



The curious case of 'twin' Fast Radio Bursts

*CORTEX
(NWA.1160.18.316),
Funded by NWO*

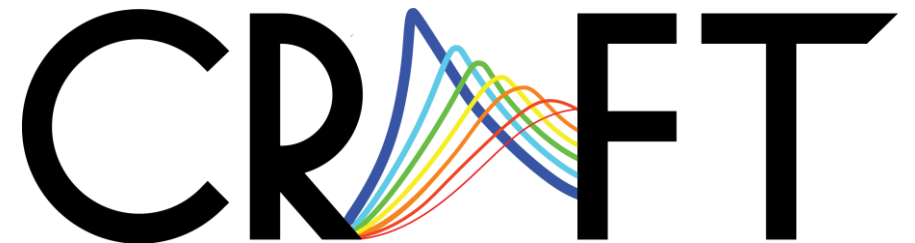
Apurba Bera

ASTRON (The Netherlands Institute for Radio Astronomy)

ICRAR / Curtin University



International
Centre for
Radio
Astronomy
Research



Fast Radio Bursts (FRBs)

- Mysterious and intense bursts of radio emission
- Short-duration ($< \sim$ tens of milliseconds)

Lorimer et al. 2007, Thornton et al. 2013

- Extragalactic origin \rightarrow useful probe of cosmology

Chatterjee et al. 2017, Bannister et al. 2019, Macquart et al. 2020

- Only a small fraction ($< 5\%$) are known to repeat
 - no pulsar-like periodicity
- Unknown source and emission mechanism
 - strongly magnetized neutron stars ?

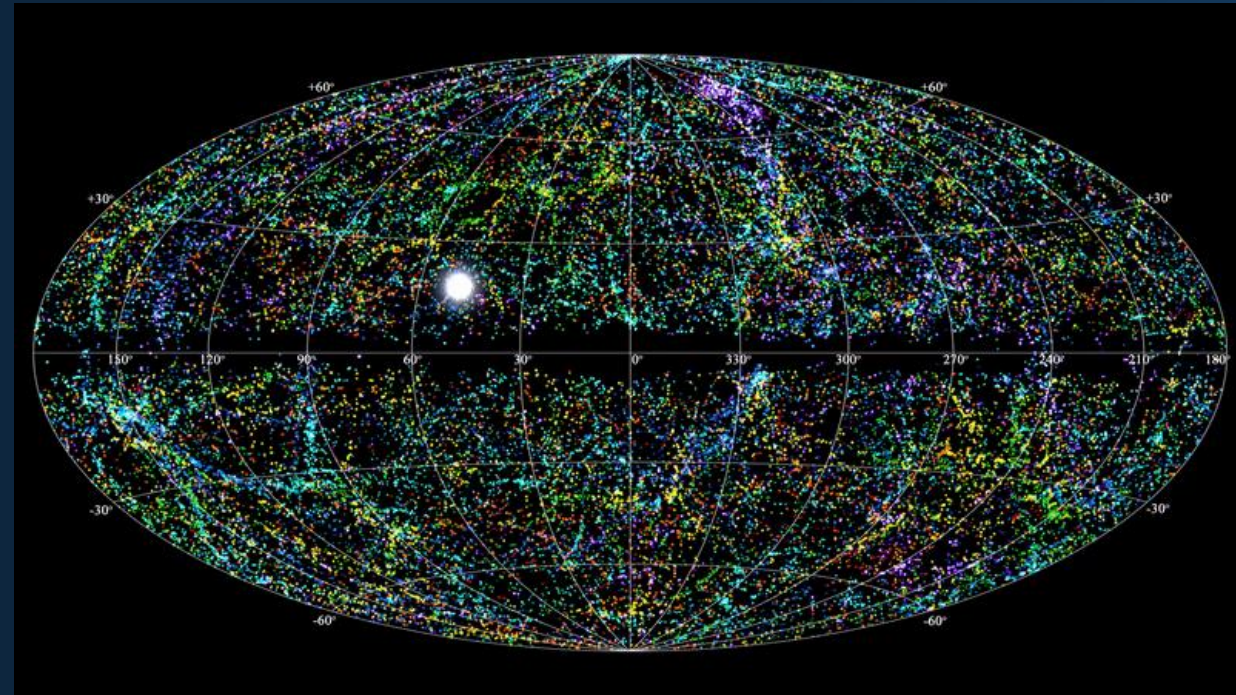


Image credit:
T. Jarrett (IPAC/Caltech)
B. Saxton (NRAO/AUI/NSF)

Shannon et al. 2018, CHIME/FRB collaboration 2021, Petroff et al. 2022

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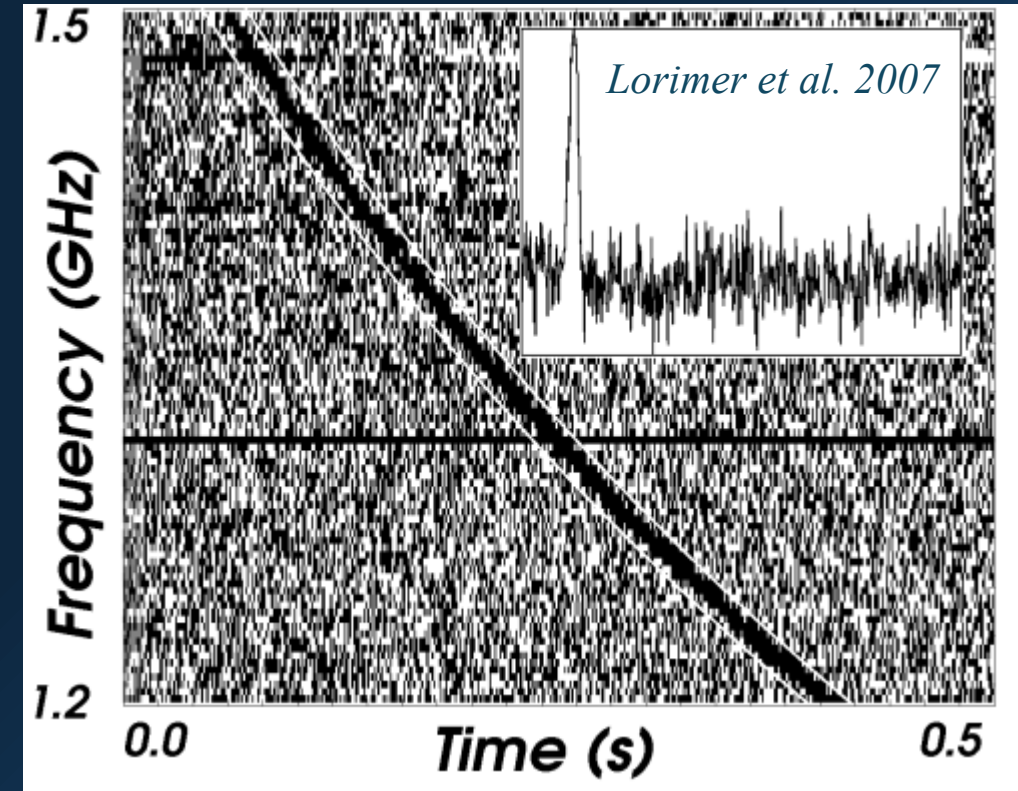
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The “Lorimer Burst”
*detected with the Murriyang
(Parkes) radio telescope*

Shannon et al. 2018, CHIME/FRB collaboration 2021, Petroff et al. 2022

The Commensal Real-time ASKAP Fast Transients survey



CRAFT team (50+ members) spans 10+ countries and 20+ institutes

The Australian SKA Pathfinder

- Located at Inyarrimanha Ilgari Bundara,
the CSIRO Murchison Radio-astronomy Observatory
- Phased Array Feed (PAF)
- Allows 36 individual beams on the sky
- Enables 30 square degrees of view!



The CRAFT survey with ASKAP

- Commensal Real-time ASKAP Fast Transient (CRAFT) Survey

- Realtime search for dispersed radio signals
- Detection triggers recording of raw voltage data

Bannister et al. 2017, Hotan et al. 2021

Qiu et al. 2023, Shannon et al. 2025

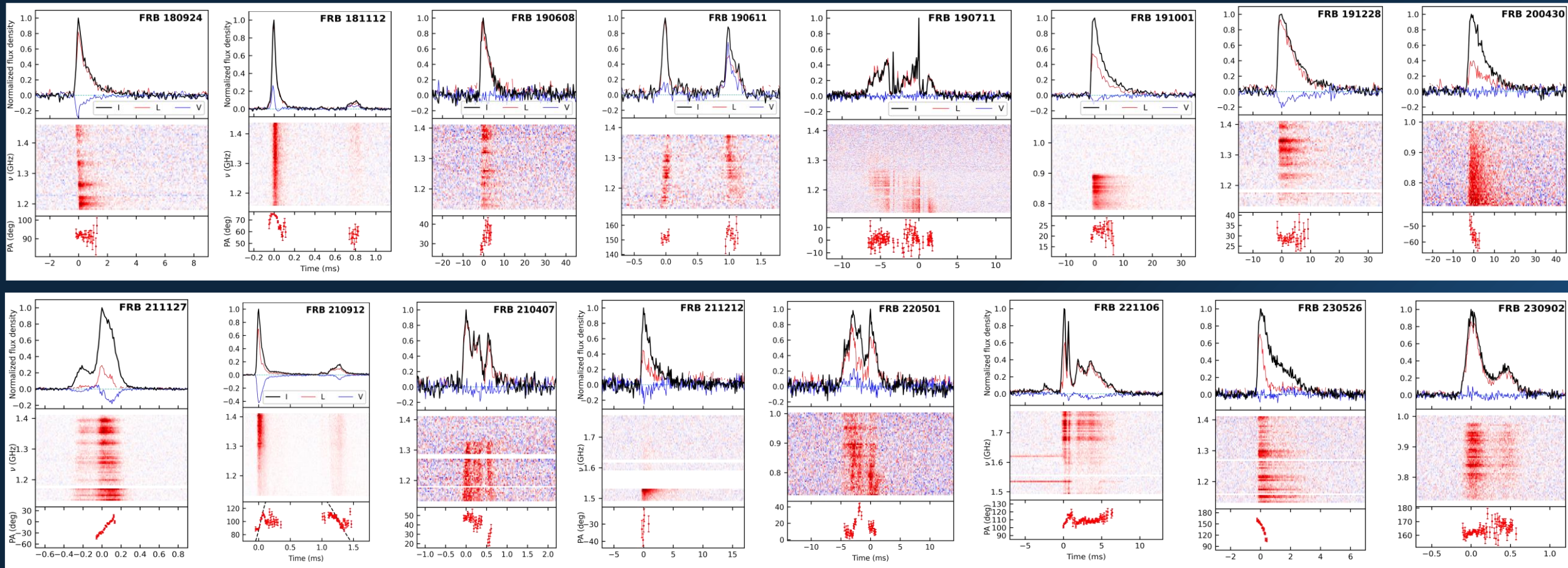
- CRAFT Effortless Localisation and Enhanced Burst Inspection (CELEBI) pipeline

- Offline correlation and imaging → sub-arcsecond localization of FRBs
- Phase coherent beamforming → full polar high time resolution output
- Reveals FRB emission properties at time scales down to ~ 3 ns !



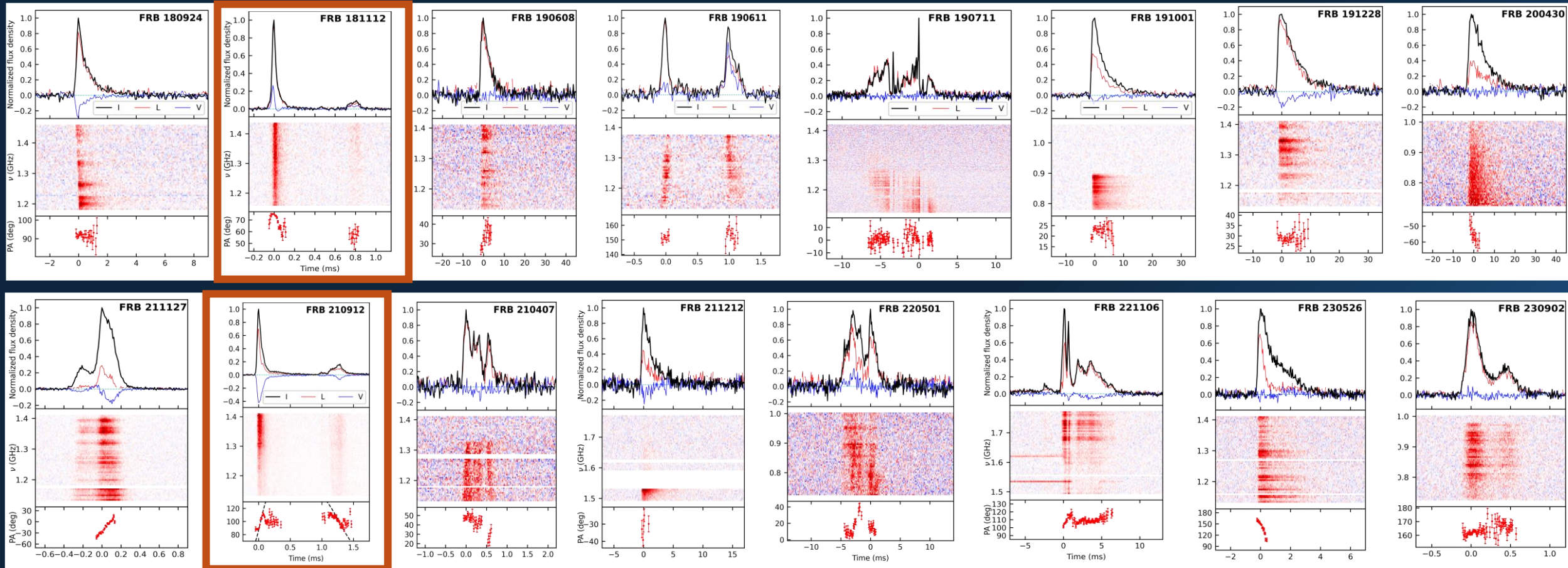
Scott et al. 2023

Diverse properties of Fast Radio Bursts



- Complex time-frequency structure and polarization behaviour ...
- Propagation effects (scattering, scintillation)

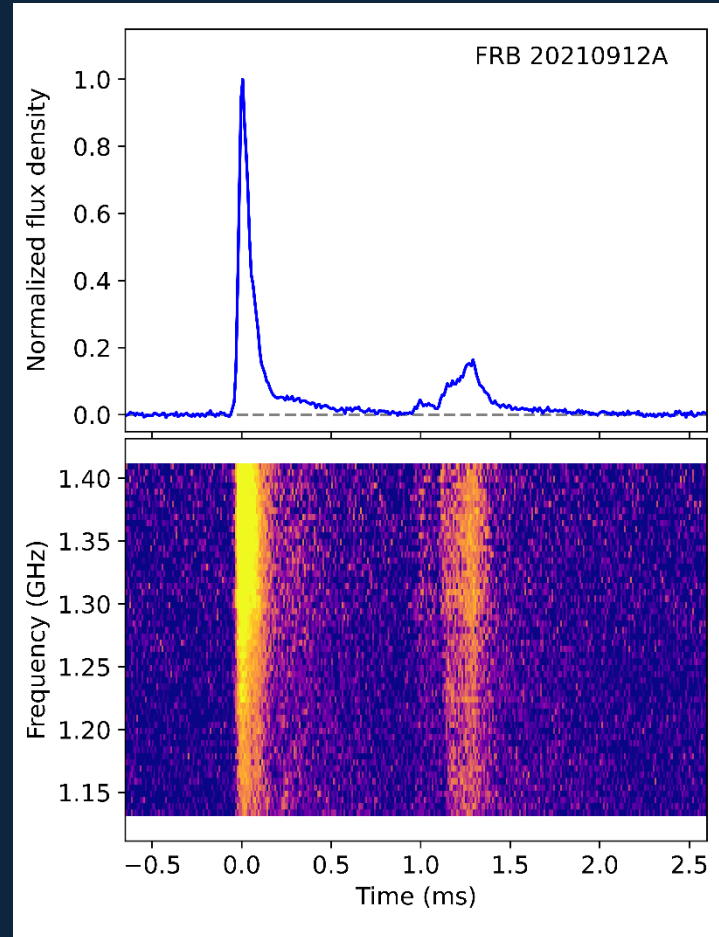
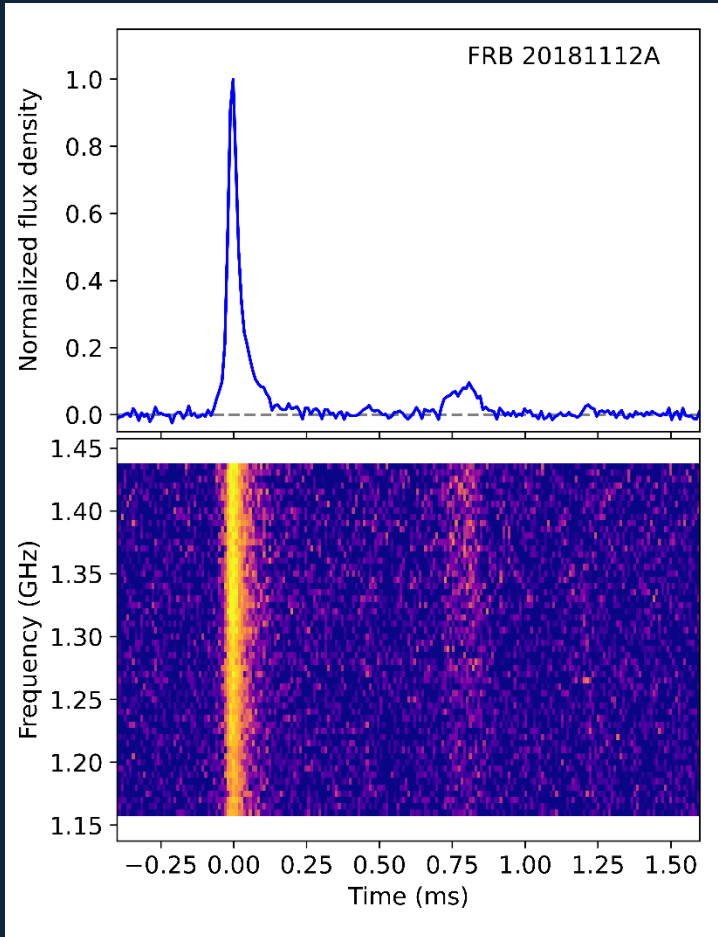
Diverse properties of Fast Radio Bursts



Scott et al. 2025

- Complex time-frequency structure and polarization behaviour ... wait !

Twin Fast Radio Bursts !



21h49m23.6s
-52d58'15.4"

R. A.
Decl.

23h23m10.4s
-30d24'19.2"

589

DM
($pc\ cm^{-3}$)

1234

26 ± 3

Fluence
(Jy ms)

70 ± 2

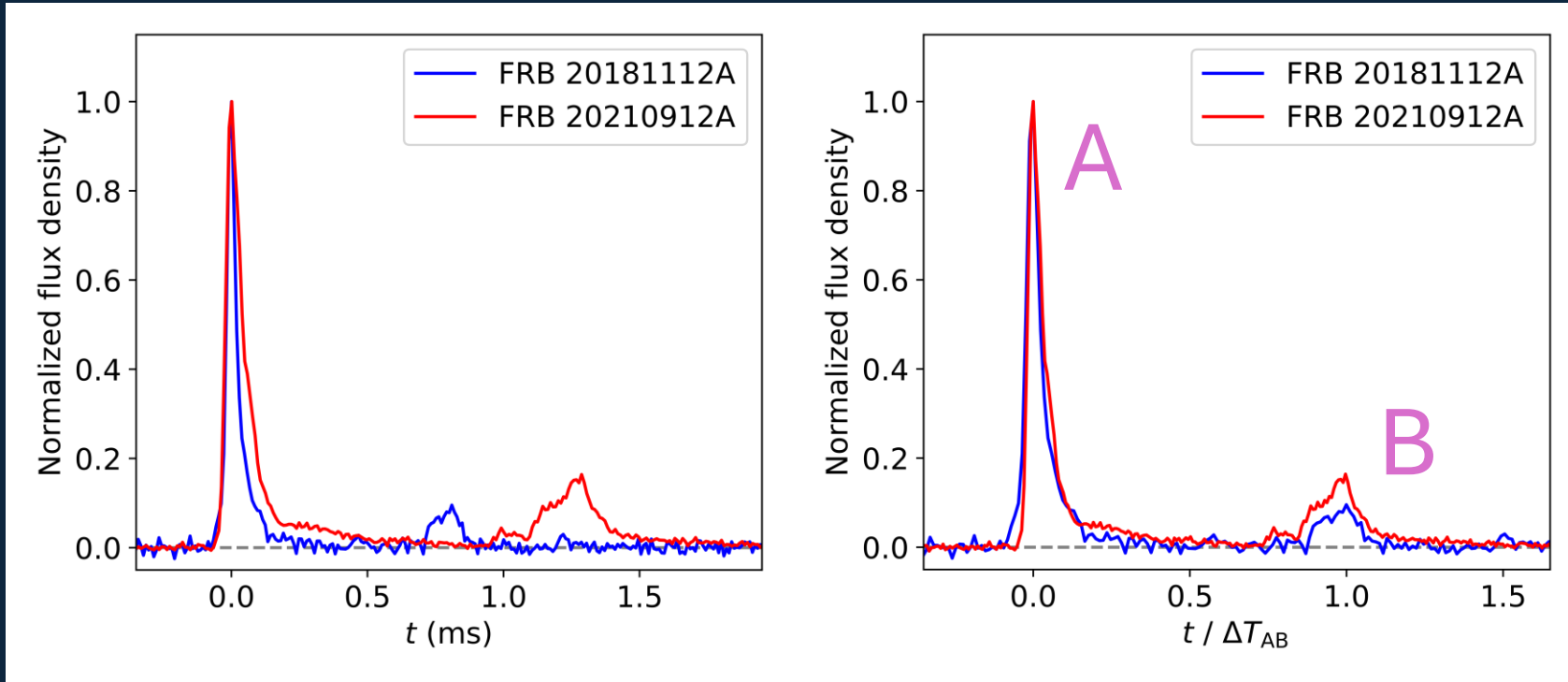
0.4755

Redshift

Unknown !

*Prochaska et al. 2019, Cho et al. 2020,
Marnoch et al. 2023, Bera et al. 2024*

Burst profiles & emission time-scales



$$0.31 \pm 0.02$$

$$FWHM_A / FWHM_B$$

$$0.32 \pm 0.01$$

$$0.81 \pm 0.06$$

$$\text{Peak separation } (\Delta T / \text{ms})$$

$$1.27 \pm 0.11$$

$$(A) 0.046 \pm 0.003$$

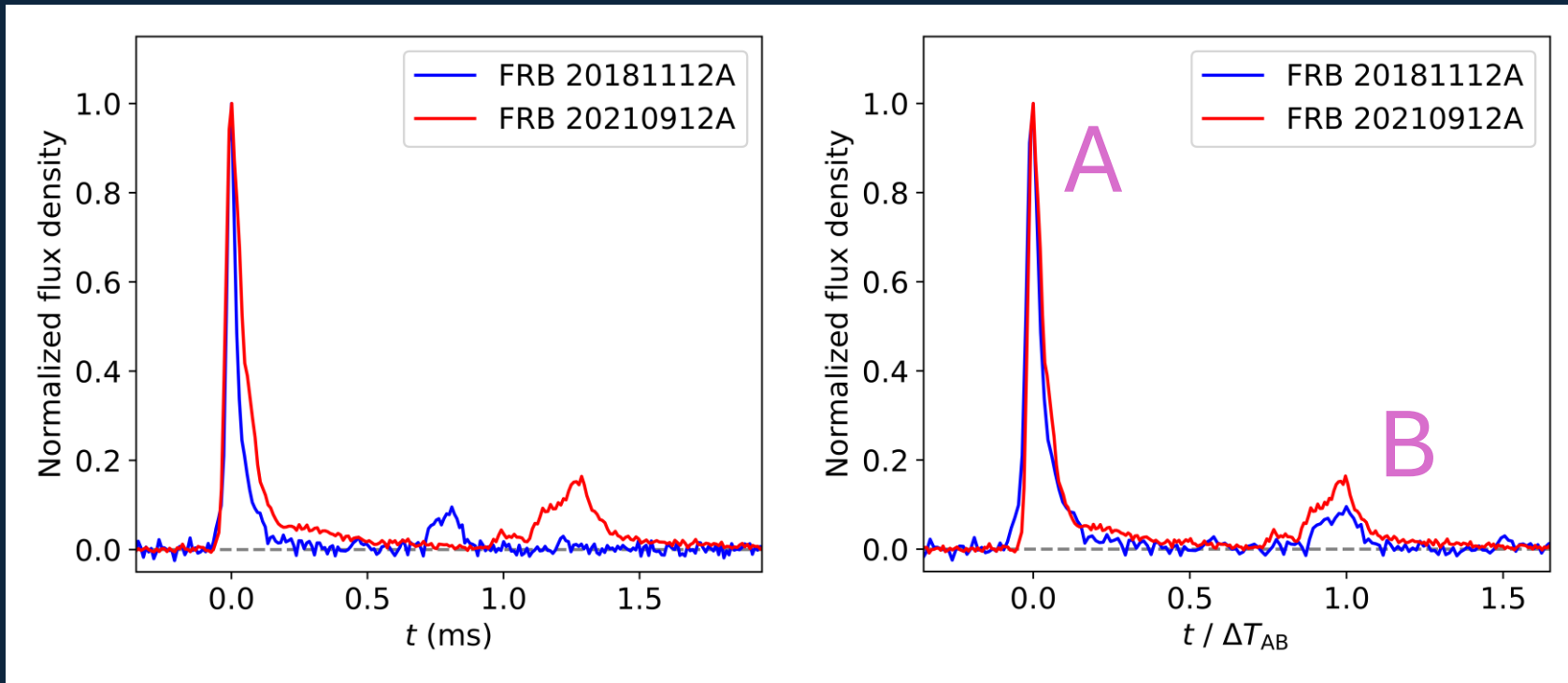
$$FWHM / \Delta T$$

$$(A) 0.052 \pm 0.005$$

$$(B) 0.148 \pm 0.007$$

$$(B) 0.16 \pm 0.01$$

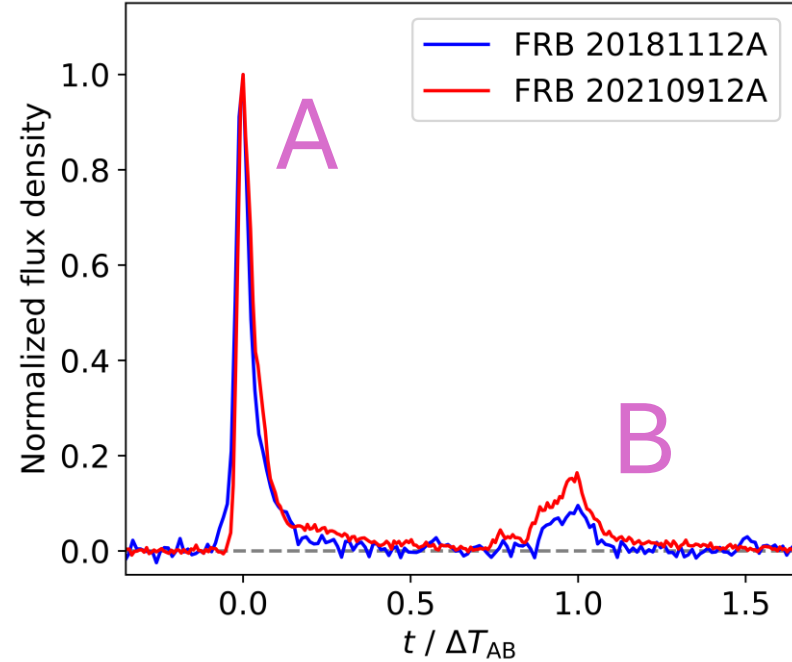
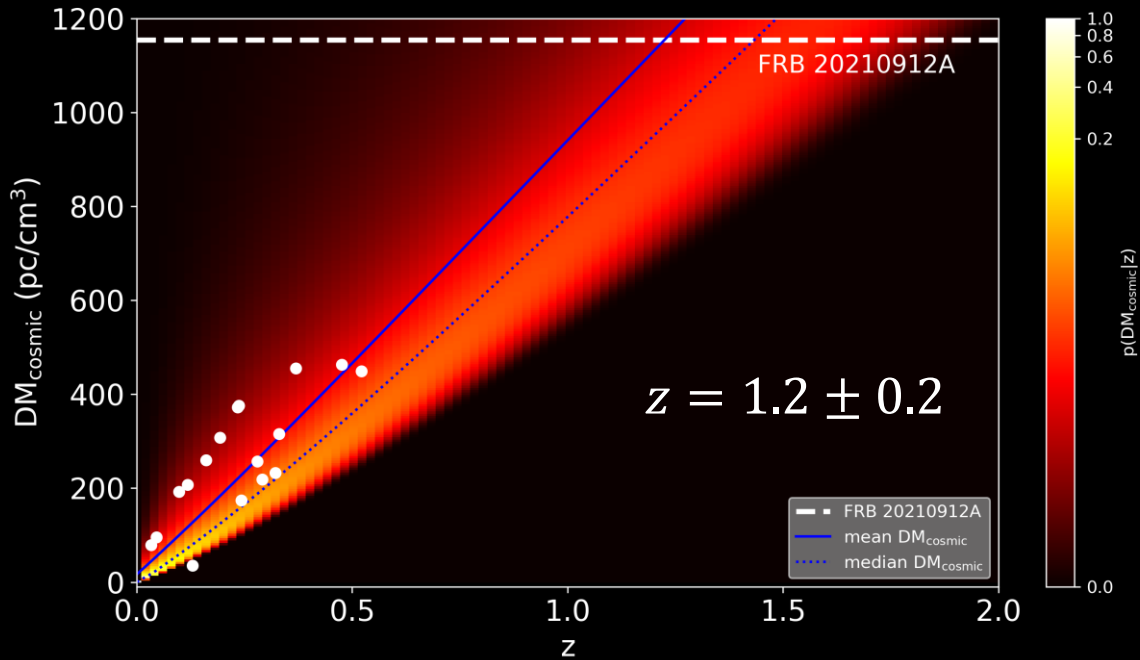
Burst profiles & emission time-scales



Identical rest-frame time-scales **if**

$$z(FRB\ 20210912) = 1.33 !$$

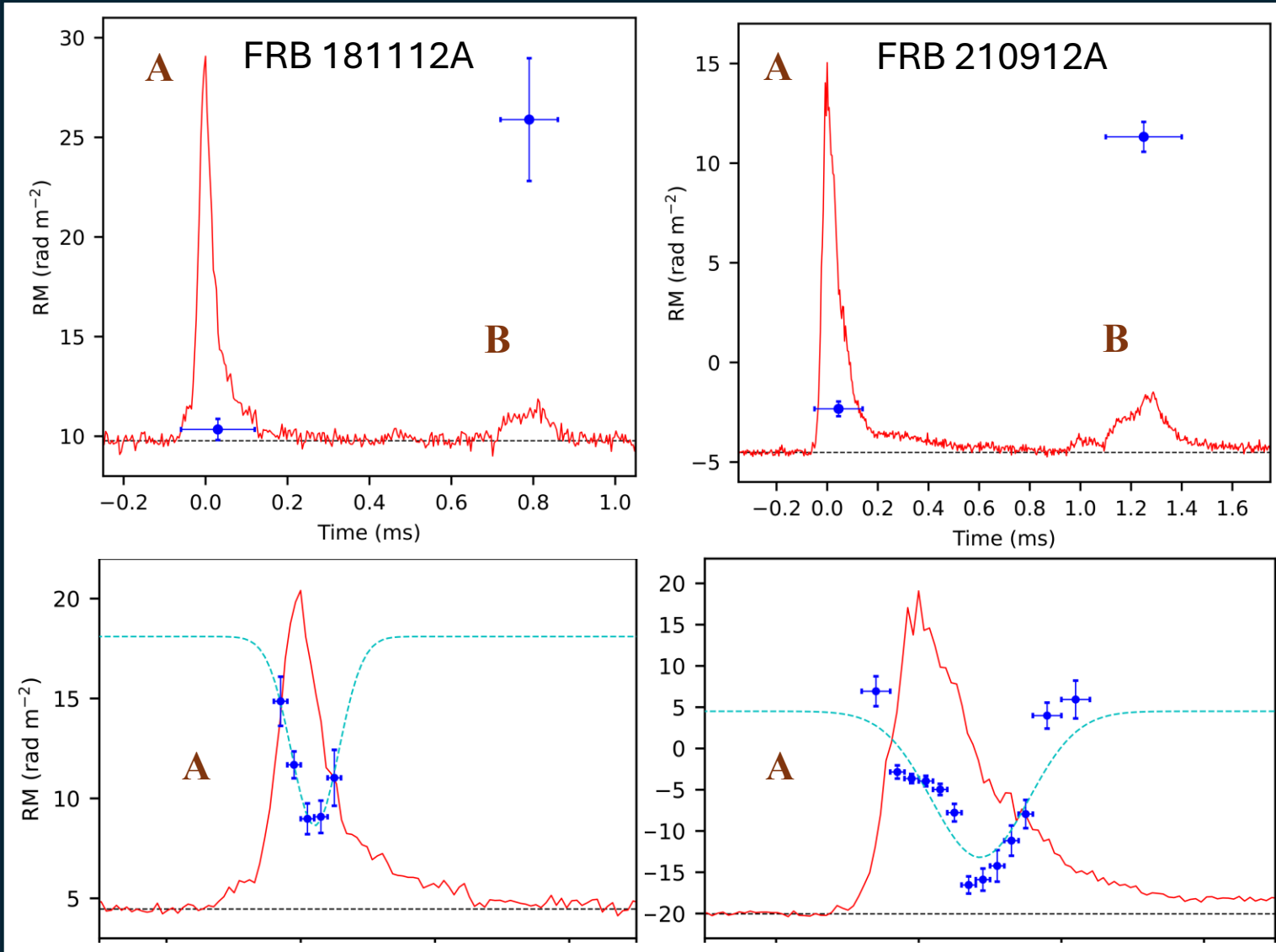
Burst profiles & emission time-scales



Identical rest-frame time-scales **if** $z(\text{FRB 20210912}) = 1.33$!

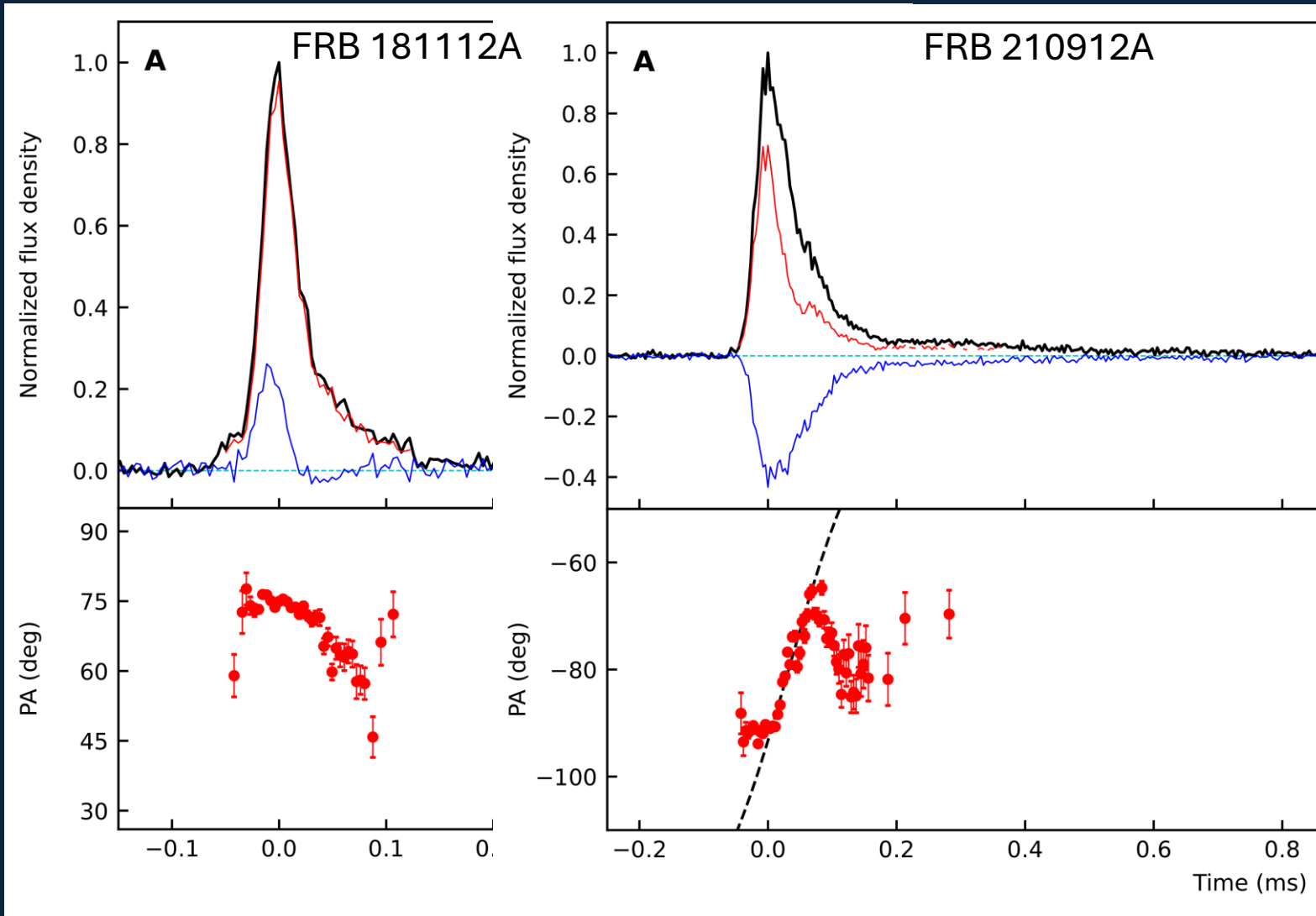
Consistent with cosmological redshift – DM relation !

Twins, but not Identical...



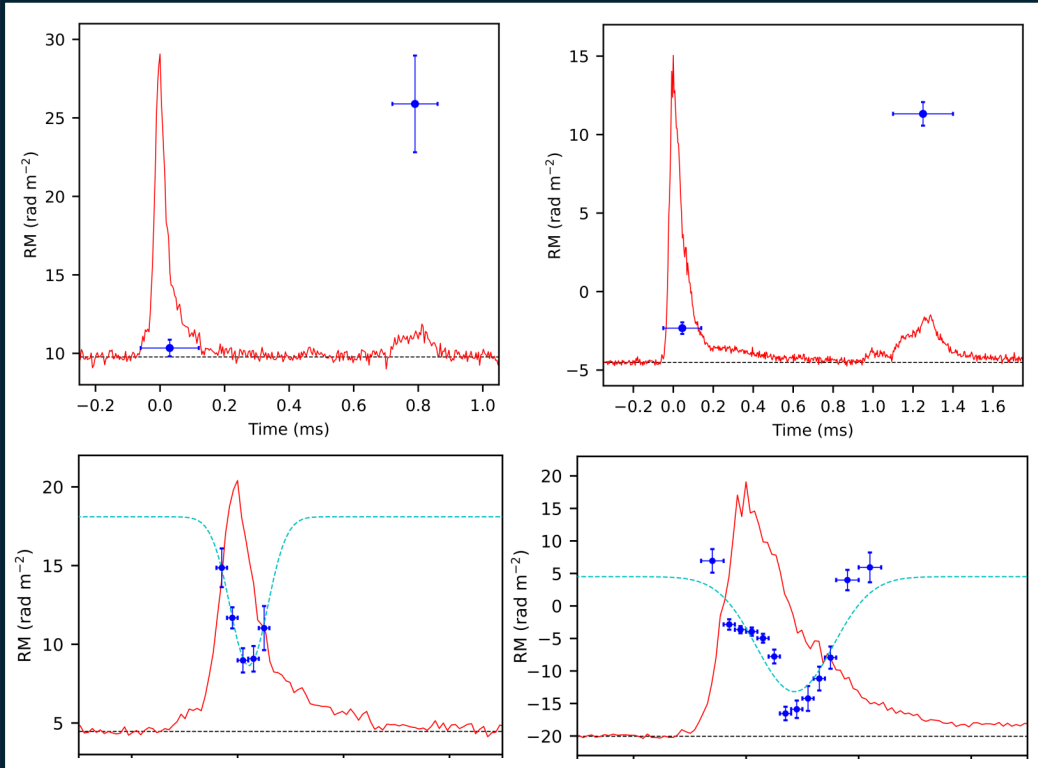
- Different Rotation measure (RM) in two sub-bursts
- $\Delta \text{RM}_{\text{AB}} \approx 15 \text{ rad m}^{-2}$
- Short time-scale ($\sim 0.01 \text{ ms}$) variation of RM
 - similar in both FRBs

Twin FRBs – Polarization profiles

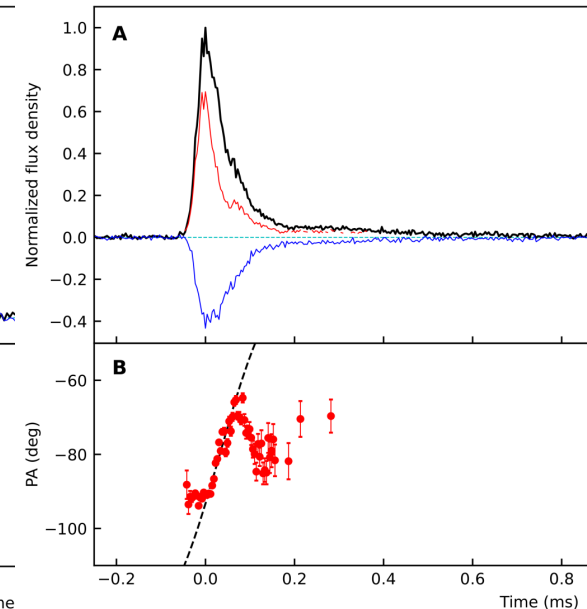
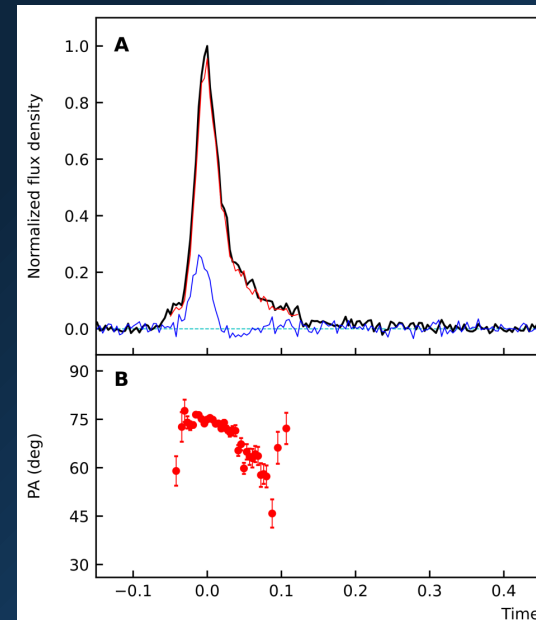
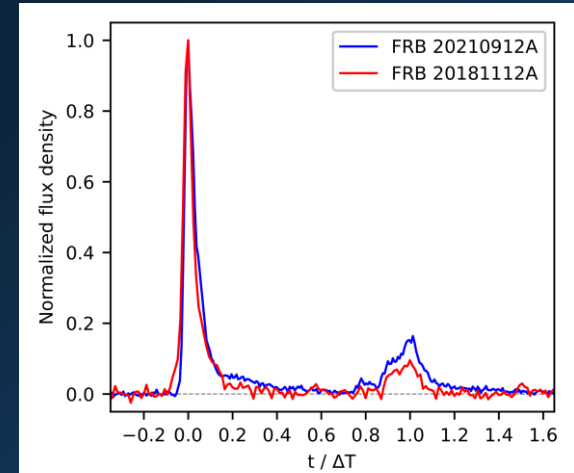


- Time varying fraction of linear and circular polarization
- Position Angle (PA) variation across primary sub-burst
- Similar in both FRBs

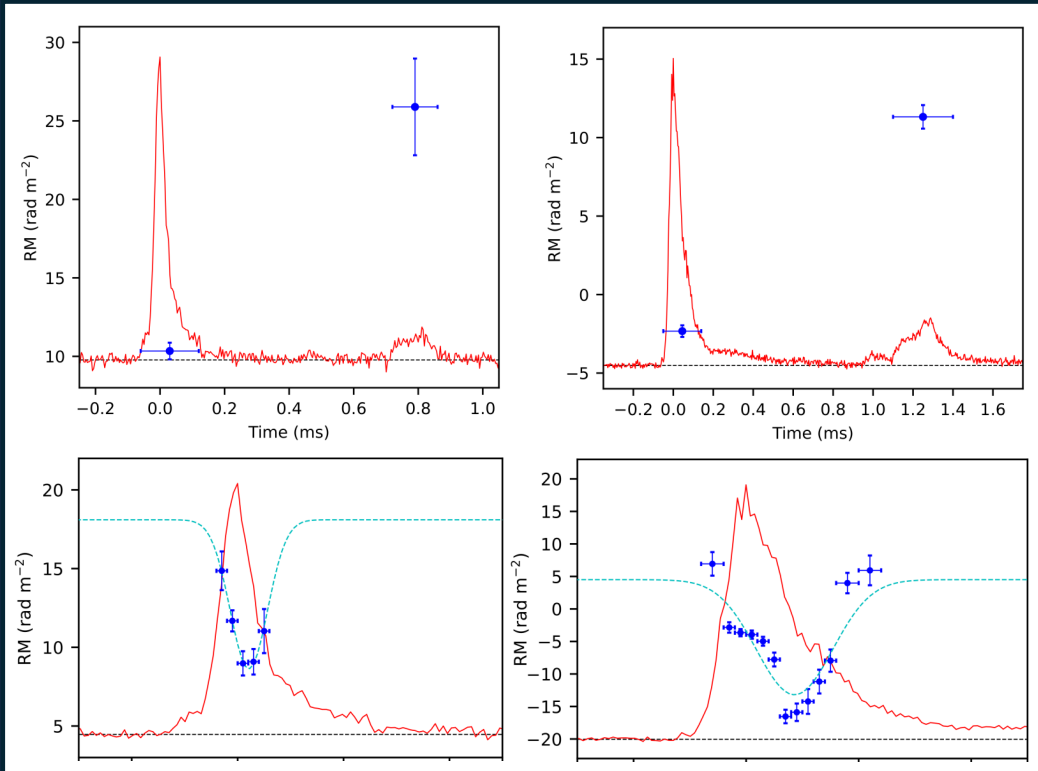
Twin FRBs – A random coincidence ?



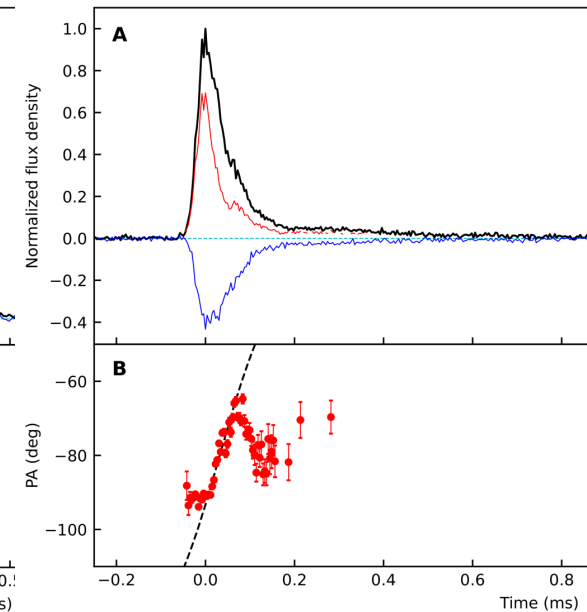
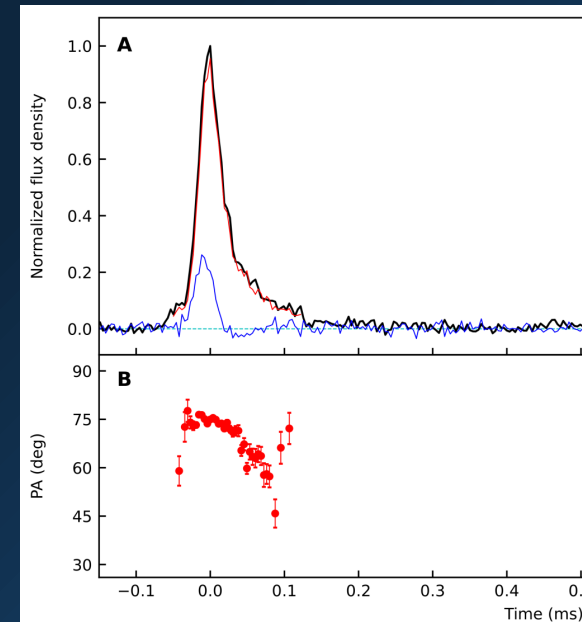
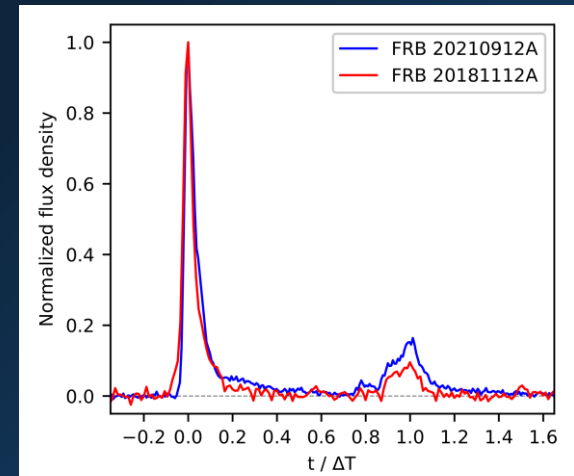
- Similarities are hard to quantify !
- Difficult to estimate probability of random coincidence



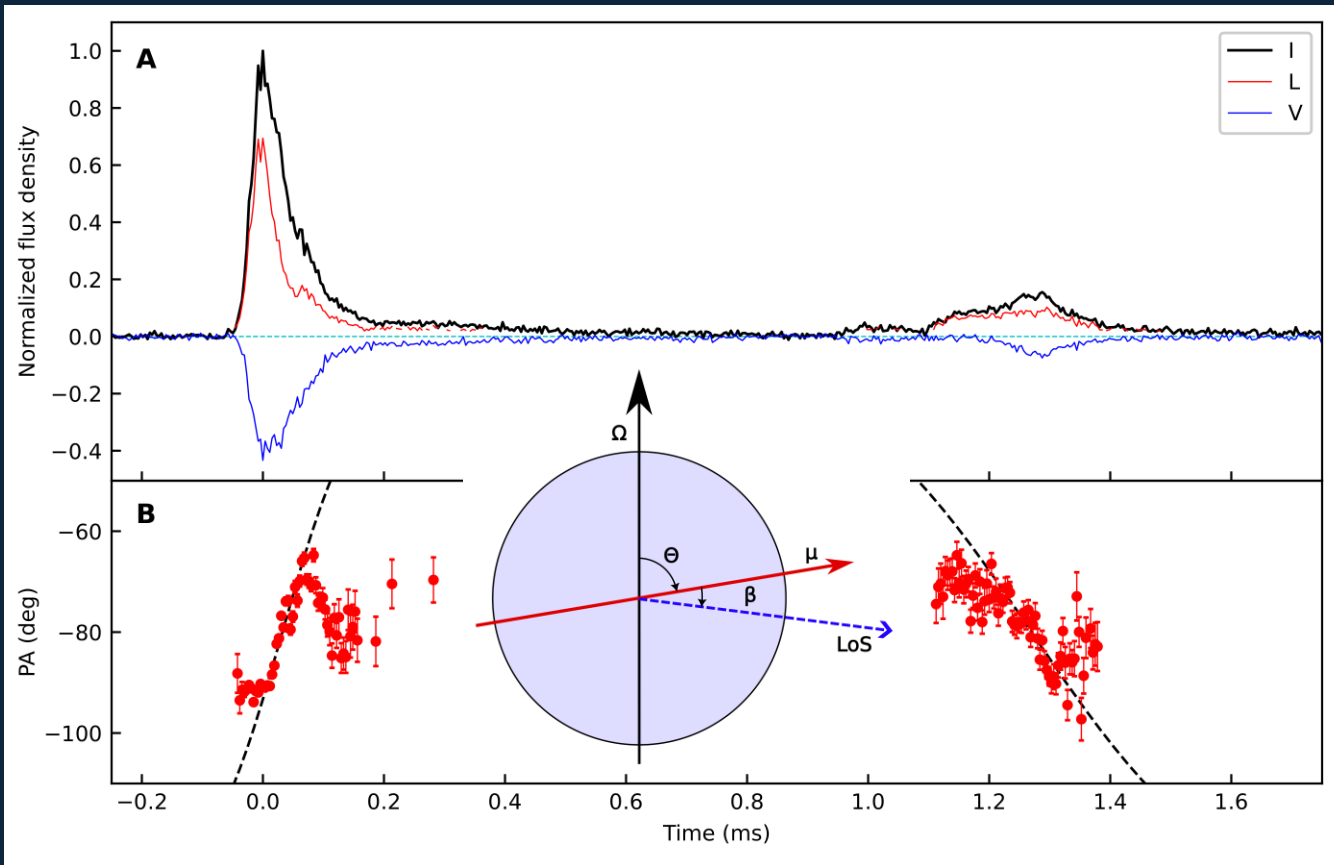
Or... similar progenitors ?



- What could be their progenitors ?



PA variation in FRB 210912A



- Rotating Vector Model for PA variation
 - PA traces projection of magnetic field on to the sky plane

Radhakrishnan & Cooke 1969

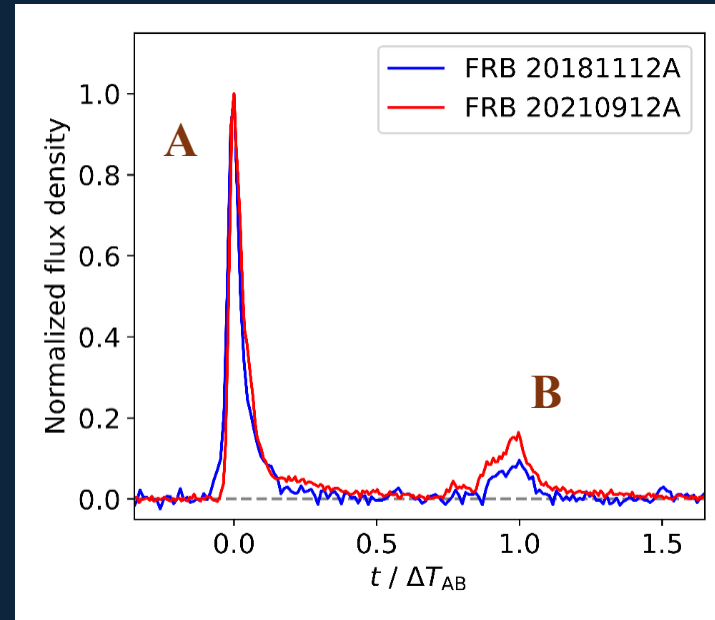
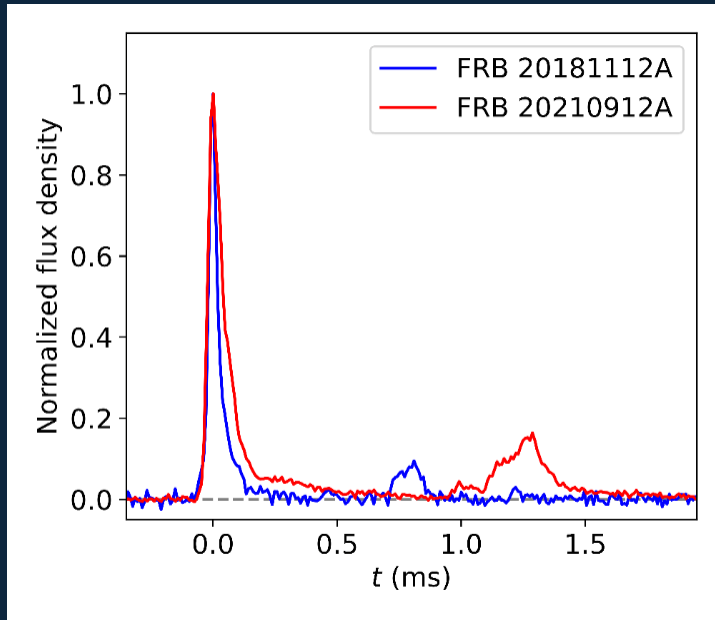
- PA variation can be broadly described by simple (dipolar) RVM with

magnetic obliquity $\rightarrow \theta \approx 59^\circ$

impact angle $\rightarrow \beta \approx 17^\circ$

Two neutron stars with rest-frame spin periods of ≈ 1.1 ms ?

Rest frame emission time-scales ?



• FRB 210912A

Host galaxy not yet found !

z-DM relation $\longrightarrow z \sim 1.2 \pm 0.2$

Intrinsic time scales would be identical if

$$z = 1.33 !$$

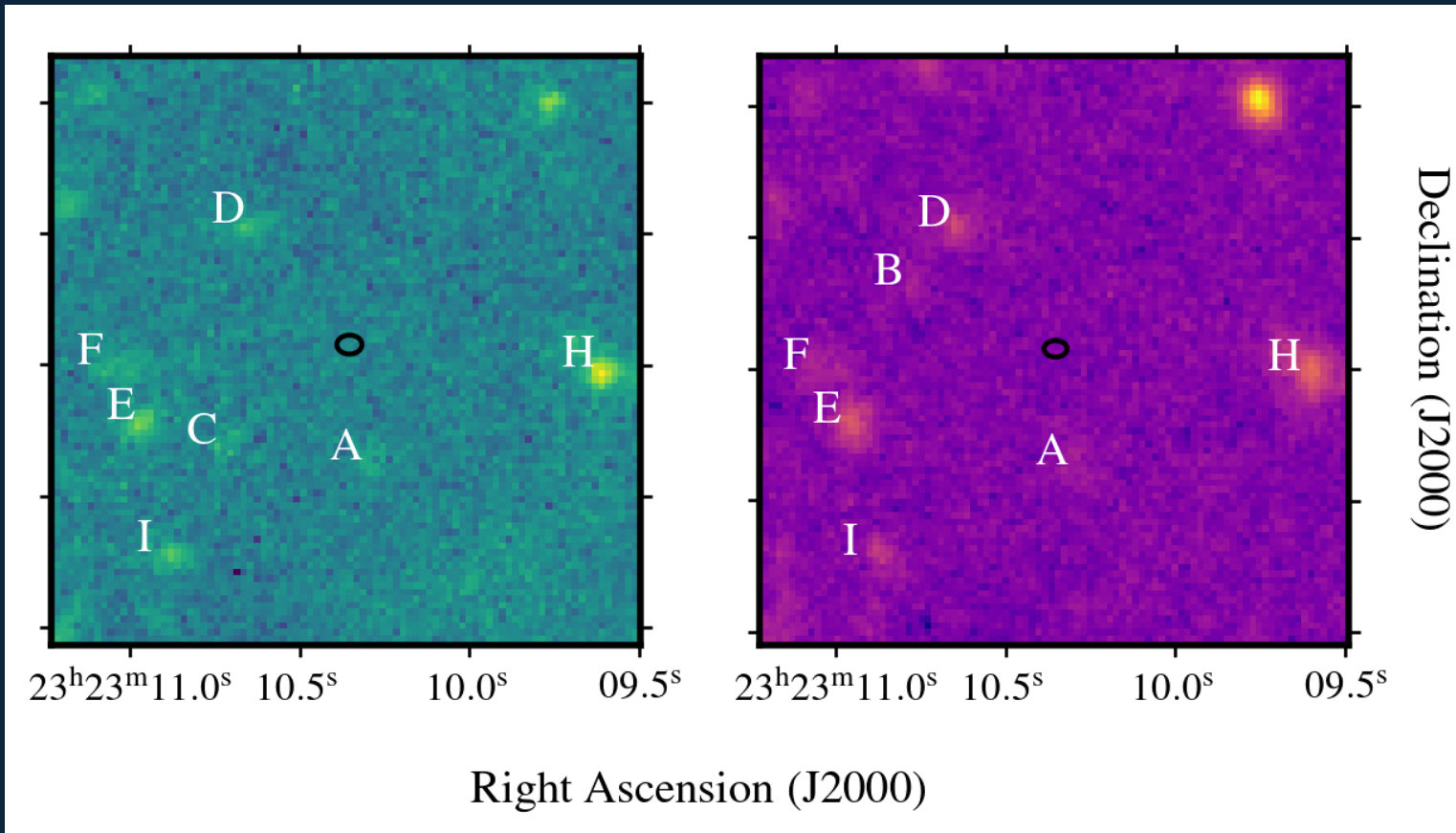
Find the host galaxy of FRB 210912A and measure its redshift !

Bera et al. 2024

Marnoch et al. 2023

Macquart et al. 2020

Finding the host galaxy...



Not found in very deep VLT image

$R > 26.7$ and $K > 24.9$ (5σ limits)

JWST image

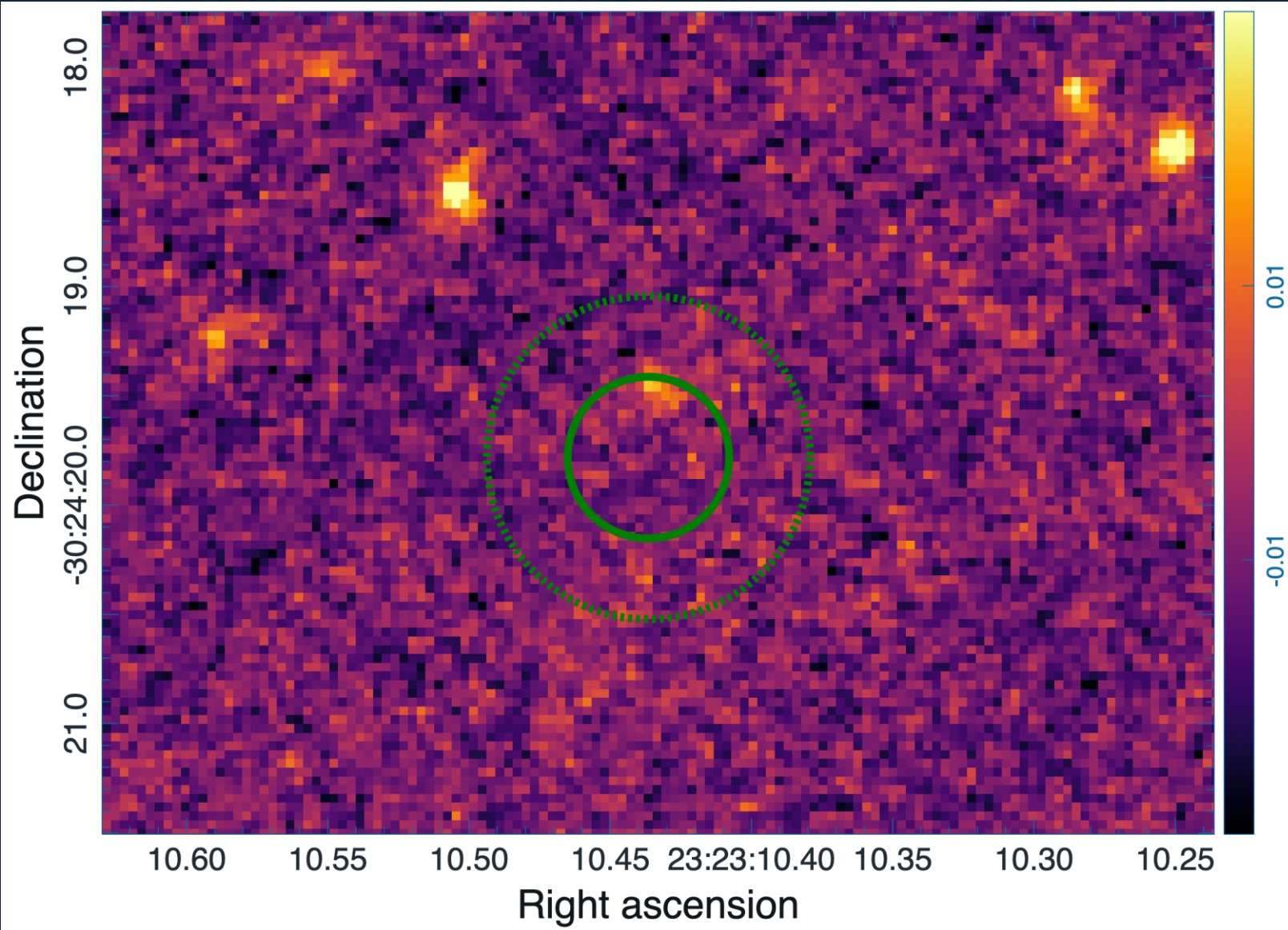
JWST NIRCAM
(F200W+F322W)

Localization

Solid $\rightarrow 1\sigma$ Dashed $\rightarrow 2\sigma$

Faintest FRB host so far !

Waiting for spectroscopy
results...



*Thanks to
Adam Deller
Lalitwadee Kawinwanichakij
Themiya Nanayakkara
Karl Glazebrook*

JWST image

JWST NIRCAM
(F200W+F322W)

Localization

Solid $\rightarrow 1\sigma$ Dashed $\rightarrow 2\sigma$

Faintest FRB host so far !

If it gives

$$z = 1.33 \dots$$

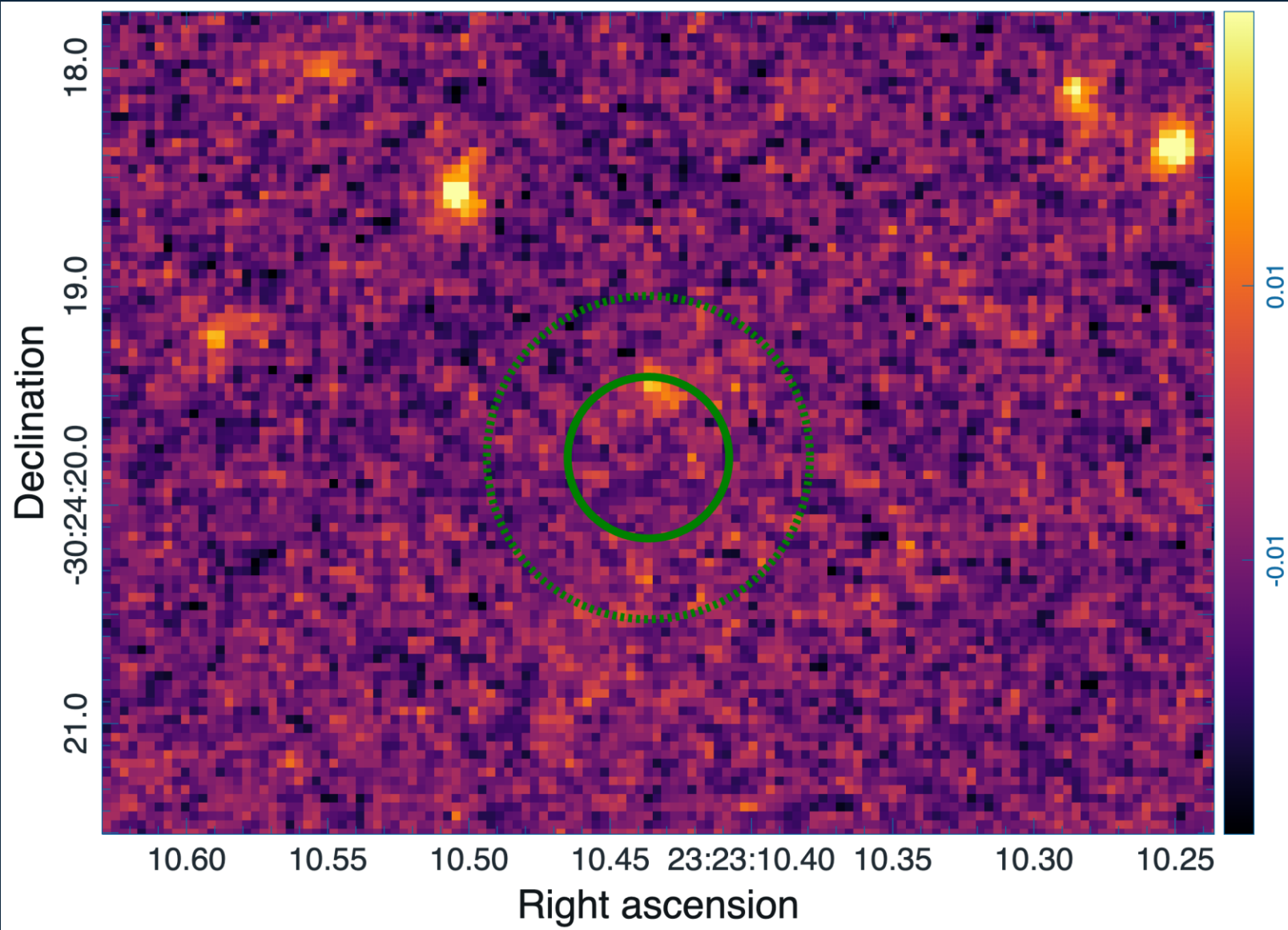
Thanks to

Adam Deller

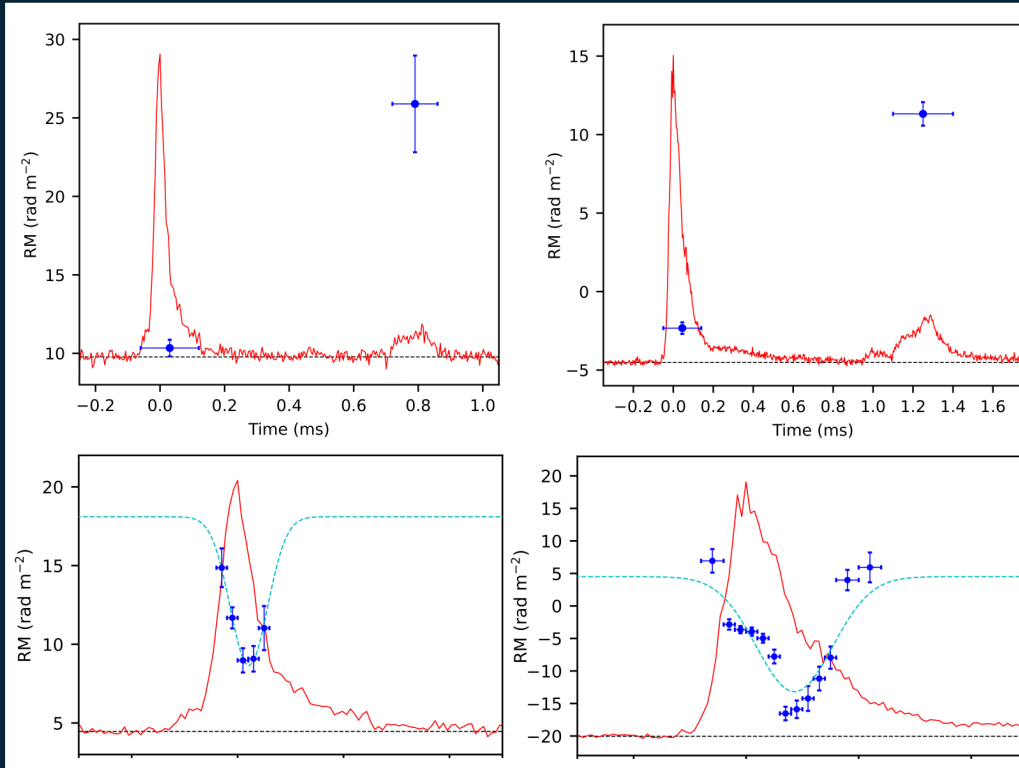
Lalitwadee Kawinwanichakij

Themiya Nanayakkara

Karl Glazebrook

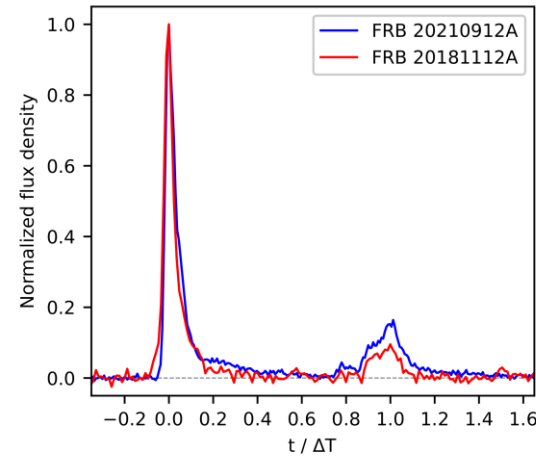


Twin FRBs – 1.1 ms-period neutron stars?



- A class of FRBs originating from fast rotating neutron stars?

Bera et al. 2024



- Faster than the fastest rotating known millisecond pulsar

Hessels et al. 2006

- Close to shortest possible spin period of neutron stars !

Haskell et al. 2018

