Smooth Causal Patches for AdS Black Holes

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This talk is based on

S.R, "Smooth Causal Patches for AdS Black Holes", arXiv:1604.03095.

Also based on previous work with Kyriakos Papadodimas (CERN & Groningen), Souvik Banerjee (Groningen) and Jan-Willem Bryan (Groningen) and work in progress with Sudip Ghosh (ICTS-TIFR).



• The context for this talk is the Information Paradox. In its modern avatar, this turns into the question:

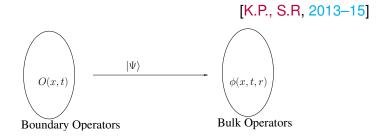
"Can Holography describe the BH Interior?"

[Mathur, Almheiri, Marolf, Polchinski, Sully, Stanford, 2009-2015]

• This version is not restricted to evaporating BHs, but also applies to thermodynamically stable large black holes in AdS.

Context

• Resolution: Paradox can be completely resolved using a state-dependent map between interior bulk observables and boundary observables.



Is this consistent? Or does it violate the linearity of QM?



 Marolf and Polchinski suggested that state-dependent constructions of the BH interior violate a general rule of statistical mechanics

"Low energy excitations in a large thermal system have small effects on observables"

 I will show that if these violations are unobservable due to causality as manifested in properties of AdS position-space correlators.

Outline



- 2 Holography and the BH Interior
- 3 The Paradox of Low Energy Excitations

4 Resolution



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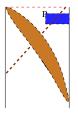
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- 4 Resolution
- 5 Summary and Open Questions

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The Old Information Paradox



• In the shaded patch, physics is independent of details of collapse.

$$\langle a^{\dagger}_{\omega} a_{\omega'} \rangle = rac{e^{-eta \omega}}{1 - e^{-eta \omega}} \delta(\omega - \omega')$$

Suggests that for different inputs, we get the same output.

Input A
$$--- > \begin{bmatrix} Black \\ Hole \end{bmatrix} - -- > \begin{bmatrix} Black \\ Radiation \end{bmatrix}$$

Input B - - -
$$\gg$$
 Black Body Radiation

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Smooth Causal Patches

Resolution to the Old Information Paradox

- Very small corrections of the order of e^{-S} can restore unitarity.
 [Maldacena, 2001]
- Pure density matrix in a very large system can mimic a thermal density matrix to extreme accuracy

$$\operatorname{Tr}\left(\rho_{\mathsf{pure}}\boldsymbol{A}_{\alpha}\right) = \frac{1}{\mathcal{Z}}\operatorname{Tr}\left(\boldsymbol{e}^{-\beta H}\boldsymbol{A}_{\alpha}\right) + \mathsf{O}\left(\boldsymbol{e}^{-\frac{S}{2}}\right),$$

for a large class of observables A_{α} .

Another way to state this is

$$\rho_{\text{pure}} = \frac{1}{\mathcal{Z}} e^{-\beta H} + e^{-S} \rho_{\text{corr}}; \quad \rho_{\text{pure}}^2 = \rho_{\text{pure}}$$

Path Integral Perspective

- Effective field theory insufficient to control such corrections.
- A semi-classical spacetime is a saddle point of the QG path-integral.

$$\mathcal{Z} = \int e^{-S} \mathcal{D} g_{\mu\nu}$$

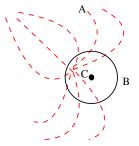
- Perturbative effective field theory (used to derive the Hawking answer) is an asymptotic series expansion of this path-integral.
- Non-perturbatively, the notion of local spacetime breaks down.

Complementarity

• For example, in a very high-point correlator

$$\langle \phi(x_1) \dots [\phi(x_{\text{out}}), \phi(x_{\text{in}})] \dots \phi(x_S) \rangle \neq 0.$$

 Hilbert space does not factorize into far-away region and near-horizon region.



• Concrete example in empty AdS.

[S. Banerjee, J.W. Bryan, K. Papadodimas, S. R. ,2016]

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Old Information Paradox: Slogan

Hawking's calculation is not precise enough to lead to a paradox.

Outline



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Modern Information Paradox

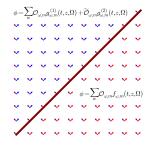
Can the black hole interior be described holographically?

This may look different from the information paradox. But it is still Unitarity vs Effective Field Theory

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Inside a Black Hole

 To describe a local field inside the black hole, we need both left and right movers.



- Can think of $\widetilde{\mathcal{O}}_{\omega,m}$ as modes that have bounced off the origin (Hawking)
- Can also think of them as modes coming from left asymptotic region of the eternal black hole.

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Smoothness of the Horizon and Two-Point correlators

Smoothness of the horizon \leftrightarrow KMS condition for $\mathcal{O}_{\omega,m}$ and $\widetilde{\mathcal{O}}_{\omega,m}$

$$\begin{split} \langle \Psi | \widetilde{\mathcal{O}}_{\omega,m} \widetilde{\mathcal{O}}_{\omega',m'}^{\dagger} | \Psi \rangle &= \langle \Psi | \mathcal{O}_{\omega,m} \mathcal{O}_{\omega',m'}^{\dagger} | \Psi \rangle = \frac{1}{1 - e^{-\beta\omega}} \delta(\omega - \omega') \delta_{mm'} \mathcal{C}_{\omega,m} \\ \langle \Psi | \widetilde{\mathcal{O}}_{\omega,m} \mathcal{O}_{\omega',m'}^{} | \Psi \rangle &= \mathcal{C}_{\omega,m} \frac{e^{\frac{-\beta\omega}{2}}}{1 - e^{-\beta\omega}} \delta(\omega - \omega') \delta_{mm'} \\ \langle [\mathcal{O}_{\omega,m}, \mathcal{O}_{\omega',m'}^{\dagger}] \rangle &= \mathcal{C}_{\omega,m} \delta(\omega - \omega') \delta_{mm'} \end{split}$$

These must hold in typical states if typical states correspond to smooth-horizons.

How does one describe the $\widetilde{\mathcal{O}}_{\omega,m}$ in the CFT?

Unusual Properties of Mirror Modes

• From analysis of large diffeomorphisms, we find

$$[H,\widetilde{\mathcal{O}}_{\omega}]=\omega\widetilde{\mathcal{O}}_{\omega}$$

• Effective field theory requires the KMS condition

$$\langle \Psi | \widetilde{\mathcal{O}}_{\omega} \widetilde{\mathcal{O}}_{\omega}^{\dagger} | \Psi
angle = oldsymbol{e}^{eta \omega} \langle \Psi | \widetilde{\mathcal{O}}_{\omega}^{\dagger} \widetilde{\mathcal{O}}_{\omega} | \Psi
angle$$

If the black hole state is approximately thermal

$$egin{aligned} &\langle \Psi | \widetilde{\mathcal{O}}_{\omega} \widetilde{\mathcal{O}}_{\omega}^{\dagger} | \Psi
angle pprox rac{1}{Z(eta)} \mathrm{Tr}(e^{-eta H} \widetilde{\mathcal{O}}_{\omega} \widetilde{\mathcal{O}}_{\omega}^{\dagger}) = rac{1}{Z(eta)} e^{-eta \omega} \mathrm{Tr}(e^{-eta H} \widetilde{\mathcal{O}}_{\omega}^{\dagger} \widetilde{\mathcal{O}}_{\omega}) \ &pprox e^{-eta \omega} \langle \Psi | \widetilde{\mathcal{O}}_{\omega}^{\dagger} \widetilde{\mathcal{O}}_{\omega} | \Psi
angle? \end{aligned}$$

using equivalence of microcanonical and canonical ensembles, cyclicity of trace and commutator with Hamiltonian.

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The Little Hilbert Space

- $\bullet \ |\Psi\rangle \equiv \text{Black Hole Microstate}$
- Little Hilbert Space: all possible effective field theory excitations of $|\Psi\rangle$

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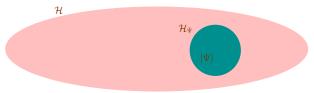
$$\mathcal{H}_{\Psi} = \mathcal{A}|\Psi\rangle,$$

$$\mathcal{A} = \operatorname{span}\{\mathcal{O}_{\omega_1}, \ \mathcal{O}_{\omega_1}\mathcal{O}_{\omega_2}, \dots, \mathcal{O}_{\omega_1}\mathcal{O}_{\omega_2}\dots \mathcal{O}_{\omega_K}\}.$$

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with





• $H_{\Psi} = H_{\text{code}}$ in QEC discussions.

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Definition of $\widetilde{\mathcal{O}}_{\omega}$

• Define $\widetilde{\mathcal{O}}_{\omega}$ precisely within H_{Ψ}

$$SA_lpha|\Psi
angle=A^\dagger_lpha|\Psi
angle$$

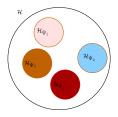
and

$$\widetilde{\mathcal{O}}_{\omega} = \mathcal{S} e^{rac{eta H}{2}} \mathcal{O}_{\omega} e^{-rac{eta H}{2}} \mathcal{S}$$

[KP, SR, 2013]

- This is closely related to the isomorphism used in Tomita-Takesaki theory.
- φ(t, Ω, λ) constructed using this *Õ*_ω is a linear operator on H_Ψ and has the correct effective field theory correlators

State-dependence



- no single linear operator $\widetilde{\mathcal{O}}_{\omega}$ behaves correctly in all H_{ψ} .
- State-dependence means that we must use different operators O_ω in different H_ψ.

Consistency of State-Dependence

- State-dependence resolves all paradoxes that suggested black-hole firewalls.
- Also leads to a precise description of ER=EPR.
- "Does interior exist in holography?" = "Is state-dependence consistent?"
- **Rest of the talk:** A paradox invented by Marolf and Polchinski (2015) to test consistency of state-dependence and its resolution.

Outline





3 The Paradox of Low Energy Excitations

4 Resolution



4 **A b b b b b b**

A Theorem in Statistical Mechanics

• A typical state in a system with many degrees of freedom is thermal for coarse-grained probes:

$$\langle \Psi | A_{\alpha} | \Psi \rangle = rac{1}{Z(\beta)} \operatorname{Tr}(e^{-\beta H} A_{\alpha}) + O\left(rac{1}{\sqrt{S}}\right),$$

• Consider a low energy excitation

$$\langle \Psi | U^{\dagger} H U | \Psi \rangle - \langle \Psi | H | \Psi \rangle = \delta E.$$

• Theorem:

$$\langle \Psi | U^{\dagger} A_{lpha} U | \Psi
angle - \langle \Psi | A_{lpha} | \Psi
angle \leq 2 \sqrt{eta \delta E} \sigma_{lpha}$$

 $(\sigma_{lpha}^{2}=\langle A_{lpha}^{2}
angle -\langle A_{lpha}
angle^{2})$

Sketch of Proof

$$U : \mathcal{H}_E \to \mathcal{H}_{E+\delta E},$$

 $\dim(\mathcal{H}_E) = e^{S(E)}, \quad \dim(\mathcal{H}_{E+\delta E}) = e^{S(E)+\beta E}$

So, decompose a typical state $|\Psi_{E+\delta E}\rangle \in \mathcal{H}_{E+\delta E}$,

$$|\Psi_{E+\delta E}
angle = (1 - rac{eta\delta E}{2})U|\Psi_E
angle + (eta\delta E)^{rac{1}{2}}|\Psi_{ ext{orth}}
angle + O\left((eta\delta E)^{rac{3}{2}}
ight).$$

Ensemble at higher energy has same temperature

$$\langle \Psi_{E+\delta E} | A_{\alpha} | \Psi_{E+\delta E} \rangle = \langle \Psi_{E} | A_{\alpha} | \Psi_{E} \rangle + O\left(\frac{1}{\sqrt{S}}\right).$$

Therefore,

$$\begin{split} \langle \Psi_{E+\delta E} | \mathbf{A}_{\alpha} | \Psi_{E+\delta E} \rangle &- \langle \Psi_{E} | \mathbf{U}^{\dagger} \mathbf{A}_{\alpha} \mathbf{U} | \Psi_{E} \rangle \\ &= \delta \langle \mathbf{A}_{\alpha} \rangle = \sqrt{\beta \delta E} \left(\langle \Psi_{E} | \mathbf{U}^{\dagger} \mathbf{A}_{\alpha} | \Psi_{\text{orth}} \rangle + \langle \Psi_{\text{orth}} | \mathbf{A}_{\alpha} \mathbf{U} | \Psi_{E} \rangle \right) + \mathcal{O} \left(\beta \delta E \right). \end{split}$$

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Low Energy Excitations

$\delta \langle \mathbf{A}_{\alpha} \rangle \leq 2 \sqrt{\beta \delta E} \sigma_{\alpha}$

Low energy excitations have small effects: it is impossible to definitively excite a thermal system with energy less than kT.

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Marolf-Polchinski Paradox

With

$$\mathcal{O}_{\omega} = \int_{-T}^{T} \mathcal{O}(t) \boldsymbol{e}^{i\omega t} dt, \quad \boldsymbol{A}_{\alpha} = \mathcal{O}_{\omega} \widetilde{\mathcal{O}}_{\omega}.$$

we need, for a smooth horizon,

$$\langle \Psi | \mathcal{O}_{\omega} \widetilde{\mathcal{O}}_{\omega} | \Psi
angle = C_{\omega} rac{e^{-rac{eta \omega}{2}}}{1 - e^{-eta \omega}}$$
 $(C_{\omega} = \langle [\mathcal{O}_{\omega}, \mathcal{O}_{\omega}^{\dagger}]
angle.)$

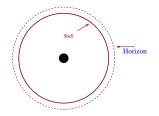
• Take
$$U_{\mathsf{MP}} = e^{i heta rac{\mathcal{O}_\omega \mathcal{O}_\omega^\dagger}{\mathcal{C}_\omega}}$$

$$\langle \Psi | U_{\mathsf{MP}}^{\dagger} \mathcal{O}_{\omega} \widetilde{\mathcal{O}}_{\omega}^{\dagger} U_{\mathsf{MP}} | \Psi
angle = e^{i heta} \mathcal{C}_{\omega} rac{e^{-rac{eta \omega}{2}}}{1 - e^{-eta \omega}}$$

• But
$$\delta E \propto \beta \theta^2 (\delta \omega)^2 \propto \beta \theta^2 / T^2$$
.

Appears to violate $\delta \langle A_{\alpha} \rangle \leq 2\sqrt{\beta \delta E} \sigma_{\alpha}$

Intuition: Firewalls near the Horizon

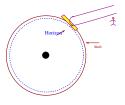


- For every black hole with an empty interior, consider a configuration with a ultra-relativistic shell close to the horizon.
- Binding energy with BH cancels rest+kinetic energy of shell ⇒ increase in AdM energy is small.
- Paradox suggests that $\delta \langle A_{\alpha} \rangle \leq 2\sqrt{\beta \delta E} \sigma_{\alpha}$ is violated unless typical states are firewalls.

Marolf-Polchinski, 2015

Significance of the Infalling Observer

• Paradox only inside the horizon: outside observer cannot distinguish shell from the Unruh effect.



- Technical version: operators outside the horizon given by HKLL construction; automatically obey $\delta \langle A_{\alpha} \rangle \leq 2\sqrt{\beta \delta E} \sigma_{\alpha}$.
- Interior ops state-dependent; inequality not guaranteed.

M-P paradox is a test of whether state-dependence leads to observable violations of standard rules of statistical mechanics.

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Resolution

- Start with an equilibrium state $|\Psi\rangle$.
- Create the excitation actively by modifying the boundary Hamiltonian

$$H_{\rm CFT} = H_{\rm CFT} + J(t)\mathcal{O}^J(t)$$

Then, as a result of a remarkable property of position space AdS correlators

$$\delta \langle \mathbf{A}_{lpha}
angle \leq \mathbf{2} \sqrt{eta \delta \mathbf{E}} \sigma_{lpha}$$

No observer can start with an equilibrium state; excite it and jump into the horion to observe the violation.

Resolution: Alternate language

We can prove a stronger result

If the unitary *U* and the observable A_{α} fit in the same causal patch, we can prove $\delta \langle A_{\alpha} \rangle \leq 2\sqrt{\beta \delta E} \sigma_{\alpha}$. Therefore violations are unobservable!

Sources and Causal Patches

Deform

$$H_{\rm CFT} = H_{\rm CFT} + J(t)\mathcal{O}^J(t)$$

Then, bulk observables modified to

$$\phi^{J}(t, r_{*}, \Omega) = \overline{\mathcal{T}} \{ e^{i \int_{\vartheta}^{t+r_{*}} J(x)\mathcal{O}(x)dx} \} \phi(t, r_{*}, \Omega) \mathcal{T} \{ e^{-i \int_{\vartheta}^{t+r_{*}} J(x)\mathcal{O}(x)dx} \}.$$

Only the part of the source in the causal past of the bulk point affects the field there.

Therefore the differences

$$\langle \phi^J(t_1, r_{*1}, \Omega_1) \dots \phi^J(t_n, r_{*n}, \Omega_n) \rangle - \langle \phi(t_1, r_{*1}, \Omega_1) \dots \phi(t_n, r_{*n}, \Omega_n) \rangle$$

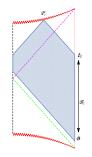
with pts in same causal patch obey constraints of statistical mechanics.

Causal Patches

$$U=e^{i\int_{\vartheta}^{t_{\mathcal{C}}}A(t)dt}.$$

A(t) is a simple boundary operator. With x_i in causal patch of t_c

 $\delta = \langle \Psi | U^{\dagger} \phi(x_1) \dots \phi(x_n) U | \Psi \rangle - \langle \Psi | \phi(x_1) \dots \phi(x_n) | \Psi \rangle$



Most general test of inequality possible within constraints of causality.

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Summary of Resolution

For this observable

$$\delta \leq 2\sqrt{\beta\delta E}\sigma.$$

as a result of a very non-trivial property of position-space AdS correlators.

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So, no observer or army of observers can detect a violation of the inequality by doing experiments that obey the constraints imposed by bulk causality.

Sketch of Proof

From the definition of mirror operators

$$\widetilde{\phi}_{r}(t,r_{*},\Omega)\mathsf{A}_{lpha}U|\Psi
angle=\mathsf{A}_{lpha}Ue^{rac{-eta H}{2}}\widehat{\phi}(t,r_{*},\Omega)e^{rac{eta H}{2}}|\Psi
angle.$$

where $\widehat{\phi}(t, r_*, \Omega)$ is an ordinary operator.

Main technical step is

$$[\widehat{\phi}(t_1, r_{*1}, \Omega_1), \mathcal{O}(t_2, \Omega_2)] = \mathbf{0}.$$

where $\mathcal{O}(t_2, \Omega_2)$ is a boundary operator and point 1 is behind the horizon but in the same causal patch as point 2.

Sketch of Proof

So

$$\langle \Psi | U A_{\alpha} \widetilde{\phi}_{r}(t, r_{*}, \Omega) U^{\dagger} | \Psi \rangle = \langle \Psi | U A_{\alpha} e^{-\frac{\beta H}{2}} \widehat{\phi}(t, r_{*}, \Omega) e^{\frac{\beta H}{2}} U^{\dagger} | \Psi \rangle + \mathcal{O}(\beta \delta E).$$

But then

$$\begin{split} \langle \Psi | U A_{\alpha} \widetilde{\phi}_{r}(t, r_{*}, \Omega) U^{\dagger} | \Psi \rangle &- \langle \Psi | A_{\alpha} \widetilde{\phi}_{r}(t, r_{*}, \Omega) | \Psi \rangle \\ &= \langle \Psi | U A_{\alpha} e^{-\frac{\beta H}{2}} \widehat{\phi}(t, r_{*}, \Omega) e^{\frac{\beta H}{2}} U^{\dagger} | \Psi \rangle - \langle \Psi | A_{\alpha} e^{-\frac{\beta H}{2}} \widehat{\phi}(t, r_{*}, \Omega) e^{\frac{\beta H}{2}} | \Psi \rangle \\ &+ O(\beta \delta E). \end{split}$$

Correlator on right is just an ordinary correlator; automatically obeys $\delta \leq 2\sqrt{\beta\delta E}\sigma$.

Therefore, the correlator of mirror operators in the same causal patch as U also obeys the constraints of statistical mechanics.

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Resolution of the Marolf-Polchinski Paradox: Summary

- In frequency space, the state-dependence of mirror operators is manifested as an anomalously large change in correlators under low-energy excitations.
- When we consider position space observables in the same causal patch as the excitation, these anomalous transformations cancel.
- So, the anomalous properties of the interior state-dependent operators are not visible in any physical experiment.

Outline



- 2 Holography and the BH Interior
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- The modern information paradox can be rephrased as a question about the existence of CFT operators dual to bulk fields.
- Can be resolved using a state-dependent map between boundary and bulk fields.
- Question is whether state-dependence is consistent.

Summary

 Marolf-Polchinski suggested that state-dependence contradicts a general rule of statistical mechanics.

$$\langle \Psi | U^{\dagger} A_{lpha} U | \Psi
angle - \langle \Psi | A_{lpha} | \Psi
angle \leq 2 \sqrt{eta \delta E} \sigma_{lpha}$$

"low-energy excitations have small effects."

- Here, we argued that this paradox is unobservable.
- If we consider boundary excitations *U* and local bulk observables A_{α} in the same causal patch as *U* then the inequality is obeyed.
- Non-trivial property of position-space AdS correlators.

Open Question

- This shows that it is possible to describe the BH interior holographically and consistently.
- But, why is this the "correct description"?
- Requires a dynamical understanding of why the bulk-observer measures the fields he does.
- Analogous to the Unruh-de Witt answer for why a particular observer uses a particular definition of "particle number"; here we want to explain why a specific CFT operator is the correct local bulk field φ.

Appendix

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Smooth Causal Patches

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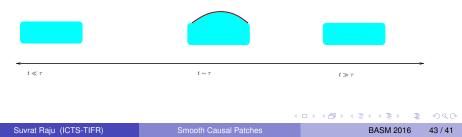
Autonomously Excited States

Consider

$$|\Psi^{\mathsf{ne}}
angle = U(au)|\Psi
angle,$$

as an autonomously excited state.

- On the boundary, think of |Ψ^{ne}⟩ as a fluid about to undergo a spontaneous excitation around τ.
- What does an infalling observer in $|\Psi^{ne}\rangle$ experience?

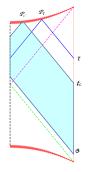


Paradox in Spontaneously Excited States

If an observer can prepare $U(\tau)|\Psi\rangle$ and compare with $|\Psi\rangle$

$$\delta = \langle \Psi | U(\tau)^{\dagger} \phi(\mathbf{x}_1) \dots \phi(\mathbf{x}_n) U(\tau) | \Psi \rangle - \langle \Psi | \phi(\mathbf{x}_1) \dots \phi(\mathbf{x}_n) | \Psi \rangle$$

where x_i are in a causal patch but τ is beyond the causal patch, then he would observe a violation of $\delta \leq 2\sqrt{\beta\delta E}\sigma$.



(Would also observe a violation of the second law of thermodynamics!).

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Smooth Causal Patches

Proposal of Causal Patch Complementarity Write

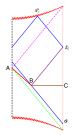
$$U(\tau) = U^{\mathcal{C}} \widehat{V}^{\mathcal{C}},$$

where

$$U^{\mathcal{C}} = e^{i\int_{artheta}^{t_{\mathcal{C}}} A_{\gamma}(t)dt}$$

and, $\forall A_{\alpha}(t)$ with $t < t_{\mathcal{C}}$,

 $\langle \Psi | U(\tau)^{\dagger} A_{\alpha}(t) U(\tau) | \Psi
angle = \langle \Psi | U^{\mathcal{C}} A_{\alpha}(t) U^{\mathcal{C}} | \Psi
angle$



A b

Proposal of Causal Patch Complementarity

Fields appropriate for causal patch corresponding to $t_{\mathcal{C}}$ satisfy

$$\langle \Psi | U^{\dagger} \phi_{\mathcal{C}}(x_1) \dots \phi_{\mathcal{C}}(x_n) U | \Psi \rangle = \langle \Psi | U^{\mathcal{C}^{\dagger}} \phi(x_1) \dots \phi(x_n) U^{\mathcal{C}} | \Psi \rangle,$$

Infalling observer is only sensitive to "part of the boundary excitation in the same patch as the observer"

Consequences: Causal Patch Complementarity

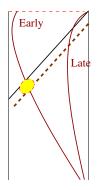
Ensures that

$$\delta = \langle \Psi | U^{\dagger} \phi_{\mathcal{C}}(x_1) \dots \phi_{\mathcal{C}}(x_n) U | \Psi \rangle - \langle \Psi | \phi_{\mathcal{C}}(x_1) \dots \phi_{\mathcal{C}}(x_n) | \Psi \rangle$$

satisfies $\delta \leq 2\sqrt{\beta \delta E}\sigma$ for all *U* provided $x_1, \ldots x_n$ are in a single causal patch.

- Correlators of $\phi_{\mathcal{C}}$ and the HKLL ϕ agree in all equilibrium states and in all equilibrium states excited with a source.
- For spontaneously excited states, this modifies HKLL outside the horizon.

Example: Stanford-Shenker State



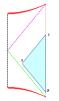
Observer outside infers that the early observer had a trans-Planckian collision. But early observer perceives a smooth geometry.

['t Hooft, Susskind, Thorlacius, Uglum, Kiem, Verlinde², 85–95]

Teleological Property of HKLL Construction

 In empty AdS, HKLL construction is causal. But, in the presence of a black hole, HKLL construction is teleological

$$\phi_{\mathsf{HKLL}}(x) = \int O(t) K(t, x) dt$$



- Fields φ_{HKLL} and φ_C differ in their response to future excitations in spontaneously excited states.
- How does observer at x "know" whether he should use $\phi_{\rm HKLL}$ or $\phi_{\rm caus}$?