

On chiral MHD

Yuji Hirono

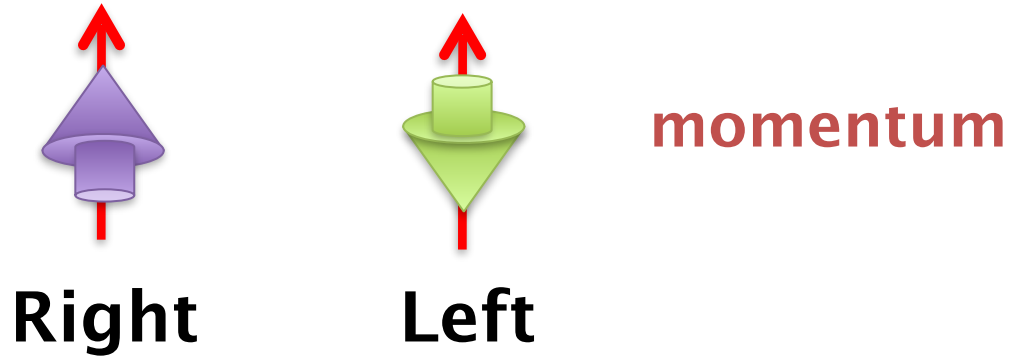
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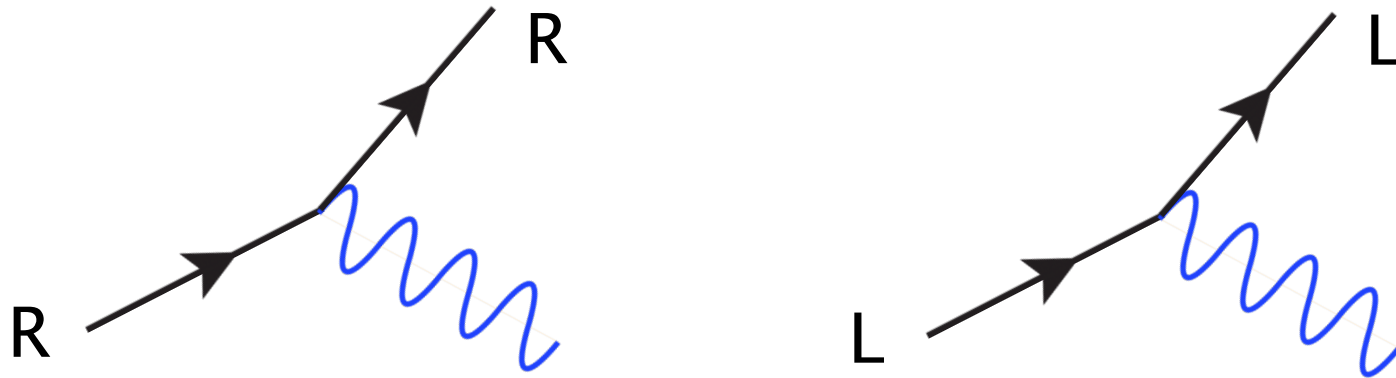
Outline

- Chiral MHD
- Higher-form symmetries
& Nambu-Goldstone modes

Chiralities of massless fermions



- Chirality does not change through interaction classically



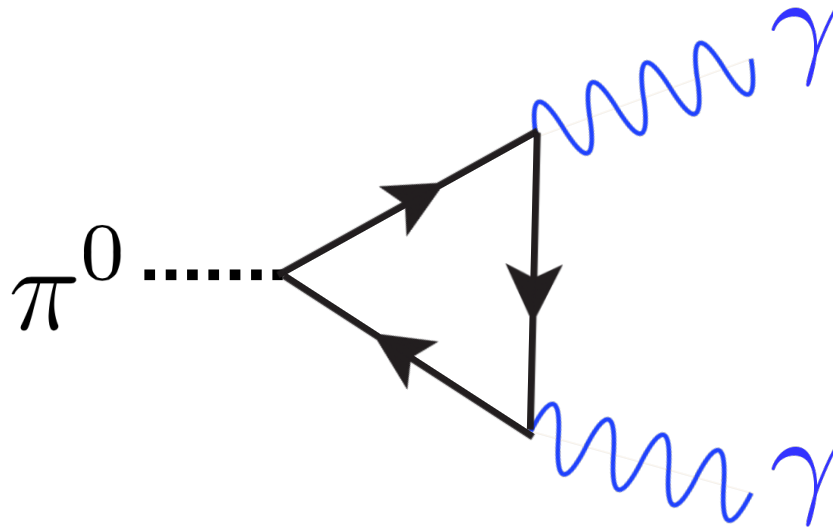
- Quantum effects breaks the chirality conservation

Chiral anomaly [Adler, Bell-Jackiw '69]

Chiral anomaly

$$j_A^\mu = j_R^\mu - j_L^\mu \quad \text{Axial current}$$

$$\partial_\mu j_A^\mu = C_A \mathbf{E} \cdot \mathbf{B} \quad C_A = \frac{e^2}{2\pi^2}$$



Chiral Magnetic Effect (CME)

[Kharzeev-Warringa-McLerran '07]

$$\mathbf{j}_{\text{CME}} = C_A \mu_A \mathbf{B}$$

- Macroscopic transport
- Dissipationless (no heat production)
- Transport coefficient is **universal**
- Where does it happen?
 - Heavy-ion collisions/Early Universe
 - Dirac/Weyl semimetals

Chiral magnetic effect in ZrTe_5

Qiang Li^{1*}, Dmitri E. Kharzeev^{2,3*}, Cheng Zhang¹, Yuan Huang⁴, I. Pletikosić^{1,5}, A. V. Fedorov⁶, R. D. Zhong¹, J. A. Schneeloch¹, G. D. Gu¹ and T. Valla^{1*}

TOPOLOGICAL MATTER

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Evidence for the chiral anomaly in the Dirac semimetal Na_3Bi

Jun Xiong,¹ Satya K. Kushwaha,² Tian Liang,¹ Jason W. Krizan,² Max Hirschberger,¹ Wudi Wang,¹ R. J. Cava,² N. P. Ong^{1*}

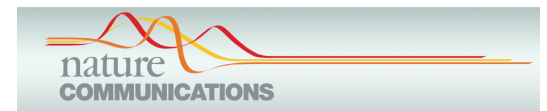
Giant negative magnetoresistance induced by the chiral anomaly in individual Cd_3As_2 nanowires

Cai-Zhen Li^{1*}, Li-Xian Wang^{1*}, Haiwen Liu², Jian Wang^{2,3}, Zhi-Min Liao^{1,3} & Da-Peng Yu^{1,3}



Signatures of the Adler–Bell–Jackiw chiral anomaly in a Weyl fermion semimetal

Cheng-Long Zhang^{1*}, Su-Yang Xu^{2*}, Ilya Belopolski^{2*}, Zhujun Yuan^{1*}, Ziquan Lin³, Bingbing Tong¹, Guang Bian², Nasser Alidoust², Chi-Cheng Lee^{4,5}, Shin-Ming Huang^{4,5}, Tay-Rong Chang^{2,6}, Guoqing Chang^{4,5}, Chuang-Han Hsu^{4,5}, Horng-Tay Jeng^{6,7}, Madhab Neupane^{2,8,9}, Daniel S. Sanchez², Hao Zheng², Junfeng Wang³, Hsin Lin^{4,5}, Chi Zhang^{1,10}, Hai-Zhou Lu¹¹, Shun-Qing Shen¹², Titus Neupert¹³, M. Zahid Hasan² & Shuang Jia^{1,10}



EM fields



Chiral fluid

Chiral MHD

[Boyarsky-Fröhlich-Ruchayskiy '15]

[Yamamoto '16]

[Brandenburg et. al. '17]

[Masada-Kotake-Takiwaki-Yamamoto '18]

[Hattori-Hirono-Yee-Yin '19]

...

Physical systems:

- Heavy-ion collisions
- Early Universe
- Neutrino matter in core-collapse supernovae
- Weyl/Dirac semimetals?

Hydrodynamics

- Low-energy theory based on conservation law & derivative expansion
- Degrees of freedom:
 - **Conserved densities**: particle number, energy, momentum, etc.

$$\{n, e, \boldsymbol{v} \cdots\}$$

- **Nambu-Goldstone modes**
- “Hydrodynamic variables”
- Time evolution is given by the conservation law

$$\partial_t n + \nabla \cdot \boldsymbol{j} = 0$$

Magnetohydrodynamics (MHD)

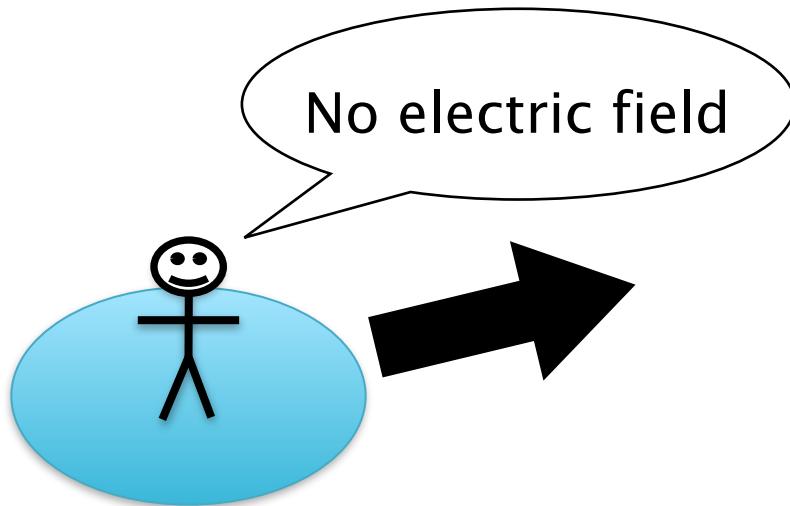
- Hydro variables $\{e(x), u^\mu(x), B^\mu(x)\}$

$$E^\mu := F^{\mu\nu} u_\nu \quad B^\mu := \tilde{F}^{\mu\nu} u_\nu \quad \tilde{F}^{\mu\nu} := \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} F_{\rho\sigma}$$

- EOM

$$\partial_\mu T_{\text{tot}}^{\mu\nu} = 0 \quad \partial_\mu \tilde{F}^{\mu\nu} = 0$$

Ideal MHD: no electric field in the fluid frame



$$E_{(0)}^{\mu} = 0$$

In the lab-frame variable,

$$\mathbf{E} + \mathbf{v} \times \mathbf{B} = 0$$

Corresponds to large conductivity

Constitutive relation for ideal MHD

$$T_{\text{fluid}(0)}^{\mu\nu} = (e + p)u^\mu u^\nu - p\eta^{\mu\nu}$$

$$T_{\text{EM}(0)}^{\mu\nu} = \mathbf{B}^2 \left[u^\mu u^\nu - b^\mu b^\nu - \frac{1}{2}\eta^{\mu\nu} \right]$$

$$B^\mu = |\mathbf{B}|b^\mu \quad b_\mu b^\mu = -1$$

$$F_{(0)}^{\mu\nu} = \epsilon^{\mu\nu\rho\sigma} u_\rho B_\sigma$$

Chiral MHD

- Hydro variables $\{e(x), u^\mu(x), B^\mu(x), n_A(x)\}$

$$E^\mu := F^{\mu\nu} u_\nu \quad B^\mu := \tilde{F}^{\mu\nu} u_\nu \quad \tilde{F}^{\mu\nu} := \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} F_{\rho\sigma}$$

- EOM

$$\partial_\mu T_{\text{tot}}^{\mu\nu} = 0 \quad \partial_\mu \tilde{F}^{\mu\nu} = 0 \quad \partial_\mu j_A^\mu = C E_\mu B^\mu$$

- Requirement of the 2nd law of thermodynamics leads to a constitutive relation including CME

[Hattori-Hirono-Yee-Yin '19]

Constitutive relation for ideal CMHD

$$T_{\text{fluid}(0)}^{\mu\nu} = (e + p)u^\mu u^\nu - p\eta^{\mu\nu}$$

$$T_{\text{EM}(0)}^{\mu\nu} = \mathbf{B}^2 \left[u^\mu u^\nu - b^\mu b^\nu - \frac{1}{2}\eta^{\mu\nu} \right]$$

$$B^\mu = |\mathbf{B}|b^\mu \quad b_\mu b^\mu = -1$$

$$F_{(0)}^{\mu\nu} = \epsilon^{\mu\nu\rho\sigma} u_\rho B_\sigma$$

$$j_{\text{A}(0)}^\mu = n_{\text{A}} u^\mu$$

CME doesn't play any role

First order in derivative expansion

$$T_{\text{fluid}(1)}^{\mu\nu} = \zeta \Delta^{\mu\nu} \partial \cdot u + 2\eta \nabla^{<\mu} u^{\nu>}$$

$$\sigma E_{(1)}^\mu = \epsilon^{\mu\nu\alpha\beta} u_\nu \partial_\alpha B_\beta + \epsilon^{\mu\nu\alpha\beta} u_\nu B_\alpha \partial_\beta \ln T - \underbrace{\sigma_B B^\mu}_{\text{CME}}$$

σ : electric conductivity

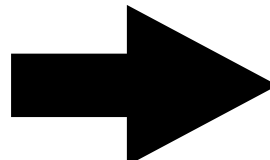
$$\sigma_B := C_A \mu_A$$

First order in derivative expansion

EM tensor and field strength are modified accordingly

$$T_{\text{EM}(1)}^{\mu\nu} = 2\epsilon^{(\mu\alpha\beta}u^{\nu)}E_{\alpha(1)}B_{\beta}$$

$$\tilde{F}_{(1)}^{\mu\nu} = \epsilon^{\mu\nu\rho\sigma}E_{(1)\rho}u_{\sigma}$$


$$\partial_{\mu} \left[T_{\text{tot}(0)}^{\mu\nu} + T_{\text{tot}(1)}^{\mu\nu} \right] = 0$$
$$\partial_{\mu} \left[\tilde{F}_{(0)}^{\mu\nu} + \tilde{F}_{(1)}^{\mu\nu} \right] = 0$$

Force from \mathbf{B}

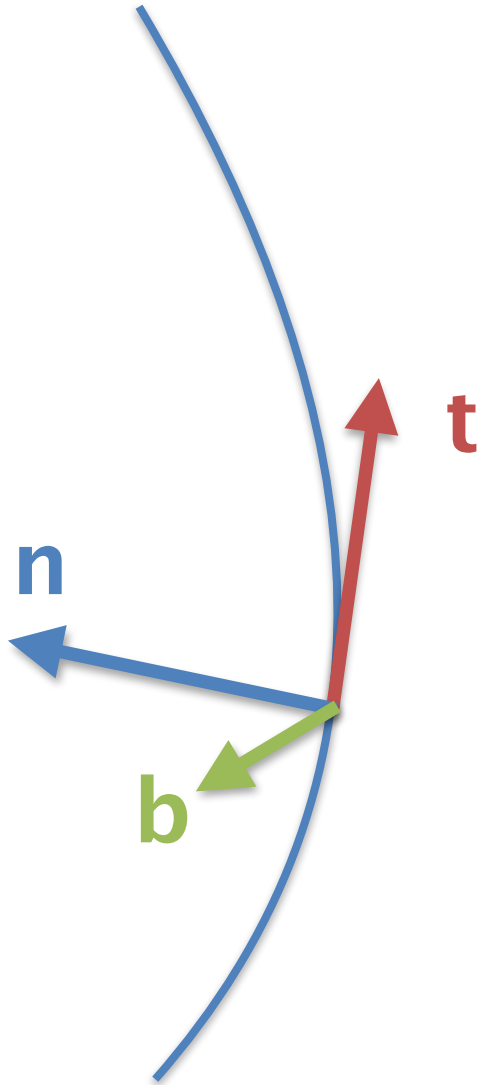
Stress tensor

$$T_{ij} = B_i B_j - \frac{1}{2} B^2 \delta_{ij}$$

$$\text{Force} = \partial_i T_{ij}$$

$$\mathbf{F} = B^2 \kappa \mathbf{n} - \nabla_{\perp} \frac{1}{2} B^2$$

κ : curvature of \mathbf{B} field line



Force from \mathbf{B}

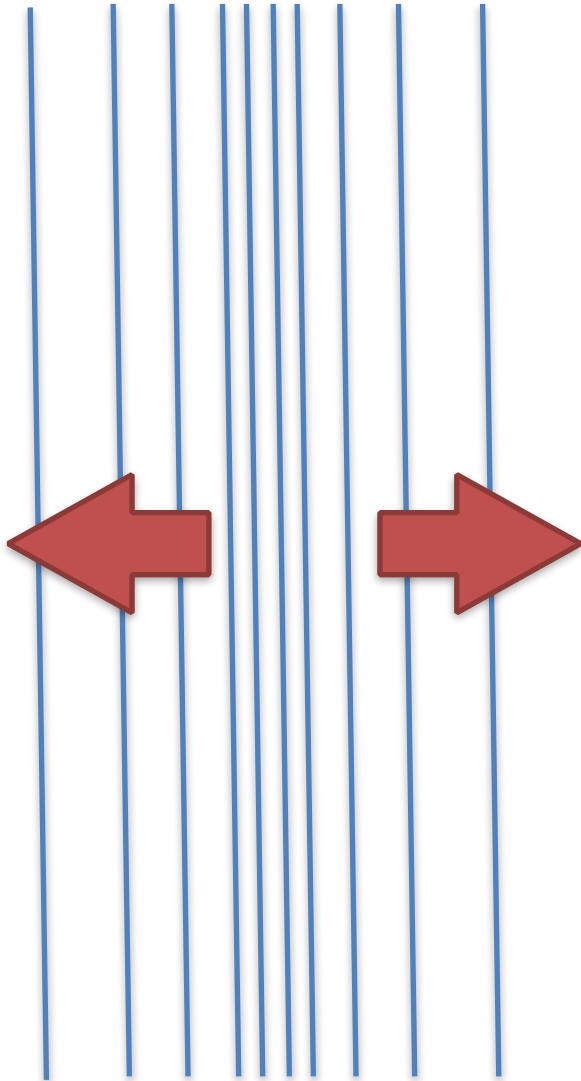
Stress tensor

$$T_{ij} = B_i B_j - \frac{1}{2} B^2 \delta_{ij}$$

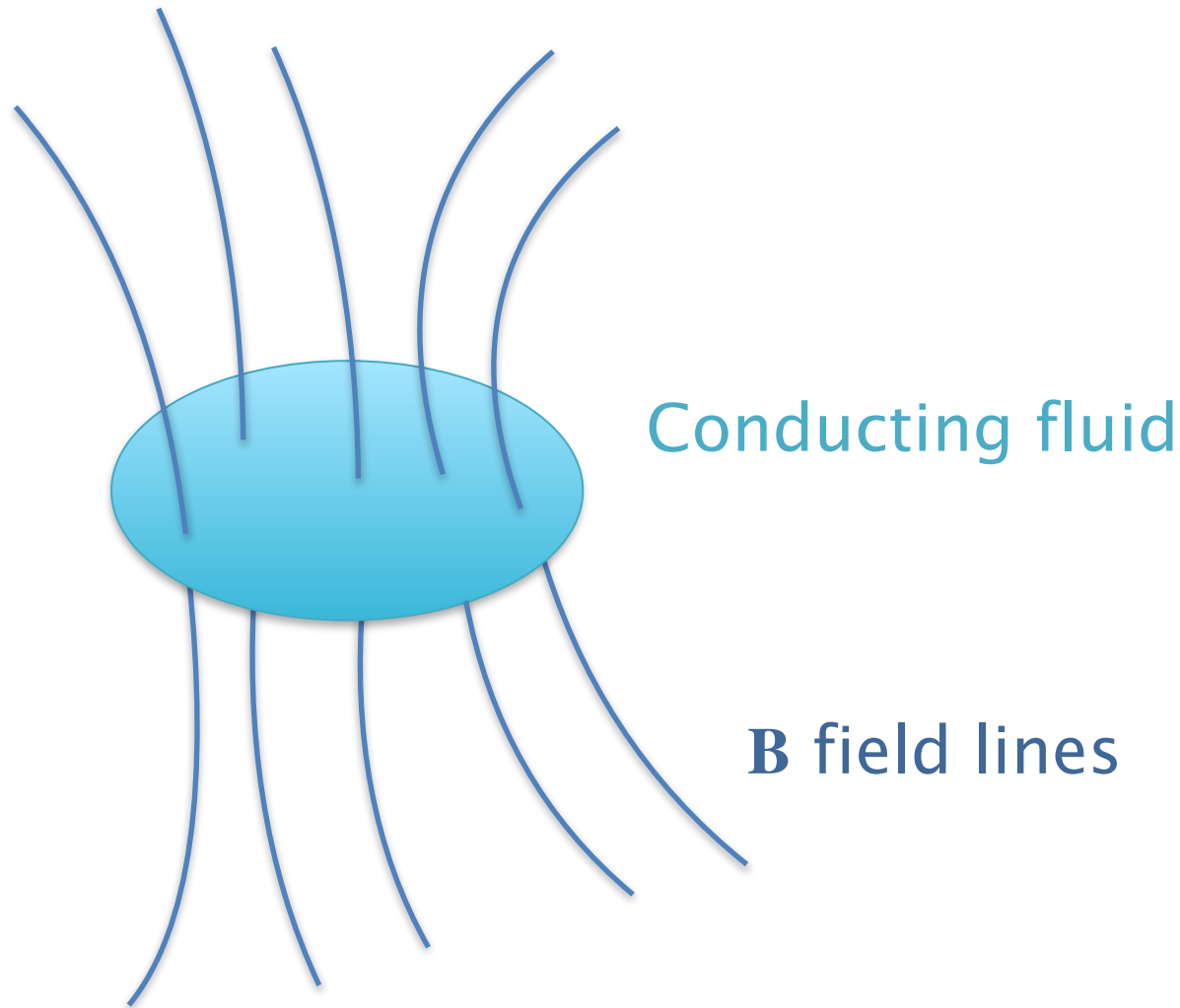
$$\text{Force} = \partial_i T_{ij}$$

$$\mathbf{F} = B^2 \kappa \mathbf{n} - \nabla_{\perp} \frac{1}{2} B^2$$

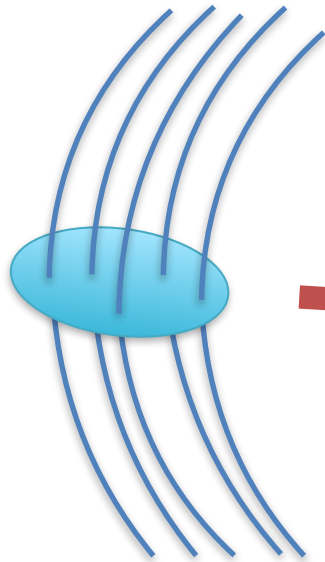
κ : curvature of \mathbf{B} field line



Frozen in of \mathbf{B} field lines



Alfven wave



B field lines

 force

Dispersion $\omega = \pm v_A k_{||}$

Momentum along B



$$v_A^2 = \frac{B^2}{e + p + B^2}$$

Alfven velocity

Alfven wave: with dissipation

Dispersion relation

$$\omega = \pm v_A k_{||} - \frac{i}{2} [\bar{\eta} + \lambda] k^2$$

.....
Damping

$$\lambda = \frac{1}{\sigma}$$

σ : electric conductivity

$$\bar{\eta} \equiv \frac{\eta}{e + p + \mathbf{B}^2}$$

η : shear viscosity

Alfven wave in chiral MHD

Including CME,

$$\omega = \pm v_A k_{||} - \frac{i}{2}(\bar{\eta} + \lambda)k_{||}^2 + \underbrace{\frac{i}{2}s\epsilon_B k_{||}}_{\text{CME}}$$

s indicates the helicity of the mode

$$\epsilon_B := \frac{C_A \mu_A}{\sigma}$$

$$i\mathbf{k} \times \underbrace{\mathbf{e}^{(s)}}_{\text{helicity eigenstate}} = s k \mathbf{e}^{(s)}$$

helicity eigenstate

Alfven wave in chiral MHD

Instability appears
in positive(or negative)-helicity mode

If $\sigma_B > 0$,

- Negative helicity modes decay
- Positive helicity modes with $k < k_c$ exponentially grow

$$k_c = \frac{\sigma_B}{1 + \bar{\eta}\sigma}$$

.....

dissipative effects tame the instability

Bianchi identity

$$\partial_\mu \tilde{F}^{\mu\nu} = 0$$

- This can be regarded as a conservation law of charges associated with a $U(1)$ 1-form symmetry

Higher-form symmetries

[Gaiotto-Kapustin-Seiberg-Willet '15]

- Usual symmetry: $\psi_i(x) \mapsto R_i^j(g) \psi_j(x)$
- Higher-form symmetry
 - Charged objects are **extended**
 - “ p -form symmetry” \rightarrow charged obj. is p -dimensional



0-form symmetry



1-form symmetry

- $U(1)^{[p]}$ symmetry: $W(C_p) \rightarrow e^{i\alpha} W(C_p)$

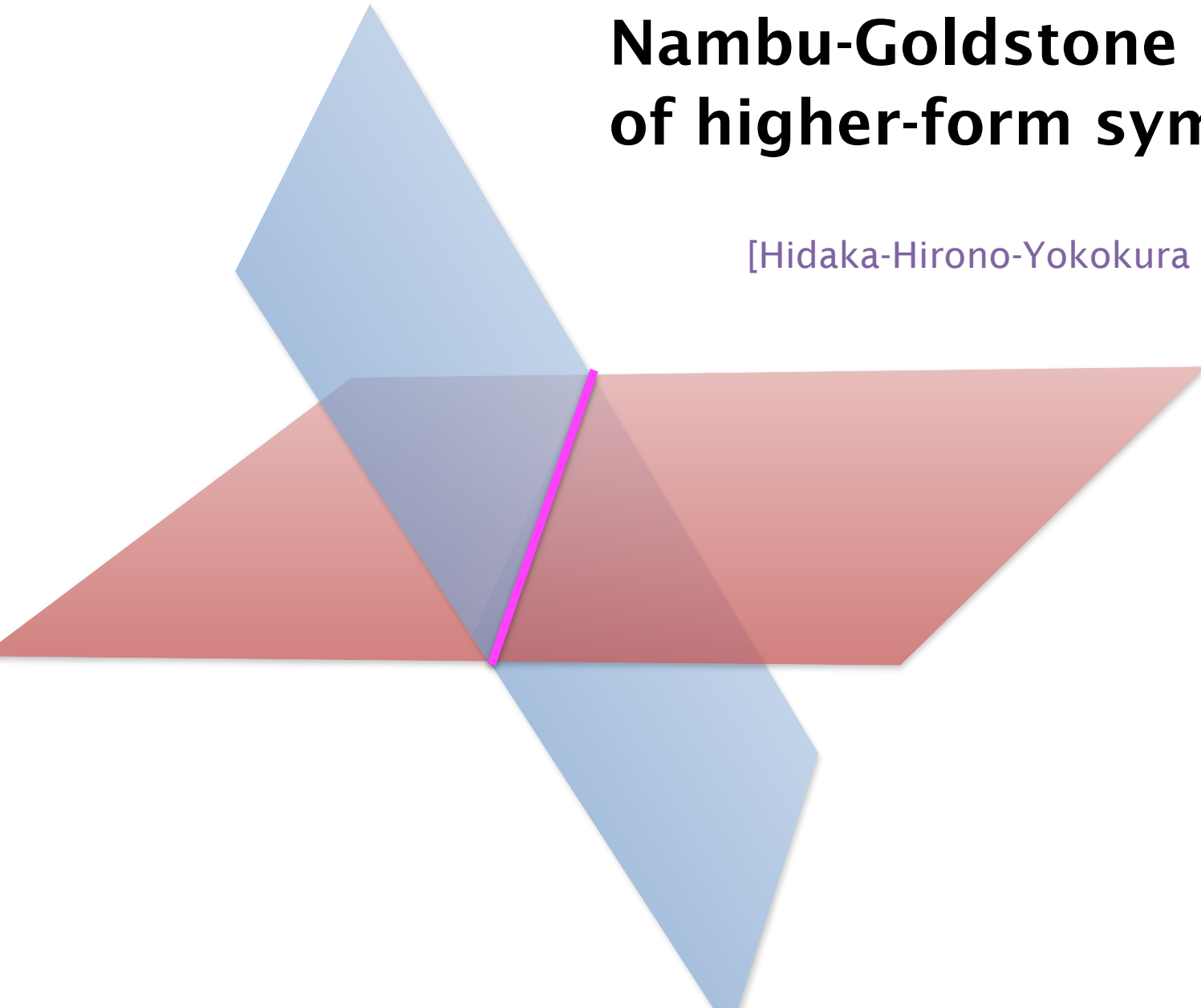
Higher-form symmetries

[Gaiotto-Kapustin-Seiberg-Willet '15]

- Provides us with a unifying viewpoint
 - Classification of phases
 - Certain **topological orders** are understood as a consequence of broken discrete HFS
 - Nambu-Goldstone theorem
 - 't Hooft anomalies
 - Constraints on the low-energy physics
- Formulation of MHD [Grozdanov-Hofman-Iqbal '17], ...

Counting the number of Nambu-Goldstone modes of higher-form symmetries

[Hidaka-Hirono-Yokokura '20]




Spontaneous symmetry breaking \rightarrow low-energy d.o.f

- Examples of Nambu-Goldstone (NG) modes
 - pions in QCD
 - magnons
 - superfluid phonons (He^4 , He^3 , atomic BEC)
 - lattice phonons
 - Kelvin (vibration of a vortex string)
 - photons
 - ...

Number of NG modes

- (Internal) symmetry breaking $G \rightarrow H$
- Lorentz invariance

$$N_{\text{NG}} = N_{\text{BG}}$$


of massless NG modes # of broken generators

- Dispersion relation: $\omega = ck$ $k = |\mathbf{k}|$

Magnets: Heisenberg model $H = J \sum_{\langle i,j \rangle} \vec{s}_i \cdot \vec{s}_j$

- $J > 0$: Anti-ferromagnetism

$$N_{\text{NG}} = 2 \quad \omega \propto k$$



- $J < 0$: Ferromagnetism

$$N_{\text{NG}} = 1 \quad \omega \propto k^2$$



Symmetry breaking pattern is common $SO(3) \rightarrow SO(2)$

$$N_{\text{BG}} = 2$$

Thermodynamic properties: heat capacity

- When $\omega \sim k^n$

heat capacity behaves as

$$C \sim T^{\frac{d}{n}}$$

d : spatial dimension

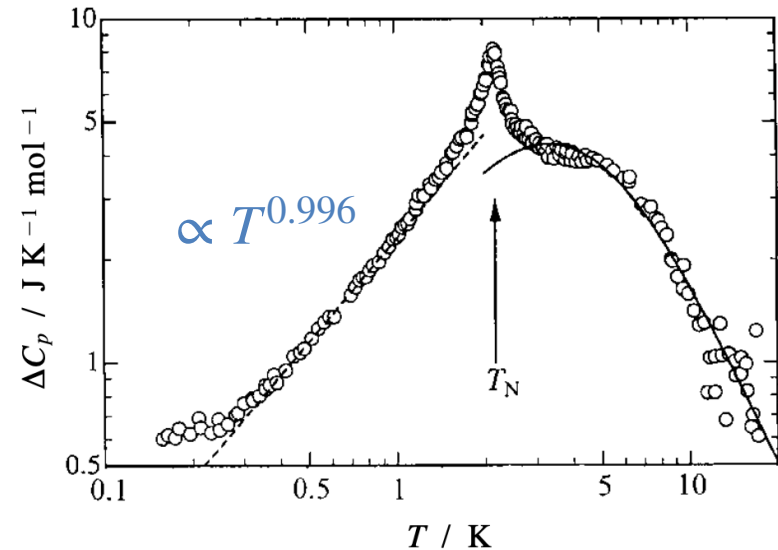


Fig.9 Excess heat capacities of $\text{MnCu(obbz)} \cdot 5\text{H}_2\text{O}$ around the antiferromagnetic phase transition temperature T_N . Solid line indicates the theoretical heat capacity curve estimated by the high-temperature series expansion for $S=2$ one-dimensional ferromagnetic Heisenberg model with $J/k_B=0.75$ K. The broken straight line shows the heat capacity due to the spin-wave excitation.³⁸⁾

What's the difference?

- Anti-ferro: $\langle \hat{s}_z \rangle = 0$
- Ferro: $\langle \hat{s}_z \rangle = \langle [\hat{s}_x, \hat{s}_y] \rangle \neq 0$
- $\langle [Q_\alpha, Q_\beta] \rangle = 0$ for $\forall \alpha, \beta \rightarrow N_{\text{NG}} = N_{\text{BG}}$

[Schafer-Son-Stephanov-Toublan-Verbaarschot '01]

Number of NG modes

- Non-relativistic systems

$$N_{\text{NG}} = N_{\text{BG}} - \frac{1}{2} \text{rank} \langle [Q_\alpha, Q_\beta] \rangle$$



of massless NG modes



of broken generators

[Watanabe-Brauner '11]

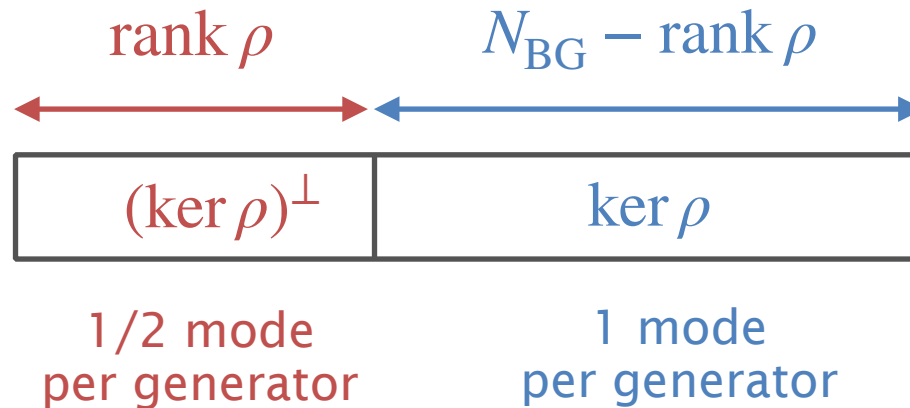
[Watanabe-Murayama'12] [Hidaka'12]

- Dispersion relation
 - Type A: $\omega \sim k$
 - Type B: $\omega \sim k^2$

Derivation by effective field theories

- Up to 2nd order in fields & derivatives,

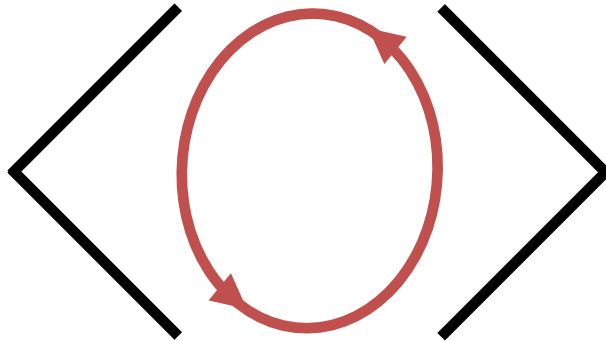
$$\mathcal{L} = \rho_{AB} \pi^A \dot{\pi}^B + \frac{1}{2} G_{AB} \dot{\pi}^A \dot{\pi}^B - \frac{1}{2} \bar{G}_{AB} \partial_i \pi^A \partial^i \pi^B$$



$$N_{\text{NG}} = \frac{1}{2} \cdot \text{rank } \rho + 1 \cdot (N_{\text{BG}} - \text{rank } \rho) = N_{\text{BG}} - \frac{1}{2} \text{rank } \rho$$

NG modes for higher-form symmetries

- Spontaneous breaking of a continuous HFS \rightarrow massless particle


$$\simeq 1$$

Perimeter law
or Coulomb law

- Breaking of $U(1)_e^{[1]}$ 1-form symmetry \rightarrow photons
- Questions
 - How many NG modes?
 - Dose $\langle [Q_\alpha, Q_\beta] \rangle$ affect the number of NG modes?
 - cf. Non-commuting 0-form and 1-form charges

[Sogabe-Yamamoto'19]

Generalized formula to count NG modes

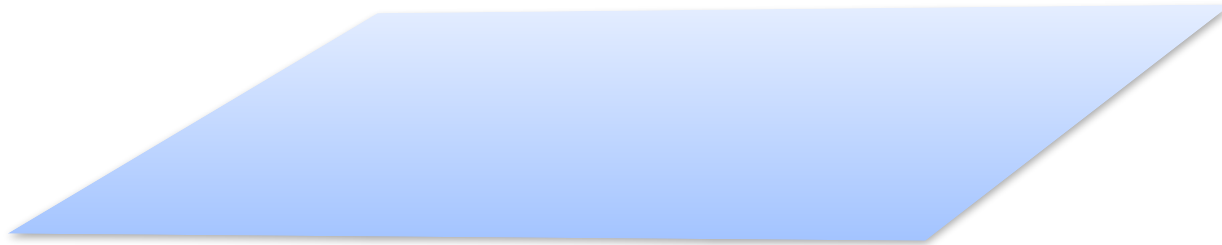
$$N_{\text{NG}} = \sum_A {}^{D-2}C_{p_A} - \frac{1}{2} \text{rank} \langle [Q_\alpha, Q_\beta] \rangle$$

[Hidaka-Hirono-Yokokura '20]

- SSB of internal symmetries (that can include higher-form)
- D -dim Minkowski space $\mathbb{R}^{1,D-1}$
 - Lorentz symmetry doesn't have to be there
 - Translational symmetry is intact
- “ A ” is a label for a p_A -form symmetry

Generalized formula to count NG modes

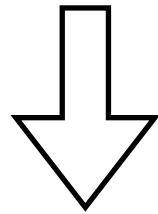
- Symmetry generators Q_α should be enumerated taking into account **how to place them**
 - For 0-form symmetries, $Q(V)$ is $(D - 1)$ -dimensional
 - For a p -form symmetry, the generator is $D - p - 1$ -dimensional
 - Ex) 1-form-symmetry in $(3+1)$ -dim. \rightarrow gen. is 2-dim.



$$U(e^{i\alpha}, S) = e^{i\alpha Q^{[1]}(S)}$$

Generalized formula to count NG modes

$$N_{\text{NG}} = \sum_A {}^{D-2}C_{p_A} - \frac{1}{2} \text{rank} \langle [Q_\alpha, Q_\beta] \rangle$$



When all the symmetries are 0-form

$$N_{\text{NG}} = N_{\text{BG}} - \frac{1}{2} \text{rank} \langle [Q_\alpha, Q_\beta] \rangle$$

Reproduce the previous result

[Watanabe-Brauner '11]

[Watanabe-Murayama'12] [Hidaka'12]

Generalized formula to count NG modes

$$N_{\text{NG}} = \sum_A {}^{D-2}C_{p_A} - \frac{1}{2} \text{rank} \langle [Q_\alpha, Q_\beta] \rangle$$

- Photons in $(3 + 1)$ -dimensions
- Broken symmetry: $U(1)_e^{[1]}$

$$N_{\text{NG}} = {}^{D-2}C_p = 2 \qquad D = 4, \quad p = 1$$

Ex) Photons under a θ gradient

[Yamamoto '16]

- Lagrangian

$$L = -\frac{1}{2e^2} f \wedge \star f - \frac{\theta}{2\pi} f \wedge f$$

- $d\theta = C = C_i dx^i \neq 0$
- Broken symmetry: $U(1)_e^{[1]}$
 - Conserved charge: $Q^{[1]}(S)$
- Charge commutators $\langle [Q^{[1]}(S_1), Q^{[1]}(S_2)] \rangle \propto \int_{S_1 \cap S_2} d\theta \neq 0$

Ex) Photons under a θ gradient

S_1

$$\langle [Q^{[1]}(S_1), Q^{[1]}(S_2)] \rangle \propto \int_{S_1 \cap S_2} d\theta \neq 0$$

S_2

Ex) Photons under a θ gradient

- Independent generators: $\{Q^{[1]}(S_x), Q^{[1]}(S_y), Q^{[1]}(S_z)\}$
 S_i is a 2-dim surface perpendicular to i -axis

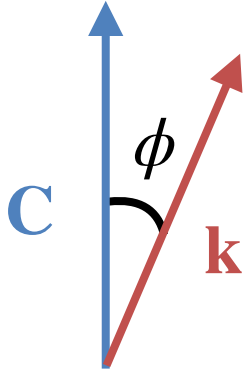
- Charge commutators

$$M_{ij} = \langle [Q^{[1]}(S_i), Q^{[1]}(S_j)] \rangle \propto \epsilon_{ijk} C_k$$

- Number of NG modes

$$N_{\text{NG}} = D-2 C_1 - \frac{1}{2} \text{rank } M_{ij} = 2 - \frac{1}{2} \cdot 2 = 1$$

Ex) Photons under a θ gradient



Dispersion relation

$$\omega^2(k) = \begin{cases} C^2 + (2 - \sin^2 \phi)k^2 + O(k^4) & \text{massive} \\ \sin^2 \phi k^2 + \frac{\cos^4 \phi}{C^2} k^4 + O(k^6) & \text{massless} \end{cases}$$

$$C := |\mathbf{C}|$$

- $\omega \sim k^2$ when $\mathbf{C} \parallel \mathbf{k}$
- $\omega \sim k$ for other angles
- No generic properties about dispersion relations for higher-form symmetries?

Derivation: effective field theories

- EFT of NG modes
 - Write down terms consistent with the symmetries
 - Derivative expansion
- EFT for broken higher-form symmetries
 - Extension of the coset construction

Summary

- Chiral MHD
 - Low-energy theory of chiral matter & EM fields
 - Formation of helical \mathbf{B} configurations
- Higher-form symmetries & Nambu-Goldstone modes
 - Formula to count the number
 - EFT of NG modes for higher-form symmetries
 - Extension of the coset construction