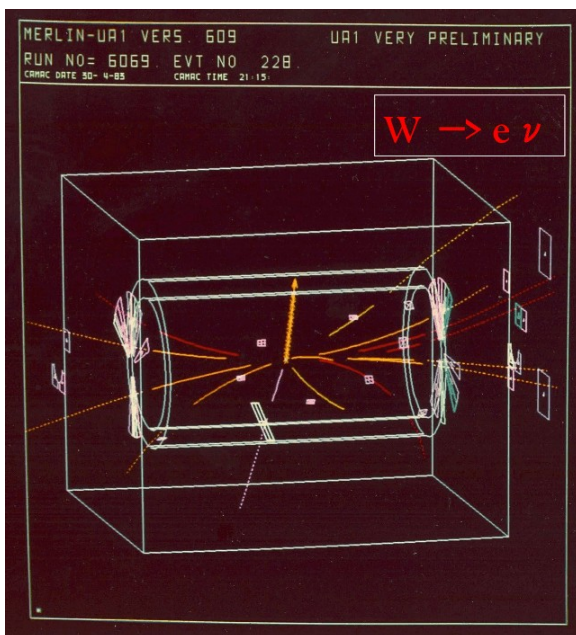
large muon chambers

end cap detectors (the bouchons)

forward and very forward detectors (up to 1° from the beams)



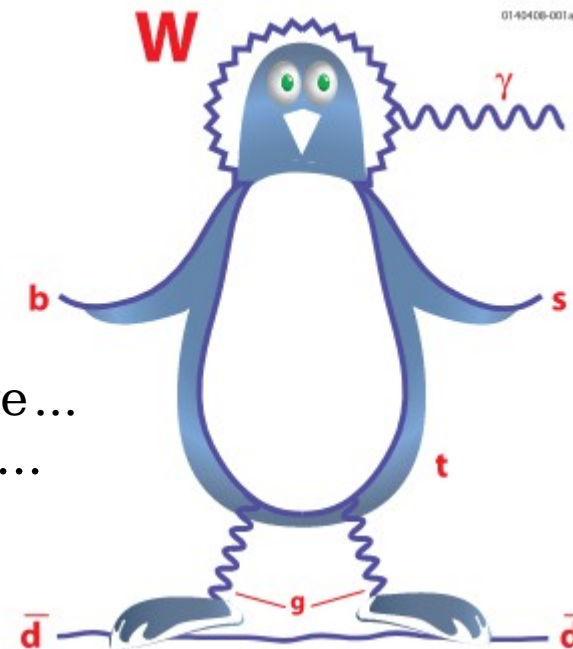
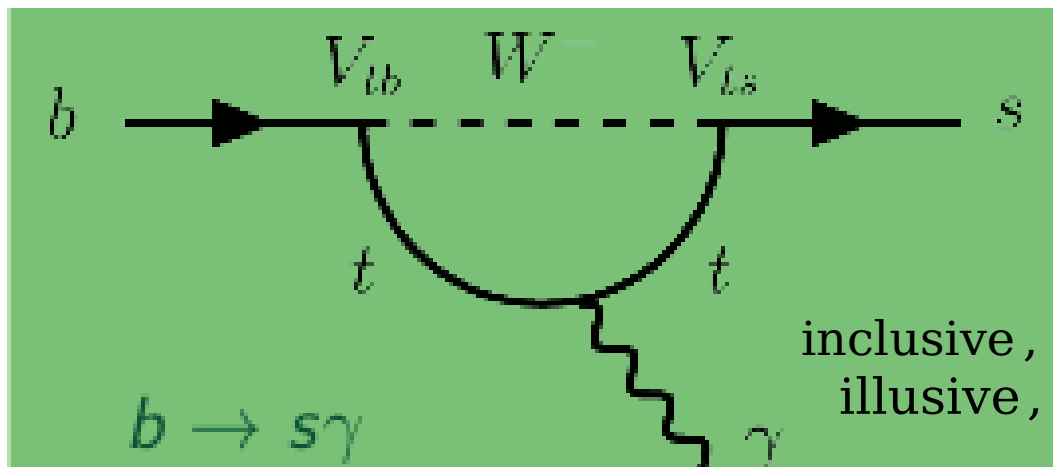
W and Z bosons discoveries



Rubbia and van der Meer were promptly awarded the 1984 Nobel Prize in Physics.

Radiative B decays

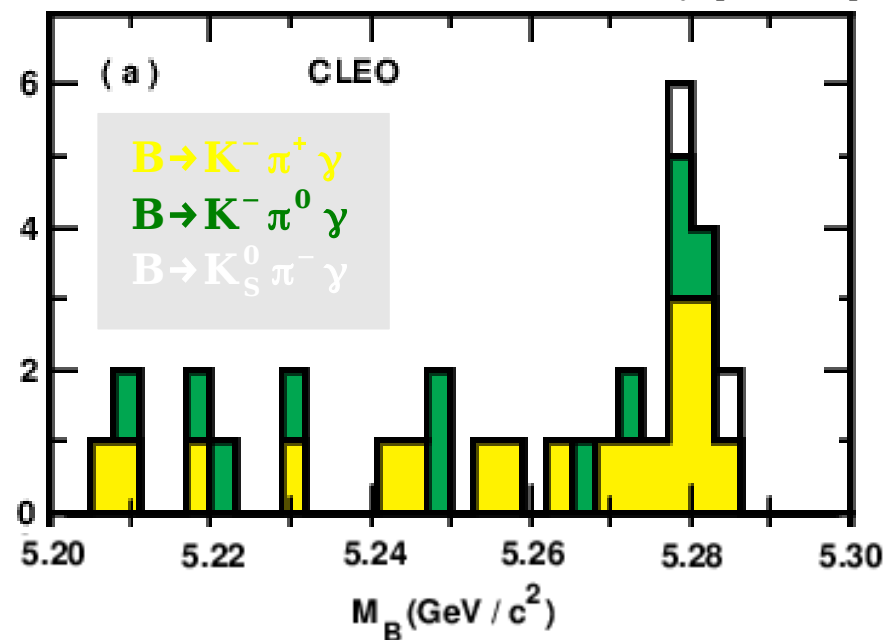
artist's view ... of the penguin diagram



CLEO

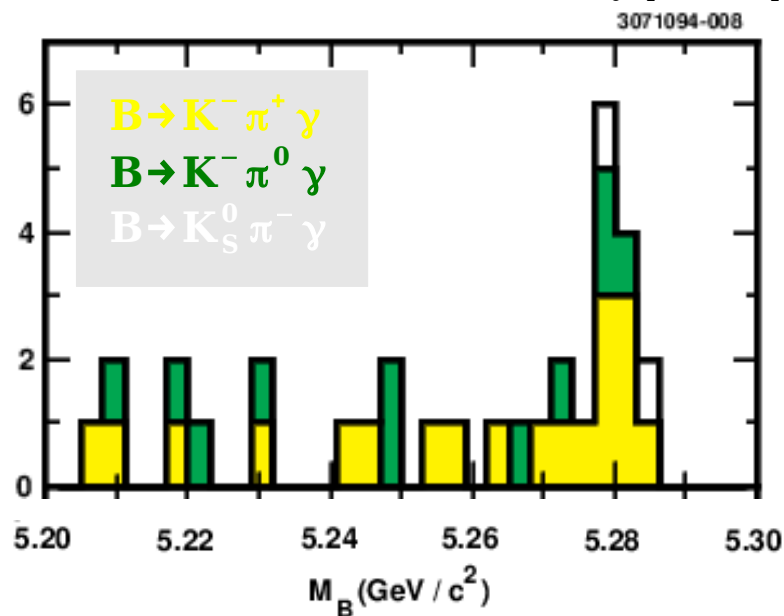
- 1975: "South Area Experiment" group conceives CLEO
- 1979: First data collected
- 1980: B meson discovered
- 1983: Ds meson discovered
- 1986: CLEO II detector with CsI calorimeter installed
- 1989: $b \rightarrow u$ transitions discovered
- 1993: $b \rightarrow s$ penguin decays discovered
- 1995: CLEO II.V with silicon vertex detector installed
- 1999: CLEO III with RICH installed
- 2003: CLEO-c data collection started
- 2004: hc discovered and D^+ meson decay constant measured
- 2008: Running ends on March 3rd
- 2009: 500th paper published

CLEO observation of $B \rightarrow K^* \gamma$ [1993]



$B \rightarrow K^* \gamma$ measurements

CLEO observation of $B \rightarrow K^* \gamma$ [1993]



$B \rightarrow K^* \gamma$

means $B^+ \rightarrow K^{*+} \gamma$ or $B^0 \rightarrow K^{*+} \gamma$

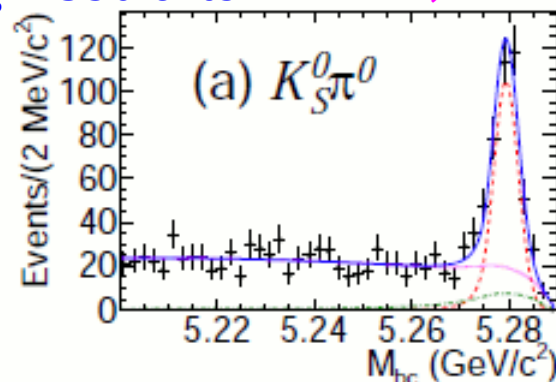
$$\text{Br}(B^0 \rightarrow K^{*0} \gamma) = (41.7 \pm 1.2) \times 10^{-6}$$

$$\text{Br}(K^{*0} \rightarrow K^+ \pi^-) \sim 2/3, \text{Br}(K^{*0} \rightarrow K^0 \pi^0) \sim 1/3$$

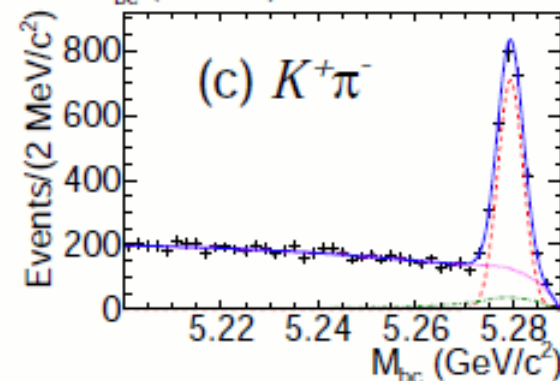
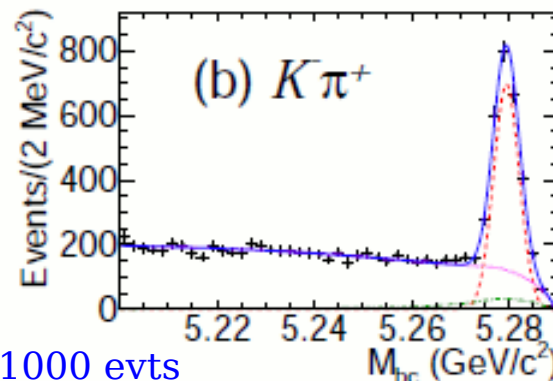
$$\text{Br}(B^+ \rightarrow K^{*+} \gamma) = (39.2 \pm 1.3) \times 10^{-6}$$

$$\text{Br}(K^{*+} \rightarrow K^0 \pi^+) \sim 2/3, \text{Br}(K^{*+} \rightarrow K^+ \pi^0) \sim 1/3$$

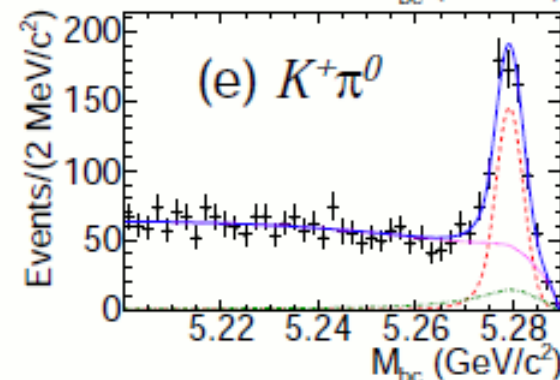
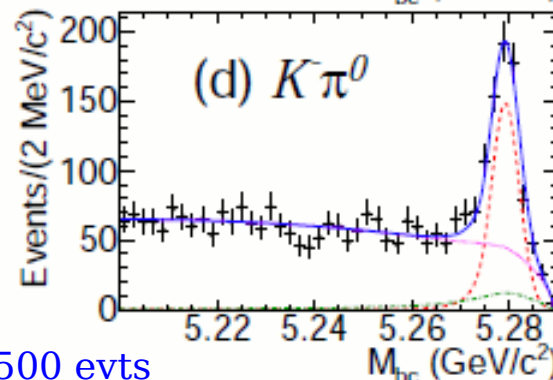
$N_s \sim 350$ evts Belle, submitted to PRL



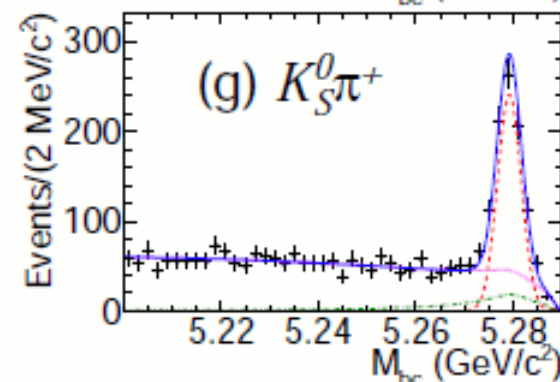
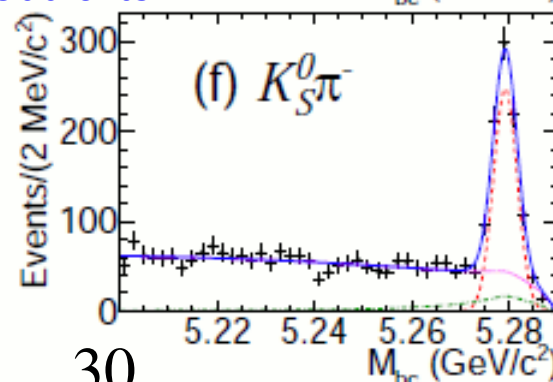
$N_s \sim 4500$ evts



$N_s \sim 1000$ evts



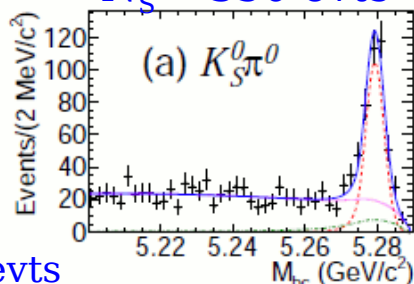
$N_s \sim 1500$ evts



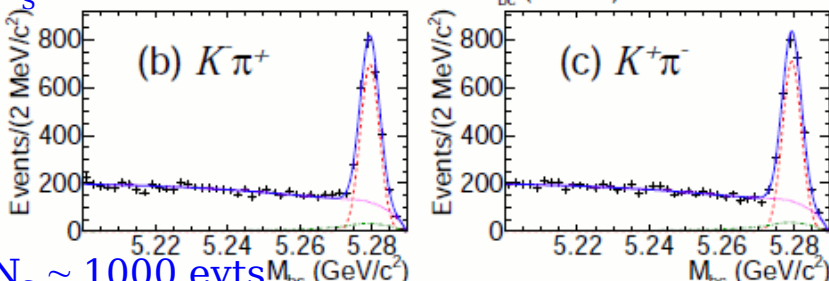
$B \rightarrow K^* \gamma$ measurements

simultaneous fit of 4 final states
 \Rightarrow extraction of BF's....

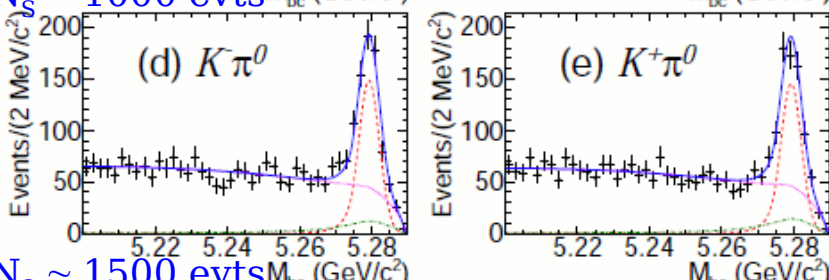
$N_s \sim 350$ evts



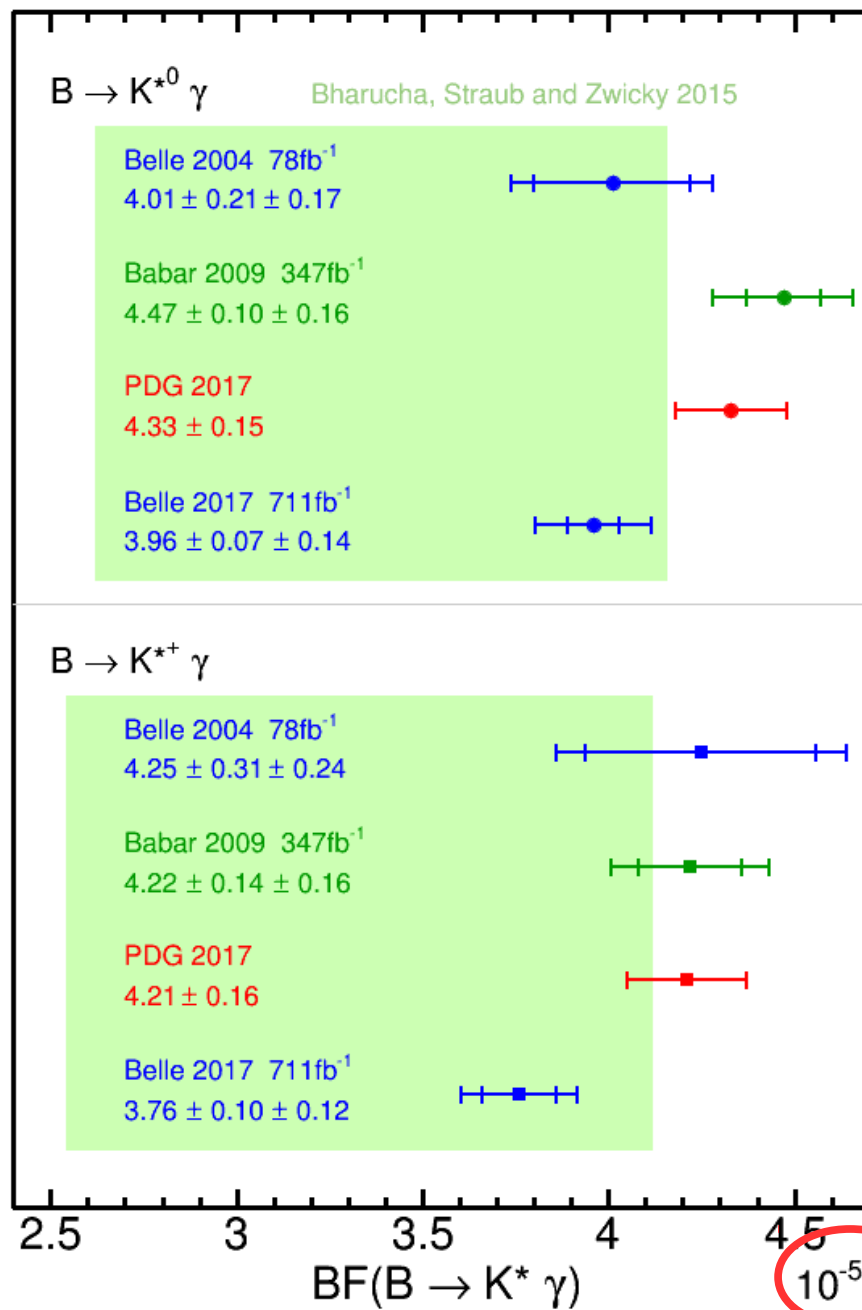
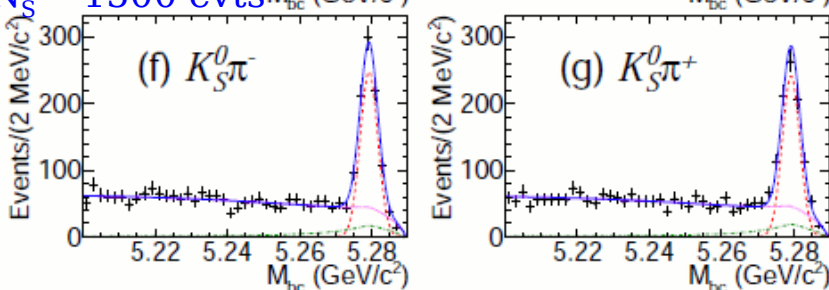
$N_s \sim 4500$ evts



$N_s \sim 1000$ evts



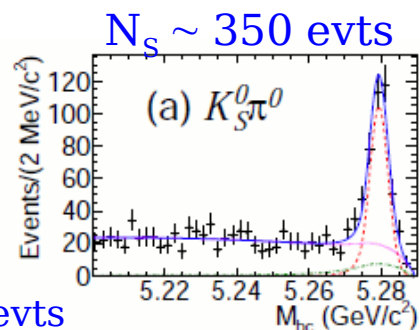
$N_s \sim 1500$ evts



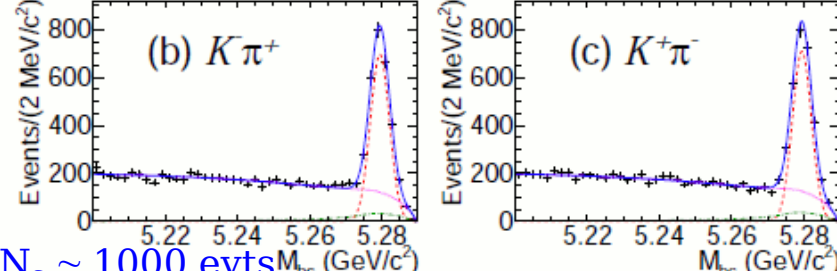
but uncertainty in the hadronization process limits the ability to predict individual exclusive rates from first principles of the theory

$B \rightarrow K^* \gamma$ measurements

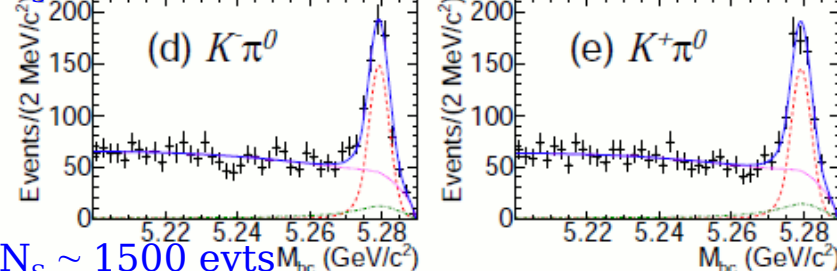
simultaneous fit of 4 final states
 \Rightarrow extraction of BFs, Δ_{0+} , A_{CP} , ΔA_{CP} ...



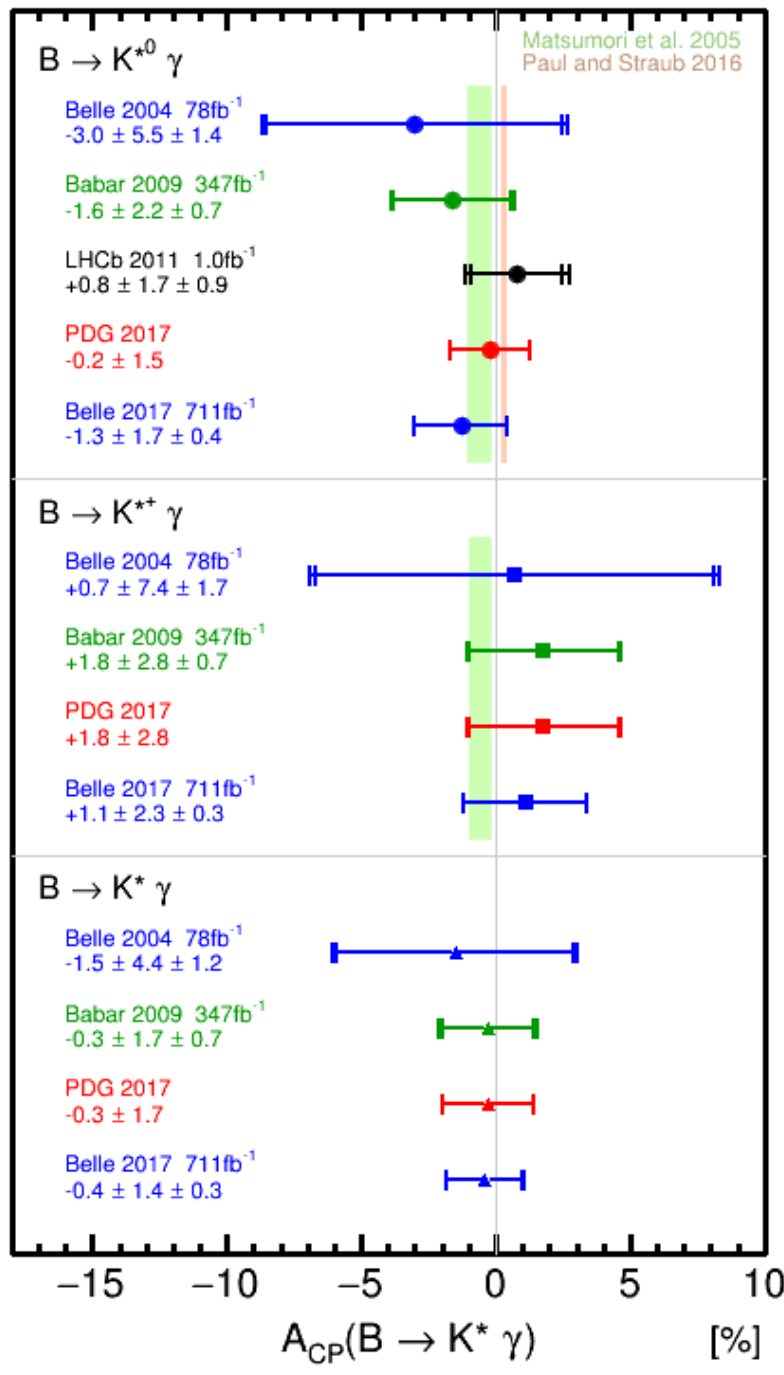
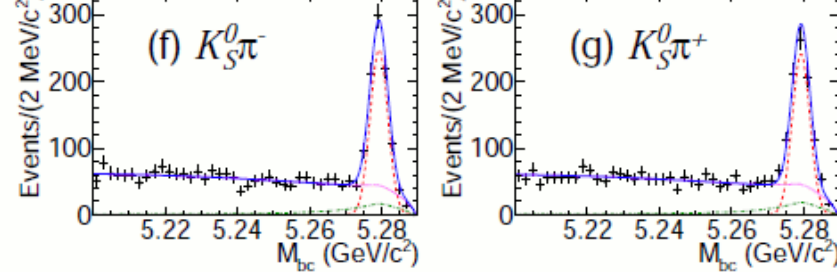
$N_s \sim 4500$ evts



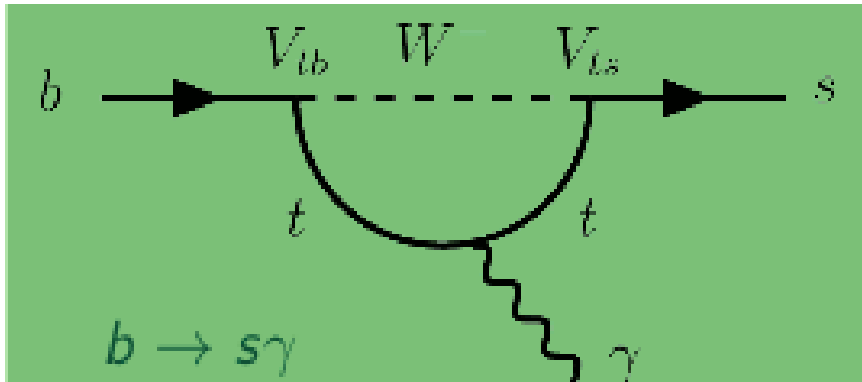
$N_s \sim 1000$ evts



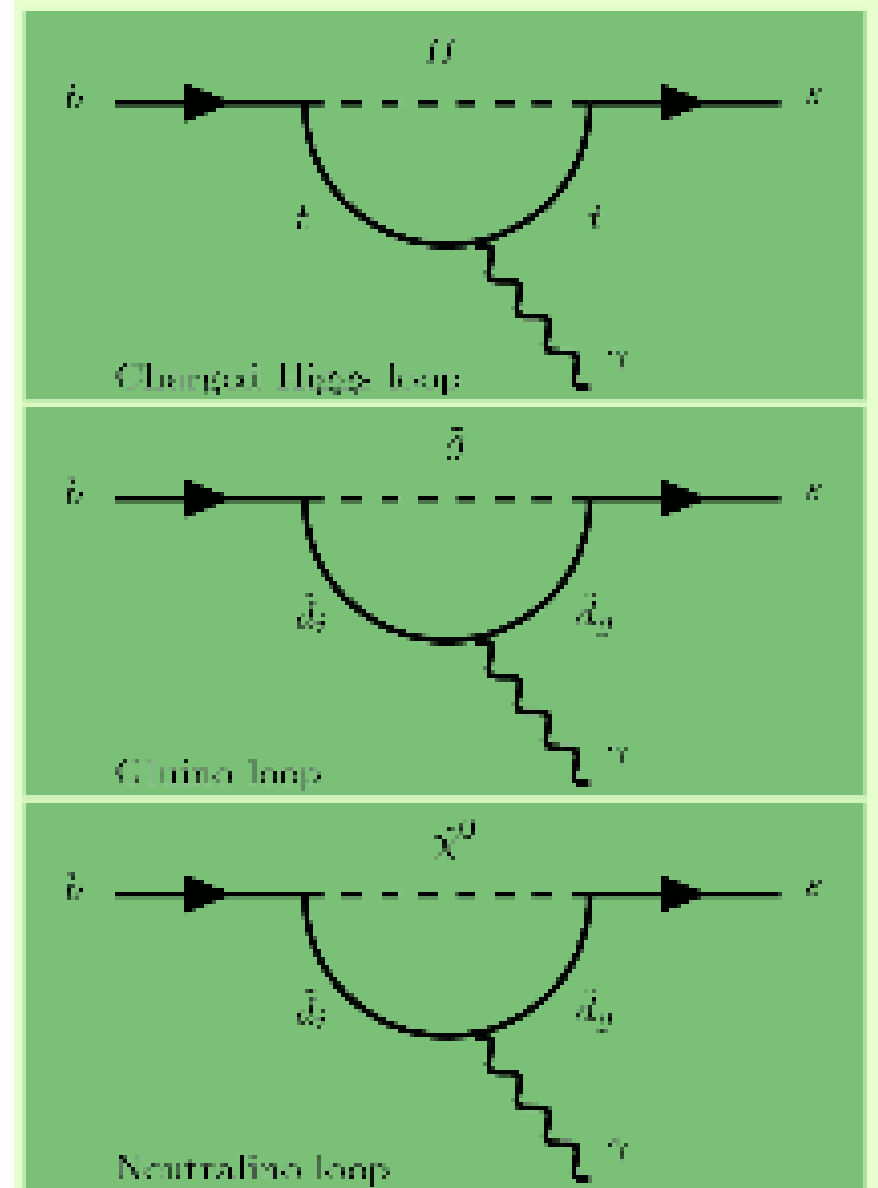
$N_s \sim 1500$ evts



$b \rightarrow s \gamma$



- Amplitude $\propto V_{ts} |C_7|$
- First penguin ever observed (93)
- Experiment:
 - $B \simeq 3 \cdot 10^{-4}$
- SM: $B = (3.36 \pm 0.23) \cdot 10^{-4}$
 - [Misiak et al., hep-ph/0609232]
 - \Rightarrow [Misiak et al, arXiv:1503.01789]
- Strong constraint on New Physics



exclusive vs inclusive

easier experimentally
difficult theoretically

difficult experimentally
easier theoretically (closer to $b \rightarrow s \gamma$)

$b \rightarrow s \gamma$ SM branching fraction

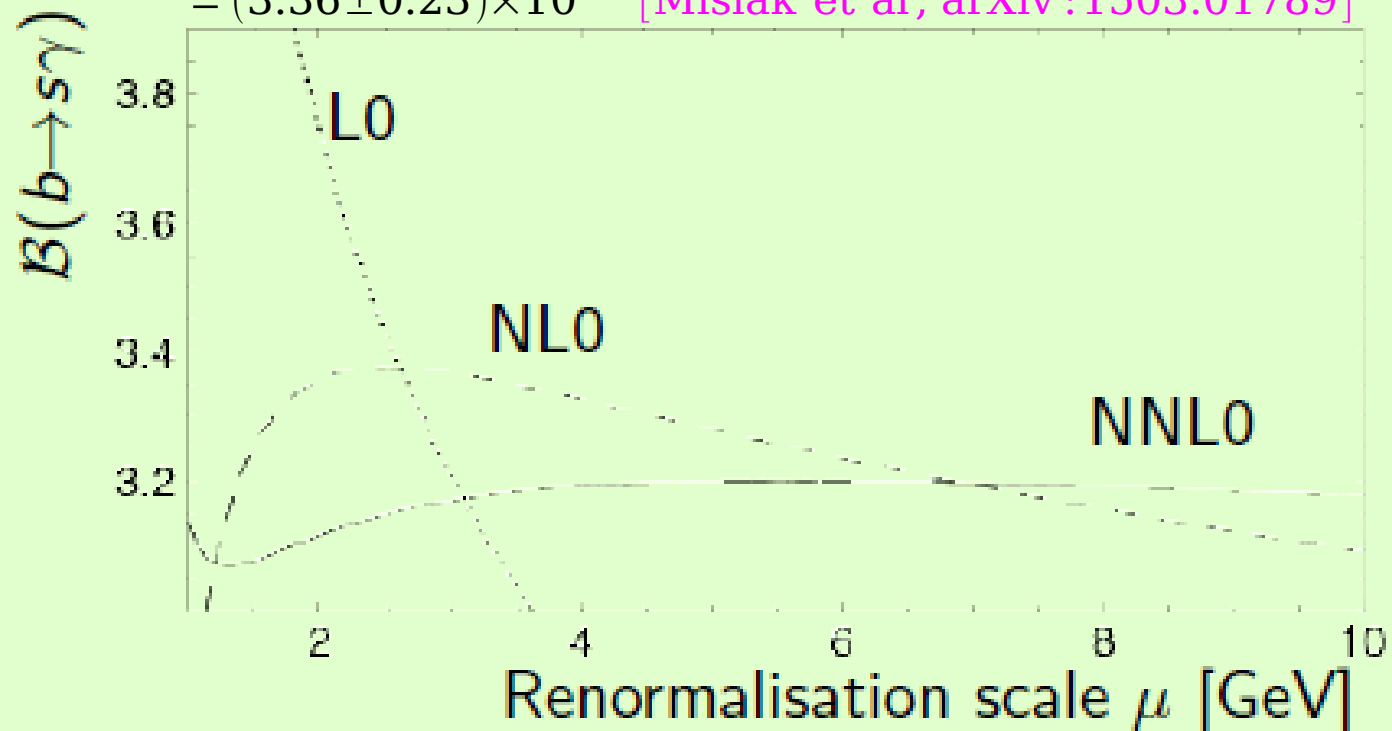
[Misiak et al, PRL 98, 02202, 2007]

- From effective Hamiltonian one gets the BF
- Uncertainties due to m_b and m_c : normalise to $b \rightarrow ce\nu$ and $b \rightarrow ue\nu$ [Misiak & Steinhauser, NPB764:62,2007]
- $b \rightarrow s\gamma$ branching fraction calculated at all NNLO orders in 2006

$$\mathcal{B}(B \rightarrow X_s \gamma)_{E_\gamma > 1.6 \text{ GeV}} = \overline{(3.15 \pm 0.23) \cdot 10^{-4}}$$

$$= (3.36 \pm 0.23) \times 10^{-4} \quad [\text{Misiak et al, arXiv:1503.01789}]$$

✓ BF very stable
versus μ



How to estimate the branching fraction $b \rightarrow s \gamma$?

Semi-inclusive (sum-of-exclusive)



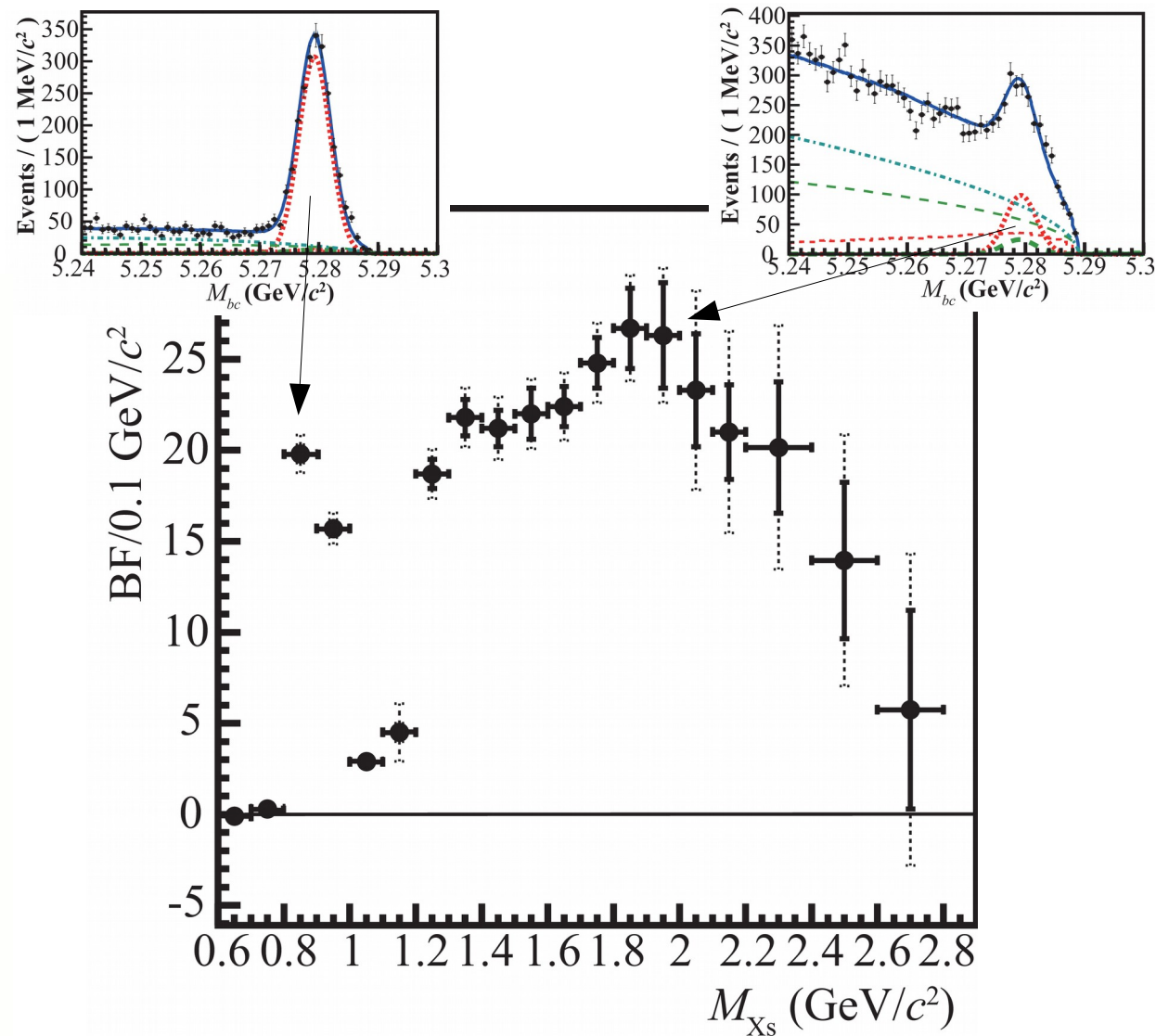
[772 MBB]

[arXiv:1411.7198]

38 modes

$M_{X_s} < 2.8 \text{ GeV}/c^2$, $E^* > 1.9 \text{ GeV}$

Mode ID	Final State	Mode ID	Final State
1	$K^+ \pi^-$	20	$K_S^0 \pi^+ \pi^0 \pi^0$
2	$K_S^0 \pi^+$	21	$K^+ \pi^+ \pi^- \pi^0 \pi^0$
3	$K^+ \pi^0$	22	$K_S^0 \pi^+ \pi^- \pi^0 \pi^0$
4	$K_S^0 \pi^0$	23	$K^+ \eta$
5	$K^+ \pi^+ \pi^-$	24	$K_S^0 \eta$
6	$K_S^0 \pi^+ \pi^-$	25	$K^+ \eta \pi^-$
7	$K^+ \pi^- \pi^0$	26	$K_S^0 \eta \pi^+$
8	$K_S^0 \pi^+ \pi^0$	27	$K^+ \eta \pi^0$
9	$K^+ \pi^+ \pi^- \pi^-$	28	$K_S^0 \eta \pi^0$
10	$K_S^0 \pi^+ \pi^+ \pi^-$	29	$K^+ \eta \pi^+ \pi^-$
11	$K^+ \pi^+ \pi^- \pi^0$	30	$K_S^0 \eta \pi^+ \pi^-$
12	$K_S^0 \pi^+ \pi^- \pi^0$	31	$K^+ \eta \pi^- \pi^0$
13	$K^+ \pi^+ \pi^+ \pi^- \pi^-$	32	$K_S^0 \eta \pi^+ \pi^0$
14	$K_S^0 \pi^+ \pi^+ \pi^- \pi^-$	33	$K^+ K^+ K^-$
15	$K^+ \pi^+ \pi^- \pi^- \pi^0$	34	$K^+ K^- K_S^0$
16	$K_S^0 \pi^+ \pi^+ \pi^- \pi^0$	35	$K^+ K^+ K^- \pi^-$
17	$K^+ \pi^0 \pi^0$	36	$K^+ K^- K_S^0 \pi^+$
18	$K_S^0 \pi^0 \pi^0$	37	$K^+ K^+ K^- \pi^0$
19	$K^+ \pi^- \pi^0 \pi^0$	38	$K^+ K^- K_S^0 \pi^0$



Semi-inclusive (sum-of-exclusive)

38 modes

$M_{X_s} < 2.8 \text{ GeV}/c^2$, $E^* > 1.9 \text{ GeV}$

**possible but large systematics
(difficult to estimate/trust)**

Mode ID	Final State	Mode ID	Final State
1	$K^+\pi^-$	20	$K_S^0\pi^+\pi^0\pi^0$
2	$K_S^0\pi^+$	21	$K^+\pi^+\pi^-\pi^0\pi^0$
3	$K^+\pi^0$	22	$K_S^0\pi^+\pi^-\pi^0\pi^0$
4	$K_S^0\pi^0$	23	$K^+\eta$
5	$K^+\pi^+\pi^-$	24	$K_S^0\eta$
6	$K_S^0\pi^+\pi^-$	25	$K^+\eta\pi^-$
7	$K^+\pi^-\pi^0$	26	$K_S^0\eta\pi^+$
8	$K_S^0\pi^+\pi^0$	27	$K^+\eta\pi^0$
9	$K^+\pi^+\pi^-\pi^-$	28	$K_S^0\eta\pi^0$
10	$K_S^0\pi^+\pi^+\pi^-$	29	$K^+\eta\pi^+\pi^-$
11	$K^+\pi^+\pi^-\pi^0$	30	$K_S^0\eta\pi^+\pi^-$
12	$K_S^0\pi^+\pi^-\pi^0$	31	$K^+\eta\pi^-\pi^0$
13	$K^+\pi^+\pi^+\pi^-\pi^-$	32	$K_S^0\eta\pi^+\pi^0$
14	$K_S^0\pi^+\pi^+\pi^-\pi^-$	33	$K^+K^+K^-$
15	$K^+\pi^+\pi^-\pi^-\pi^0$	34	$K^+K^-K_S^0$
16	$K_S^0\pi^+\pi^+\pi^-\pi^0$	35	$K^+K^+K^-\pi^-$
17	$K^+\pi^0\pi^0$	36	$K^+K^-K_S^0\pi^+$
18	$K_S^0\pi^0\pi^0$	37	$K^+K^+K^-\pi^0$
19	$K^+\pi^-\pi^0\pi^0$	38	$K^+K^-K_S^0\pi^0$

Mode Category	Definition	Mode ID	Data
1	$K\pi$ without π^0	1,2	4.2 ± 0.4
2	$K\pi$ with π^0	3,4	2.1 ± 0.2
3	$K2\pi$ without π^0	5,6	14.5 ± 0.5
4	$K2\pi$ with π^0	7,8	24.0 ± 0.7
5	$K3\pi$ without π^0	9,10	8.3 ± 0.8
6	$K3\pi$ with π^0	11,12	16.1 ± 1.8
7	$K4\pi$	13-16	11.1 ± 2.8
8	$K2\pi^0$	17-22	14.4 ± 3.5
9	$K\eta$	23-32	3.2 ± 0.8
10	$3K$	33-38	2.0 ± 0.3



[772 MBB]

[arXiv: 1411.7198]

(for $E_y^* > 1.9 \text{ GeV}$)



[471 MBB]

[arXiv: 1207.2520]

$$\left\{ \begin{array}{l} B(B \rightarrow X_s \gamma) = (3.51 \pm 0.17 \pm 0.33) \times 10^{-4} \\ B(B \rightarrow X_s \gamma) = (3.29 \pm 0.19 \pm 0.48) \times 10^{-4} \end{array} \right.$$

36

[syst: cross-feed, peaking BG, X_s fragmentation]

$B \rightarrow X_s \gamma$ spectrum

- $b \rightarrow s \gamma$ is a 2-body decay. The energy of the photon in the b quark frame is

$$E_\gamma = \frac{m_b}{2} \left(1 - \frac{m_s^2}{m_b^2} \right) \simeq \frac{m_b}{2}$$

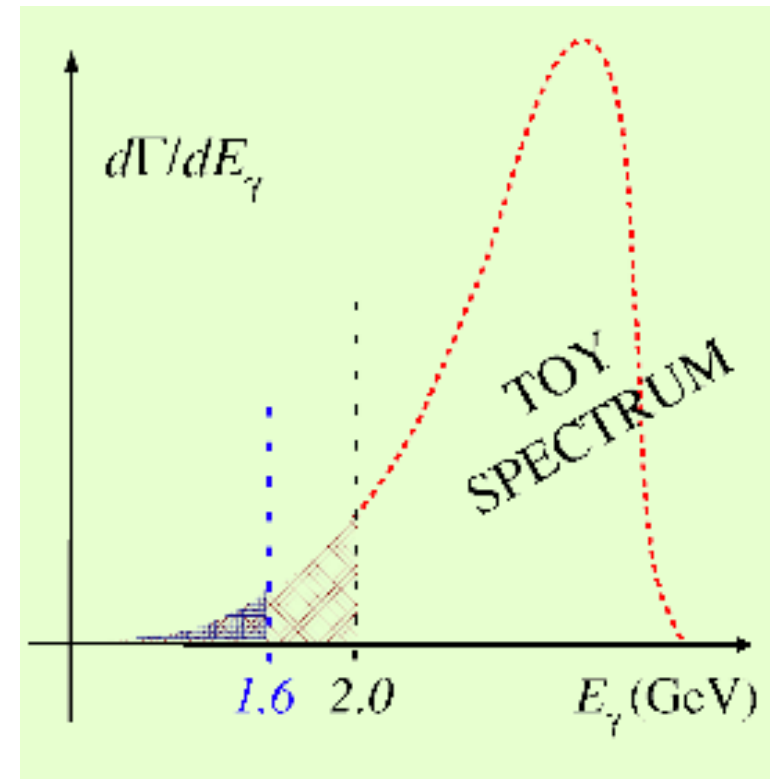
- But we measure $B \rightarrow X_s \gamma$ and in the B meson, the b quark is moving which smears the energy spectrum

→ Mean $\sim \frac{m_B}{2}$

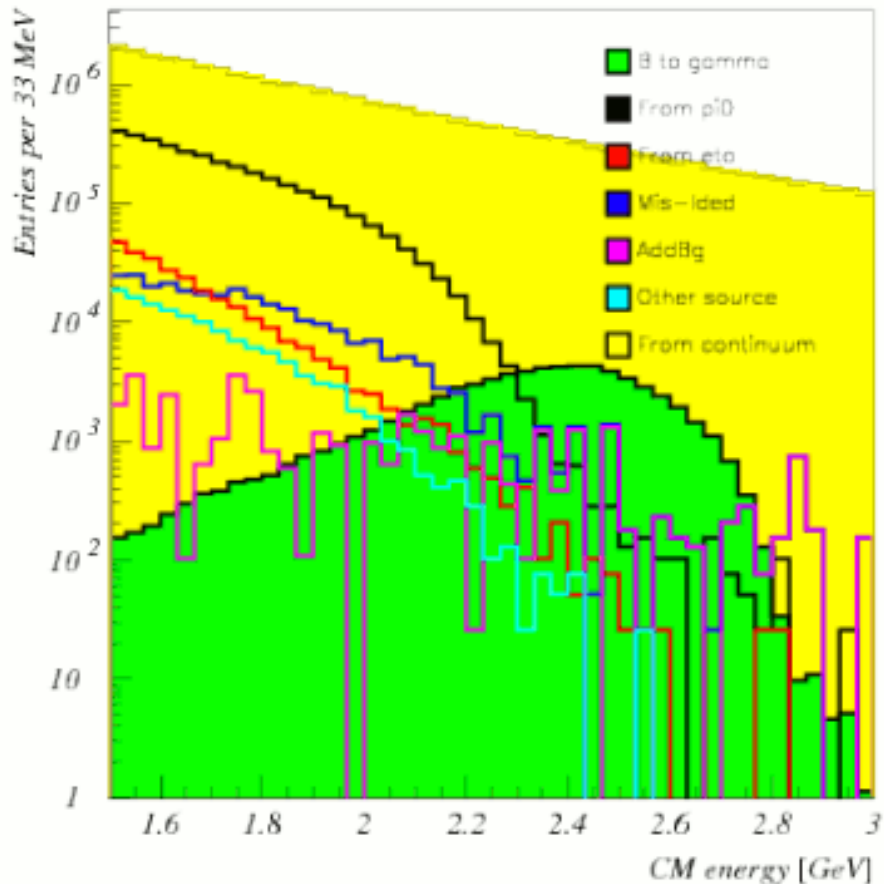
→ Width \sim Fermi motion in B meson

- The BF is calculated for some energy cutoff (1.6 GeV). For other cutoffs E_0 apply [Misiak et al, (2007)]

$$\left(\frac{B(E_\gamma > E_0)}{B(E_\gamma > 1.6 \text{ GeV})} \right) \simeq 1 + 0.15 \frac{E_0}{1.6 \text{ GeV}} - 0.14 \left(\frac{E_0}{1.6 \text{ GeV}} \right)^2$$



$b \rightarrow s \gamma$ spectrum at Belle



One would like to measure the photon energy spectrum in $b \rightarrow s \gamma$ decays

- Be unbiased: only look at the γ
- B mesons only decay to γ via $b \rightarrow s \gamma$
- But there are indirect γ from π^0 and η in $B\bar{B}$ events
- ...and a lot more indirect π^0 and η in non- $B\bar{B}$ events

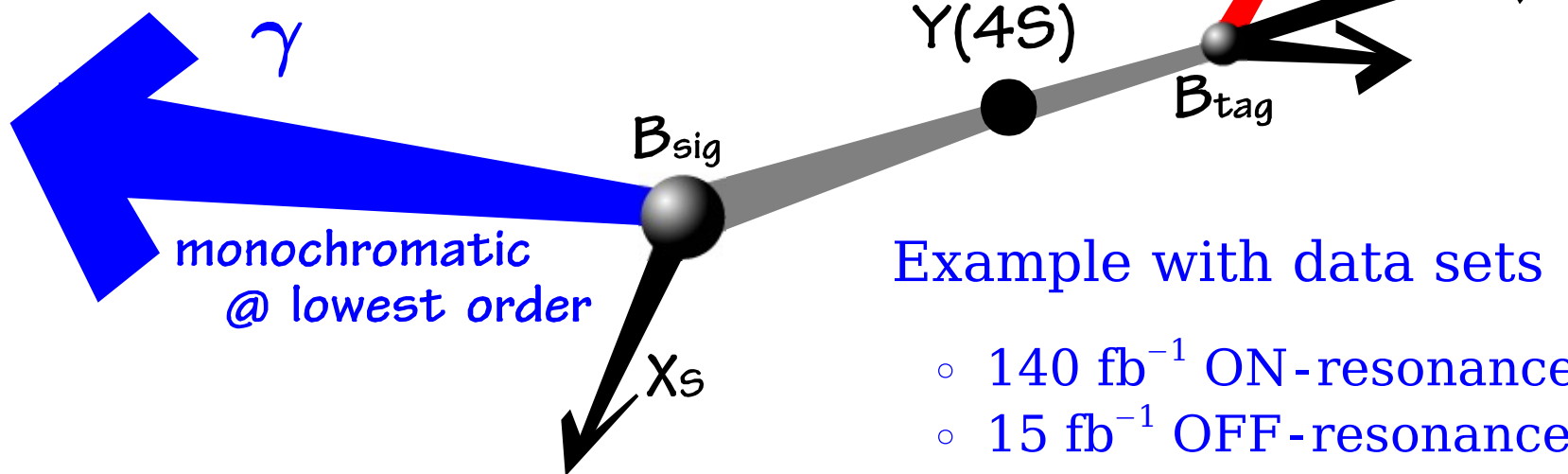
⇒ Lots of background at low energy

$b \rightarrow s \gamma$ spectrum at Belle

inclusive $B \rightarrow X_s \gamma$ measurement

untagged

lepton tag: background suppression, low stat



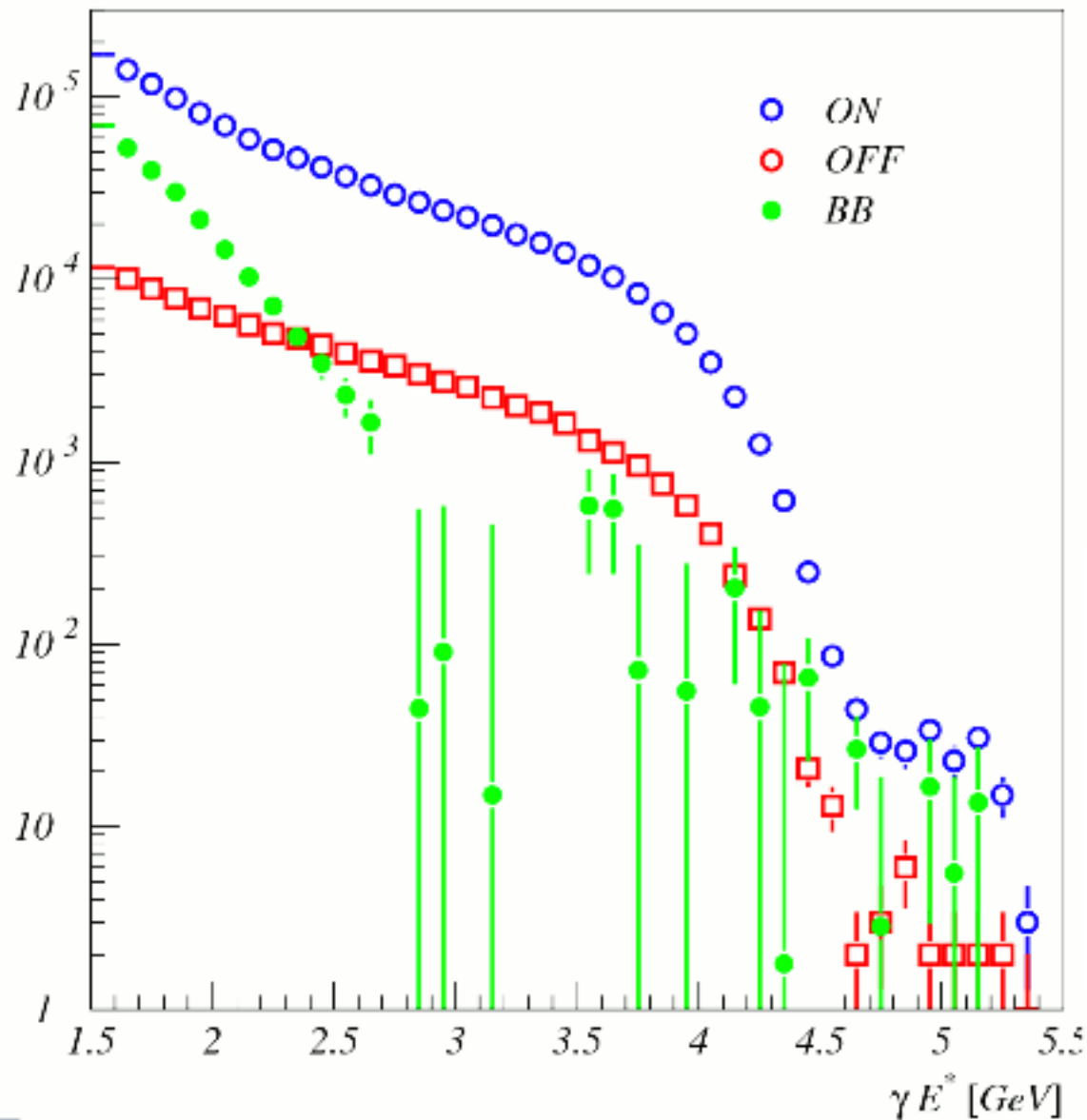
Example with data sets

- 140 fb^{-1} ON-resonance
- 15 fb^{-1} OFF-resonance

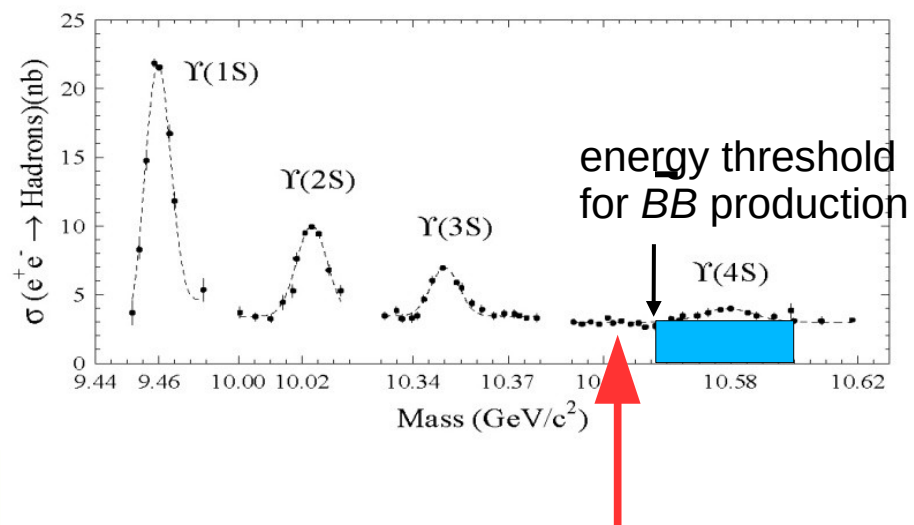
Event selection:

- No kinematic constraints
- Only a high energy photon measured in $\Upsilon(4S)$ rest frame
- Lower E_γ threshold (1.7 GeV)
- Hadronic events with isolated photon(s) in ECL. $E^* > 1.5 \text{ GeV}$.
- Veto γ from π^0 and η
- Apply event shape cuts to suppress continuum background.

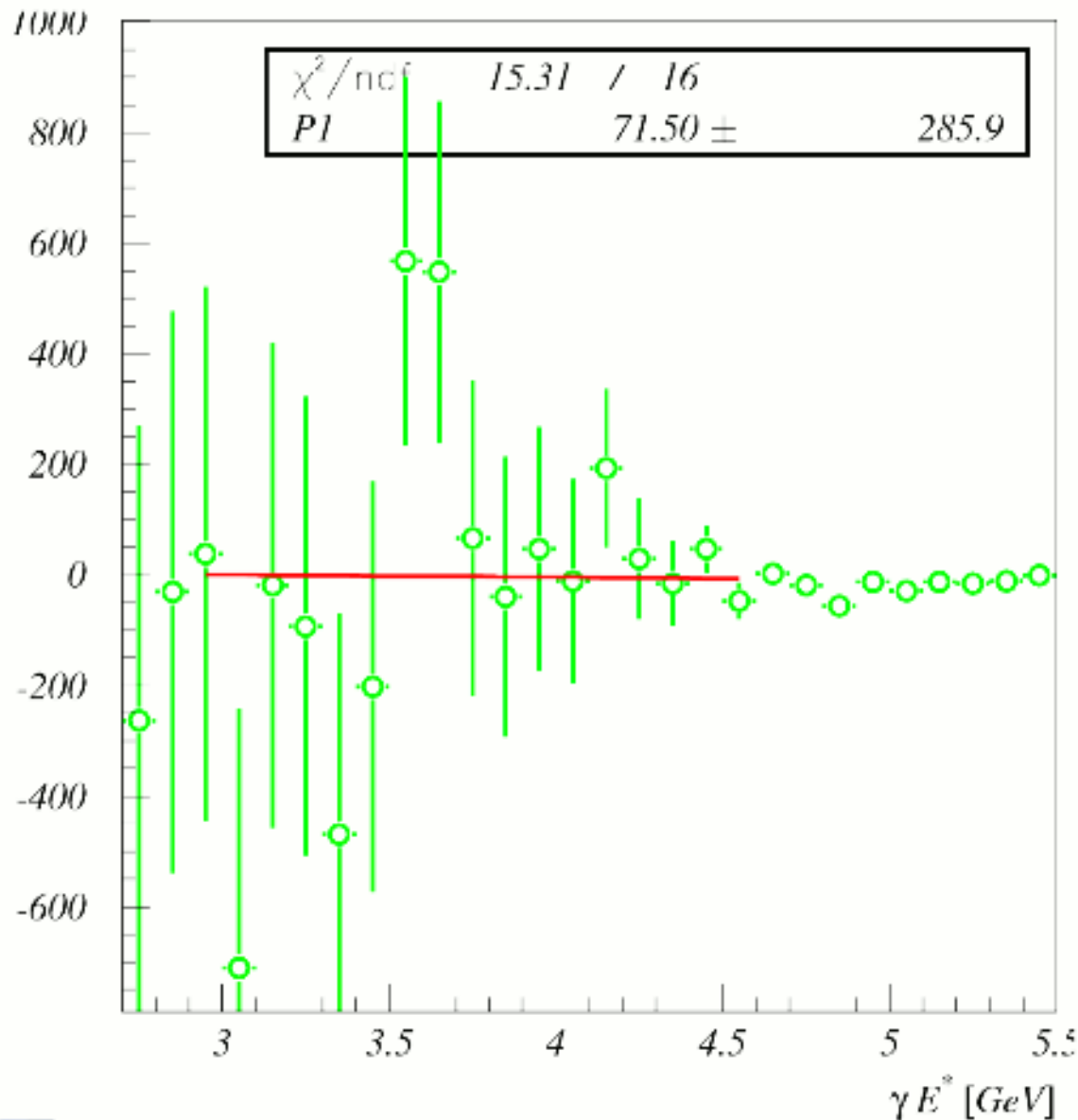
The spectrum



OFF-resonance data is scaled according to luminosities and subtracted from ON-resonance data



The spectrum



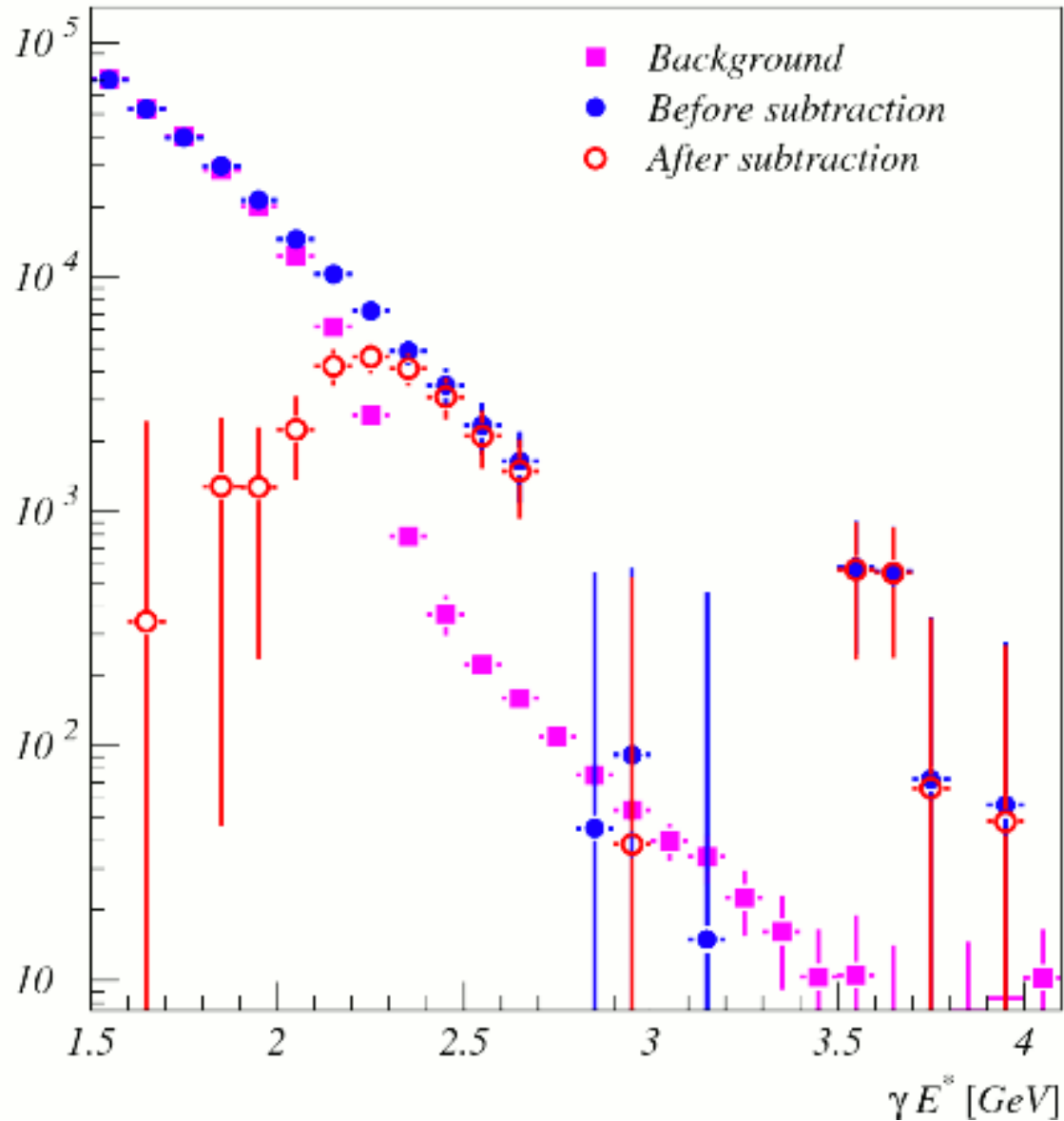
Endpoint check:

Photons from e^+e^- collisions can have an energy up to 5 GeV

But not if they come from a B decay. The kinematic limit is $E^* = m_B/2$.

No significant deviation from 0 observed

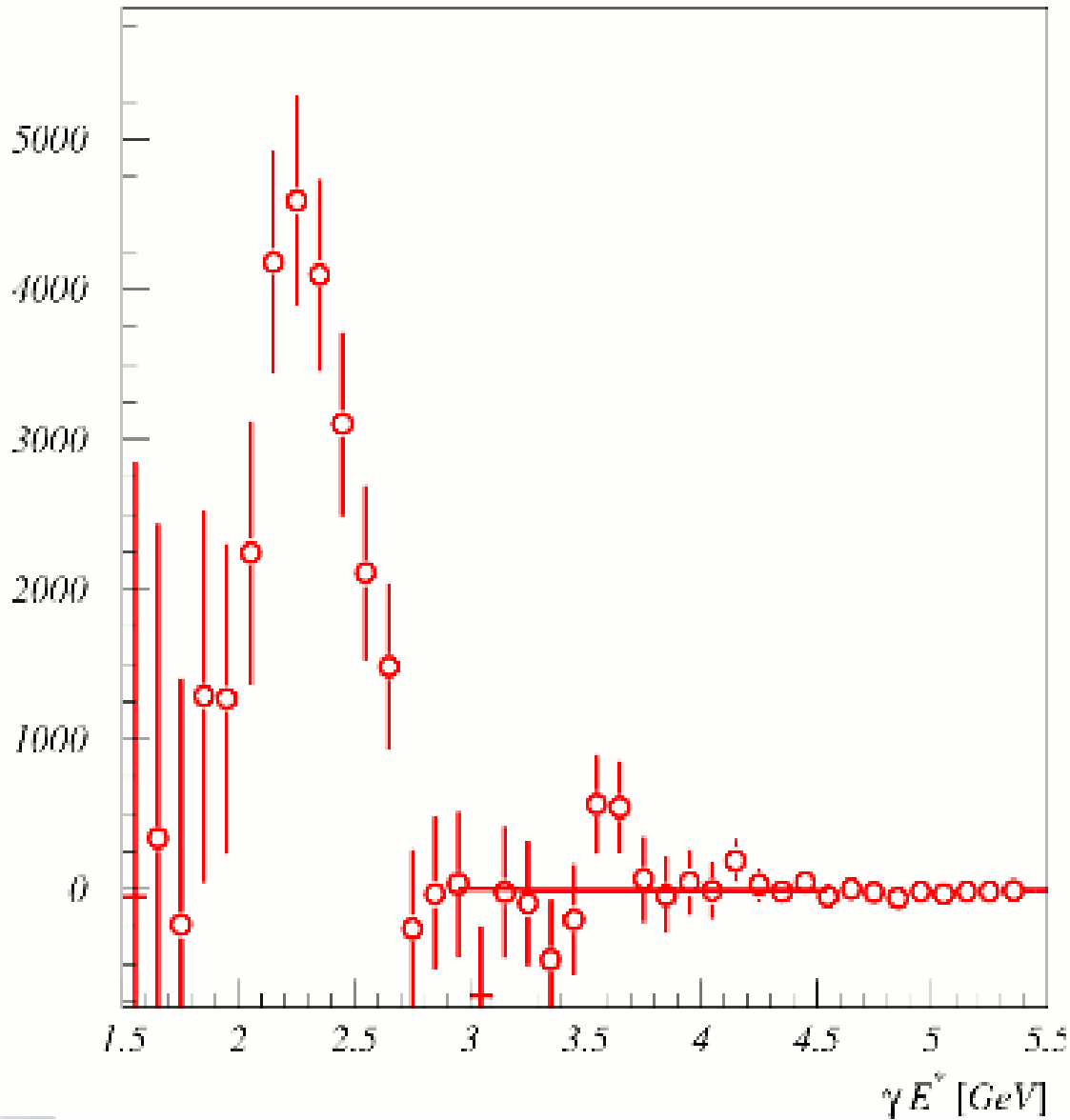
The spectrum



$B\bar{B}$ subtraction:

Using measured π^0 and η spectra and some efficiency-corrected MC.

The spectrum

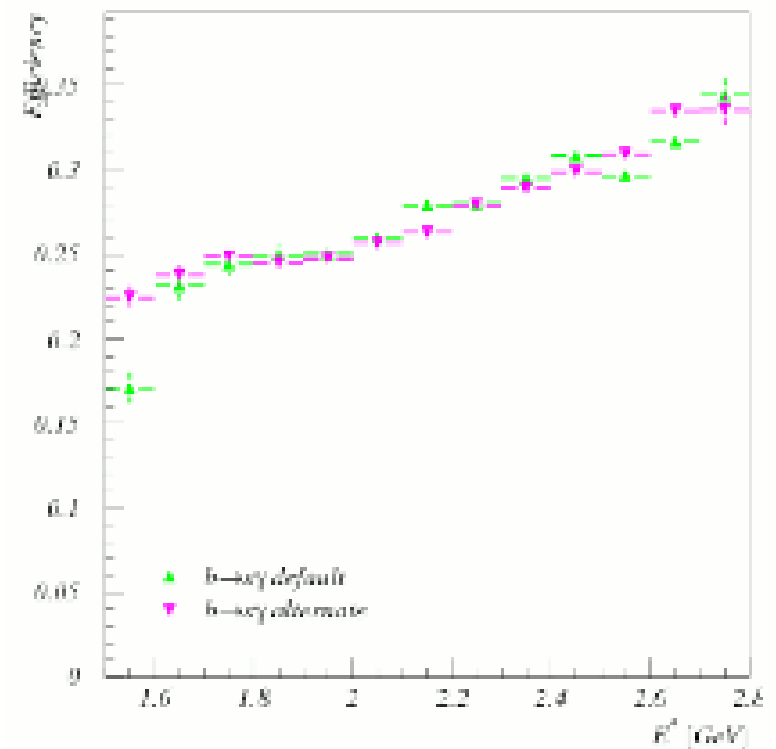
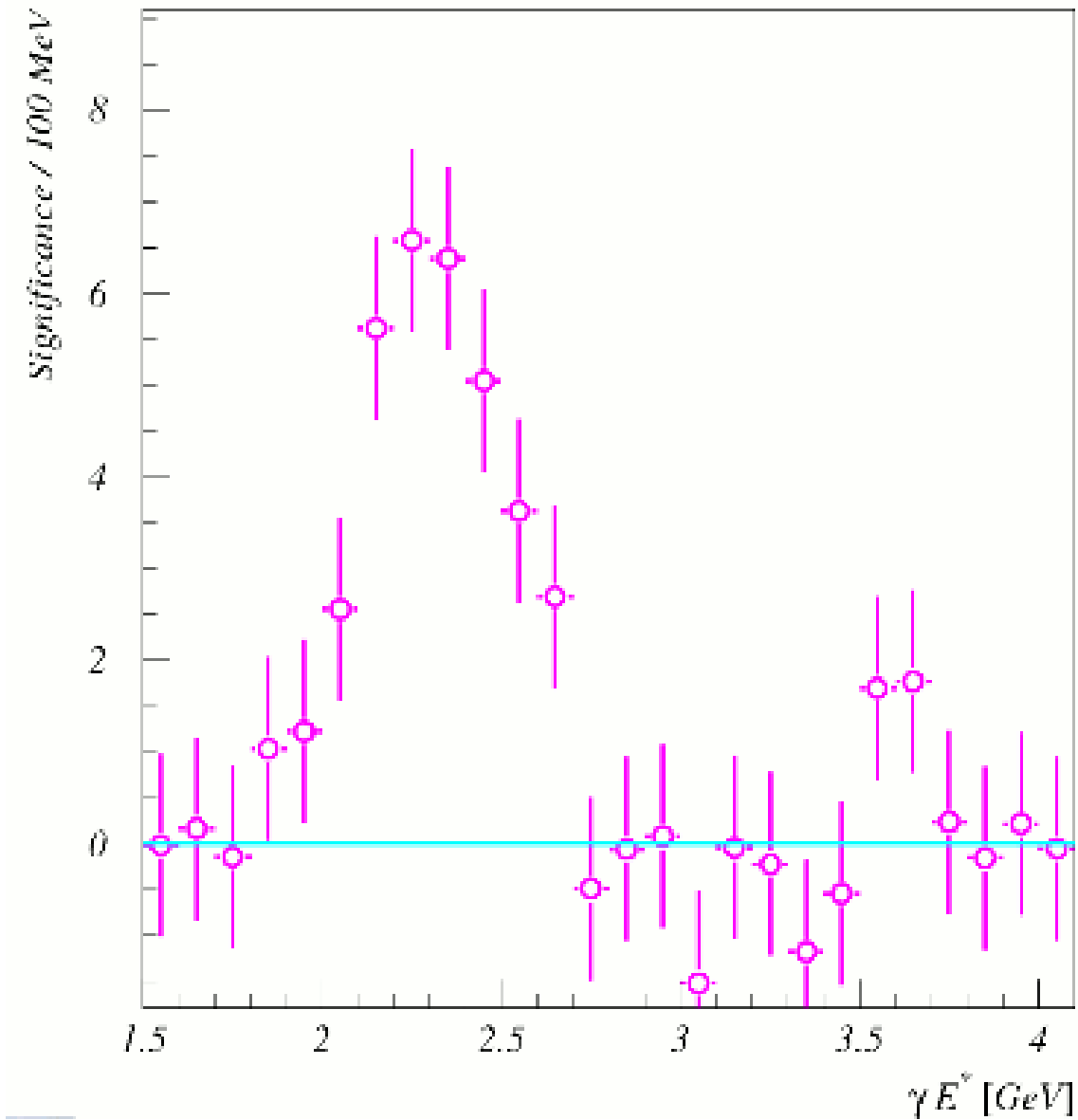


Raw spectrum after all
cuts and background
corrections

Signal yield:
 24100 ± 2200 events

The spectrum

Efficiency corrected spectrum

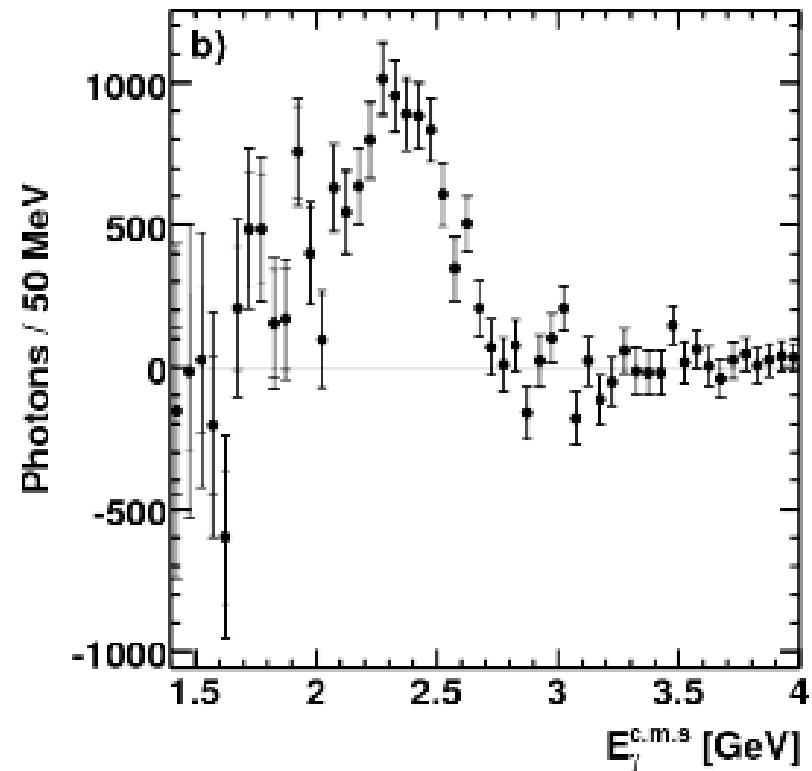
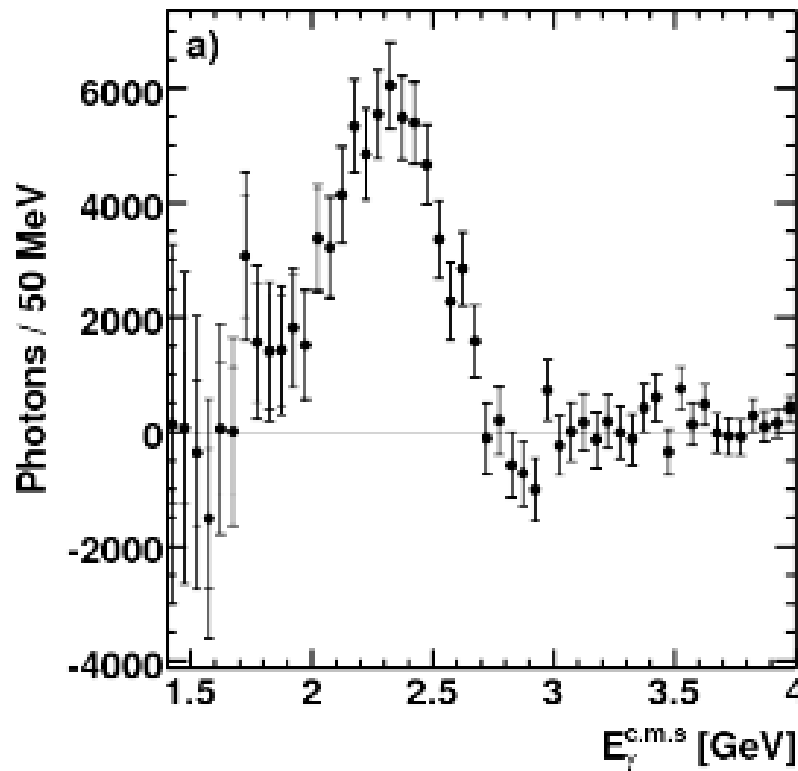


$X_s \gamma$ inclusive

PRL 103, 241801 (2009)

arXiv:0907.1384

Lower E_γ threshold (1.7 GeV) \Rightarrow 97% of the spectrum !



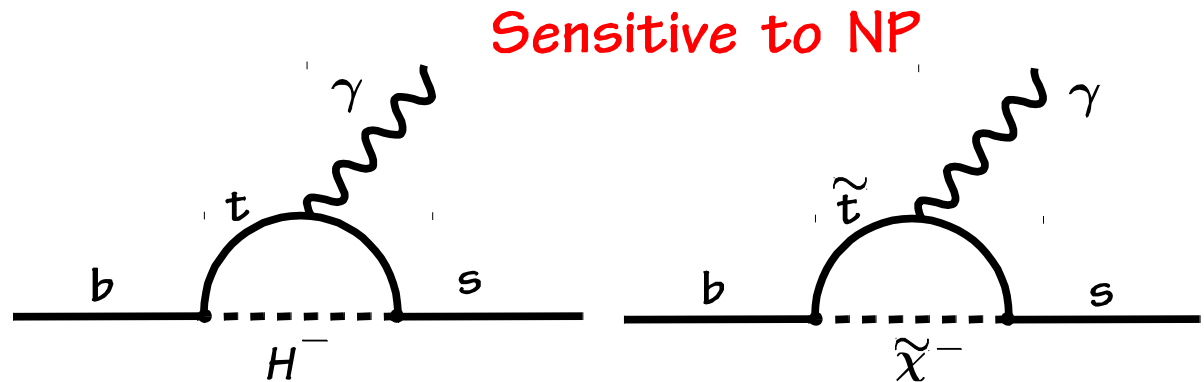
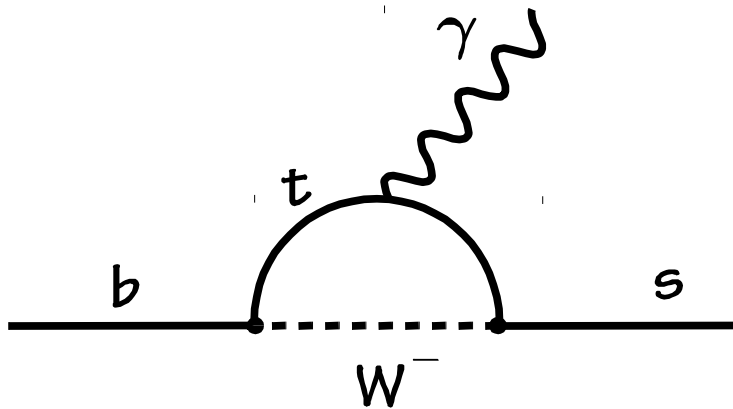
$$B(B \rightarrow X_s \gamma) = (3.45 \pm 0.15 \pm 0.40) \times 10^{-4} \quad (\text{for } E_\gamma^* > 1.7 \text{ GeV})$$

$$B(B \rightarrow X_s \gamma) = (3.21 \pm 0.15 \pm 0.29 \pm 0.08) \times 10^{-4} \quad (\text{for } E_\gamma^* > 1.8 \text{ GeV})$$

$$B(B \rightarrow X_s \gamma) = (3.06 \pm 0.41 \pm 0.26) \times 10^{-4} \quad (\text{for } E_\gamma^* > 2.0 \text{ GeV})$$

- Most precise measurement of $B(B \rightarrow X_s \gamma)$ (lowest E_γ^* threshold)
- Crucial input for global fit to extract $|V_{ub}|$ and $B \rightarrow X_s \gamma$ decay rate
- B is given for E_γ thresholds: 1.7, 1.8, 1.9, 2.0 GeV
- Systematic error is dominated by off-resonance subtraction !

$B \rightarrow X_s \gamma$ as an illustration



NNLO SM calculation:

$$B_{SM}(B \rightarrow X_s \gamma) = (3.36 \pm 0.23) \times 10^{-4}$$

(for $E_\gamma > 1.6$ GeV)

M. Misiak et al.

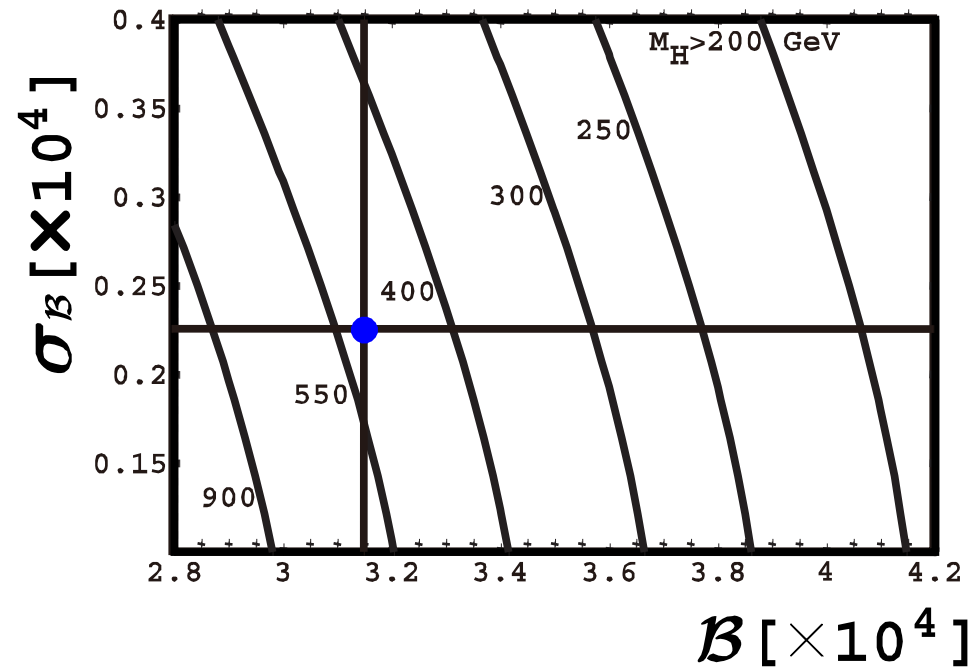
[arXiv:1503.01789]

(central value increased by
6.4% compared to 2007 value)

PRL 98, 022002 (2007)

The lower γ energy threshold, the smaller the model uncertainties in SM, but the larger background in measurement

Charged Higgs (2HDM Type II) bound
(up- and down-type quarks couple to separate doublets)



$B \rightarrow X_s \gamma$

WA: $B(B \rightarrow X_s \gamma) = (3.49 \pm 0.20) \times 10^{-4}$ (for $E_\gamma > 1.6$ GeV)

vs

SM: $B(B \rightarrow X_s \gamma) = (3.36 \pm 0.23) \times 10^{-4}$ (for $E_\gamma > 1.6$ GeV)

[Misiak et al, arXiv:1503.01789]

Charged Higgs bound (2HDM TypeII): $M_{H^\pm} > 400$ GeV @ 95% C.L.

[arXiv:1706.07414]

