



TATA INSTITUTE OF FUNDAMENTAL RESEARCH

## **ICTS Astrophysical Relativity Seminar**

**Title** : Systematic bias on the inspiral-merger-ringdown consistency test due to

neglect of orbital eccentricity

**Speaker** : Sajad Bhat (Chennai Mathematical Institute)

**Date** : Wednesday, 17<sup>th</sup> August, 2022

**Time** : 03:00 pm (IST)

**Abstract**: The inspiral-merger-ringdown (IMR) consistency test checks the consistency

of the final mass and final spin of a binary black hole merger remnant, independently inferred via the inspiral and merger-ringdown parts of the waveform. As binaries are expected to be nearly circularized when entering the frequency band of ground-based detectors, tests of general relativity (GR) currently employ quasi-circular waveforms. We quantify the effect of residual orbital eccentricity on the IMR consistency test. We find that eccentricity causes a significant systematic bias in the inferred final mass and spin of the remnant black hole at an orbital eccentricity (defined at 10Hz) of  $e0 \ge 0.1$  in the LIGO band (for a total binary mass in the range  $65 - 200 \,\mathrm{M}\odot$ ). For binary black holes observed by Cosmic Explorer (CE), the systematic bias becomes significant for  $e0 \ge 0.015$  (for  $200 - 600 \,\mathrm{M}\odot$  systems). This eccentricity-induced bias on the final mass and spin leads to an apparent inconsistency in the IMR consistency test, manifesting as a false violation of GR. Hence, eccentric corrections to waveform models are important for constructing a robust test of GR,especially for 3rd-generation (3G) detectors.

**Venue** : **Online:** Please click on the below link to join the seminar

 $\underline{https://icts-res-in.zoom.us/j/81533329666?pwd=VGlLUUNtc3RTSU5ob3RqaXdXcFJpZz09}$ 

Meeting ID: 815 3332 9666

Passcode: 171722