

How close are shell models to the 3D Navier–Stokes equations?

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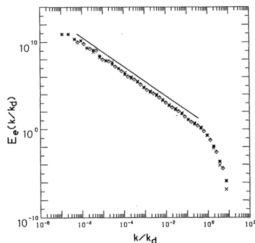


Shell models of turbulence

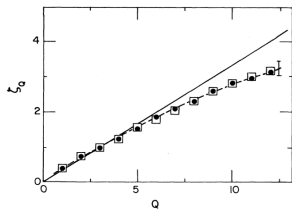
Sabra shell model (L'vov, Podivilov, Pomyalov, Procaccia, *Phys. Rev. E*, 1998)

$$\dot{u}_j = i(ak_{j+1}u_{j+1}^*u_{j+2} + bk_ju_{j+1}u_{j-1}^* - ck_{j-1}u_{j-1}u_{j-2}) - \nu k_j^2 u_j + f_j$$

$$k_j = k_0 \lambda^j, \quad \lambda > 1, \quad j = 1, 2, \dots$$



Yamada, Ohkitani, *J. Phys. Soc. Japan* (1987)



$$\langle |u_j|^Q \rangle \sim k_j^{-\zeta_Q}$$

Jensen, Paladin, Vulpiani, *Phys. Rev. A* (1991)

Frisch, *Turbulence: The legacy of AN Kolmogorov* (1995)

Bohr, Jensen, Paladin, Vulpiani, *Dynamical systems approach to turbulence* (1998)

Biferale, *Annu. Rev. Fluid Mech.* (2003)

Verma, *Energy transfers in fluid flows* (2019)

Mathematical analysis of shell models

Constantin, Levant, and Titi *Physica D* (2006); *J. Stat. Phys.*, (2007)

- ▶ Global regularity for weak and strong solutions
- ▶ Exponential decay of $|u_j|$ with k_j
- ▶ Finite-dimensional global attractor and inertial manifold

Stochastically forced

Barbato, Barsanti, Bessaih, Flandoli *J. Stat. Phys.*, (2006)

Inviscid

Constantin, Levant, and Titi *Phys. Rev. E* (2007)

Barbato, Flandoli, Morandin *Proc. Am. Math. Soc.* (2011), *Trans. Am. Math. Soc.* (2011), *Ann. Appl. Probab.* (2011)

A comparison between shell models and the 3D NSEs

How the powers of the velocity derivatives depend on Re ?

3D Navier–Stokes equations

Gibbon *J. Nonlin. Sci.* (2019)

$$\langle \|\nabla^n \mathbf{u}\|_{2m}^{\alpha_{n,m}} \rangle_T \leq c_{n,m} L^{-1} \nu^{\alpha_{n,m}} Re^3$$

$$V = [0, L]^3, \quad \alpha_{n,m} = 2m/[2m(n+1) - 3]$$

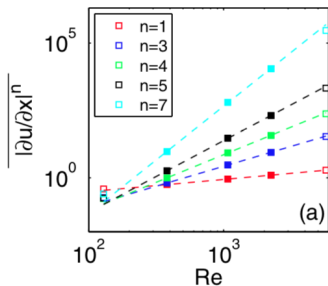
$$1 \leq n, 1 \leq m \leq \infty \quad \text{or} \quad n = 0, 3 \leq m \leq \infty$$

Shell models

?

Velocity gradients in turbulence

$$\langle |\nabla \mathbf{u}|^p \rangle \sim Re^{\chi_p}$$



Schumacher et al. *NJP* (2007)

Multifractal model:

Nelkin *Phys. Rev. A* (1990)

Benzi, Biferale, Paladin, Vulpiani, Vergassola *Phys. Rev. Lett.* (1991)

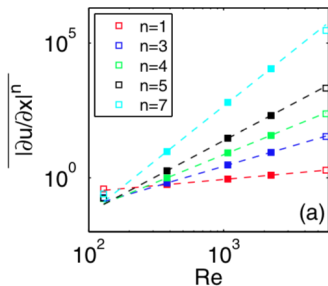
DNS: Schumacher, Sreenivasan, Yakhot *New J. Phys.* (2007); *PNAS* (2014)

Burgers equation: Chakraborty, Frisch, Pauls, Ray *Phys. Rev. E* (2012)

Velocity gradients in turbulence

$$\langle |\nabla \mathbf{u}|^p \rangle \sim Re^{\chi_p}$$

$$\langle \langle |\nabla^n \mathbf{u}|^p \rangle_V^{\alpha_{n,p}} \rangle_T$$



Schumacher et al. *NJP* (2007)

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$$k_j = k_0 \lambda^j, \lambda > 1, j = 1, 2, \dots$$

Energy

$$E(t) = \sum_{j=1}^{\infty} |u_j(t)|^2$$

Energy conservation: $a + b + c = 0$

Initial energy: $E(0) < \infty$

Sabra shell model

$$\dot{u}_j = i(ak_{j+1}u_{j+1}^*u_{j+2} + bk_ju_{j+1}u_{j-1}^* - ck_{j-1}u_{j-1}u_{j-2}) - \nu k_j^2 u_j + f_j$$

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Initial energy: $E(0) < \infty$

Forcing

$$f_j = F\phi_j \text{ with } \phi_j = 0 \text{ if } j < j_f \text{ and } j > j_{\max}$$

Characteristic and maximum wave numbers: $k_f = k_0 \lambda^{j_f}$ and $k_{\max} = k_0 \lambda^{j_{\max}}$

Sabra shell model

$$\dot{u}_j = i(ak_{j+1}u_{j+1}^*u_{j+2} + bk_ju_{j+1}u_{j-1}^* - ck_{j-1}u_{j-1}u_{j-2}) - \nu k_j^2 u_j + f_j$$
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r.m.s. velocity

$$U = \sqrt{\langle E \rangle_T} \text{ with } \langle \cdot \rangle_T = \frac{1}{T} \int_0^T \cdot dt$$

$$Re = \frac{U}{\nu k_f}$$

$$Gr = \frac{|F|}{\nu^2 k_f^3}$$

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$$\dot{u}_j = i(ak_{j+1}u_{j+1}^*u_{j+2} + bk_ju_{j+1}u_{j-1}^* - ck_{j-1}u_{j-1}u_{j-2}) - \nu k_j^2 u_j + f_j$$
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Energy $E(t) = \sum_{j=1}^{\infty} |u_j(t)|^2$

Energy conservation: $a + b + c = 0$

Initial energy: $E(0) < \infty$

Forcing $f_j = F\phi_j$ with $\phi_j = 0$ if $j < j_f$ and $j > j_{\max}$

Characteristic and maximum wave numbers: $k_f = k_0 \lambda^{j_f}$ and $k_{\max} = k_0 \lambda^{j_{\max}}$

r.m.s. velocity $U = \sqrt{\langle E \rangle_T}$ with $\langle \cdot \rangle_T = \frac{1}{T} \int_0^T \cdot dt$

$$Re = \frac{U}{\nu k_f}$$

$$Gr = \frac{|F|}{\nu^2 k_f^3}$$

Numerics $a = 1, b = c = -1/2, k_0 = 2^{-4}, \lambda = 2$

$$f_j = F\delta_{j,1}, F = 5 \times 10^{-3}(1 + i), 10^{-7} \leq \nu \leq 6 \times 10^2, 8 \leq N \leq 27$$

slaved Adams–Bashforth [Pisarenko, Biferale, Courvoisier, Frisch, Vergassola *Phys. Fluids* (1993)]

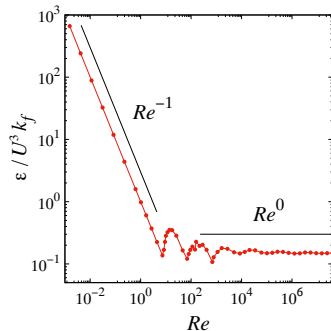
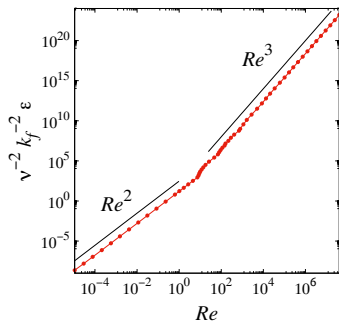
- ▶ Adapted from the study of the 3D Navier–Stokes equations
Doering & Gibbon *Applied Analysis of the Navier–Stokes equations*, 1995
- ▶ Use the estimates for the lower-order derivatives to obtain estimates for the higher-order ones

Energy dissipation rate

$$\epsilon = \nu \left\langle \sum_{j=1}^{\infty} k_j^2 |u_j|^2 \right\rangle_T$$

$$\epsilon \leq \nu^3 k_f^4 (c_1 Re^2 + c_2 Re^3)$$

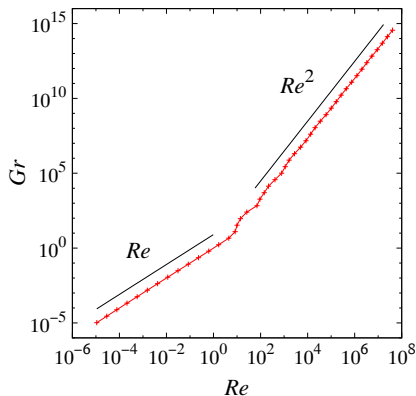
The constants c_1, c_2 depend on a, b, c, λ and the 'shape' of the forcing, but are uniform in $\nu, k_0, k_f, k_{\max}, |F|$.



3D Navier-Stokes equations \rightarrow Doering and Foias *J. Fluid Mech.* (2005)

Grashof vs Reynolds relation

$$Gr \leq c_1 Re + c_2 Re^2$$



3D Navier–Stokes equations → Doering and Foias *J. Fluid Mech.* (2005)

Higher-order derivatives

$$H_n = \sum_{j=0}^{\infty} k_j^{2n} |u_j|^2, \quad n \geq 1$$

$$H_0 = E$$
$$\langle H_1 \rangle_T = \epsilon/\nu$$

Ladder inequalities

$$\frac{1}{2} \frac{dH_n}{dt} \leq -\nu H_{n+1} + c_n H_n \sup_{j \geq 1} (k_j |u_j|) + H_n^{1/2} \Phi_n^{1/2} \quad (n \geq 0)$$

where $\Phi_n = \sum_{j=0}^{\infty} k_j^{2n} |f_j|^2$

Higher-order derivatives

$$H_n = \sum_{j=0}^{\infty} k_j^{2n} |u_j|^2, \quad n \geq 1$$

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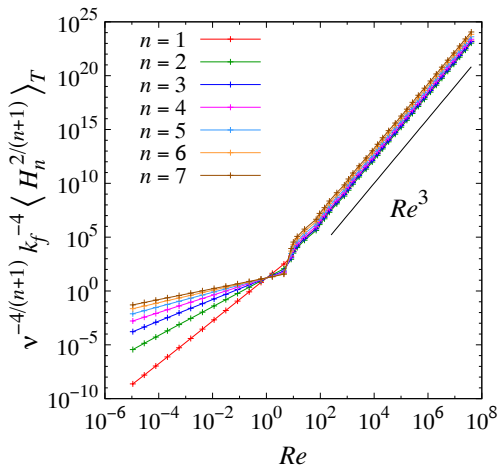
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where $\Phi_n = \sum_{j=0}^{\infty} k_j^{2n} |f_j|^2$

3D Navier-Stokes equations → Bartuccelli, Doering, Gibbon *Nonlinearity* (1991)

$$H_n = \|\nabla^n \mathbf{u}\|_2^2, \quad \frac{1}{2} \dot{H}_n \leq -\nu H_{n+1} + c_n H_n \|\nabla \mathbf{u}\|_{\infty} + H_n^{1/2} \Phi_n^{1/2}, \quad n \geq 1$$

Higher-order velocity derivatives



$$H_n = \sum_{j=1}^{\infty} k_j^{2n} |u_j|^2$$

For $n \geq 1$ and $Re \gg 1$

$$\langle H_n^{2/(n+1)} \rangle_T \leq c_n \nu^{4/(n+1)} k_f^4 Re^3$$

Higher-order velocity derivatives

Shell model

$$H_n = \sum_{j=1}^{\infty} k_j^{2n} |u_j|^2$$

$$\left\langle H_n^{\frac{2}{n+1}} \right\rangle_T \leq c_n \nu^{\frac{4}{n+1}} k_f^4 Re^3$$

3D Navier–Stokes

Foias, Guillopé, Temam
Comm. Part. Differ. Equat. (1981)

$$H_n = \|\nabla^n \mathbf{u}\|_2^2 = \int_V |\nabla^n \mathbf{u}|^2 \, d\mathbf{x}$$

$$\left\langle H_n^{\frac{2}{2n-1}} \right\rangle_T \leq c_n L^{-1} \nu^{\frac{2}{2n-1}} Re^3$$

In the shell model, the higher-order derivatives of the velocity display a weaker dependence on Re compared to the 3D Navier–Stokes equations

Absorbing ball for H_n

Alternative ladder inequality

$$\begin{aligned}\frac{1}{2} \frac{dH_n}{dt} &\leq -\frac{\nu}{2} H_{n+1} + \frac{d_n}{\nu} H_n \sup_j |u_j|^2 + H_n^{\frac{1}{2}} \Phi_n^{\frac{1}{2}} \\ &\leq -2H_n \left[\frac{\nu}{2} \frac{H_n^{\frac{1}{n}}}{H_0^{\frac{1}{n}}} - \frac{d_n}{\nu} H_0 - \frac{\Phi_n^{\frac{1}{2}}}{H_n^{\frac{1}{2}}} \right]\end{aligned}$$

By using $\Phi_n = C_n \nu^4 k_f^{2n+6} Gr^2$ and

$$\overline{\lim}_{t \rightarrow \infty} H_0 \leq \nu^2 \left(\frac{k_f}{k_1} \right)^4 k_f^2 Gr^2 \quad (\text{Levant, Constantin, Titi, } \textit{Physica D}, 2006)$$

$$\overline{\lim}_{t \rightarrow \infty} H_n \leq \nu^2 k_f^{2(n+1)} \left[c_n Gr^{2(n+1)} + d_n Gr^2 \right]$$

Higher-order moments

Shell model

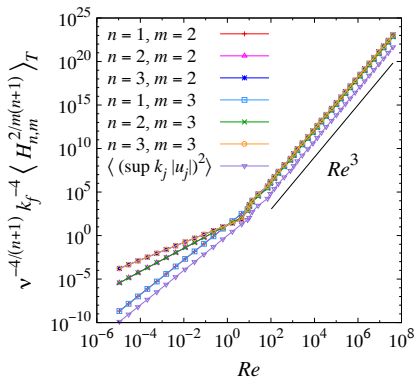
$$H_{n,m} = \sum_{j=1}^{\infty} (k_j^n |u_j|)^{2m}, \quad H_{n,1} = H_n$$

$$H_{n,m} \leq H_n^m$$

For $n, m \geq 1$ and $Re \gg 1$

$$\left\langle H_{n,m}^{\beta_{n,m}^{\text{sh}}} \right\rangle_T \leq c_n \nu^{\frac{4}{n+1}} k_f^4 Re^3$$

$$\beta_{n,m}^{\text{sh}} = \frac{2}{m(n+1)}$$



Higher-order moments

Shell model

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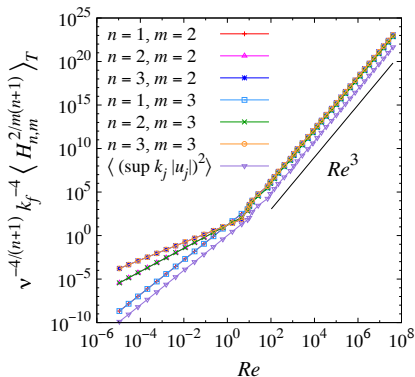
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$$\beta_{n,m}^{\text{sh}} = \frac{2}{m(n+1)}$$

$$\left\langle \left(\sup_j k_j^n |u_j| \right)^{\frac{4}{n+1}} \right\rangle \leq \hat{c}_n \nu^{\frac{4}{(n+1)}} k_f^4 Re^3$$



Higher-order moments

Shell model

$$H_{n,m} = \sum_{j=1}^{\infty} k_j^{2nm} |u_j|^{2m}$$

$$H_{n,m} \leq H_n^m$$

For $n, m \geq 1$ and $Re \gg 1$

$$\langle H_{n,m}^{\text{sh}} \rangle_T \leq c_n \nu^{\frac{4}{n+1}} k_f^4 Re^3$$

$$\beta_{n,m}^{\text{sh}} = \frac{2}{m(n+1)}$$

3D Navier–Stokes

$$H_{n,m} = \|\nabla^n \mathbf{u}\|_{2m}^{2m} = \int_V |\nabla^n \mathbf{u}|^{2m} dx$$

Gibbon *J. Nonlin. Sci.* (2019)

For $1 \leq n, 1 \leq m \leq \infty$ and $Re \gg 1$

$$\langle H_{n,m}^{\text{ns}} \rangle_T \leq c_n L^{-1} \nu^{2m\beta_{n,m}^{\text{ns}}} Re^3$$

$$\beta_{n,m}^{\text{ns}} = \frac{1}{2m(n+1) - 3}$$

$$\beta_{n,m}^{\text{sh}} > \beta_{n,m}^{\text{ns}} \text{ for all } n > 1$$

Velocity gradient averaged NSEs

Shell model

$$\frac{1}{2} \dot{H}_n \leq -\nu H_{n+1} + c_n H_n \sup_j (k_j |u_j|) + H_n^{1/2} \Phi_n^{1/2}$$

3D Navier–Stokes

$$\frac{1}{2} \dot{H}_n \leq -\nu H_{n+1} + c_n H_n \|\nabla \mathbf{u}\|_\infty + H_n^{1/2} \Phi_n^{1/2}$$

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$$\sup_{j \geq 1} (k_j |u_j|) \leq \left[\sum_{j=1}^{\infty} (k_j |u_j|)^2 \right]^{\frac{1}{2}} = H_1^{\frac{1}{2}}$$

3D Navier–Stokes

$$\frac{1}{2}\dot{H}_n \leq -\nu H_{n+1} + c_n H_n \|\nabla \mathbf{u}\|_{\infty} + H_n^{1/2} \Phi_n^{1/2}$$

$$\|\nabla \mathbf{u}\|_{\infty} \leq c_n H_{n+1}^{\frac{3}{4n}} H_1^{\frac{2n-3}{4n}}$$

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$$\|\nabla \mathbf{u}\|_{\infty} \leq c_n H_{n+1}^{\frac{3}{4n}} H_1^{\frac{2n-3}{4n}}$$

Velocity gradient averaged Navier–Stokes equations

Assumption: $\|\nabla \mathbf{u}\|_{\infty} \approx c L^{-3/2} \|\nabla \mathbf{u}\|_2 = c \langle |\nabla \mathbf{u}|^2 \rangle_V^{1/2}$

$$\frac{1}{2}\dot{H}_n \leq -\nu H_{n+1} + c_n L^{-3/2} H_n H_1^{1/2} + H_n^{1/2} \Phi_n^{1/2}$$

$$\beta_{n,m}^{\text{sh}} = \beta_{n,m}^{\text{vga}} \text{ for all } n, m \geq 1$$

Moments of the energy

3D NSEs (Gibbon, *JNLS*, 2019)

$$H_{0,m} = \|\mathbf{u}\|_{2m}^{2m} = \int_V |\mathbf{u}|^{2m} d\mathbf{x} \quad \left\langle H_{0,m}^{\frac{1}{2m-3}} \right\rangle_T \leq c_m L^{-1} \nu^{\frac{2m}{2m-3}} Re^3 \quad (3 < m \leq \infty)$$

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3D NSEs (Gibbon, *JNLS*, 2019)

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VGA NSEs

$$\|\mathbf{u}\|_{2m} \leq cL \|\nabla \mathbf{u}\|_{2m} \leq cL^{\frac{2m+3}{2m}} \|\nabla \mathbf{u}\|_{\infty}$$

$$\|\nabla \mathbf{u}\|_{\infty} \approx cL^{-3/2} \|\nabla \mathbf{u}\|_2 \implies \|\mathbf{u}\|_{2m} \leq cL^{-\frac{m-3}{2m}} \|\nabla \mathbf{u}\|_2$$

$$\left\langle H_{0,m}^{\frac{1}{m}} \right\rangle_T \leq c_m L^{-\frac{2m-3}{m}} \nu^2 Re^3 \quad (1 \leq m \leq \infty)$$

Moments of the energy

3D NSEs (Gibbon, JNLS, 2019)

$$H_{0,m} = \|\mathbf{u}\|_{2m}^{2m} = \int_V |\mathbf{u}|^{2m} \mathbf{d}\mathbf{x} \quad \left\langle H_{0,m}^{\frac{1}{2m-3}} \right\rangle_T \leq c_m L^{-1} \nu^{\frac{2m}{2m-3}} Re^3 \quad (3 < m \leq \infty)$$

VGA NSEs

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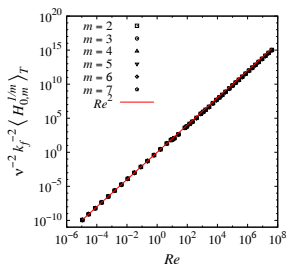
Shell model

$$H_{0,m} = \sum_{j=1}^{\infty} |u_j|^{2m} \leq H_0^m$$

$$\left\langle H_{0,m}^{\frac{1}{m}} \right\rangle_T \leq \nu^2 k_f^2 Re^2$$

$$(1 \leq m \leq \infty)$$

Equivalent to assuming $\|\mathbf{u}\|_{2m} \approx c_m L^{\frac{3(m-1)}{2m}} \|\mathbf{u}\|_2$



Shell models as NSEs “on a point”

Navier–Stokes equations in d dimensions (Gibbon *EPL*, 2020)

$$\left\langle H_{n,m}^{\beta_{n,m,d}^{\text{ns}}} \right\rangle_T \leq c_n L^{-1} \nu^{2m\beta_{n,m,d}^{\text{ns}}} Re^3$$

$$\beta_{n,m,d}^{\text{ns}} = \frac{(4-d)}{2m(n+1)-d}$$

Shell models as NSEs “on a point”

Navier–Stokes equations in d dimensions (Gibbon *EPL*, 2020)

$$\left\langle H_{n,m}^{\beta_{n,m,d}^{\text{ns}}} \right\rangle_T \leq c_n L^{-1} \nu^{2m\beta_{n,m,d}^{\text{ns}}} Re^3$$

$$\beta_{n,m,d}^{\text{ns}} = \frac{(4-d)}{2m(n+1)-d}$$

Shell model

$$\beta_{n,m}^{\text{sh}} = \frac{2}{m(n+1)} = \lim_{d \rightarrow 0} \beta_{n,m,d}^{\text{ns}}$$

Conclusions

- ▶ The estimates of the energy dissipation rate in terms of Re are the same in the shell model and in the 3D NSEs
- ▶ However, the estimates of the velocity and its higher-order derivatives display a weaker dependence on Re than in the 3D NSEs
- ▶ The estimates of the velocity derivatives for the shell model are equivalent to those corresponding to a velocity gradient averaged version of the 3D NSEs

D. Vincenzi & J.D. Gibbon, arXiv:2009.01583