

Collective motion – a universal phenomenon in nature

BI-SW: http://www.youtube.com/watch?v=V71hz9wNsgs

SHIBUYA:

http://www.youtube.com/watch?v=QXtOdSgf6Ic&feature=BFa&list=PL881859D3DB1AFA9F&lf=results_video

Ped: http://arxiv.org/abs/cond-mat/9805074

FI: http://www.youtube.com/watch?v=clgHEhziUxU

Computer Simulations of Pedestrian Dynamics and Trail Formation D. Helbing, P. Molnar, F. Schweitzer

(Submitted on 6 May 1998 (v1), last revised 7 May 1998 (this version, v2))

Abstract: A simulation model for the **dynamic behaviour of pedestrian crowds** is mathematically formulated in terms of a **social force model**, that means, pedestrians behave in a way as if they would be subject to an acceleration force and <u>to repulsive forces</u> describing the reaction to borders and other pedestrians. The computational simulations presented yield many realistic results that can be compared with video films of pedestrian crowds. Especially, they show the self-organization of <u>collective behavioural patterns</u>. By assuming that pedestrians tend to choose routes that are frequently taken the above model can be extended to an active walker model of trail formation. The topological structure of the evolving trail network will depend on the disadvantage of building new trails and the durability of existing trails. Computer simulations of trail formation indicate to be a valuable tool for designing systems of ways which satisfy the needs of pedestrians best. An example is given for a non-directed trail network.

For related work see this http URL Statistical Mechanics (cond-mat.stat-mech); Pattern Formation and Solitons (nlin.PS) Pages 229-234 in: Evolution of Natural Structures (Sonderforschungsbereich Journal reference: 230, Stuttgart, 1994) Cite as:arXiv:cond-mat/9805074v2 [cond-mat.stat-mech] 'SOCIAL' repulsive force:

$$\vec{f}_{\alpha\beta}(\vec{r}_{\alpha} - \vec{r}_{\beta}) = -\nabla_{\vec{r}_{\alpha}} V_{\beta}[b(\vec{r}_{\alpha} - \vec{r}_{\beta})].$$
(5)

The sum over the repulsive potentials V_{β} defines the *interaction potential* which influences the behaviour of each pedestrian:

$$V_{\rm int}(\vec{r},t) := \sum_{\beta} V_{\beta} \{ b[\vec{r} - \vec{r}_{\beta}(t)] \} \,. \tag{6}$$



b = *distance to neighbor*

Repulsive 'social' force:

$$\vec{f}_{\alpha\beta}(\vec{r}_{\alpha} - \vec{r}_{\beta}) = -\nabla_{\vec{r}_{\alpha}} V_{\beta}[b(\vec{r}_{\alpha} - \vec{r}_{\beta})].$$
(5)

The sum over the repulsive potentials V_{β} defines the *interaction potential* which influences the behaviour of each pedestrian:



How do we recognize COLLECTIVE MODES in Nuclei ??

(Here: Nuclear Eigenstates, not, e.g., collective flow in collisions)

Recipe: Quantization of a classical collective system and search for the respective characteristic patterns in the nuclear spectral response

for example: rotational spectra

Classical rigid rotation: $\mathbf{E}_{\mathbf{L}} = \mathbf{L}^2 / 2\Theta$ (continuous) Quantized system: $\mathbf{E}_{\mathbf{I}} = \hbar^2 \mathbf{I} (\mathbf{I} + \mathbf{1}) / 2\Theta$ (discrete eigenstates) $\mathbf{E}_{\mathbf{V}} = \hbar^2 \mathbf{2} (\mathbf{I} - \mathbf{1}) / \Theta$ or $\Delta \mathbf{E}_{\mathbf{V}} = \text{const.}$



• T. Lauritsen et al., Phys. Rev. Lett. 89 (2002) 282501

Vibrational nuclear states:

Surface (β -, γ - , octupole) vibrations

Density oscillations ('giant resonances ')

However:

Not much exp. information on two- (higher-) phonon states available

Quantized harmonic oscillator: $E_n = n E_o$ or $\Delta E = E_o$ ('phonon'!)

Rotatinal nuclear spectra and surface vibrations discussed already, see M. Carpenter, P. Regan, P. Reiter ...

this lecture restricted to <u>(specific aspects of)</u> giant resonances !



Lecture I: <u>Giant Resonances (GR) – Introduction, Basics</u>

Phenomenology, properties, main facts Classification schemes Theoretical description – concepts What we (might) learn form GR's Tools (reactions) for GR excitation History – milestones

Lecture II-IV: <u>Giant Resonances – selected topics</u>

Spin-flip resonances : Gamow-Teller puzzle Isoscalar GMR and GDR : nuclear matter (in)compressibility

Multi-phonon Giant Resonances

Multipole Strength in (neutron-rich) 'exotic' nuclei Pygmy resonances, soft modes Symmetry energy Neutronstar – neutronskin and pygmy

Outlook

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Giant Resonances

Fundamental High-Frequency Modes of Nuclear Excitation

> M. N. HARAKEH and A. van der WOUDE



OXFORD SCIENCE PUBLICATIONS

Nuclear response to external field , at moderate energy (ω) and momentum (q) transfer



Photoabsorption (GDR)



Berman and Fulz, Rev. Mod. Phys. 47 (1975) 47



Inelastic α scattering (GMR + GQR)



FIG. 2: (a) Horizontal-position spectrum of the $^{112}Sn(\alpha, \alpha')$ reaction at 0°. The hatched region is background events. (b) Background-free spectrum

Courtesy of U. Garg



Classification of giant resonances (macroscopic view)

Four fluids in the nucleus, characterized by **spin** and **isospin**

neutron
$$(T_z = 1/2)$$

proton $(T_z = -1/2)$

Spin up (
$$\sigma_z = \frac{1}{2}$$
) - down ($\sigma_z = -\frac{1}{2}$)

<u>Accordingly, 4 basic modes :</u>

Electric isoscalar Magnetic isoscalar Magnetic isoscalar Magnetic isovector Magnetic Ma

rich pattern of GR's :



isovector

isovector

Nomenclature:

- L (λ) multipolarity S (σ) – spin T (τ) – isospin Δ S, Δ T – spin-, isospin transfer
- $E\lambda$ electric transition
- $M\lambda$ magnetic transition
- IS (is) isoscalar (p-n in phase)IV (iv) – isovector (p-n out of phase)

GMR,GDR,GQR... – giant monopole resonance

e.g. *ivSGDR* = *isovector spinflip giant dipole resonance*

Microscopic understanding of the GR

GR created by the action of a <u>one-body transition operator</u> onto the ground state of a nucleus:

$$|\Psi_{GR}^{\lambda,\sigma,\tau}\rangle = O^{\lambda,\sigma,\tau} |\Psi_{g.s.}\rangle$$

The transition operator contains a multipole-type operator

(e.g. $\Sigma_i r_i^{\lambda} Y_m^{\lambda}$)

and/or may involve appropriate spin and/or isospin operator



How are such operators realized in nuclear reaction experiments ?

i.e., how do we relate nuclear reaction cross section ↔ strength distribution ??

consider a (naive) shell model (harmonic osc.):



Excited states from p-h excitations within / across major shells N

remember: Parity alternates with N, thus, for multipolarity, e.g., $\Delta L = 1 \implies \Delta N = 1$ $\Delta L = 2 \implies \Delta N = 0 \text{ or } 2 \text{ etc...}$ however, exp. value: 13 MeV !!

GR's are understood as a

<u>Coherent superposition</u> of <u>many</u> 1p-1h excitations across major shells. <u>Strongly shifted</u> in energy by <u>residual interactions.</u>

Low-lying E2: rotation, surface vibration

Theoretical description of giant resonances – microscopic approach

Keywords:

Level of sophistication



QRPA quasi-particle RPA

S-,E-,C-RPA second-, extended, continuum RPA

Theoretical description of giant resonances – microscopic approach

Keywords:



S-,E-,C-RPA second-, extended, continuum RPA Eigenstates of single-particle Hamiltonian ($\mathbf{H} = \Sigma_i \mathbf{h}_i$) with energy : ε_i thus, the 1p-1h states | \mathbf{ni}^{-1} > with excitation energies : $\varepsilon_n - \varepsilon_i$

The 1p-1h states | ni^{-1} > form a complete orthonormal set, thus

with a (simplified) residual interaction V ($\mathbf{H} = \Sigma_i \mathbf{h}_i + \mathbf{V}$), obtain the 'perturbed' particle-hole energies:

$$\mathbf{E} = \mathbf{\varepsilon}_{\mathbf{n}} - \mathbf{\varepsilon}_{\mathbf{i}} + \Sigma_{\mathbf{mj}} < \mathbf{ni}^{-1} | \mathbf{V} | \mathbf{mj}^{-1} > c_{\mathbf{mj}} / c_{\mathbf{ni}}$$

(notice: ground state remains untouched)



Further (over-) simplifying by adopting all e_k to be equal, apparently <u>only one state is shifted up (down)</u> maximally and

most importantly: <u>it receives 100% of the transition probability ('the winner takes it all' -ABBA)</u> (i.e. the sum of the transition probability between the unperturbed states) ==> the collective state !



Theoretical description of giant resonances – microscopic approach

Keywords:





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1ph dipole strength gathers in narrow energy band

looks "resonance like" but is composed of several different 1ph states

RPA dipole strength collected in resonance peak (and somewhat fragmented to 1ph ...)

Watch green curve only !

dipole strength [arb.units]

Theoretical description of giant resonances – microscopic approach



Theoretical description of giant resonances – microscopic approach

Keywords:



WIDTH of GDR in heavy nuclei



closed shell nuclei: $\Gamma \sim 4-5$ MeV giant resonances – WIDTH and Damping



GDR vibration period T_s : $E_{GDR} = hv = h / T_s$ Heisenberg - lifetime τ : $\Gamma_{GDR} = h / 2\pi \tau$

$$\tau^{208} Pb \implies \tau = 0.5 T_s$$

strongly damped motion !

giant resonances - WIDTH - damping

Apparent GR width due to -

- fragmention of strength ('Landau damping')
- particle decay

GR as a 1p-1h state should decay predominantly into particular hole states in the (A-1) nucleus

However: in heavier nuclei mainly statistical decay observed !

==> coupling to 2p-2h.....



'spreading width '

Damping / Width of giant resonances :

Strength fragmentation (Landau damping)Deformation splitting (open-shell nuclei)Direct particle decay (lighter nuclei): Γ^{\uparrow} escape widthStatistical (compound nucleus) decay: Γ^{\downarrow} spreading width

What do we (hope to) learn from GR's ?

unperturbed particle-hole states \leftrightarrow mean field collective state \leftrightarrow interaction between ph states

→ learn about <u>effective in-medium interactions</u>

(can be differentiated .. ΔL , ΔT , ΔS)

GR's are density vibrations → learn about nuclear bulk properties :
 Finite matter (i.e. nuclei): radii, deformations, skins, clusters ...(GEOMETRY)
 Infinite matter: symmetry energy, compressibility etc.
 restricted, however, to around saturation density and (near) zero temperature

reaction and decay rates (astrophysical interests)

test of effective forces

GDR strength distribution for several Skyrme forces

180 SV-bas ²⁰⁸Pb dipole strength [arb.units] 160 SLy6 Skl3 GDR centroid energy 140 exp 120 **20** GDR in ²⁰⁸Pb pygmy 100 / 18 80 ŝν SIV SII 60 $E_1 \ ({ m MeV})$ 16 Ο 40 Ŝka SIII 14 exp. SKM 20 SKI2 SKM* SVL SGII SKI5 0 12 14 8 10 12 excitation energy [MeV] 10 2 8 10 0

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Collective resonance excitations (example ²⁰⁸*Pb*)



Systematic variation of Skyrme forces





Sum rules may be derived from algebraic relations between transition operators and (powers of) the Hamiltonian

Energy-weighted sum rules are model independent to the extent that

- nn interactions do not depend on the particle momenta (electric isoscalar transitions)
- and charge-exchange components are absent (electric isovector transitions)



some Sum rules

GDR:

TRK : $\int \sigma(E) dE \sim NZ/A$ $\int \sigma(E)/E dE \sim NZ/A < r^2 >_p$ Migdal: $\int \sigma(E)/E^2 dE \sim R^2 A / E_{sym}$

Cluster: TRK = TRK₁ + TRK₂ + TRK₁₂

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other sum rules:
Ikeda: S_{(GT)} - S_{+}(GT) = 3(N-Z)
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S_(IVSDR) - S₊(IVSDR) ~ N
$$\langle r^2 \rangle_n - Z \langle r^2 \rangle_p$$

etc..

 $\sigma(E) ~~ E \cdot dB(E1, 0 \rightarrow E) / dE$

History

- Milestones

1937 - Bothe and Gentner



300 keV proton 'beam'

- \rightarrow Li(p, γ)
- \rightarrow γ -radiation onto sample S, e.g. ⁶³Cu
- \rightarrow radioactivity from (γ ,n) measured
- " vielleicht spielen hier Resonanz -
- verhältnisse eine entschiedene Rolle "

History

- Milestones

1937 - Bothe and Gentner

1947 – Balwin and Klaiber GDR in γ-absorption spectra

19717" - isGQR Pitthan and Walcher

1977 – isoscalar monopole

- 1980... GR in hot nuclei
- 1985... M1 scissor mode
- 1985... multiphonon GDR, GQR
- ~ 2000 dipole strength in unstable nuclei