

# Numerical computation of Schwarzschild Quasi Normal Modes

1. Integrate the Regge-Wheeler equation

$$\frac{d^2 R}{dr_\star^2} + [\omega^2 - V(r)]R = 0 \quad (1)$$

where  $r_\star = r + 2M \log[\frac{r}{2M} - 1]$  is the tortoise coordinate, and the potential  $V(r)$  is given by:

$$V(r) = \left(1 - \frac{2M}{r}\right) \left(\frac{\ell(\ell+1)}{r^2} - \frac{6M}{r^3}\right). \quad (2)$$

using a direct integration method. Compute the real and imaginary part of the QNMs for different values of the multipole number  $\ell \geq 2$ . Find the QNMs by also employing a second numerical method, like the continued fraction method, i.e. by solving the following equation

$$0 = \beta_0 - \frac{\alpha_0 \gamma_1}{\beta_1 -} \frac{\alpha_1 \gamma_2}{\beta_2 -} \dots, \quad (3)$$

where the coefficients  $(\alpha_n, \beta_n, \gamma_n)$  read:

$$\alpha_n = n(2 - 4iM\omega) - 4iM\omega + n^2 + 1, \quad (4a)$$

$$\beta_n = -l(l+1) + 32M^2\omega^2 - n(2 - 16iM\omega) + 8iM\omega - 2n^2 + 3, \quad (4b)$$

$$\gamma_n = -16M^2\omega^2 - 8iMn\omega + n^2 - 4. \quad (4c)$$