Quantum-to-classical transition of the primordial perturbations



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We've been "doing physics for many thousands of years" and only discovered quantum 100 years ago

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path of discovery	A. Albrecht @ ICTS Sep 2, 2020	year 13

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Cosmic inflation is both OVER and UNDER rated as a solution to these problems

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"Many worlds" or "Everett"

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I) Some key elements of quantum physics

- Quantum coherence
- Quantum entanglement
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 - •Discuss the roles of the above phenomena
- III) Some examples

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Quantum coherence

 The state of quantum system is described by a "wavefunction" which assigned a quantum amplitude to each element of a basis which spans the full Hilbert space:

$$|\psi\rangle = \sum_{i} \alpha_{i} |i\rangle$$

• If $|\Psi_1\rangle$ and $|\Psi_2\rangle$ are possible states of the system, then so is

$$\left|\psi\right\rangle = a_{1}\left|\psi_{1}\right\rangle + a_{2}\left|\psi_{2}\right\rangle$$



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- The "coherence" of a quantum state only means something in relation to what measurement you are going to make.
- If $\psi(x)$ corresponds to classical wavepacket, and you only ever expect to measure the particle in a wavepacket basis, then the particle simply moves from here to there, and there is no particular meaning to "quantum coherence".

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Quantum coherence

 The double slit experiment is a striking example of quantum coherence because there is a large discrepancy between the diffracted quantum state of the particle and the localized basis in which the screen makes the measurement.



https://www.nature.com/articles/35089156/figures/1

Quantum coherence

- The double slit experiment is a striking example of quantum coherence because there is a large discrepancy between the diffracted quantum state of the particle and the localized basis in which the screen makes the measurement.
- If instead of a screen making localized measurements, you had an apparatus that could measure the particle in a "double slit state", the narrative would be much more mundane, and the concept of "quantum coherence" would not be needed.



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As with the topic of coherence, the degree things "look quantum" depends on the questions you ask

Expectation value of observable in s:

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Quantum entanglement

Interactions can create entanglement $|\psi\rangle_{tot} =$

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 $\left|\psi\right\rangle_{tot} = \frac{1}{\sqrt{2}} \left|\psi_{1}\right\rangle_{b} \left|\psi_{1}\right\rangle_{\gamma} + \frac{1}{\sqrt{2}} \left|\psi_{2}\right\rangle_{b} \left|\psi_{2}\right\rangle_{\gamma}$ Entangled state after collision



 $|\psi_1\rangle_{\gamma} \neq \sqrt{2}$

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Entangled state after collision

In the example of the photon striking the ball, the Schrödinger cat initial state evolved into a <u>classical</u> <u>mixture of localized packets</u>, thanks to the *locality of the interactions and the locality of the photon state*

> (need $_{b}\langle_{i}\psi|\psi_{j}\rangle_{b}\approx\delta_{ij}$ which is natural for ball separation much greater than that photo coherence length)

Quantum entanglement

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One could imagine probing the "Schrödinger cat" nature of the initial ball state with a coherent plane wave, suitably measure after the collision.

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Which states are "prepared"

A. Albrecht @ ICTS Sep

In what basis you make measurements

Zurek 2009 https://arxiv.org/abs/0903.5082



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Quantum state of the CMB: a two-mode

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Inflation and squeezed quantum states ALBRECHT, FERREIRA, JOYCE, AND PROKOPEC



FIG. 1. Phase space trajectories for a classical upsidedown harmonic oscillator. The presence of one growing and one decaying solution produces a "squeezing" effect even at the classical level. The circular region shown evolves with time into the squeezed shape above it.

https://arxiv.org/abs/astro-ph/9303001

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Quantum state of the CMB: a two-mode supered s

The wave functional of curvature perturbations takes the form $\Psi[\zeta]=\prod_{{\bf k}\in {\bf R}^{3+}}\Psi(\zeta_{{\bf k}},\zeta_{-{\bf k}})$

 $\Psi_R(\zeta_{\mathbf{k}}, \zeta_{-\mathbf{k}}) = \frac{1}{\sqrt{\pi}} e^{-(\zeta_{\mathbf{k}} - \zeta_{-\mathbf{k}})^2 / (4R^2)} e^{-R^2(\zeta_{\mathbf{k}} + \zeta_{-\mathbf{k}})^2 / 4}$

Two mode vacuum squeezed state

classical distribution

Evolution is essentially quadratic

See J. Martin talk

Cosmological evolution
 ←→ "squeezing"→
 enhanced classicality

As long as you "agree to" only ever measure the system in classical wavepackets to sep 2, 2020

FIG. 1. Phase space trajectories for a classical upsidedown harmonic oscillator. The presence of one growing and one decaying solution produces a "squeezing" effect even at the classical level. The circular region shown evolves with time into the squeezed shape above it

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https://arxiv.org/abs/astro-ph/9303

See D. Wands talk

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And, there are a relentless array of *local* interactions that will cause decoherence:

- Particles produced during reheating
- The CMB
- Cosmic dust
- Starlight



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"One photon will do it"



In fact, the actual universe is much like the "pendulum in a dark room" thought experiment:

- An early phase that is technically quantum, but can be modeled as a classical ensemble
- Followed by a phase "quantum Darwinism" class decoherence



In fact, the actual universe is much like the "pendulum in a dark" room" thought experiment:

- An early phase that is technically quantum, but can be modeled as a classical ensemble
- Followed by a phase "quantum Darwinism" class decoherence
- Observations made by sampling a very small part of the environment





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- However, classically equivalent early evolution, plus a vast array of subsequent decoherence processes seems to be working against you.

OUTLINE

I) Some key elements of quantum physics

- Quantum coherence
- Quantum entanglement
- Quantum Darwinism
- Wavefunction as a classically evolving distribution

II) The actual Universe

• Discuss the roles of the above phenomena

III) Some examples

IV) Conclusions

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- Maldacena 2015 <u>https://arxiv.org/abs/1508.01082</u> (Constructs an *extremely exotic* inflation model specially designed to give observable quantum effects.)
- Quantum Entanglement in Cosmology workshop (IPMU May 2019) <u>https://indico.ipmu.jp/event/300/overview</u>
- J. Martin talk here and at IPMU workshop
- Kanno et al Noise and decoherence induced by gravitons <u>https://arxiv.org/abs/2007.09838</u>
- Parikh et al https://arxiv.org/abs/2005.07211 (with story)

How to characterize nonclassicality?

Qunatum optics: Count photon numbers in a given state



Distribution of photons in classical theory is always super-Poissonian and F > 1

contraposition

Sub-Poissonian distribution or F < 1 must be a signature of nonclassicality

S. Kanno @ IPMU 2019 https://indico.ipmu.jp/event/300/overview

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Kanno: Let's apply this test to primordial gravitons

4

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AA: How do you count individual gravitons?

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Inflation and squeezed quantum states

ALBRECHT, FERREIRA, JOYCE, AND PROKOPEC

https://arxiv.org/abs/astro-ph/9303001

fies the WKB criterion for classicality. This is equivalent to the WKB classicality at late times in an inverted harmonic oscillator studied by Guth and Pi in Ref. [11]. The point of this section is to explain that the apparently very quantum mechanical squeezed state is in fact classical in the sense with which cosmologists are familiar. That the truly quantum mechanical features of these states which are probed, for example, in quantum optics might have cosmological implications is a fascinating claim but one which has no substance at present. In Sec. VIII we sum-

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Sub-Poissonian distribution or F < 1 must be a signature

The "relativity of quantum coherence"

The properties of the noise – its amplitude, power spectrum, etc – are calculable and depend on the state. We find that for the vacuum state or a coherent state, the fluctuations in the arm length are extremely small and almost certainly undetectable, as foreseen by Dyson. But for thermal states – such as from cosmology or evaporating black holes – the noise is significantly enhanced. Most favorably, if the gravitational field is in a squeezed state, as predicted by some inflationary models, the fluctuations in the arm length can be enhanced by an exponential of the squeezing parameter, and are potentially detectable.

$(\Delta n)^2$

Detection of this fundamental noise would constitute direct evidence for the quantization of gravity and the existence of gravitons.

contraposition

Sub-Poissonian distribution or F < 1 must be a signature of nonclassicality

Parikh et al: Don't need to measure individual gravitons, just look at noise in (future) gravity wave detectors. <u>https://arxiv.org/abs/2005.07211</u>

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My goal is to equip you for your own path of discovery The universe works hard to hide its quantum nature, but also we are not so good at seeking it out

I'll give examples of new ideas that were just published this year