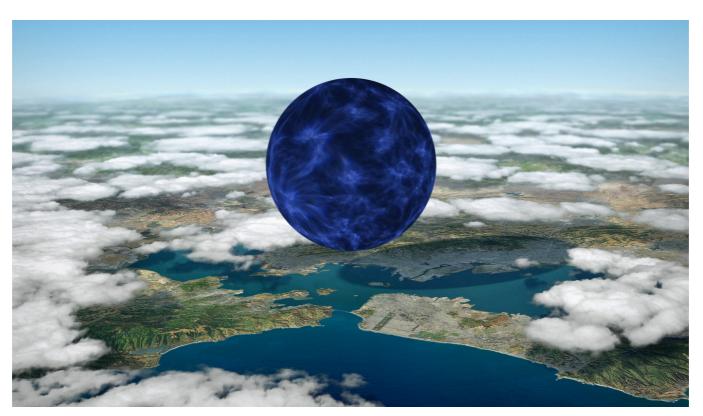




The neutron stars phenomenon



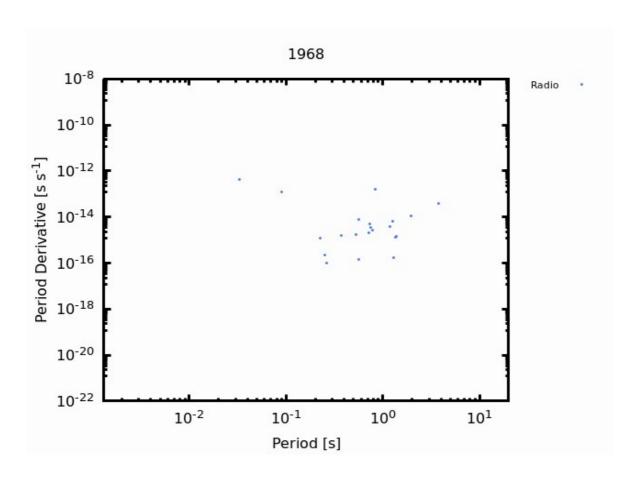
Credit: NASA







Why do we study the X-ray Universe?



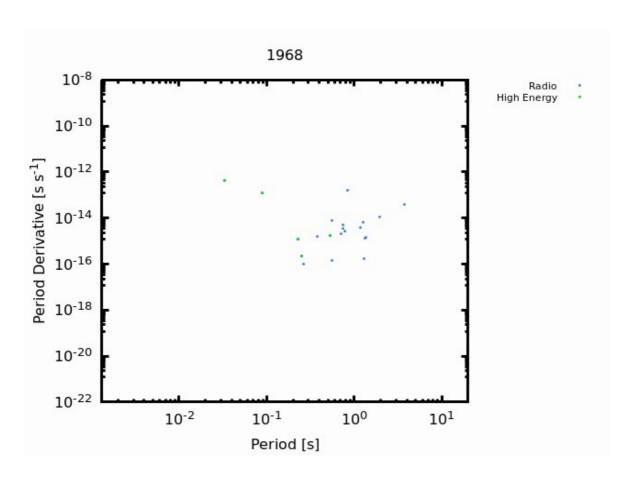
ATNF Pulsar Catalog (Manchester et al. 1995), version 1.50.







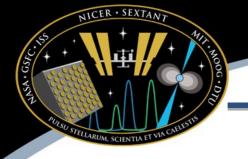
Why do we study the X-ray Universe?



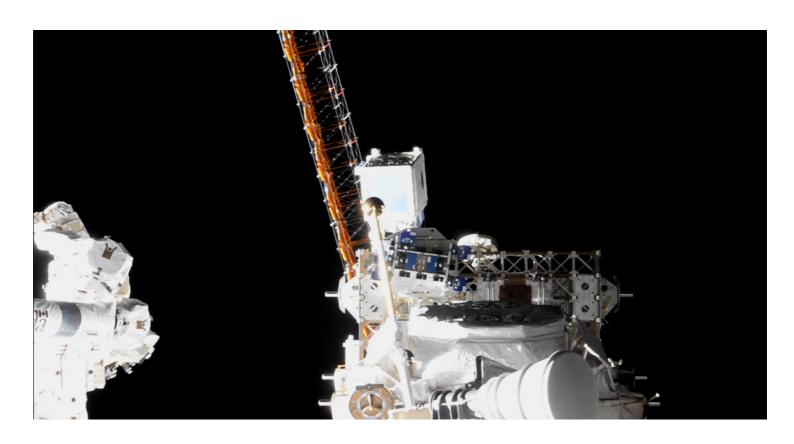
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NICER in operation on ISS



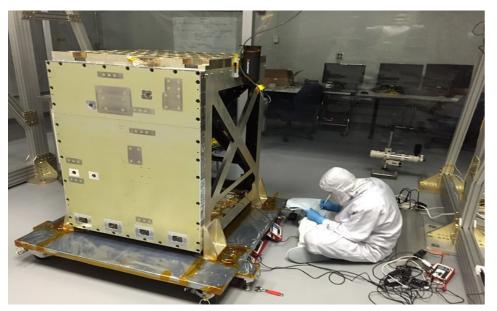
Credit: NASA







NICER Launch: June 3, 2017



Credit: NASA

X-ray instrument mounted on ISS

sensitivity range: 0.2 - 12 keV

peak collecting area: 1900 cm² (1.5 keV)

launch: June 3, 2017

science operations: July 17, 2017







NICER Launch: June 3, 2017



Credit: NASA)

X-ray instrument mounted on ISS

sensitivity range: 0.2 - 12 keV

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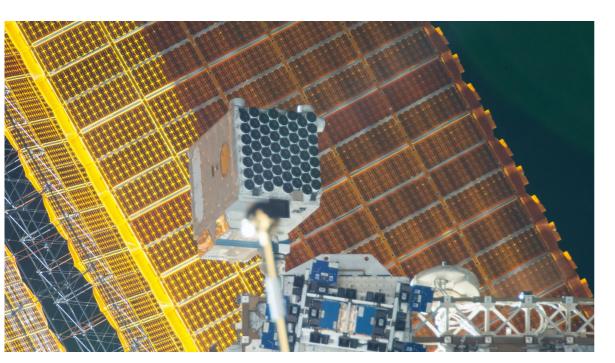
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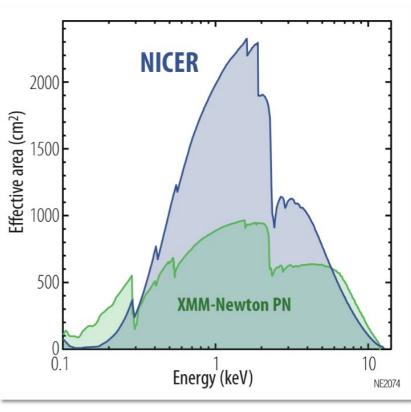




How NICER works (in a nutshell)



Credit: NASA



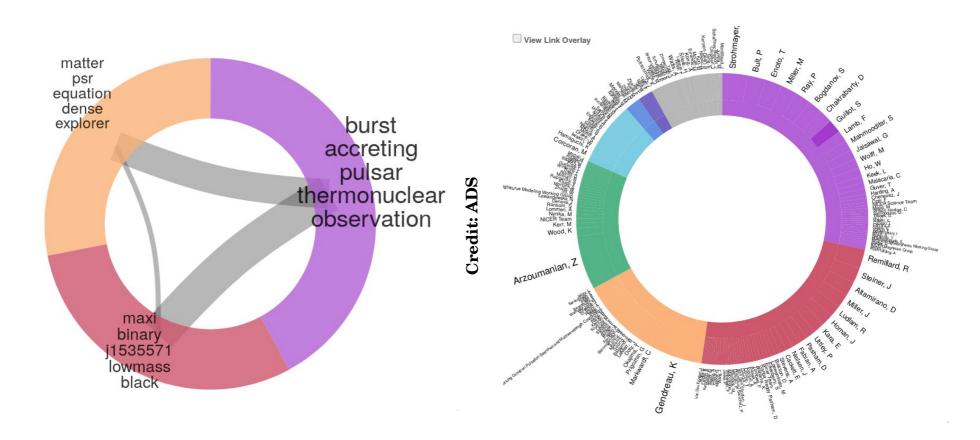
Credit: K. Genderau







What happened since then?



Referred papers: 65 (Science Team: 57; Others: 8); ATels: 70







NICER Precision Timing Group

Andrea Lommen (Chair), Zaven Arzoumanian, Julia Deneva, Teruaki Enoto, Matthew Kerr, Natalia Lewandowska, Walid Majid, Craig Marquardt, Scott Ransom, Paul Ray, Kent Wood, The Haverford Student Team: **Sydney Dorman,** Zaynab Ghazi, Lauren Lugo, Sergio Montano, Lizzie Spano, Dominik Rowan, Jesse Zeldes







THE ASTROPHYSICAL JOURNAL, 874:160 (13pp), 2019 April 1 © 2019. The American Astronomical Society. All rights reserved.

https://doi.org/10.3847/1538-4357/ab0966



High-precision X-Ray Timing of Three Millisecond Pulsars with NICER: Stability Estimates and Comparison with Radio

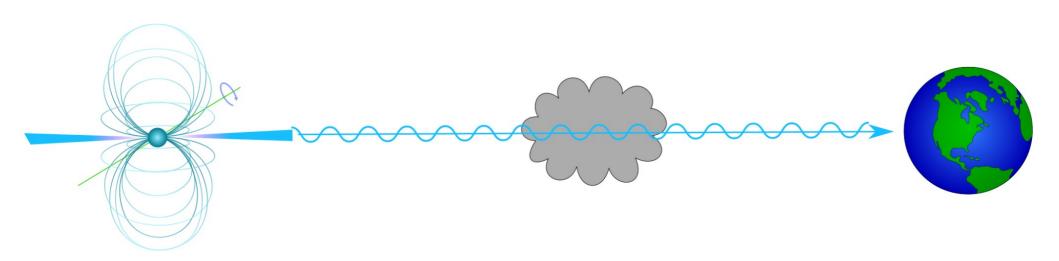
```
J. S. Deneva<sup>1</sup>, P. S. Ray<sup>2</sup>, A. Lommen<sup>3</sup>, S. M. Ransom<sup>4</sup>, S. Bogdanov<sup>5</sup>, M. Kerr<sup>2</sup>, K. S. Wood<sup>6</sup>, Z. Arzoumanian<sup>7</sup>,
K. Black<sup>8</sup>, J. Doty<sup>9</sup>, K. C. Gendreau<sup>7</sup>, S. Guillot<sup>10,11</sup>, A. Harding<sup>8</sup>, N. Lewandowska<sup>12,13</sup>, C. Malacaria<sup>14,15</sup>, C. B. Markwardt<sup>8</sup>, S. Price<sup>8</sup>, L. Winternitz<sup>8</sup>, M. T. Wolff<sup>2</sup>, L. Guillemot<sup>16,17</sup>, I. Cognard<sup>16,17</sup>, P. T. Baker<sup>12,13</sup>, H. Blumer<sup>12,13</sup>, P. R. Brook<sup>12,13</sup>,
H. T. Cromartie<sup>18</sup>, P. B. Demorest<sup>19</sup>, M. E. DeCesar<sup>20</sup>, T. Dolch<sup>21</sup>, J. A. Ellis<sup>22</sup>, R. D. Ferdman<sup>23</sup>, E. C. Ferrara<sup>8</sup>, E. Fonseca<sup>24</sup>,
N. Garver-Daniels<sup>12,13</sup>, P. A. Gentile<sup>12,13</sup>, M. L. Jones<sup>12,13</sup>, M. T. Lam<sup>12,13</sup>, D. R. Lorimer<sup>12,13</sup>, R. S. Lynch<sup>25</sup>, M. A. McLaughlin<sup>12,13</sup>, C. Ng<sup>26,27</sup>, D. J. Nice<sup>20</sup>, T. T. Pennucci<sup>28</sup>, R. Spiewak<sup>29</sup>, I. H. Stairs<sup>26</sup>, K. Stovall<sup>19</sup>, J. K. Swiggum<sup>30</sup>,
                                                                        S. J. Vigeland<sup>30</sup>, and W. W. Zhu<sup>31</sup>
                                     George Mason University, Resident at the Naval Research Laboratory, Washington, DC 20375, USA
                                                            <sup>2</sup> U.S. Naval Research Laboratory, Washington, DC 20375, USA
                                                                       Haverford College, Haverford, PA 19041, USA
                                                         <sup>4</sup> National Radio Astronomy Observatory, Charlottesville, VA, USA
                                             Columbia Astrophysics Laboratory, Columbia University, New York, NY 10027, USA
                                                  Praxis, Resident at the Naval Research Laboratory, Washington, DC 20375, USA
                                     <sup>7</sup> X-Ray Astrophysics Laboratory, NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA
                                                           NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA
                                                                    Noqsi Aerospace Ltd., Billerica, MA 01821, USA
                                          <sup>10</sup> IRAP, CNRS, 9 avenue du Colonel Roche, BP 44346, F-31028 Toulouse Cedex 4, France
                                                       <sup>11</sup> Université de Toulouse, CNES, UPS-OMP, F-31028 Toulouse, France
                            <sup>12</sup> Department of Physics and Astronomy, West Virginia University, P.O. Box 6315, Morgantown, WV 26506, USA
          <sup>13</sup> Center for Gravitational Waves and Cosmology, West Virginia University, Chestnut Ridge Research Building, Morgantown, WV 26505, USA
                                        <sup>14</sup> ST12 Astrophysics Branch, NASA Marshall Space Flight Center, Huntsville, AL 35812, USA
                                                    Universities Space Research Association, NSSTC, Huntsville, AL 35805, USA
                <sup>16</sup> Laboratoire de Physique et Chimie de l'Environnement et de l'Espace, LPC2E, CNRS-Université d'Orléans, F-45071 Orléans, France
                                       Station de Radioastronomie de Nançay, Observatoire de Paris, CNRS/INSU, F-18330 Nançay, France
                                   <sup>18</sup> University of Virginia, Department of Astronomy, P.O. Box 400325, Charlottesville, VA 22904, USA
                                            National Radio Astronomy Observatory, 1003 Lopezville Road, Socorro, NM 87801, USA
                                                        <sup>20</sup> Department of Physics, Lafayette College, Easton, PA 18042, USA
                                           <sup>21</sup> Department of Physics, Hillsdale College, 33 E. College Street, Hillsdale, MI 49242, USA
```



Deneva et al. (2019)







$$\Delta t = \mathcal{D} \cdot \frac{DM}{f^2}$$

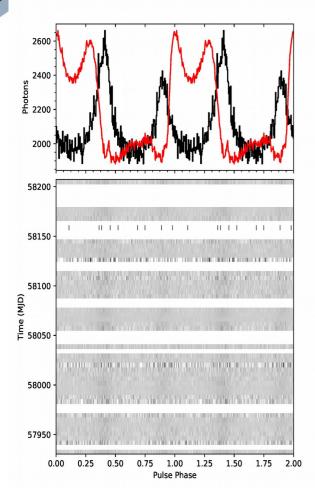
NICER data: taken at infinite frequencies

=> immune to ISM influence!



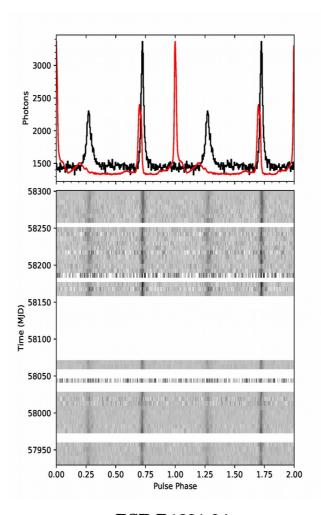




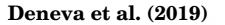


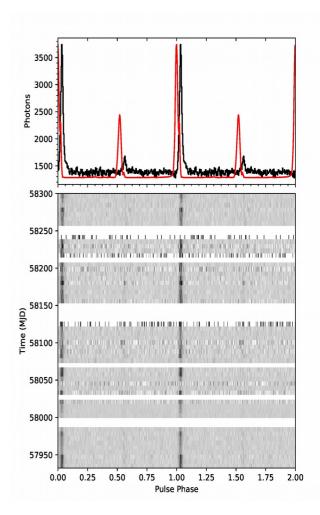
PSR J0218+4232 (0.80 - 6.20 keV)





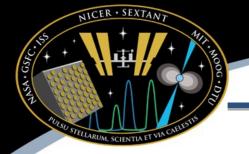
PSR B1821-24 (1.0 - 5.50 keV)

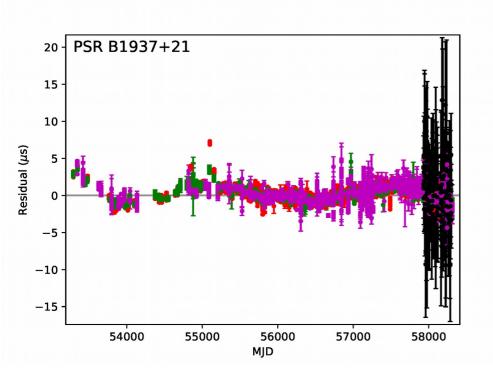


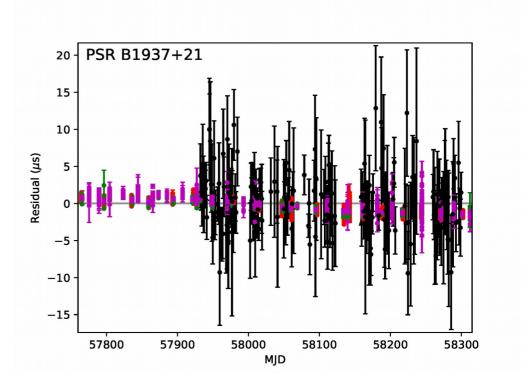


PSR B1937+21 (1.15 - 5.55 keV)









Deneva et al. (2019)







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A NICER VIEW OF SPECTRAL AND PROFILE EVOLUTION FOR THREE X-RAY EMITTING MILLISECOND PULSARS

Dominick M. Rowan,¹ Zaynab Ghazi,¹ Lauren Lugo,¹ Elizabeth Spano,¹ Andrea Lommen,¹ Alice Harding,² Christo Venter,³ Renee Ludlam,^{4,5} Paul S. Ray,⁶ Matthew Kerr,⁶ Zaven Arzoumanian,² Slavko Bogdanov,⁷ Julia Deneva,⁸ Sebastien Guillot,^{9,10} Natalia Lewandowska,^{11,12} Craig B. Markwardt,² Scott Ransom,¹³ Teruaki Enoto,¹⁴ Kent S. Wood,¹⁵ and Keith C. Gendreau²

¹ Haverford College, 370 Lancaster Ave, Haverford PA, 19041, USA

² Astrophysics Science Division, NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA

³ Centre for Space Research, North-West University, Private Bag X6001, Potchefstroom 2520, South Africa

⁴ Cahill Center for Astronomy and Astrophysics, California Institute of Technology, Pasadena, CA 91125, USA

⁵ Einstein Fellow

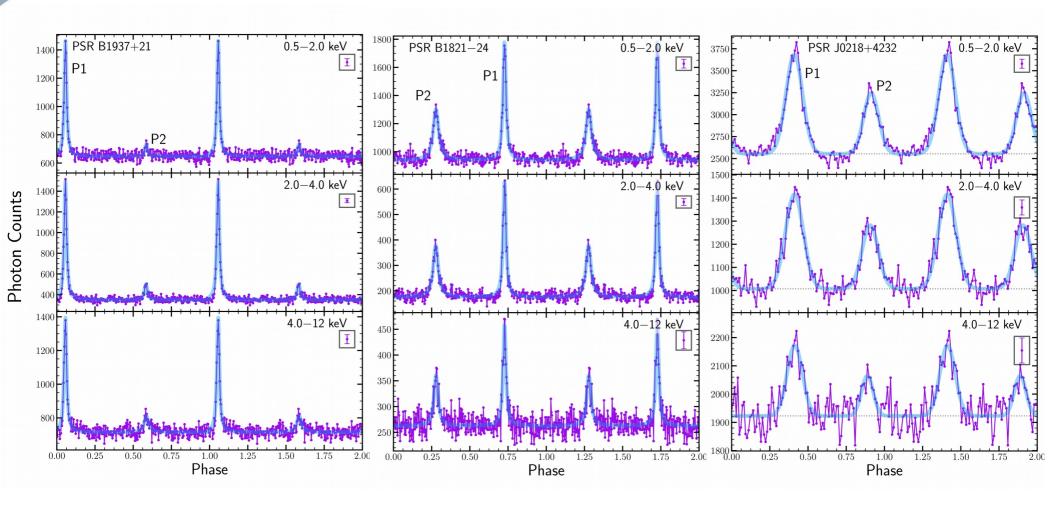
⁶ U.S. Naval Research Laboratory, Washington, DC 20375, USA
 ⁷ Columbia Astrophysics Laboratory, Columbia University, 550 West 120th Street, New York, NY, 10027, USA
 ⁸ George Mason University, resident at Naval Research Laboratory, Washington, DC 20375, USA
 ⁹ IRAP, CNRS, 9 avenue du Colonel Roche, BP 44346, F-31028 Toulouse Cedex 4, France
 ¹⁰ Université de Toulouse, CNES, UPS-OMP, F-31028 Toulouse, France







2 years of NICER data

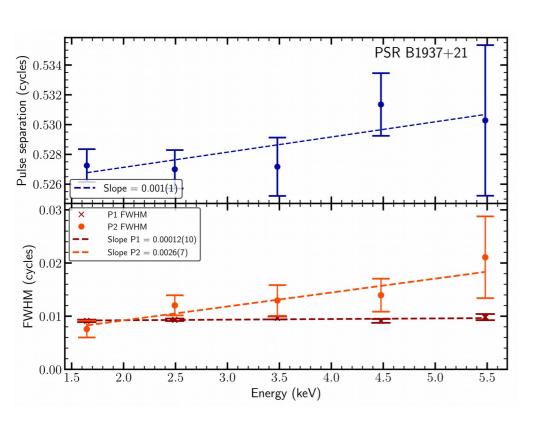


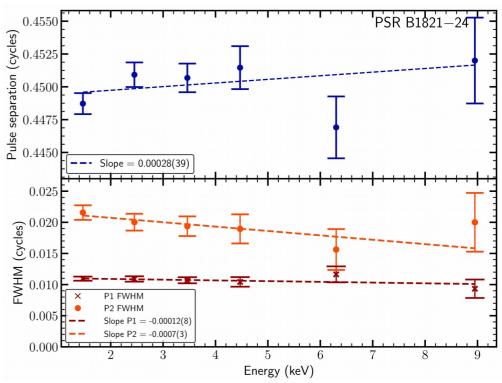






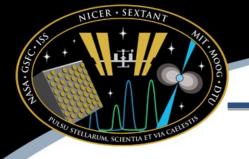
Evolution of Pulse Separation & FWHM



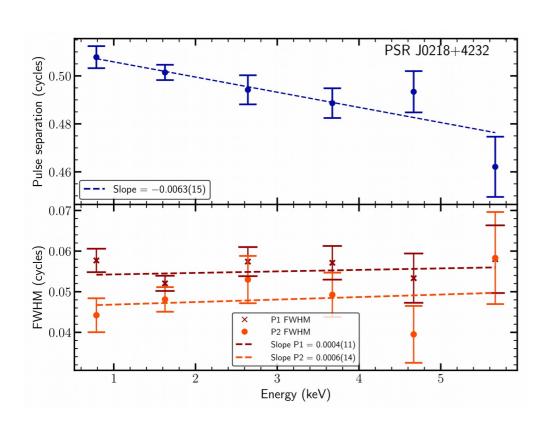








Evolution of Pulse Separation & FWHM



Results:

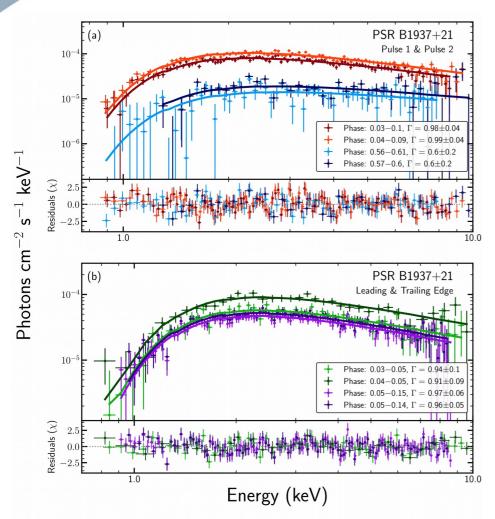
J0218: pulse component separation decreases with increasing energy (> 3 sigma confidence)

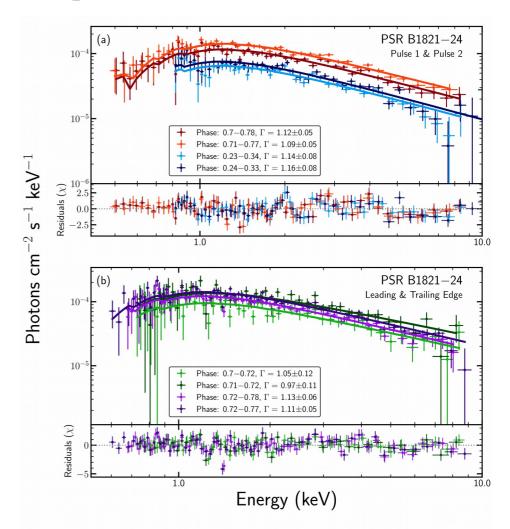






Phase-resolved spectra



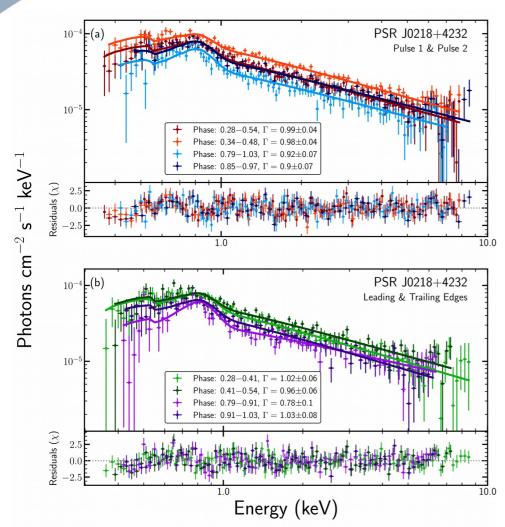








Phase-resolved spectra



Results:

- -updated model parameters: single single absorbed power-law model
- -no evidence of additional black body emission component (B1937 & B1821)
- -J0218: additional component at energies < 1keV
- -B1937: different P1 & P2 emission photon indices (> 2 sigma level)







NICER Lightcurve Modeling Group

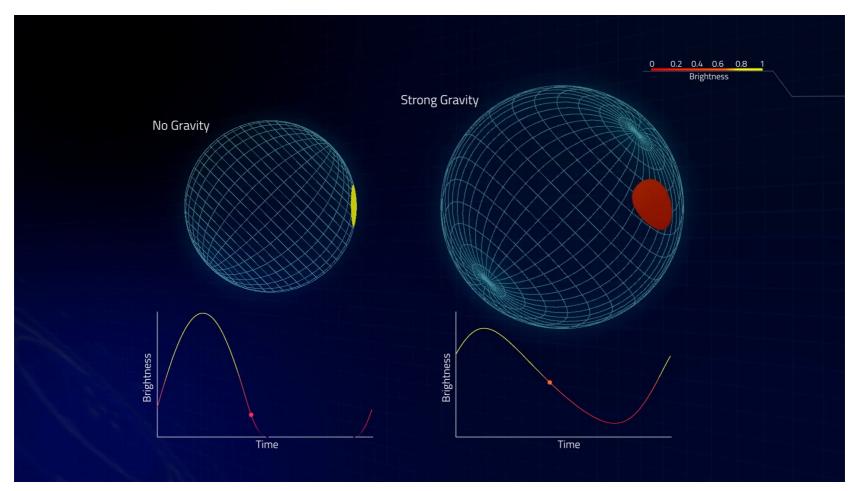
Slavko Bogdanov (Chair), Zaven Arzoumanian, Anna Bilous, Deepto Chakrabarty, Alexander Dittmann, Keith Gendreau, Sebastien Guillot, Alice Harding, Wynn Ho, Frederick Lamb, Jim Lattimer, Renee Ludlam, Simin Mahmoodifar, M. Coleman Miller, Sharon Morsink, Paul Ray, Thomas Riley, Tod Strohmayer, Anna Watts, Michael Wolff, **Kent Wood**







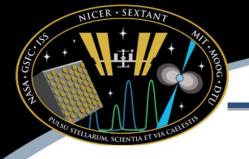
Mass-radius Relation



Credit: Morsink/Moir/Arzoumanian/NASA GSFC







DRAFT VERSION DECEMBER 13, 2019 Typeset using IATEX default style in AASTeX62

CONSTRAINING THE NEUTRON STAR MASS-RADIUS RELATION AND DENSE MATTER EQUATION OF STATE WITH NICER. I. THE MILLISECOND PULSAR X-RAY DATA SET

SLAVKO BOGDANOV, ¹ SEBASTIEN GUILLOT, ^{2,3} PAUL S. RAY, ⁴ MICHAEL T. WOLFF, ⁴ DEEPTO CHAKRABARTY, ⁵ WYNN C. G. HO, ^{6,7} MATTHEW KERR, ⁴ FREDERICK K. LAMB, ^{8,9} ANDREA LOMMEN, ⁶ RENEE M. LUDLAM, ^{10,*} REILLY MILBURN, ⁶ SERGIO MONTANO, ⁶ M. COLEMAN MILLER, ¹¹ MICHI BAUBÖCK, ¹² FERYAL ÖZEL, ¹³ DIMITRIOS PSALTIS, ¹³ RONALD A. REMILLARD, ⁵ THOMAS E. RILEY, ¹⁴ JAMES F. STEINER, ¹⁵ TOD E. STROHMAYER, ¹⁶ ANNA L. WATTS, ¹⁴ KENT S. WOOD, ¹⁷ JESSE ZELDES, ⁶ TERUAKI ENOTO, ¹⁸ TAKASHI OKAJIMA, ¹⁹ JAMES W. KELLOGG, ²⁰ CHARLES BAKER, ²⁰ CRAIG B. MARKWARDT, ²¹ ZAVEN ARZOUMANIAN, ²¹ AND KEITH C. GENDREAU²¹

¹ Columbia Astrophysics Laboratory, Columbia University, 550 West 120th Street, New York, NY, 10027, USA

² IRAP, CNRS, 9 avenue du Colonel Roche, BP 44346, F-31028 Toulouse Cedex 4, France

³ Université de Toulouse, CNES, UPS-OMP, F-31028 Toulouse, France

⁴ Space Science Division, U.S. Naval Research Laboratory, Washington, DC 20375, USA

⁵MIT Kavli Institute for Astrophysics and Space Research, Massachusetts Institute of Technology, 70 Vassar Street, Cambridge, MA 02139, USA

⁶ Department of Physics and Astronomy, Haverford College, 370 Lancaster Avenue, Haverford, PA 19041, USA
 ⁷ Mathematical Sciences, Physics and Astronomy, and STAG Research Centre, University of Southampton, Southampton SO17 1BJ, UK
 ⁸ Center for Theoretical Astrophysics and Department of Physics, University of Illinois at Urbana-Champaign, 1110 West Green Street, Urbana, IL 61801-3080, USA

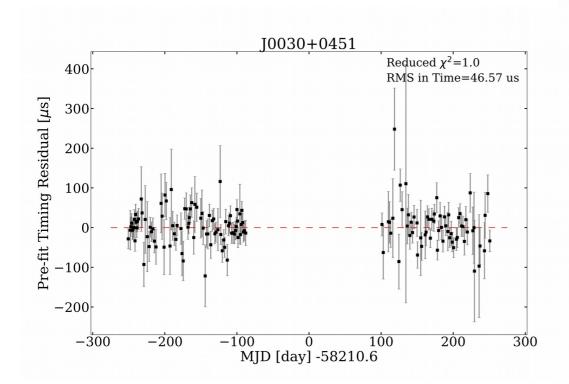
⁹Department of Astronomy, University of Illinois at Urbana-Champaign, 1002 West Green Street, Urbana, IL 61801-3074, USA
¹⁰Cahill Center for Astronomy and Astrophysics, California Institute of Technology, Pasadena, CA 91125, USA
¹¹Department of Astronomy and Joint Space-Science Institute, University of Maryland, College Park, MD 20742-2421, USA
¹²Max Planck Institut für Extraterrestrische Physik, Gießenbachstr. 1, D-85737 Garching, Germany

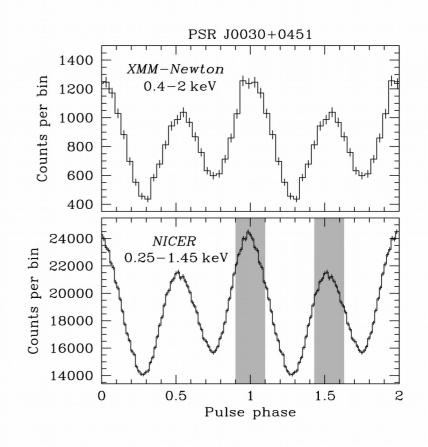


arXiv:1912.05706 Bogdanov et al. (2019)

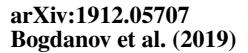






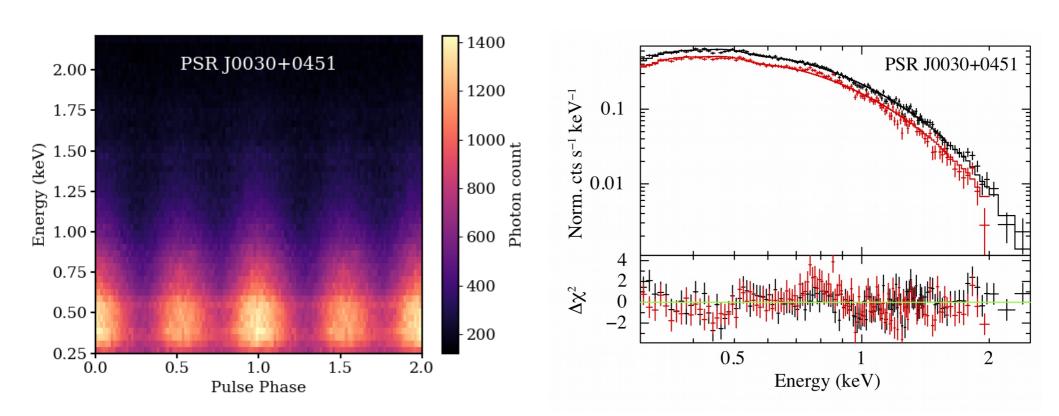














arXiv:1912.05706 Bogdanov et al. (2019)





DRAFT VERSION DECEMBER 13, 2019 Typeset using LATEX default style in AASTeX62

CONSTRAINING THE NEUTRON STAR MASS-RADIUS RELATION AND DENSE MATTER EQUATION OF STATE WITH *NICER*. II. EMISSION FROM HOT SPOTS ON A RAPIDLY ROTATING NEUTRON STAR

SLAVKO BOGDANOV,¹ FREDERICK K. LAMB,^{2,3} SIMIN MAHMOODIFAR,⁴ M. COLEMAN MILLER,⁵ SHARON M. MORSINK,⁶ THOMAS E. RILEY,⁷ TOD E. STROHMAYER,⁴ ALBERT K. TUNG,⁶ ANNA L. WATTS,⁷ ALEXANDER J. DITTMANN,⁵ DEEPTO CHAKRABARTY,⁸ SEBASTIEN GUILLOT,^{9,10} ZAVEN ARZOUMANIAN,¹¹ AND KEITH C. GENDREAU¹¹

¹ Columbia Astrophysics Laboratory, Columbia University, 550 West 120th Street, New York, NY 10027, USA

² Center for Theoretical Astrophysics and Department of Physics, University of Illinois at Urbana-Champaign, 1110 West Green Street, Urbana, IL 61801-3080, USA

³Department of Astronomy, University of Illinois at Urbana-Champaign, 1002 West Green Street, Urbana, IL 61801-3074, USA

⁴Astrophysics Science Division and Joint Space-Science Institute, NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA

⁵Department of Astronomy and Joint Space-Science Institute, University of Maryland, College Park, MD 20742-2421, USA

⁶Department of Physics, University of Alberta, Edmonton, AB T6G 2G7, Canada

Anton Pannekoek Institute for Astronomy, University of Amsterdam, Science Park 904, 1090GE Amsterdam, the Netherlands
 MIT Kavli Institute for Astrophysics and Space Research, Massachusetts Institute of Technology, 70 Vassar Street, Cambridge, MA 02139, USA

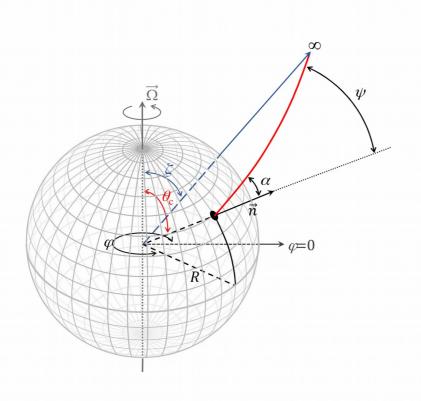
⁹IRAP, CNRS, 9 avenue du Colonel Roche, BP 44346, F-31028 Toulouse Cedex 4, France
 ¹⁰Université de Toulouse, CNES, UPS-OMP, F-31028 Toulouse, France
 ¹¹X-Ray Astrophysics Laboratory, NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA



arXiv:1912.05707 Bogdanov et al. (2019b)



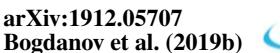




Verification tests:

- -Schwarzschild + Doppler approximations
- -Oblate Schwarzschild approximations
- -comparison with numerical calculations
- => Good agreement between code outputs!









DRAFT VERSION DECEMBER 13, 2019 Typeset using IATEX default style in AASTeX62

A NICER VIEW OF PSR J0030+0451: MILLISECOND PULSAR PARAMETER ESTIMATION

T. E. RILEY, A. L. WATTS, S. BOGDANOV, P. S. RAY, R. M. LUDLAM, S. GUILLOT, C. Z. ARZOUMANIAN, C. L. BAKER, A. V. BILOUS, D. CHAKRABARTY, K. C. GENDREAU, A. K. HARDING, U. C. G. HO, LATTIMER, S. M. MORSINK, AND T. E. STROHMAYER,

¹Anton Pannekoek Institute for Astronomy, University of Amsterdam, Science Park 904, 1090GE Amsterdam, the Netherlands

²Columbia Astrophysics Laboratory, Columbia University, 550 West 120th Street, New York, NY 10027, USA

³Space Science Division, U.S. Naval Research Laboratory, Washington, DC 20375, USA

⁴Cahill Center for Astronomy and Astrophysics, California Institute of Technology, Pasadena, CA 91125, USA

⁵Einstein Fellow

⁶IRAP, CNRS, 9 avenue du Colonel Roche, BP 44346, F-31028 Toulouse Cedex 4, France ⁷Université de Toulouse, CNES, UPS-OMP, F-31028 Toulouse, France.

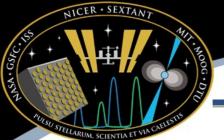
⁸X-Ray Astrophysics Laboratory, NASA Goddard Space Flight Center, Code 662, Greenbelt, MD 20771, USA
 ⁹Applied Engineering and Technology Directorate, NASA Goddard Space Flight Center, Code 592, Greenbelt, MD 20771, USA
 ¹⁰MIT Kavli Institute for Astrophysics and Space Research, Massachusetts Institute of Technology, Cambridge, MA 02139, USA
 ¹¹Astrophysics Science Division, NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA
 ¹²Department of Physics and Astronomy, Haverford College, 370 Lancaster Avenue, Haverford, PA 19041, USA

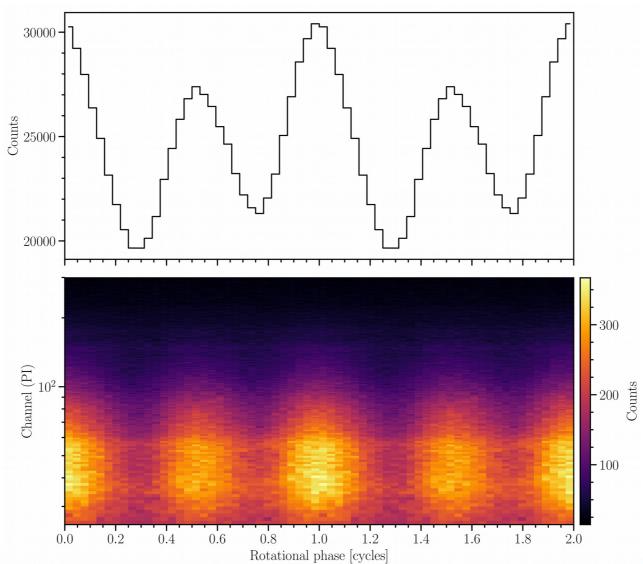
¹³ Mathematical Sciences, Physics and Astronomy, and STAG Research Centre, University of Southampton, Southampton, SO17 1BJ, UK

¹⁴ Department of Physics and Astronomy, Stony Brook University, Stony Brook, NY 11794-3800, USA





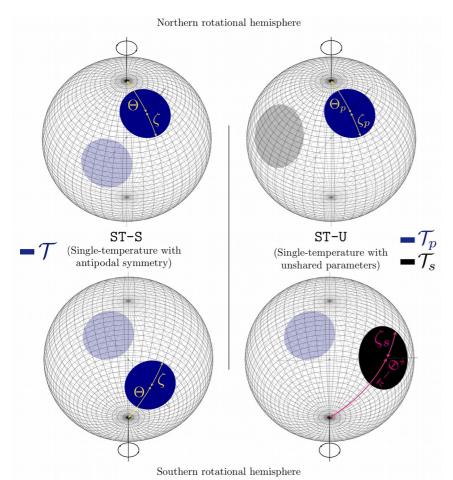


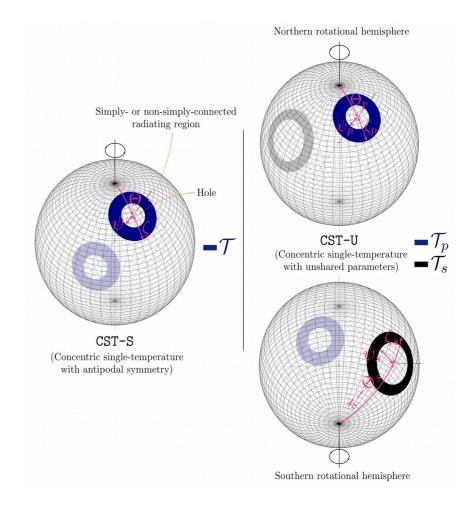








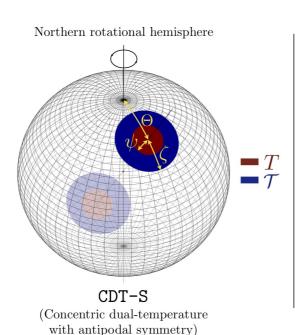


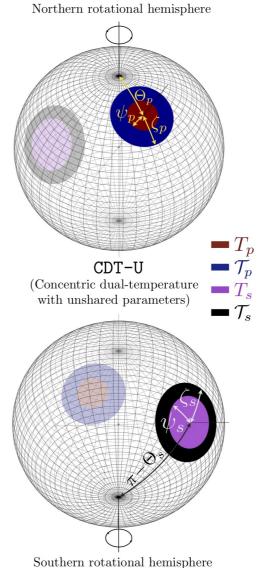


















Summary of results:

- -Bayesian parameter estimation of mass & equatorial radius
- => mass ~ 1.34 M_{Sun}
- => radius ~ 12.71 km
- -in favor: models with hot spots in same hemisphere
- => more parameters needed for successful description!
- -disfavored: models with simply connected single temperature hot spots







Summary

- NICER science operations ongoing since July 2017
- lots of interesting science projects published already
 - => https://heasarc.gsfc.nasa.gov/docs/nicer/nicer_results.html

NICER Precision Timing Group:

- 1 year of timing results published!
- 2 year of X-ray pulsar data published
 - => profile evolution paper

NICER Light curve Modeling Group:

several new results published



