

# Gravity-related alterations of non-relativistic quantum theory

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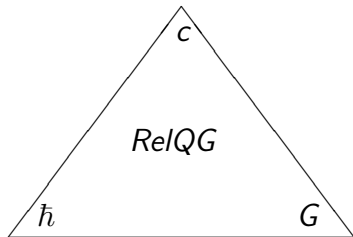
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# Why Newtonian Quantum Gravity? 1980's

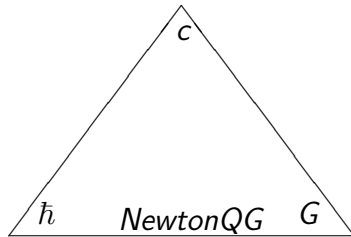
Quantum or Gravity is bottle-neck of Quantum Gravity?

That's just Gravity!



perturbativ QFT  
 $Q-\mathcal{G}^{(3)}$  dynamics  
 loops, strings, ... ???

1980s-: **That's Quantum!**



Bottle-neck: Schrödinger Cat  
 Let's go non-relativistic, first!  
 Hope: new QM comes from NQG

# Spatial mass density $\hat{\rho}(\mathbf{r})$ , and resolution $\sigma$

Forget full energy-momentum  $\hat{T}_{ab}$ ! Use  $\hat{T}_{00}/c^2 = \hat{\rho}$

Non-relativistic spatial mass distribution:

$$\hat{\rho}(\mathbf{r}) = \sum_k m_k \delta_\sigma(\mathbf{r} - \hat{\mathbf{x}}_k)$$

$\nearrow$  mass density  
 $\nwarrow$  spatial resolution  $\sigma$   
 $\delta_\sigma$ : Gaussian of width  $\sigma$

What are the sources of gravity?

	nuclei	...	bulk matter
$\sigma$ resolution	$10^{-12} \text{ cm}$	...	$10^{-5} \text{ cm}$

No evidence!

Spatial resolution  $\sigma$ : **burning and lasting issue**

# Standard Schrödinger Eq. with Newton gravity

That's standard QM with standard Newtonian pair-potential:

$$|\dot{\Psi}\rangle = -\frac{i}{\hbar} \left( \hat{H} - G \int \frac{\hat{\rho}(\mathbf{r})\hat{\rho}(\mathbf{s})}{|\mathbf{r} - \mathbf{s}|} d\mathbf{r}d\mathbf{s} \right) |\Psi\rangle$$

Justified in classical limit only.

Quantum regime - **tests are (currently) impossible:**

$G$  is too small, massive d.o.f. would be influenced.

But massive d.o.f. decohere immediately by environment.

Open issue:

	nuclei	...	bulk matter
$\sigma$ resolution	$10^{-12} cm$	...	$10^{-5} cm$

Classical limit is insensitive to  $\sigma$ , **quantum tests are missing.**

We are free to speculate: to **alter QM**, to **choose**  $\sigma$ .

# Semiclassical Schrödinger-Newton Equation 1984

Quantized matter in classical Newton-potential:

$$|\dot{\Psi}\rangle = -\frac{i}{\hbar} \left( \hat{H} - G \int \frac{\hat{\rho}(\mathbf{r}) \langle \hat{\rho}(\mathbf{s}) \rangle_{\Psi}}{|\mathbf{r} - \mathbf{s}|} d\mathbf{r} d\mathbf{s} \right) |\Psi\rangle$$

Non-linearity generates **self-interaction**. For c.o.m.  $\hat{\mathbf{x}}$  (at  $\Delta x \ll \sigma$ ):

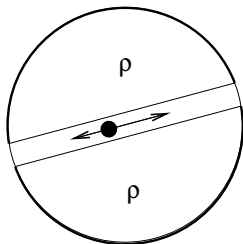
$$|\dot{\Psi}\rangle = -\frac{i}{\hbar} \left( \hat{H} + \frac{1}{2} M \omega_G^2 (\hat{\mathbf{x}} - \langle \hat{\mathbf{x}} \rangle_{\Psi})^2 \right) |\Psi\rangle$$

Solitonic solutions: **localized stationary states for bulk c.o.m.!**  
**No decoherence, no wave function collapse**, of course!

Localization heavily depends on gravity's spatial resolution  $\sigma$ :

		nuclei	...	bulk matter
$\sigma$	resolution	$10^{-12} \text{ cm}$	...	$10^{-5} \text{ cm}$
$\omega_G$		$10^3 / \text{s}$	...	$10^{-3} / \text{s}$

The strong-effect-edge may become falsifiable in coming years.

$\omega_G$ : Classical Newton oscillator frequency 2007

[Named as "SN-frequency" (2013-):  
incorrect name,  $\omega_G$  has no  $\hbar$ !]

$$\omega_G = \sqrt{\frac{4\pi}{3} G \rho_0}$$

$$\frac{\omega_G \text{ in nuclear matter}}{\omega_G \text{ in common matter}} \sim 10^6$$

		nuclei	...	bulk matter
$\sigma$	resolution	$10^{-12} \text{ cm}$	...	$10^{-5} \text{ cm}$
$\omega_G$		$10^3 / \text{s}$	...	$10^{-3} / \text{s}$

SN self-interaction  $\frac{1}{2} M \omega_G^2 (\hat{\mathbf{x}} - \langle \hat{\mathbf{x}} \rangle)^2$  becomes  **$10^{12}$ -times stronger** if nucleus-size resolution is taken.

# G-related spontaneous decoherence (DP) 1987

Master equation to decohere mass-density superpositions:

$$\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \frac{G}{2\hbar} \int [\hat{\rho}(\mathbf{r}), [\hat{\rho}(\mathbf{s}), \hat{\rho}]] \frac{d\mathbf{r}d\mathbf{s}}{|\mathbf{r} - \mathbf{s}|}$$

Fast **decoherence of bulk d.o.f.**. For c.o.m.  $\hat{\mathbf{x}}$  (at  $\Delta x \ll \sigma$ ):

$$\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \frac{1}{2\hbar} M \omega_G^2 [\hat{\mathbf{x}}, [\hat{\mathbf{x}}, \hat{\rho}]]$$

↙ decoherence rate

Universal **heating of all d.o.f.**:  $\dot{\epsilon} = \frac{1}{2} \hbar \omega_G^2 / d.o.f.$

	nuclei	...	bulk matter
$\sigma$ resolution	$10^{-12} cm$	...	$10^{-5} cm$
$\omega_G$ decoh. rate	$10^3/s$	...	$10^{-3}/s$
$\dot{\epsilon}$ heating rate	$10^{-21} erg/s/d.o.f.$	...	$10^{-33} erg/s/d.o.f.$

At strong-effect-edge heating is  $10\mu K/s$ , much too high! **Ruled out?**



# LISA Pathfinder experiment 2016

$$\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \frac{1}{2\hbar} M \omega_G^2 [\hat{\mathbf{x}}, [\hat{\mathbf{x}}, \hat{\rho}]]$$

↙ decoherence rate

Free falling test mass  $M = 2\text{kg}$  on satellite

C.o.m. acceleration noise  $\leq 5.2\text{fms}^{-2}/\sqrt{\text{Hz}}$  in range  $0.7 - 20\text{mHz}$ .

**Most precise control of c.o.m. ever!**

A most direct test of spontaneous decoherence.

Seems to rule out DP strong-edge version (finest resolution  $10^{-12}\text{cm}$ ).

	nuclei	...	bulk matter
$\sigma$ resolution	$10^{-12}\text{cm}$	...	$10^{-5}\text{cm}$
$\omega_G$ decoh. rate	$10^3/\text{s}$	...	$10^{-3}/\text{s}$
$\dot{\epsilon}$ heating rate	$10^{-21}\text{erg/s/d.o.f.}$	...	$10^{-33}\text{erg/s/d.o.f.}$

LISA 2016:  $\sigma \not\ll 4 \cdot 10^{-12}\text{cm}$ .

Helou-Slagmolen-McClelland-YanbeiChen, arXiv:1606.03637.

# Quantum-gravity spontaneous decoherence? 1983

Hawking 1983: unitarity is lost due to space-time fluctuations (instantons) on Planck scale

$$\hat{\rho} \rightarrow \hat{\$}\hat{\rho} \neq \hat{S}\hat{\rho}\hat{S}^\dagger$$

Banks-Susskind-Peskin 1984: master equation to detail Hawking's  $\hat{\$}$ . Violation of conservations laws (cf. **heating**):

$$\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \int [\hat{Q}(\mathbf{r}), [\hat{Q}(\mathbf{s}), \hat{\rho}]] h(\mathbf{r} - \mathbf{s}) d\mathbf{r} d\mathbf{s}$$

D 1987: unitarity is lost due to gravitational fluctuations much before the Planck scale, same structure as BSP's:

$$\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \frac{G}{2\hbar} \int [\hat{\varrho}(\mathbf{r}), [\hat{\varrho}(\mathbf{s}), \hat{\rho}]] \frac{d\mathbf{r} d\mathbf{s}}{|\mathbf{r} - \mathbf{s}|}$$

Reznik&Oppenheim 2009: to conserve total momentum, decohere relational variables:

$$\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \frac{1}{2} \sum_{k,l} \gamma_{kl} [\hat{\mathbf{x}}_k - \hat{\mathbf{x}}_l, [\hat{\mathbf{x}}_k - \hat{\mathbf{x}}_l, \hat{\rho}]]$$

# Semiclassical gravity + DP collapse 2016

Master equation to decohere mass-density superpositions:

$$\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \frac{G}{2\hbar} \int [\hat{\varrho}(\mathbf{r}), [\hat{\varrho}(\mathbf{s}), \hat{\rho}]] \frac{d\mathbf{r}d\mathbf{s}}{|\mathbf{r} - \mathbf{s}|}$$

Straightforward underlying mechanism: spontaneous monitoring of  $\hat{\varrho}$ :

$$\text{monitoring of } \hat{\varrho} \Rightarrow \left\{ \begin{array}{l} \text{stochastic Q-trajectory } \Psi(t) \\ \text{measured signal } \varrho(\mathbf{r}, t) \end{array} \right\}$$

$$\varrho(\mathbf{r}, t) = \langle \hat{\varrho}(\mathbf{r}) \rangle_{\Psi(t)} + \text{noise}$$

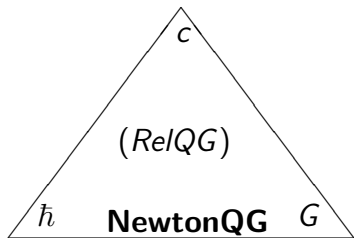
Let's source Newton field by  $\varrho(\mathbf{r}, t)$ ! Tilloy&D 2016 (cf. D 1990):

$$\dot{\hat{\rho}} = -\frac{i}{\hbar} \left[ \hat{H} - G \int \hat{\varrho}(\mathbf{r}) \hat{\varrho}(\mathbf{s}) \frac{d\mathbf{r}d\mathbf{s}}{|\mathbf{r} - \mathbf{s}|}, \hat{\rho} \right] - \frac{G}{\hbar} \int [\hat{\varrho}(\mathbf{r}), [\hat{\varrho}(\mathbf{s}), \hat{\rho}]] \frac{d\mathbf{r}d\mathbf{s}}{|\mathbf{r} - \mathbf{s}|}$$

Newtonian pair-potential emerges, self-interaction & non-linearity gone, DP decoherence doubles.

# Why Newtonian Quantum Gravity?

1980's-2010's-...



- **Standard QM with pair-potential:** too small, far from testability.
- **Non-standard SNE, DP, ...:** New effects depend heavily on  $\sigma$ . In strong-effect regime  $\sigma \sim 10^{-12} cm$  SN self-attraction, DP-heating become testable in coming years.

If Schrödinger Cats are **bottle-neck** of Quantum Gravity:  
 Newtonian Quantum Gravity might be the key first.  
 G-related Newtonian alterations of QM are still uncomplete.  
 They can yet encode a piece of truth.  
 Experiments will decide if NQG concerns new physics at all.