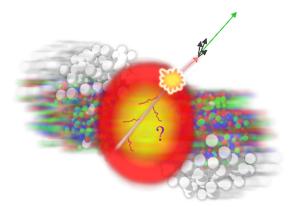
Jet evolution in a dense QCD medium

#### Edmond Iancu IPhT Saclay & CNRS



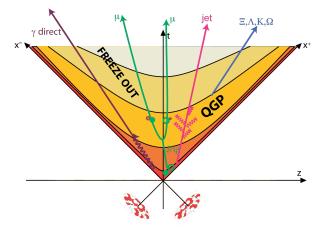
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### Outline

- Motivation: Hard Probes at RHIC and the LHC
  - observables related to jet quenching
  - the general physical picture
- Jets in the vacuum
  - radiation formation time
  - angular ordering
  - double logarithmic approximation
- Jet quenching in perturbative QCD
  - transverse momentum broadening
  - medium-induced radiation
  - multiple branching, energy loss, wave turbulence
- Adding vacuum-like radiation
  - the full, Markovian, picture ( $\Rightarrow$  Monte-Carlo)
  - in-medium jet fragmentation

### Hard probes in heavy ion collisions

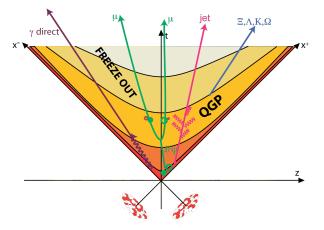
• The QGP phase "lives" for about 10 fm, that is, .... 10<sup>-23</sup> seconds !



- How to measure such an ephemeral form of matter ?
- Use energetic partons, or jets as "thermometers"

### Hard probes in heavy ion collisions

• The QGP phase "lives" for about 10 fm, that is, ....  $10^{-23}$  seconds !



- Hard partons, photons, leptons created at early times :  $au \lesssim 0.1 \; {\rm fm/c}$
- Interact with the surrounding medium on their way to the detectors

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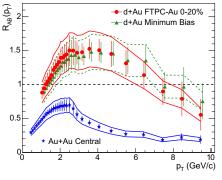
Jet evolution in a dense medium

#### Nuclear modification factor for hadrons

• Ratio of particle yield in AA and pp scaled by the number of binary collisions

$$R_{AA} \equiv \frac{1}{A^{4/3}} \frac{\mathrm{d}N_{pA}/\mathrm{d}^2 p_{\perp} \mathrm{d}\eta}{\mathrm{d}N_{pp}/\mathrm{d}^2 p_{\perp} \mathrm{d}\eta}$$

• Compare d+Au and Au+Au at RHIC at midrapidities ( $\eta \sim 0$ )



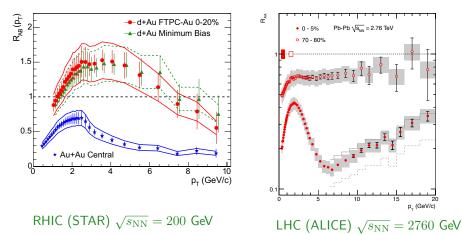
RHIC (STAR):  $\sqrt{s_{\rm NN}} = 200 \text{ GeV}$ 

- d+Au: Cronin peak
- multiple scattering off the dense nuclear target
- "initial state effect": gluon saturation
- Au+Au: suppression at all  $p_T$ 's
- "final state effect" : energy loss

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### Nuclear modification factor for hadrons

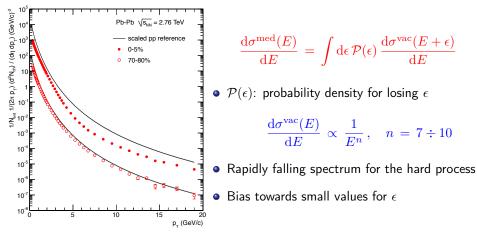
• A similar pattern emerges at the LHC energies



- The amount of suppression increases with the centrality of the collision
  - central collisions (head-on Au+Au scattering) look denser

### Energy loss

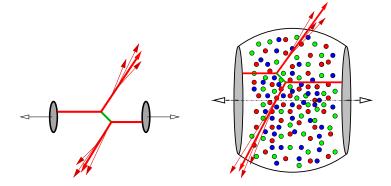
- Partons can lose energy via interactions in the plasma
- Hadrons measured with a given energy E have been produced with  $E + \epsilon$



• Even a small  $\epsilon$  may imply strong suppression

## Jet quenching

• Hard partons are typically created in pairs which propagate back-to-back in the transverse plane

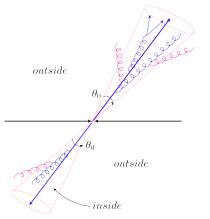


- 'Jet': 'leading particle' + 'products of fragmentation'
- *AA* collisions: jet propagation and fragmentation can be modified by the surrounding medium: 'jet quenching'

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#### Jets in practice (cf. lectures by Michael Spira)

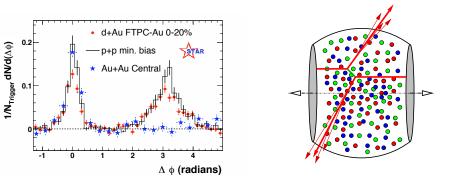
- Experimentally, jets are constructed by grouping together hadrons which propagate at nearby angles
- The jet opening angle  $\theta_0$  (a.k.a. R,  $\theta_{\rm jet}$ ,  $\bar{\theta}$  ...) is the same for both jets



• Medium modifications refer both to the jets and to the outer regions

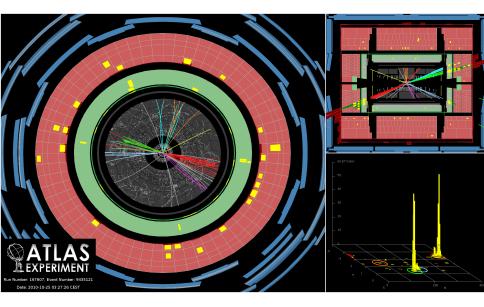
### Di-hadron azimuthal correlations at RHIC

- Distribution of pairs of particles w.r.t. the relative azimuthal angle  $\Delta\Phi$
- Compare p+p, d+Au and Au+Au, all at midrapidities  $(\eta \sim 0)$

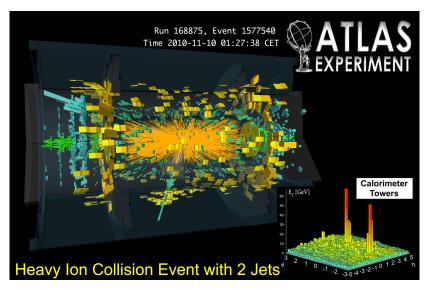


- Peak at  $\Delta \Phi \sim 0$ : both hadrons belong to a same jet
- Peak at  $\Delta \Phi \sim \pi$ : they belong to two back-to-back jets
- Au+Au: no peak at  $\Delta \Phi \sim \pi$  ! One of the jets has "melted" into the medium

## LHC: Di-jets in p+p collisions

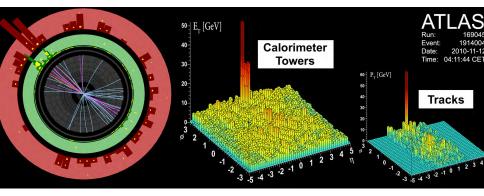


### Jets in peripheral Pb+Pb collisions



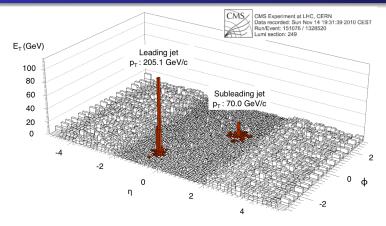
• Jets in peripheral AA collisions look very much like in pp collisions

## "Mono-jets" in Pb+Pb collisions



- Central Pb+Pb: 'mono-jet' events
- The secondary jet can barely be distinguished from the background:  $E_{T1} \ge 100$  GeV,  $E_{T2} > 25$  GeV

### Di-jet asymmetry at the LHC



- Huge difference between the energies of the two jets
- The missing energy is found in the underlying event:
  - many soft ( $p_{\perp} < 2$  GeV) hadrons propagating at large angles
- Very different from the usual jet fragmentation pattern in the vacuum

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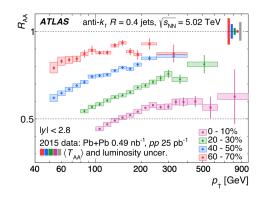
Jet evolution in a dense medium

## The nuclear modification factor for jets

• The jet yield in Pb+Pb collisions normalized by p+p times the average nuclear thickness function  $\langle T_{AA}\rangle$ 

$$R_{AA} \equiv \frac{\frac{1}{N_{\text{evt}}} \left. \frac{\mathrm{d}^2 N_{\text{jet}}}{\mathrm{d} p_T \mathrm{d} y} \right|_{AA}}{\langle T_{AA} \rangle \frac{\mathrm{d}^2 \sigma_{\text{jet}}}{\mathrm{d} p_T \mathrm{d} y} \right|_{pp}}$$

- different centrality bins
- stronger suppression for more central collisions



- Naturally interpreted as a consequence of energy loss inside the medium
- $R_{AA}$  is almost flat at very high  $p_T$ : energy loss increases with  $p_T$

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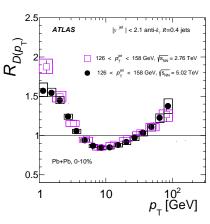
Jet evolution in a dense medium

### Jet fragmentation function

• Energy distribution of the hadrons inside the jet

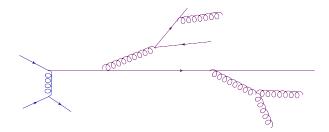
$$D(\omega) \equiv \omega \frac{\mathrm{d}N}{\mathrm{d}\omega}$$
$$= \int_0^R \mathrm{d}\theta \ \omega \frac{\mathrm{d}N}{\mathrm{d}\theta \mathrm{d}\omega}$$

- $\omega \equiv p_T$  of a hadron inside the jet
- ratio of FFs in Pb+Pb and p+p
- enhancement at low energies:  $p_T \ll p_T^{\rm jet}$  ...
- $\bullet$  ... and at relatively high ones:  $p_T \sim p_T^{\rm jet}$
- slight suppression at intermediate energies
  - This is an example of intra-jet medium-induced modifications



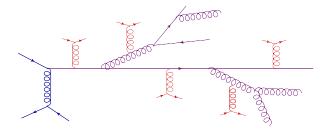
### Medium-induced jet evolution

- The leading particle (LP) is produced by a hard scattering
- It subsequently evolves via radiation (branchings) ...



### Medium-induced jet evolution

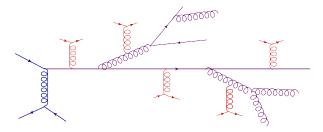
- The leading particle (LP) is produced by a hard scattering
- It subsequently evolves via radiation (branchings) ...



• ... and via collisions off the medium constituents

### Medium-induced jet evolution

- The leading particle (LP) is produced by a hard scattering
- It subsequently evolves via radiation (branchings) ...



- ... and via collisions off the medium constituents
- Collisions can have several effects
  - transfer energy and momentum between the jet and the medium
  - trigger additional radiation ("medium-induced")
  - wash out the color coherence (destroy interference pattern)

#### Jets in the vacuum (cf. lectures by M. Spira)

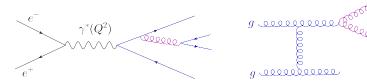
• A hard scattering generates a time-like system (positive virtuality  $Q^2 > 0$ )

$$Q^{2} \equiv (p_{1} + p_{2})^{2} = 2p_{1} \cdot p_{2} = 2E_{1}E_{2}(1 - \cos\theta) = 4E_{c.o.m}^{2}$$

• ... which can decay into a pair of on-shell partons (at tree-level)

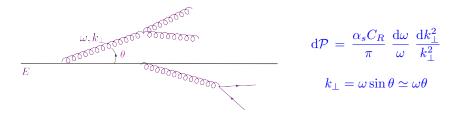


• ... or into other time-like partons, which will in turn decay ... and so on !



# Jets in the vacuum (cont.)

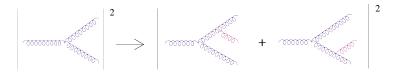
- Multiple emissions ('fragmentation') leading to a jet structure
- The differential probability for one splitting is given by bremsstrahlung :



- It favors soft ( $\omega \ll E$ ) and collinear ( $\theta \ll 1$ ) splittings
  - many soft gluons ... but they carry very little energy
  - most of the energy remains in the few partons with large  $x \equiv \omega/E$
  - small angle emissions  $k_\perp\simeq\omega\theta$  with  $\theta\ll 1\Rightarrow$  jets are collimated
- The collimation property is further enhanced by angular ordering

### Angular ordering in the vacuum

- Interference between emissions by several sources
  - successive quantum emissions are generally not independent
- One sums the amplitudes and only then one takes the modulus squared
  - the 'cross-terms' represent interference effects



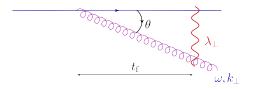
- Quantum interference complicates the "automatization" of the parton showers (e.g. Monte-Carlo event generators)
  - a priori, inconsistent with a classical probabilistic description
- Angular ordering "saves" the probabilistic picture and allows for Monte-Carlo

### **Radiation:** Formation time

• Uncertainty principle: quantum particles are delocalized

de Broglie wavelength: 
$$\lambda_{\perp} = \frac{2}{k_{\perp}} \simeq \frac{2}{\omega \theta}$$

• The gluon has been emitted when it has no overlap with its source

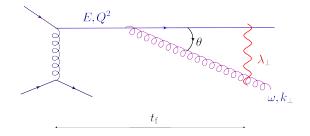


$$\Delta x_{\perp} \simeq heta \, \Delta t \, \gtrsim \, \lambda_{\perp} \, \Longrightarrow \Delta t \, \gtrsim \, t_{
m f} \, \equiv \, rac{2\omega}{k_{\perp}^2} \, \simeq \, rac{2}{\omega heta^2}$$

- "Formation time" : the time it takes to emit a gluon
- This argument universally applies to radiation: in vacuum & in the medium

### Formation time in the vacuum

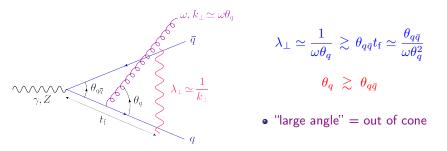
• The emission is triggered by the hard scattering



- The formation time is controlled by the parton virtuality:  $t_{
  m f} \sim 2E/Q^2$
- t<sub>f</sub> is measured from the hard scattering
  - $\bullet\,$  the gluon is emitted within a distance  $\sim t_{\rm f}$  from the scattering vertex
- In medium: additional decoherence introduced by collisions (see below)

### A colorless antenna

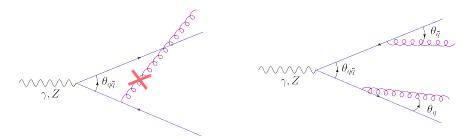
- Interference: the gluon must overlap with both sources during its formation
- Simplest case:  $e^+e^-$  annihilation in a boosted frame  $\Longrightarrow$ 
  - $\vartriangleright$  colorless quark-antiquark "antenna" with opening angle  $\theta_{q\bar{q}} \ll 1$



 Out-of-cone emissions "see" the total color charge — here zero — hence they are not permitted (destructive interference between q and q
)

### A colorless antenna

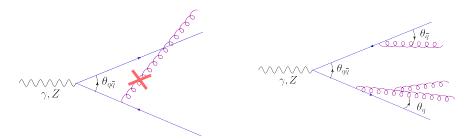
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  - $\rhd$  colorless quark-antiquark "antenna" with opening angle  $\theta_{q\bar{q}} \ll 1$



• Independent emissions by the quark and the antiquark at smaller angles  $\theta_q, \theta_{\bar{q}} < \theta_{q\bar{q}} \Longrightarrow$  "angular ordering"

### A colorless antenna

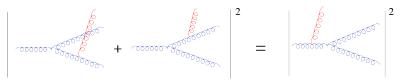
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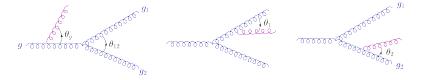
- Independent emissions by the quark and the antiquark at smaller angles  $\theta_q, \theta_{\bar{q}} < \theta_{q\bar{q}} \Longrightarrow$  "angular ordering"
- This argument iterates to the subsequent emissions

### A colored antenna

- E.g.: a color-octet antenna generated by the branching of a gluon
- A subsequent emission at  $\theta > \theta_{gg}$  "sees" the overall color charge
  - it can be formally treated as an emission by the parent gluon



• Effectively: independent emissions with angular ordering

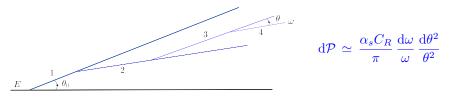


• Quantum emissions with interferences pprox classical branchings with AO

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### **Double-logarithmic approximation**

- Parton showers in the vacuum: successive emissions are ordered in
  - energy  $(\omega_i < \omega_{i-1})$ , by energy conservation
  - angle  $(\theta_i < \theta_{i-1})$ , by color coherence



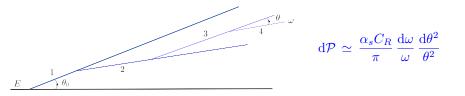
- Log enhancement for soft ( $\omega \ll E$ ) and collinear ( $\theta \ll 1$ ) gluons
- Double-logarithmic approximation (DLA): strong double ordering

 $E \gg \omega_1 \gg \omega_2 \gg \cdots \gg \omega$  &  $\theta_0 \gg \theta_1 \gg \theta_2 \gg \cdots \gg \theta$ 

•  $\theta_0$ : the maximal angle as set by the first emission

### **Double-logarithmic approximation**

- Parton showers in the vacuum: successive emissions are ordered in
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  - angle  $(\theta_i < \theta_{i-1})$ , by color coherence



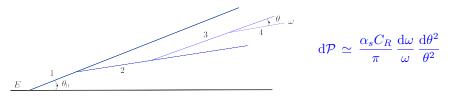
• The double-differential distribution of gluons in the  $(\omega, \theta)$  phase-space:

$$\frac{\mathrm{d}^2 N}{\mathrm{d}\omega \mathrm{d}\theta^2} \simeq \frac{\bar{\alpha}}{\omega \theta^2} \sum_{n \ge 0} \bar{\alpha}^n \left[ \frac{1}{n!} \left( \ln \frac{E}{\omega} \right)^n \right] \left[ \frac{1}{n!} \left( \ln \frac{\theta_0^2}{\theta^2} \right)^n \right]$$

•  $n \ge 0$  intermediate emissions that are not measured: additional sources for the measured gluon with  $(\omega, \theta)$ 

### **Double-logarithmic approximation**

- Parton showers in the vacuum: successive emissions are ordered in
  - energy  $(\omega_i < \omega_{i-1})$ , by energy conservation
  - angle  $(\theta_i < \theta_{i-1})$ , by color coherence



• The double-differential distribution of gluons in the  $(\omega, \theta)$  phase-space:

$$\omega heta^2 rac{\mathrm{d}^2 N}{\mathrm{d}\omega \mathrm{d} heta^2} \,=\, ar{lpha} \mathrm{I}_0 \left( 2 \sqrt{ar{lpha} \ln rac{E}{\omega} \ln rac{ heta_0^2}{ heta^2}} 
ight) \propto \, \exp \left\{ 2 \sqrt{ar{lpha} \ln rac{1}{x} \ln rac{ heta_0^2}{ heta^2}} 
ight\}$$

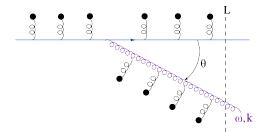
• exponential growth, like for the parton distribution, but no saturation: after being emitted, partons move away from each other

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Jet evolution in a dense medium

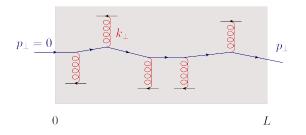
# Jet quenching in pQCD

- Focus on the parton shower, use a simple description for the medium itself
- Weakly coupled quark-gluon plasma in thermal equilibrium at temperature T
  - jet quenching is biased towards (semi)hard momentum transfers
  - see also lectures by Peter Petreczky and Tony Rebhan for the QGP
- The basic physical mechanism for jet-medium interactions: elastic collisions
- Successive collisions are independent
  - $\bullet\,$  mean free path between 2 successive collisions  $\gg$  duration of a collision



### Transverse momentum broadening

• An energetic quark acquires a transverse momentum  $p_{\perp}$  via collisions in the medium, after propagating over a distance L



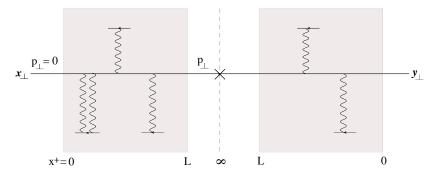
- A random walk in transverse momentum:  $\langle p_{\perp}^2 
  angle \simeq \hat{q}L$
- $\hat{q}$ : the "jet quenching parameter" (a medium transport coefficient)
- $\hat{q}L$  plays exactly the same role as the saturation momentum  $Q_s^2(L)$  for a "large nucleus" with longitudinal width equal to L
- $\bullet~$  Quasi-independent scattering centers in the plasma  $\Longleftrightarrow~$  MV model

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Jet evolution in a dense medium

# **Dipole picture**

• Direct amplitude (DA)  $\times$  Complex conjugate amplitude (CCA)

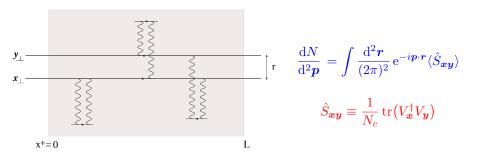


• Scattering can be computed in the eikonal approximation: Wilson lines

$$V^{\dagger}(\boldsymbol{x}) = \operatorname{P}\exp\left\{ig\int \mathrm{d}x^{+}A_{a}^{-}(x^{+},\boldsymbol{x})t^{a}
ight\}$$

• Two such Wilson lines (DA imes CCA)  $\Longrightarrow$  a q ar q color dipole

## Dipole picture (cont.)



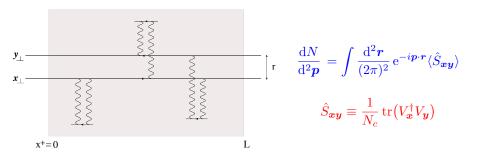
 Average over a Gaussian distribution of color charges representing the quasi-free thermal quarks and gluons:

$$\langle \rho_a(x^+, \boldsymbol{x}) \rho_b(y^+, \boldsymbol{y}) \rangle = g^2 \nu \, \delta^{ab} \delta(x^+ - y^+) \delta^{(2)}(\boldsymbol{x} - \boldsymbol{y})$$

•  $\nu = (C_F n_q + N_c n_g)/N_g \sim T^3$ : color-weighted density of quarks & gluons

• The same as the MV problem with  $\mu^2 
ightarrow g^2 
u L$ 

### Dipole picture (cont.)



• Classical Yang-Mills solution covariant gauge:  $-\nabla_{\perp}^2 A_a^-(x) = \rho_a(x^+, x)$ 

$$\left\langle A_a^-(x^+, \mathbf{k}) A_b^-(y^+, -\mathbf{k}) \right\rangle_0 = \delta_{ab} \delta(x^+ - y^+) \frac{g^2 \nu}{(\mathbf{k}^2 + m_D^2)^2}$$

• Infrared behavior at small- $k_{\perp}$  is regulated by Debye screening

$$m_D^2 = rac{N_c + N_f/2}{3} g^2 T^2$$
 (cf. lecture by Tony Rebhan)

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## The dipole *S*-matrix

• The same as the MV problem with  $\mu^2 o g^2 
u L$  and  $\Lambda_{
m QCD} o m_D$ 

$$S(r) \,=\, \exp\left\{-rac{1}{4}\,L\hat{q}(1/r^2)\,m{r}^2
ight\}$$

• The scale-dependent "jet quenching parameter"  $\hat{q}(Q^2)$ :

$$\hat{q}(Q^2) \equiv g^4 C_F \nu \int^{Q^2} \frac{\mathrm{d}^2 \mathbf{k}}{(2\pi)^2} \frac{\mathbf{k}^2}{(\mathbf{k}^2 + m_D^2)^2} \simeq 4\pi \alpha_s^2 C_F \nu \ln \frac{Q^2}{m_D^2}$$

• A simple interpretation in kinetic theory:  $\Gamma_{\rm el} =$  the elastic collision rate

$$\hat{q}(Q^2) = \int^{Q^2} \! \mathrm{d}^2 oldsymbol{k} \, rac{\mathrm{d}\Gamma_{\mathrm{el}}}{\mathrm{d}^2 oldsymbol{k}} \, oldsymbol{k}^2$$

- $\bullet\,$  the average  $k_{\perp}^2$  acquired per unit time via elastic collisions
- logarithmic ultraviolet divergence: what sets the upper scale  $Q^2$  ?
- $\bullet\,$  In coordinate space, this is fixed by the dipole size  $r\colon\,Q^2\sim 1/r^2$

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## The jet quenching parameter

• The transverse momentum distribution (typical  $p_{\perp}$ ) :

$$\frac{\mathrm{d}N}{\mathrm{d}^2 \boldsymbol{p}} = \int \frac{\mathrm{d}^2 \boldsymbol{r}}{(2\pi)^2} \,\mathrm{e}^{-i\boldsymbol{p}\cdot\boldsymbol{r}} \,\,\mathrm{e}^{-\frac{1}{4}L\hat{q}(1/r^2)\,\boldsymbol{r}^2} \simeq \,\frac{1}{\pi Q_s^2(L)} \,\mathrm{e}^{-p_\perp^2/Q_s^2(L)}$$

• The "plasma saturation momentum"  $Q^2_s(L)$ : exponent of  $\mathcal{O}(1)$ 

$$Q_{s}^{2}(L) = L\hat{q}(Q_{s}^{2}) = 4\pi\alpha_{s}^{2}C_{F}\nu L\ln\frac{Q_{s}^{2}(L)}{m_{D}^{2}} \propto L\ln L$$

• The physical parameter  $\hat{q}$  is self-consistently determined by evaluating the function  $\hat{q}(Q^2)$  for  $Q^2=\hat{q}L$  :

$$\hat{q} = 4\pi \alpha_s^2 C_F \nu \ln \frac{\hat{q}L}{m_D^2} \sim \alpha_s^2 T^3 \ln(\alpha_s^2 TL)$$

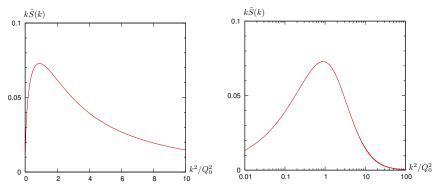
• Power-law tail at large  $p_{\perp} \gg Q_s$ : single hard scattering

$$\frac{{\rm d}N}{{\rm d}^2 {\bm p}} \,\simeq\, \int \frac{{\rm d}^2 {\bm r}}{(2\pi)^2}\, {\rm e}^{-i {\bm p}\cdot {\bm r}} \, \left\{ -\frac{1}{4} L \hat{q}(1/r^2)\, {\bm r}^2 \right\} \,\simeq\, \frac{4\pi \alpha_s^2 C_F \nu}{p_\perp^4} \label{eq:delta_eq}$$

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## $p_{\perp}$ -broadening: numerical results

- Left: the dipole Fourier transform  $k_{\perp} \tilde{\mathcal{S}}(k_{\perp}) \propto k_{\perp} (\mathrm{d}N/\mathrm{d}^2k_{\perp})$ 
  - the probability distribution for  $k_{\perp}$
  - "the dipole unintegrated gluon distribution"

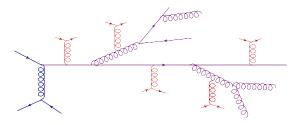


• Right: the same function, but in logarithmic units

• peaked at  $k \simeq Q_s$ , power-law tail at  $k \gg Q_s$ 

## Medium-induced radiation

• Collisions in the medium can trigger additional radiation (similarly to the original hard scattering): they provide acceleration ("virtuality")



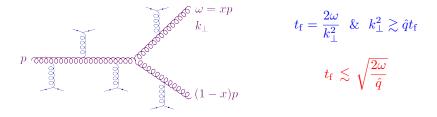
- A priori, 2 mechanisms for radiation ...
  - "vacuum-like" triggered by the original hard scattering
  - "medium-induced" associated rescattering in the medium
- ... but can one distinguish between them ?
- Look at the formation times: it takes a time  $t_{\rm f} = \frac{2\omega}{k_\perp^2}$  to emit a gluon with energy  $\omega$  and transverse momentum  $k_\perp$

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Jet evolution in a dense medium

## Medium-induced radiation (cont.)

- In the vacuum,  $\omega$  and  $k_{\perp}$  are independent kinematical variables
  - collinear radiation  $(k_{\perp} \simeq \omega \theta \rightarrow 0)$  has large formation times
- In the medium, collisions introduce a lower limit on  $k_{\perp}$  ...



- ... hence an upper limit on the formation time !
  - vacuum-like emissions:  $k_{\perp}^2 \gg \hat{q} t_{
    m f}$  , or  $t_{
    m f} \ll \sqrt{2\omega/\hat{q}}$
  - medium-induced emissions:  $k_{\perp}^2 \simeq \hat{q} t_{
    m f}$  , or  $~t_{
    m f} \simeq \sqrt{2\omega/\hat{q}}$
- In the medium, there is no genuinely collinear radiation:  $k_{\perp}^2 \geq k_{
  m f}^2 \equiv \sqrt{2\omega\hat{q}}$

## Radiation from multiple soft scattering

#### (Baier, Dokshitzer, Mueller, Peigné, Schiff; Zakharov; 1996-97)

- Medium-induced emissions can occur anywhere inside the medium
  - no correlation with the original hard process
- It is appropriate to work with the ("BDMPS-Z") emission rate :
  - differential probability for one medium-induced emission per unit phase-space per unit time

$$\frac{\mathrm{d}\mathcal{P}}{\mathrm{d}\omega\,\mathrm{d}^{2}\boldsymbol{k}\,\mathrm{d}t} \,=\, \frac{\alpha_{s}C_{R}}{\pi}\,\frac{1}{\omega}\,\frac{1}{t_{\mathrm{med}}(\omega)}\,\frac{1}{\sqrt{2\hat{q}\omega}}\,\mathrm{e}^{-\frac{k_{\perp}^{2}}{\sqrt{2\hat{q}\omega}}},\qquad t_{\mathrm{med}}(\omega)\equiv\sqrt{\frac{2\omega}{\hat{q}}}$$

- transverse momentum broadening during formation:  $\langle k_{\perp}^2 \rangle \simeq \sqrt{2 \hat{q} \omega}$
- $\omega t_{
  m med}(\omega) \propto \omega^{3/2} \Longrightarrow$  no soft or collinear logarithm
- This applies so long mean-free-path  $< t_{
  m med}(\omega) \leq L$

$$T < \omega \leq \omega_c \equiv \frac{1}{2} \hat{q} L^2$$

# Angular distribution

- Transverse momentum broadening during formation & after formation
- formation angle:

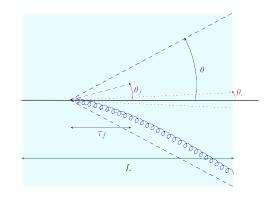
$$heta_{
m f}(\omega) \simeq rac{(2\hat{q}\omega)^{1/4}}{\omega} \simeq \left(rac{2\hat{q}}{\omega^3}
ight)^{1/4}$$

• the minimal angle

$$\theta_c = \theta_{\rm f}(\omega_c) = \frac{2}{\sqrt{\hat{q}L^3}}$$

• final angle

$$heta(\omega) \simeq rac{\sqrt{\hat{q}L}}{\omega} = heta_c rac{\omega_c}{\omega}$$



- Soft gluons  $\omega \ll \omega_c$ : small formation times  $(t_f \ll L)$  & large angles  $(\theta \gg \theta_c)$
- Gluons with angles larger than the jet opening angle  $\theta_0$  move outside the jet

## Average energy loss by the LP

• Differential probability for one emission (integrated over t and over  $k_{\perp}$ )

$$\omega \frac{\mathrm{d}\mathcal{P}}{\mathrm{d}\omega} \simeq \bar{\alpha} \frac{L}{t_{\mathrm{med}}(\omega)} \simeq \bar{\alpha} \sqrt{\frac{\omega_c}{\omega}} \qquad (\omega < \omega_c \equiv \hat{q}L^2/2)$$

• The average energy loss by a particle with energy  $E > \omega_c$ 

$$\langle \Delta E \rangle = \int^{\omega_c} \mathrm{d}\omega \; \omega \; \frac{\mathrm{d}\mathcal{P}}{\mathrm{d}\omega} \; \sim \; \alpha_s \omega_c \; \sim \; \alpha_s \hat{q} L^2$$

- integral dominated by its upper limit  $\omega = \omega_c$
- Hard emissions with  $\omega \sim \omega_c$ : probability of  $\mathcal{O}(\alpha_s)$ 
  - rare events but which take away a large energy
  - $\bullet\,$  small emission angle  $\theta_c \Rightarrow$  the energy remains inside the jet
- Irrelevant for the di-jet asymmetry and also for the typical events

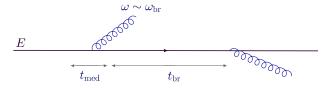
J.-P. Blaizot, E. I., Y. Mehtar-Tani, PRL 111, 052001 (2013)

$$\omega \frac{\mathrm{d}\mathcal{P}}{\mathrm{d}\omega} \simeq \bar{\alpha} \frac{L}{t_{\mathrm{med}}(\omega)} \simeq \bar{\alpha} \sqrt{\frac{\omega_c}{\omega}}$$

• When  $\omega(d\mathcal{P}/d\omega)\gtrsim 1$ , multiple branching becomes important

$$\omega \lesssim \omega_{\rm br}(L) \equiv \bar{\alpha}^2 \omega_c \quad \Longleftrightarrow \quad L \gtrsim t_{\rm br}(\omega) \equiv \frac{1}{\bar{\alpha}} t_{\rm med}(\omega)$$

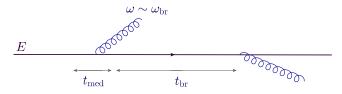
•  $\omega_{
m br}\sim ar{lpha}^2 \hat{q} L^2$  : characteristic energy for the onset of multiple branching



•  $t_{\rm br} = t_{\rm med}/\bar{\alpha}$ : typical distance between 2 successive branchings

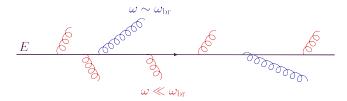
•  $t_{
m br} \gg t_{
m med}$  : successive emissions do not overlap with each other

• LHC: the leading particle has  $E \sim 100 \, {
m GeV} \gg \omega_{
m br} \sim 5 \, {
m GeV}$ 



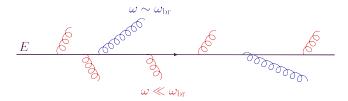
- In a typical event, the LP emits ...
  - a number of  $\mathcal{O}(1)$  of gluons with  $\omega \sim \omega_{\mathrm{br}}$

• LHC: the leading particle has  $E \sim 100 \, {
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- In a typical event, the LP emits ...
  - a number of  $\mathcal{O}(1)$  of gluons with  $\omega \sim \omega_{\mathrm{br}}$
  - a large number of softer gluons with  $\omega \ll \omega_{
    m br}$
  - the energy loss is controlled by the hardest primary emissions

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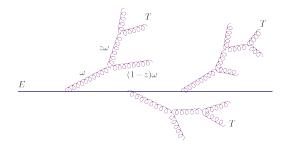


- In a typical event, the LP emits ...
  - a number of  $\mathcal{O}(1)$  of gluons with  $\omega \sim \omega_{\mathrm{br}}$
  - a large number of softer gluons with  $\omega \ll \omega_{
    m br}$
  - the energy loss is controlled by the hardest primary emissions
- In a typical event, the LP loses an energy  $\Delta E \sim \omega_{\rm br}$ 
  - albeit smaller than the average energy loss,  $\Delta E \sim \bar{\alpha} \langle \Delta E \rangle$ , this typical energy loss is more important for our purposes
- This is also the energy lost by the jet

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## **Democratic branchings**

- The primary gluons generate 'mini-jets' via democratic branchings
  - daughter gluons carry comparable energy fractions:  $z \sim 1-z \sim 1/2$



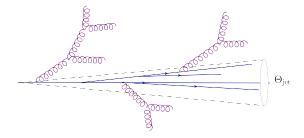
• the probability for the splitting  $\omega \to (z\omega,\,(1-z)\omega)$  :

$$\mathcal{P}(z\omega,L)\simeq rac{L}{t_{
m br}(z\omega)}\simeq ar{lpha}\,L\,\sqrt{rac{\hat{q}}{z\omega}}$$

• when  $\omega\sim\omega_{
m br},\, {\cal P}(z\omega,L)\sim 1$  independently of the value of z

## Energy loss by the jet

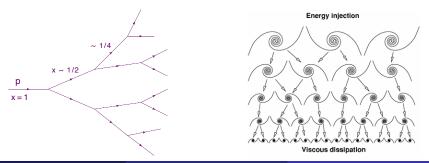
- Democratic branchings are very unusual in the context of gauge theories
  - $\bullet\,$  recall: bremsstrahlung favors asymmetric splittings with  $z\ll 1$
- Extremely efficient in redistributing the energy among softer quantas
  - a mini-jet with  $\omega \lesssim \omega_{
    m br}$  "disappears" in a time  $t_{
    m br}(\omega) \lesssim L$
  - its energy is successively transmitted to softer and softer gluons
- Soft gluons are easily deviated outside the jet cone by the elastic collisions



ullet The energy appears in many soft quanta propagating at large angles  $\checkmark$ 

#### Wave turbulence

- Democratic branchings lead to wave turbulence
  - energy flows from one parton generation to the next one, at a rate which is independent of the generation
  - $\bullet\,$  formally, it accumulates into a condensate at  $x=0\,$
  - physically, it dissipates into the medium, via thermalization
  - similar to Kolmogorov turbulence ... but much simpler: 1+1 dimensions, inverse cascade, exact solutions



## Probabilistic picture for medium-induced radiation

#### Blaizot, Dominguez, E.I., Mehtar-Tani (arXiv:1311.5823)

- Medium-induced jet evolution  $\approx$  a Markovien stochastic process
  - successive branchings are non-overlapping:  $t_{\rm br} \sim \frac{1}{\alpha_{\rm s}} t_{\rm med}$
  - interference phenomena could complicate the picture ... (in the vacuum, they lead to angular ordering)
  - ... but they are suppressed by rescattering in the medium (see below)
- Hierarchy of equations for *n*-point correlation functions ( $x \equiv \omega/E$ )

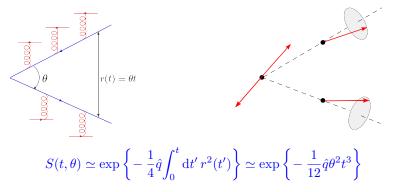
$$D(x,t) \equiv x \frac{\mathrm{d}N}{\mathrm{d}x}(t), \qquad D^{(2)}(x,x',t) \equiv xx' \frac{\mathrm{d}N_{\mathrm{pair}}}{\mathrm{d}x \,\mathrm{d}x'}(t)$$

- Analytic solutions (Blaizot, E.I., Mehtar-Tani, '13; Escobedo, E.I., '16)
- New phenomena: wave turbulence, KNO scaling, large fluctuations...
- Recent Monte-Carlo implementation (see the talk by Paul Caucal next week) (P. Caucal, E.I., G. Soyez, A.H. Mueller, to appear)

## Color antenna in the medium

#### Mehtar-Tani, Salgado, Tywoniuk (2010-11); Casalderrey-Solana, E. I. (2011)

- The two quarks undergo independent color precessions due to rescattering
- The instantaneous color state of each quark: the respective Wilson line
- Their color correlation is measured by the dipole S-matrix



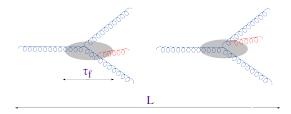
• Color coherence is lost for  $t\gtrsim t_{
m coh}\sim 1/(\hat{q} heta^2)^{1/3}$ 

## Color decoherence for medium induced emissions

• Consider an antenna generated via a medium-induced gluon splitting:

$$heta_{
m f}(\omega) \simeq \left(rac{2\hat{q}}{\omega^3}
ight)^{1/4} \Longrightarrow t_{
m coh} \sim rac{1}{(\hat{q} heta_{
m f}^2)^{1/3}} \sim \sqrt{rac{2\omega}{\hat{q}}} = t_{
m med}(\omega)$$

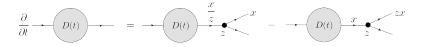
- the two daughter gluons lose their mutual color coherence already during the formation time
- not a "coincidence": decoherence via rescattering is what triggers the medium-induced radiation



### Medium-induced cascade: the gluon spectrum

#### J.-P. Blaizot, E. I., Y. Mehtar-Tani, 2013

• Kinetic equation for D(x,t) = x(dN/dx): 'gain' - 'loss'



• Exact solution with initial condition  $D(x, t = 0) = \delta(x - 1)$ 

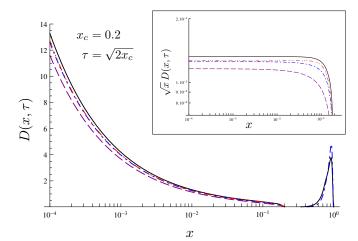
$$D(x,\tau) = \frac{\tau}{\sqrt{x}(1-x)^{3/2}} e^{-2\pi \frac{\tau^2}{1-x}}, \qquad x \equiv \frac{\omega}{E}, \ \tau \equiv \frac{t}{t_{\rm br}(E)}$$

•  $t_{\rm br}(E)$  : the lifetime of the LP until its first democratic branching

- power-law spectrum  $D \propto \frac{1}{\sqrt{x}}$  at  $x \ll 1$  for any au
- Kolmogorov fixed point: wave turbulence

### Medium-induced cascade: the gluon spectrum

• Numerical results for D(x,t) = x(dN/dx):  $E = 5\omega_c$  and  $\bar{\alpha} = 0.3$ 



- pronounced LP peak near x = 1
- power-law spectrum  $D \propto \frac{1}{\sqrt{x}}$  at  $x \ll 1$

# Energy loss at (very) large angles

- The energy leaks into the condensate at  $x = 0 \implies$  the energy fraction contained in the spectrum decreases with time:  $\int_0^1 dx D(x, \tau) = e^{-2\pi\tau^2}$
- The energy lost at very large angles = the energy missing from the spectrum

$$\Delta E = E \left( 1 - e^{-2\pi\tau^2} \right) = E \left[ 1 - e^{-2\pi\frac{\omega_{\rm br}}{E}} \right]$$

• When  $E \gg \omega_{\rm br}$ , this is independent of  $E: \Delta E \simeq 2\pi\omega_{\rm br} = \pi \bar{\alpha}_s^2 \hat{q} L^2$ 

• LHC:  $E \sim 100 \,\mathrm{GeV} \gg \omega_{\mathrm{br}} \sim 5 \div 10 \,\mathrm{GeV}$ 

• Additional, angle-dependent, contribution from soft modes in the spectrum

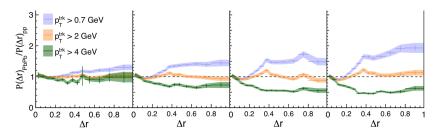
$$heta(\omega) \simeq rac{\sqrt{\hat{q}L}}{\omega} \gtrsim heta_0 \implies \omega \lesssim rac{\sqrt{\hat{q}L}}{ heta_0}$$

• A natural explanation for the di-jet asymmetry observed at the LHC

# Energy loss at (very) large angles

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$$\Delta E = E \left( 1 - e^{-2\pi\tau^2} \right) = E \left[ 1 - e^{-2\pi\frac{\omega_{\rm br}}{E}} \right]$$

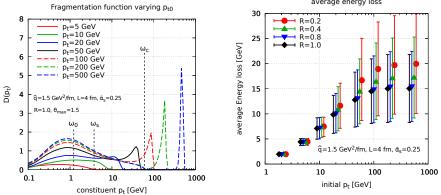


• CMS: The jet transverse momentum profile  $P(\Delta r)$  ( $p_T$  in annular rings)

• Pb+Pb vs. p+p: excess of soft particles at large angles outside the jet

### Monte Carlo: Medium-induced radiation only

- Initial parton with energy  $E \equiv p_{t0}$ , jet with opening angle R
- Left: fragmentation function (or spectrum)  $D(\omega) = \omega \frac{dN}{d\omega}$ 
  - leading-particle peak so long as  $E \gg \omega_{\rm br} \, (\equiv \omega_s)$
  - medium-induced radiation:  $D(\omega) \propto \frac{1}{\sqrt{\omega}}$  at  $\omega < \omega_c$

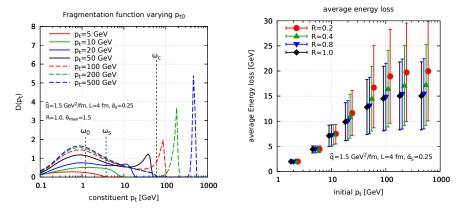


average energy loss

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## Monte Carlo: Medium-induced radiation only

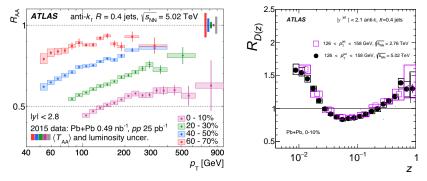
- Initial parton with energy  $E \equiv p_{t0}$ , jet with opening angle R
- Right: average energy loss by the jet + its fluctuations
  - $\Delta E$  becomes independent of E when  $E \gg \omega_{\rm br}$
  - $\bullet\,$  energy is recovered when increasing  $R\,\ldots\,$  but very slowly



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## With due respect to the vacuum

- So far: energy loss at large angles via medium-induced emissions
  - energy loss predicted to approach a constant value at high energy
- The data for  $R_{AA}$  suggest that the energy loss should increase with  $p_T$



- Medium-induced modifications of the hadron distribution inside the jet cone
- Features that are naturally explained after adding the vacuum-like radiation

## Vacuum-like emissions inside the medium

- "What do you mean by that ?!?"
- Emissions triggered by parton virtualities, not by collisions in the medium
- Can they be different from the standard parton showers in the vacuum (say, as measured in *pp* collisions) ?
- YES, THEY CAN ! They have a restricted phase-space ...
  - no collinear singularity, prompt formation times

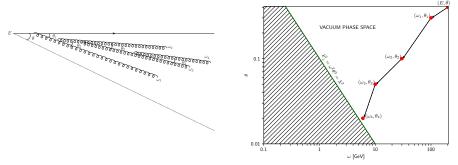
$$t_{
m f} = rac{2}{\omega heta^2} < t_{
m med} = \sqrt{rac{2\omega}{\hat{q}}} \qquad {
m or} \qquad heta > \left(rac{2\hat{q}}{\omega^3}
ight)^{1/4}$$

• They can suffer energy loss &  $p_{\perp}$ -broadening via medium-induced emissions

- they do so ... but only after formation !
- They can be affected by color decoherence due to in-medium scattering
  - is angular ordering still there ?

## Lund plot: vacuum emissions at DLA

- Each emission: a point  $(\omega_i, \theta_i)$  in the energy–emission angle phase-space
  - logarithmic units: convenient for the double logarithmic approximation
  - $\overline{\theta}$ : the maximal angle allowed for the first emission
  - E: the energy of the leading parton, hence the maximal energy allowed to the first emission

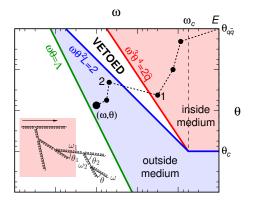


• Evolution stopped by hadronisation:  $k_{\perp} \simeq \omega \theta \gtrsim \Lambda_{\rm QCD}$ 

## Vacuum-like emissions inside the medium

#### P. Caucal, E. I., A. Mueller, G. Soyez, arXiv:1801.09703, PRL 120 (2018)

• Vacuum-like parton cascades occur relatively fast and can be factorized in time from the medium-induced cascades



- inside-medium cascades:
  - $\frac{2}{\omega\theta^2} < \sqrt{\frac{2\omega}{\hat{q}}}$
- outside-medium emissions:

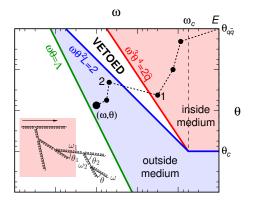
$$\frac{2}{\omega\theta^2} > L$$

- forbidden region in between
- End point of VETOED at  $\omega_c = \frac{1}{2}\hat{q}L^2$ ,  $\theta_c = 2/\sqrt{\hat{q}L^3}$ 
  - typical values:  $\hat{q} = 1 \, {\rm GeV^2/fm}, \ L = 4 \, {\rm fm}, \ \omega_c = 50 \, {\rm GeV}, \ \theta_c = 0.05$

## Vacuum-like emissions inside the medium

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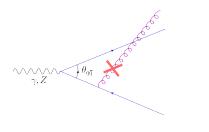
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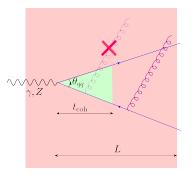
$$\frac{2}{\omega\theta^2} > L$$

- forbidden region in between
- End point of VETOED at  $\omega_c = \frac{1}{2} \hat{q} L^2 , \ \theta_c = 2/\sqrt{\hat{q} L^3}$
- Energy loss &  $p_{\perp}$ -broadening during formation are negligible

# Color (de)coherence

• In vacuum, wide angle emissions  $(\theta > \theta_{q\bar{q}})$  are suppressed by color coherence





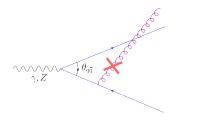
• In medium, color coherence is washed out by collisions after a time  $t_{\rm coh}$ 

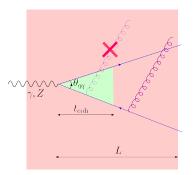
$$t_{\rm coh} = rac{1}{(\hat{q} heta_{qar{q}}^2)^{1/3}} \ll L \quad {\rm if} \quad heta_{qar{q}} \gg heta_c \simeq 0.05$$

• Angular ordering could be violated for emissions inside the medium

# Color (de)coherence

• In vacuum, wide angle emissions  $(\theta > \theta_{q\bar{q}})$  are suppressed by color coherence





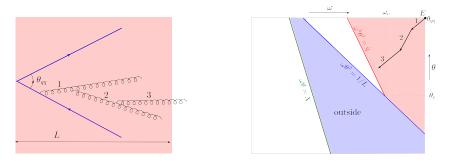
• But this is not the case for the VLEs !

$$heta > heta_{qar q} \quad \& \quad t_{
m f} = rac{2}{\omega heta^2} \ll \sqrt{rac{2\omega}{\hat q}} \implies t_{
m f} \ll t_{
m coh}$$

• Wide-angle VLEs  $(\theta > \theta_{q\bar{q}})$  are still suppressed  $\implies$  angular ordering

## Vacuum-like cascades inside the medium

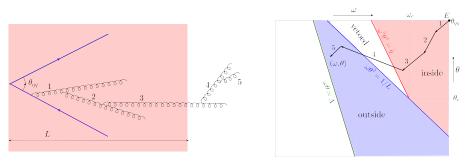
- Previous arguments extend to a sequence of angular-ordered VLEs
  - $\bullet\,$  angular ordering  $\Longrightarrow$  formation times are strongly increasing
  - $\bullet\,$  formation time for the whole cascade  $\approx\,$  that of the last gluon



- After formation, partons propagate in the medium along a distance  $\sim L$ 
  - additional sources for medium-induced radiation
  - ... and for vacuum-like emissions outside the medium

## First emission outside the medium

- The respective formation time is necessarily large:  $t_{
  m f}\gtrsim L$
- In-medium sources with  $heta \gg heta_c$  lose coherence in a time  $t_{
  m coh} \ll L$
- First outside emission can violate angular ordering
  - re-opening of the angular phase-space
  - enhanced radiation at low energies and large angles

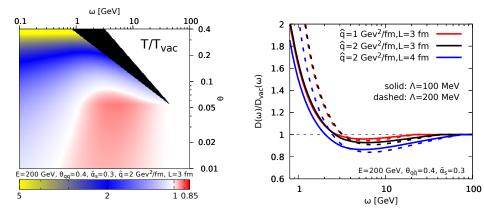


• Subsequent "outside" emissions obey angular-ordering, as usual

#### DLA as a warm up

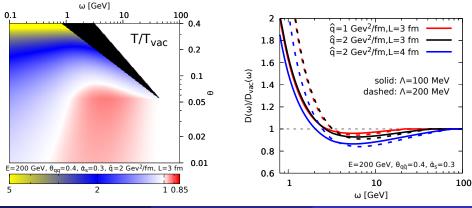
• The double-differential distribution & the fragmentation function:

$$T(\omega,\theta) \equiv \omega \theta^2 \frac{\mathrm{d}^2 N}{\mathrm{d}\omega \mathrm{d}\theta^2} \implies D(\omega) \equiv \omega \frac{\mathrm{d}N}{\mathrm{d}\omega} = \int \frac{\mathrm{d}\theta^2}{\theta^2} T(\omega,\theta^2)$$



#### DLA as a warm up

- Slight suppression at intermediate energies (from 3 GeV up to  $\omega_c$ )
  - the phase-space is reduced by the vetoed region
- Significant enhancement at low energy (below 2 GeV)
  - lack of angular ordering for the first emission outside the medium

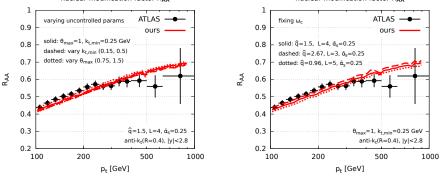


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### First Monte Carlo results: $R_{AA}$

(P. Caucal, E.I., A. H. Mueller and G. Soyez, in preparation)

- Probabilistic (Markovian) picture  $\implies$  straightforward MC implementation
- Not a fit: just a choice of "reasonable" values for the physical parameters
- Left: varying kinematical cuts:  $k_{t,\min}\equiv \Lambda$ ,  $heta_{\max}\equiv heta_{qar q}$
- Right:  $R_{AA}$  depends upon the medium only via  $\omega_c = \hat{q}L^2$

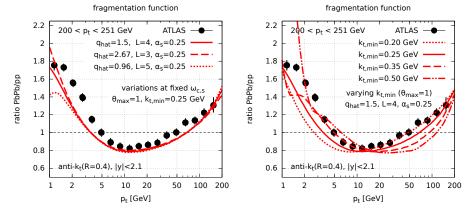


nuclear modification factor R<sub>AA</sub>

nuclear modification factor RAA

## Monte Carlo: Fragmentation function

- The same "canonical" values for the medium parameters as in the "best" description of  $R_{AA}$  (ATLAS data from arXiv:1805.05424)
- Stronger variability with respect to the "uncontrolled parameters"



• See the talk by Paul Caucal next Wednesday for more details !

## THANK YOU FOR YOUR ATTENTION !





Have a good end of stay at ICTS and a safe trip back home!

THE MYRIAD, ICTS, April 2019

Jet evolution in a dense medium