## Hadron-Hadron Interactions from Lattice QCD

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(Univ. of Tsukuba)
for HAL QCD Collaboration

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## Hadron-Hadron Interactions from Lattice QCD

## Takumi Doi <br> (Univ. of Tsukuba) for HAL QCD Collaboration

- Motivation
- Formulation for NN potential in Lattice QCD
- Extension to YN, YY potentials
- Recent progress on Three Nucleon Force (TNF)
- Summary and Outlook


## Motivation



## Understand the various phenomena from fundamental theory

- Nuclei
- Neutron star
- SuperNova

Nuclear Force is the key concept which bridges (effective) DOF in different hierarchy

## Phenomenological NN potential

( $\sim 40$ parameters to fit 5000 phase shift data)


## Nuclear Force from Experiments

- Potential is constructed so as to reproduce the NN phase shift (or, S-matrix)





## Various applications: few/ many-body

 system of nuclei, EOS of Nuclear matter..
## Nuclear Force from QCD

- First principle calculation of QCD



## Y. Nambu, "Quarks : Frontiers in Elementary Particle Physics", World Scientific (1985)

"Even now, it is impossible to completely describe nuclear forces beginning with a fundamental equation. But since we know that nucleons themselves are not elementary, this is like asking if one can exactly deduce the characteristics of a very complex molecule starting from Schroedinger equation, a practically impossible task."

## Lattice QCD as 1st principle calc



- well-defined statistical system (finite a and L)
- gauge invarinat
- fully non-perturbative


Monte-Calro simulations

Quenched QCD : neglects creation-anihilation of quark-anitiquak pair Full QCD : includes creation-anihilation of quark-anitiquak pair

## Status of Lattice QCD

- hadron masses (g.s.)

$\mathrm{Nf}=2+1$ clover fermion
$\mathrm{V}=32^{3} \mathrm{X} 64, \mathrm{~L}=2.9 \mathrm{fm}$,
$\mathrm{a}=0.09 \mathrm{fm}(1 / \mathrm{a}=2.18 \mathrm{GeV})$
$\mathrm{m} \pi(\mathrm{min})=156 \mathrm{MeV}$

PACS-CS Collab., PRD79(2009)034503

> 3\% accuracy at present

## $\perp$ Lattice QCD as 1st principle calc



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## Nuclear Force from Lattice QCD [HAL QCD strategy]

- Potential is constructed so as to reproduce the NN phase shift (or, S-matrix)
- Nambu-Bethe-Salpeter(NBS) wave function

$$
\psi(\vec{r})=\langle 0| N(\vec{x}+\vec{r}, t) N(\vec{x}, t)|2 N\rangle
$$

- Key concept: asymptotic region $\longleftrightarrow$ phase shift


$$
\left(\nabla^{2}+k_{\delta}^{2}\right) \psi(\vec{r})=0, \quad r>R
$$

- Define the potential at interaction region

$$
\left(\nabla^{2}+k_{\delta}^{2}\right) \psi(\vec{r})=\int d r^{\prime} U\left(\vec{r}, \overrightarrow{r^{\prime}}\right) \psi\left(\overrightarrow{r^{\prime}}\right), \quad r<R
$$

- Non-local, but E-independent potential

Luscher, NPB354(1991)531
C.-J.Lin et al., NPB619(2001)467 CP-PACS Coll., PRD71(2005)094504


Aoki-Hatsuda-Ishii PTP123(2010)89

## Effective Schrodinger equation with E-independent potential

## $K(\vec{x} ; E) \equiv\left(\vec{\nabla}^{2}+k^{2}\right) \psi(\vec{x} ; E) \quad$ [START] local but E-dep pot. ( $\mathbf{L}^{3} \mathrm{xL}^{3}$ dof)

(1) We assume $\psi(x ; E)$ for different E is linearly independent with each other.
(2) $\psi(x$; E) has a "left inverse" as an integration operator as

$$
E \equiv 2 \sqrt{m_{N}^{2}+k^{2}}
$$

$$
\int d^{3} x \tilde{\psi}\left(\vec{x} ; E^{\prime}\right) \psi(\vec{x} ; E)=2 \pi \delta\left(E-E^{\prime}\right)
$$

(3) $K(x ; E)$ can be factorized as

$$
\equiv m_{N} U(\vec{x}, \vec{y})
$$

$$
\begin{aligned}
K(\vec{x} ; E) & =\int \frac{d E^{\prime}}{2 \pi} K\left(\vec{x} ; E^{\prime}\right) \times \int d^{3} 3 \bar{\psi}\left(\vec{y} ; E^{\prime}\right) \psi(\vec{y} ; E) \\
& =\int d^{3} y\left\{\sum_{\alpha} \int \frac{d E^{\prime}}{2 \pi} K\left(\vec{x} ; E^{\prime}\right) \tilde{\psi}\left(\vec{y} ; E^{\prime}\right)\right\} \psi(\vec{y} ; E)
\end{aligned}
$$

(4) We are left with an effective Schrodinger equation with an E-independent potental U .

$$
\left(\vec{\nabla}^{2}+k^{2}\right) \psi(\vec{x} ; E)=m_{N} \int d^{3} y U(\vec{x}, \vec{y}) \psi(\vec{y} ; E)
$$

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Luscher, NPB354(1991)531

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$$

- Non-local, but E-independent potential
- Velocity expansion

Okubo-Marshak(1958)

$$
U\left(\vec{r}, \overrightarrow{r^{\prime}}\right)=V_{c}(r)+S_{12} V_{\text {LO }}(r)+\vec{L} \cdot \overrightarrow{S_{N L O}} V_{L S}(r)+\underset{\text { NNLO }}{\mathcal{O}\left(\nabla^{2}\right)}
$$



## Nuclear Potential (from Lat QCD)



## Tensor Potential from Lat QCD

- Tensor operator

$$
S_{12}=3\left(\vec{\sigma}_{1} \cdot \vec{r}\right)\left(\vec{\sigma}_{2} \cdot \vec{r}\right) / r^{2}-\left(\vec{\sigma}_{1} \cdot \vec{\sigma}_{2}\right)
$$

- Essential to understand the nuclei
- Responsible for deuteron binding
- Hyper nuclei binding ( $\Lambda \mathrm{N}-\mathrm{\Sigma} \mathrm{~N}$ )

- Coupled channel study in ${ }^{3} \mathrm{~S}_{1}-{ }^{-3} \mathrm{D}_{1}$ channel

$$
\begin{aligned}
& \left(H_{0}+V_{C}+V_{T} S_{12}\right) \psi=E \psi \\
& \psi=\psi_{S}+\psi_{D} \\
& \begin{array}{r}
\psi_{S}(\vec{r})=P \psi(\vec{r})=\frac{1}{24} \sum_{g \in O} \psi\left(g^{-1} \vec{r}\right) \\
\psi_{D}(\vec{r})=Q \psi(\vec{r})=(1-P) \psi(\vec{r}) \\
\\
\square \begin{array}{l}
\mathrm{P} \quad: \text { projection to L=0 } \\
\mathrm{Q}=(1-\mathrm{P}): \text { projection to } \mathrm{L}=2
\end{array} \\
\hline
\end{array} \\
& \begin{array}{l}
P\left(H_{0}+\overline{V_{C}}+V_{T} \beta_{12}\right) \psi=E P \psi \\
Q\left(H_{0}+\overline{V_{C}}+V_{T} \beta_{12}\right) \psi=E Q \psi
\end{array}
\end{aligned}
$$

## Tensor Potential from Lat QCD

- Coupled channel study in ${ }^{3} \mathrm{~S}_{1}-{ }^{3} \mathrm{D}_{1}$ channel

Wave function


Potentials


Aoki-Hatsuda-Ishii, PTP 123 (2010) 89

## Results from Full QCD Lattice




- Larger Repulsive Core
- Larger Tensor force
(N.B. 1/a is also different)


## Quark mass dependence



Lighter mass corresponds to...

- Longer interaction range
- Larger Repulsive Core
- Stronger Tensor Force
- (stronger attraction in Center Force)



## Phase shift from potential



## Phase shift from potential



## Scattering Length

## Scatt. Length

(w.f. $\rightarrow \mathrm{k}^{2} \rightarrow$ Luscher's formula)


## Attractive Scatt. Length

Further quantitative refinement in progress: precise determination of E and long-range w.f. behavior is essential

Yet, much smaller compared to the experimental values

$$
\begin{aligned}
& a_{0}\left({ }^{1} S_{0}\right) \sim 20 \mathrm{fm} \\
& a_{0}\left({ }^{3} S_{1}\right) \sim-5 \mathrm{fm}
\end{aligned}
$$

## Scattering Length

Y.Kuramashi, PTPS122(1996)153


OBEP + lattice hadron masses


EFT

## Scattering Length



## Challenge in Next-Gen Simulation



## Challenge in Next-Gen Simulation



Temporal info (Luscher's formula)

$$
\delta E=-\frac{2 \pi a_{0}}{\mu L^{3}}\left(1+c_{1} \frac{a_{0}}{L}+c_{2} \frac{a_{0}^{2}}{L^{2}}\right)+\mathcal{O}\left(L^{-6}\right)
$$

Spacial info (potential/phase shift)

$$
V(r \rightarrow \infty) \sim e^{-m \pi r} / r \quad \text { Localized! }
$$

Challenge in Next-Gen Talc
We are here

$$
\begin{array}{r}
a_{0}\left({ }^{3} S_{1}\right) \sim-5 \mathrm{fm} \\
\text { B.E. }\left({ }^{2} H\right) \sim 2 \mathrm{MeV} \\
\hline
\end{array}
$$



## Challenge in Next-Gen Simulation



## J apan＇s next gen computer

－K computer at Kobe，Japan
－10PFlops（2012）

```
次世代スーパーコンピュータ施設 完成イメージ図
```



SPARC64 ${ }^{\text {TM }}$ VIIIfx （C）fujitsu Limited
$K($ Kei $)=10^{16}=10$ Peta

## Frequently Asked Questions

## [Q1] Is potential observable ? Just give me phase shifts !

- Potential $\mathrm{U}(\mathrm{x}, \mathrm{y})$ is NOT observable, and is NOT unique. However, combination of ( $\Phi(x), U(x, y)$ ) gives observable, which is unique.
- Same situation for QM( $\Phi, \mathrm{U})$, QFT( $\Phi($ (asym $)$, vertices), EFT(eff. dof, LECs) ... Yet, we use "wave function $\Phi(x)$ " in QM, etc.
- We study potential (in addition to phase shift), because:
- Convenient framework/concept to understand the physics
- Potential is essential to study many-body systems
- c.f. QM: Matrix mechanics vs. Wave mechanics

$$
\begin{aligned}
\text { Lat } & \rightarrow \delta_{E} \rightarrow U(x, y)
\end{aligned} \rightarrow \text { many-body }
$$

- It is very difficult to calculate phase shift at high energy
- Lattice $\rightarrow$ only ground state + a few excited energy states
- Potential (hopefully) contains "useful" off-shell information
- Sys. error by velocity expansion can be checked order by orde



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$$

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## Frequently Asked Questions

## [Q2] Isrit Potential dependent on the sink operator?

- Yes, the potential is dependent on the choice of the sink operator, since Potential $U(x, y)$ is NOT observable. ( $\rightarrow$ go back to the 1st Q\&A)
- One can choose any sink opeartor, and the physical observables (at least phase shift) calculated from that potential remain same
- We choose local operator as convenient choice for the reduction formula
- Good operator $\longleftrightarrow$ small non-locality in potential
- We check the velocity expansion convergence a posteriori


## [Q3] How good is velocity expansion of potential ?

- We explicitly checked the validity of expansion using two methods:
- By Energy dependence of LO potential $\mathrm{V}_{\mathrm{C}}(\mathrm{r}) \quad$ K.Murano (HAL Collab.)
- By $L^{2}$ dependence of $V_{C}(r)$
@ Lattice2009, Lattice2010


## "Energy dependence" of LO $V_{c}(r)$ in velocity expansion



## $L^{2}$ dependence of $\mathrm{V}_{\mathrm{c}}(\mathrm{r})$ in $\mathrm{S}=0$



## Towards the prediction from Lattice QCD

- "Realistic" NN potentials have achieved quite a good precision
- $\sim 40$ parameters for $\sim 5000$ (high prec) phase shifts, $\chi^{2 / d o f ~} \sim 1$
- Hyperon-Nucleon(YN), Hyperon-Hyperon(YY) potentials
- Large uncertainties in YN, YY potentials, and theoretical predictions are highly awaited
- Huge impact on EoS in high density, Neutron Star Core / Supernova
- "Generalization" of the nuclear force
- $\rightarrow$ what is universal, what is not universal in hadron-hadron interactions ? (e.g. origin of repulsive core)
- Three Baryon Potentials
- The Lattice study of Three Nucleon Force (TNF)


## Hyperon potentials (YN, YY)

- Generalized BB force




## Hyperon potentials (YN, YY)

- Equation of State at high density


- Hyper nuclei
${ }_{\wedge}^{6} \mathbf{H e}$
- $\longleftrightarrow$ Hyperon interactions ${ }^{4} \mathrm{He}+\Lambda+\Lambda$
- $\wedge \mathrm{N}$ attraction
- $\wedge \wedge$ weak attraction
- No deeply bound H-dibaryon

$$
7.25 \pm 0.1 \mathrm{MeV}
$$



## Baryon-baryon interaction

\# $\mathrm{SU}(3) \times \mathbf{S U}(2)$ spin $\Rightarrow \mathbf{S U}(6)$ classification


- The $\mathbf{S U}(6)$ symmetry predicts a strong spin-isospin dependence of the $\Sigma \mathrm{N}$ interaction.
- It also predicts state dependences of the spin-orbit interaction.

Oka-Shimizu-Yazaki, NPA464(1987)700
M. Oka, J-PARC Hadron Salon talk (06/17/2010)

## SU(3) study <br> BB potentials

$$
\begin{gathered}
a=0.12 \mathrm{fm}, \mathrm{~L}=2 \mathrm{fm}, \\
\mathrm{~m}(\mathrm{PS})=0.84,1.01 \mathrm{GeV}
\end{gathered}
$$



## Pauli principle at work !




8s,10: strong repulsive core



8a: weak repulsive core 1s: deep attractive pocket

## H-dibaryon in the SU(3) limit world ?



| Singlet potential with a certain value of E |  | Schroedinger eq. predicts a bound stat at $\mathrm{E}<-30 \mathrm{MeV}$ |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| ${ }^{200}$ | $V^{(1)}$ | E [MeV] | Eo [MeV] | $\sqrt{\left\langle r^{2}\right\rangle}$ | [fm] |
| 200 |  | $E=-30$ | -0.018 | 24.7 |  |
| $\sum_{8}$ |  | $E=-35$ | -0.72 | 4.1 |  |
| , |  | $E=-40$ | -2.49 | 2.3 |  |

finite size effect is very large on this volume. (consistent with previous results.) simulations on larger volume is in progress.

$$
V(r)=a_{1} e^{-a_{2} r^{2}}+a_{3}\left(1-e^{-a_{4} r^{2}}\right)^{2}\left(\frac{e^{-a_{0} r}}{r}\right)^{2}
$$

## Meson-Baryon system

- KN interaction (S-wave) in I=0, 1


$I=0(u \bar{s}+u d d, d \bar{s}+u u d) \quad I=1(u \bar{s}+u u d)$

Y.Ikeda et al., (HAL QCD Collab.) arXiv: 1002.2309


# YN, YY potentials beyond SU(3) limit 



$$
m_{\pi}=701 \mathrm{MeV}
$$

Repulsive core is surrounded by attractive well.
> Large spin dependence of repulsive core
$>$ Weak tensor force
$>$ Net interaction is attractive.


## Quark mass dependence of $\mathrm{N} /$ potential



With decreasing u and d quark masses,
$>$ Repulsive core is enhanced.
> Attractive well moves to outer region.
> Small quark mass dependence of tensor potential

$$
\begin{gathered}
\mathrm{Nf}=2+1 \text { clover, } \mathrm{L}=2.9 \mathrm{fm}, \\
\mathrm{a}=0.091 \mathrm{fm}(1 / \mathrm{a}=2.18 \mathrm{GeV}) \\
\mathrm{m}(\mathrm{pi})=410-700 \mathrm{MeV}
\end{gathered}
$$



Scatt. length

c.f. NPLQCD: repulsive scatt. length ? PRD81(2010)054505

## Coupled channel study

- BB system ( $\mathrm{S}=-2, \mathrm{I}=0$ )
- small energy difference $\rightarrow$ coupled channel

$$
\left\{\begin{array}{l}
\psi_{\alpha}^{\sum \sum}=\langle 0| \Sigma(\vec{r}) \Sigma(\overrightarrow{0})\left|E_{\alpha}\right\rangle \\
\psi_{\alpha}^{N \equiv}=\langle 0| N(\vec{r}) \equiv(\overrightarrow{0})\left|E_{\alpha}\right\rangle \\
\psi_{\alpha}^{\Lambda \Lambda}=\langle 0| \wedge(\vec{r}) \wedge(\overrightarrow{0})\left|E_{\alpha}\right\rangle
\end{array}\right.
$$

$$
\begin{array}{r}
\left|E_{1}\right\rangle,\left|E_{2}\right\rangle,\left|E_{3}\right\rangle \\
(\alpha=1,2,3) \\
\text { (variational method) }
\end{array}
$$

$m_{\Sigma \Sigma}=2380 \mathrm{MeV}$

120 MeV
$m_{N \equiv}=2260 \mathrm{MeV}$
30 MeV
$m_{\wedge \wedge}=2230 \mathrm{MeV}$

Coupled channel eq.









Irreducible flavor representations begin to mix with each other with decreasing $u, d$ quark mass.
11/25/2010
N-N INteraction meeting @ TIFR

## Three Nucleon Force (TNF)

## Importance of

## Three Nucleon Force (TNF)

- Precise few-body calc:
- e.g. benchmark calc of ${ }^{4} \mathrm{He}$ by 7 methods (NN only)

|  |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- |
| Method | $\langle T\rangle$ | $\langle V\rangle$ | $E_{b}$ | $\sqrt{\left\langle r^{2}\right\rangle}$ |
| FY | $102.39(5)$ | $-128.33(10)$ | $-25.94(5)$ | $1.485(3)$ |
| CRCGV | 102.30 | -128.20 | -25.90 | 1.482 |
| SVM | 102.35 | -128.27 | -25.92 | 1.486 |
| HH | 102.44 | -128.34 | $-25.90(1)$ | 1.483 |
| GFMC | $102.3(1.0)$ | $-128.25(1.0)$ | $-25.93(2)$ | $1.490(5)$ |
| NCSM | 103.35 | -129.45 | $-25.80(20)$ | 1.485 |
| EIHH | $100.8(9)$ | $-126.7(9)$ | $-25.944(10)$ | 1.486 |

$\rightarrow \mathbf{0 . 5 \%}$ prec. for B.E.
H.Kamada et al., PRC64(2001)044001

- NN force cannot reproduce B.E.

$$
\begin{aligned}
& \text { סB.E. }=0.5-1 \mathrm{MeV} \text { for }{ }^{3} \mathrm{H} \quad \text { missing } \\
& \delta B . E .=2-4 \quad \mathrm{MeV} \text { for }{ }^{4} \mathrm{He} \quad \text { men }
\end{aligned}
$$

Attractive TNF necessary




## Importance of Three Nucleon Force (TNF)

- Saturation density/energy of nuclear matter also requires TNF
- EOS of neutron star

Repulsive TNF also necessary

- Flavor universal TNF (repulsive) ?
A.Akmal et al., PRC58(1998)1804

Takatsuka et al., PTPS174(2008)80

## Importance of Three Nucleon Force (TNF)

- The effect on the nuclear chart t.otsuka et al., PRL105(2010)032501
- anomaly in drip line and magic numbers by TNF
- Ay puzzle in N-d, N-A scatt., etc. (TNF may worsen the situation)


## Importance of Three Nucleon Force (TNF)

- The effect on the nuclear chart
- anomaly in drip line and nontrivial magic number in neutron rich nuclei by TNF

T.Otsuka et al., PRL105(2010)032501
J.D.Holt et al., arXiv: 1009.5984


Nucleosynthesis by Supernova

## Importance of Three Nucleon Force (TNF)

- Precise few-body calc: NN force cannot reproduce B.E.

$\delta B . E .=0.5-1 \mathrm{MeV}$ for ${ }^{3} \mathrm{H}$<br>$\delta B . E=2-4 \quad \mathrm{MeV}$ for ${ }^{4} \mathrm{He}$

Attractive TNF necessary


- Saturation density/energy of nuclear matter also requires TNF
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A.Akmal et al., PRC58(1998)1804

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## Three Nucleon Force (TNF)

- It is natural to expect the existence of TNF
- It is very nontrivial to determine TNF from QCD
- $2 \pi \mathrm{E}-\mathrm{TNF}$ Fujita-Miyazawa, PTP17(1957)360
- Off-energy-shell $\pi \mathrm{N}$ scatt

| $\pi^{\pi}$ | $\square$ | $\pi \Delta^{\pi}$ | - $-\frac{\pi}{}$ |
| :---: | :---: | :---: | :---: |

- EFT expansion $\rightarrow$ TNF appears at NNLO order
- Phenomenological short-range repulsion is necessary
- $2 \pi \mathrm{E}-\mathrm{TNF}$ too attractive, often suppressed (artificially) by form factor
- NB: the combination of (2NF,3NF) $\rightarrow$ observables


U.v.Kolck, PRC49(1994)2932

Epelbaum, Prog.Part.Nucl.Phys.57(06)654

## How can we tackle TNF in Lattice QCD ?

- In the case of 2N system... $\psi(\vec{r})=\langle 0| N(\vec{x}+\vec{r} ; t) N(\vec{x} ; t)|2 N\rangle$
$\rightarrow \quad\left(E-H_{0}\right) \psi(\vec{r})$

$$
|2 N\rangle=\bar{N}_{s r c}(t=0) \bar{N}_{s r c}(t=0)|0\rangle
$$

$$
=\left[V_{c}(r)+S_{12} V_{T}(r)+\cdots\right] \psi(r)
$$



- Obtain TNF through

$$
\left(E-H_{0}^{r}-H_{0}^{\rho}\right) \psi(\vec{r}, \vec{\rho})=\left[\frac{\sum_{i<j} V_{i j}\left(\vec{r}_{i j}\right)}{\widehat{ }}+V_{T N F}(\vec{r}, \vec{\rho})\right] \psi(\vec{r}, \vec{\rho})
$$

- Difficulty(1): volume factor
- 2N: naïve $O\left(L^{6}\right)$ calc $\rightarrow O\left(L^{3} \log L^{3}\right)$
- 3N: naïve $O\left(L^{9}\right)$ calc $\rightarrow O\left(L^{6} \log L^{6}\right)$
- Difficulty(2): naïve calc of quark dof grows in factorial ( $\sim N_{u}$ ! $N_{d}$ !)
- 2N: O(L³) X N wick $\times$ color/spinor loops
- 3N: O(L6) $\times \mathrm{N}_{\text {wick }} \times$ color/spinor loops


## How can we tackle TNF in Lattice QCD ? (cont'd)

- We studied the effective 2 N potential in 3 N system $\left({ }^{3} \mathrm{H}\right)$
$\left(E-H_{0}^{r}\right) \phi(\vec{r})=\left[V_{12}(\vec{r})+\delta V_{e f f}(\vec{r})\right] \phi(\vec{r})$
$=\left[V_{12}(\vec{r}) \phi(\vec{r})+\int d \vec{\rho}\left(V_{13}(\vec{r}, \vec{\rho})+V_{23}(\vec{r}, \vec{\rho})+V_{T N F}(\vec{r}, \vec{\rho}) \psi(\vec{r}, \vec{\rho})\right]\right.$
- Relatively small calc cost (yet, still much expensive than 2N)
- Good precision achieved thanks to the sum over spectator particle

- Indirect access to TNF (due to off-diag 2N), and the effect of TNF is "smeared" by spacial average with triton wave function
- Calculation for fixed 3D-configuration of 3N system
- Direct access to TNF is possible!
- $\rightarrow$ We can explore the various features of TNF (spin/isospin/spacial, etc.)
- Much more expensive calc cost (O(10-100) factor) and yet worse $\mathrm{S} / \mathrm{N}$
- We study linear setup



## How can we tackle TNF in Lattice QCD ? (cont'd)

- We studied the effective 2 N potential in 3 N system ( 34 ) $\left(E-H_{0}^{\circ}\right) \phi(t)=V_{12}(r)+\delta V_{V / V}(r) \phi(r)$
- Relatively small caic cost (yet, still much expensive than 2 N )
- Good precision achieved thanks to the sum over spectator particie
a Indirect access to TNF (due to off-diag 2 N ), and the effect of TNF is "smeared" by spacial average with triton wave function
- Calculation for fixed 3D-configuration of 3N system
- Direct access to TNF is possible!
- $\rightarrow$ We can explore the various features of TNF (spin/isospin/spacial, etc.)
- Much more expensive calc cost ( $\mathbf{O}(\mathbf{1 0 - 1 0 0})$ factor) and yet worse $\mathrm{S} / \mathrm{N}$
- We study linear setup



## Features of Linear setup for ${ }^{3} \mathrm{H}$

- Simplified coupled channel analysis possible
- The vector to 3rd particle $\vec{\rho}=\overrightarrow{0}$
- $\Rightarrow \mathrm{L}^{(1,2) \text {-pair }}=\mathrm{L}^{\text {total }}=0$ or 2 only

- $\rightarrow$ Possible bases are only three, which can be labeled by 1S0, 3S1, 3D1 for (1,2)-pair
- However, in order to determine TNF in 3x3 coupled channel, we need information of parity-odd potential
- Although (1,2)-pair is L=even, $(3,1),(2,3)$-pair have $L=$ odd components
- Partial wave expansion with different J acobi setup is impossible, since we do not have full wave function (only linear setup)
- Parity-odd potential from lattice QCD (still) in progress
- $\rightarrow 3 \times 3$ channel, but unknown $\mathrm{V}_{\mathrm{C}}^{1,5=0,0}, \mathrm{~V}_{\mathrm{C}}^{1, \mathrm{~S}=1,1}, \mathrm{~V}_{\mathrm{T}}^{1,5=1,1}, \mathrm{TNF}(\mathrm{s})$


## Solution using

## "symmetric" wave function

- Rotate the basis $\left.\left.\quad\left|\psi_{1 S_{0}}\right\rangle,\left\langle\psi_{3_{S_{1}}}\right\rangle,\left|\psi_{3_{D_{1}}}\right\rangle —\left|\psi_{S}\right\rangle| | \psi_{M}\right\rangle\right\rangle\left|\psi_{3_{D_{1}}}\right\rangle$

$$
\left|\psi_{S}\right\rangle=1 / \sqrt{2}\left(-\left|\psi_{1_{S_{0}}}\right\rangle+\left|\psi_{3_{S_{1}}}\right\rangle\right) \quad\left|\psi_{M}\right\rangle=1 / \sqrt{2}\left(+\left|\psi_{1 S_{0}}\right\rangle+\left|\psi_{3_{S_{1}}}\right\rangle\right)
$$

- We can construct the wave function in which any 2 N pair is spin/isospin anti-symmetric


$$
\begin{array}{r}
\left|\psi_{S}\right\rangle=1 / \sqrt{6}\left[-\left(p_{\uparrow} n_{\uparrow}-n_{\uparrow} p_{\uparrow}\right) n_{\downarrow}\right. \\
-\left(n_{\uparrow} n_{\downarrow}-n_{\downarrow} n_{\uparrow}\right) p_{\uparrow}
\end{array}
$$

$$
+1 / 2\left(p_{\uparrow} n_{\downarrow}+n_{\uparrow} p_{\downarrow}-p_{\downarrow} n_{\uparrow}-n_{\downarrow} p_{\uparrow}\right) n_{\uparrow}
$$

$$
\left.+1 / 2\left(p_{\uparrow} n_{\downarrow}-n_{\uparrow} p_{\downarrow}+p_{\downarrow} n_{\uparrow}-n_{\downarrow} p_{\uparrow}\right) \quad n_{\uparrow}\right] \quad(\leftarrow I=0, S=1)
$$

- $\rightarrow$ L=even for any 2N pair automatically guaranteed


## Solution using

## "symmetric" wave function

- Rotate the basis
$\left.\left.\left(\left|\psi_{1 S_{0}}\right\rangle\right),\left(\psi_{3 S_{1}}\right\rangle\right)\left|\psi_{B_{D_{1}}}\right\rangle \rightarrow\left|\psi \psi_{S}\right\rangle,\left|\psi_{M}\right\rangle\right\rangle\left|\psi_{3_{D_{1}}}\right\rangle$
$\left|\psi_{S}\right\rangle=1 / \sqrt{2}\left(-\left|\psi_{1_{S_{0}}}\right\rangle+\left|\psi_{3_{S_{1}}}\right\rangle\right) \quad\left|\psi_{M}\right\rangle=1 / \sqrt{2}\left(+\left|\psi_{1_{S_{0}}}\right\rangle+\left|\psi_{3_{S_{1}}}\right\rangle\right)$
- We can construct the wave function in which any 2 N pair is spin/isospin anti-symmetric

- $\rightarrow$ L=even for any $2 N$ pair automatically guaranteed
- $3 \times 3$ coupled channel is reduced to
- one channel with only TNF unknown
( $L^{2}$-dep ignored)
- two channels with $\mathrm{V}_{\mathrm{C}}{ }^{1, S=0,0}, \mathrm{~V}_{\mathrm{C}}{ }^{1, S=1,1}, \mathrm{~V}_{\mathrm{T}}^{1, S=1,1}$,(TNF) unknown
- $\rightarrow$ Even without parity-odd V, we can determine one TNF
- This methodology works for any fixed 3D-conf other than linear 11/25/2010


## Repulsive TNF (TNR)

- We determine TNF assuming scalar/isoscalar
- Phenomenologically introduced to reproduce saturation density/energy of nuclear matter, etc.


Plot of TNR only:
there is cancellation from TNA


## Lattice calculation setup

- $\mathrm{Nf}=2$ dynamical clover fermion + RG improved gauge configs (CP-PACS)
- 598 configs $X 16$ measurements
- beta=1.95, ( $\left.a^{-1}=1.27 \mathrm{GeV}, a=0.156 f m\right)$
- $16^{3} \times 32$ lattice, $L=2.5 f m$
- Kappa(ud) $=0.13750$
- $\mathrm{M}(\pi)=1.13 \mathrm{GeV}$
- $M(N)=2.15 \mathrm{GeV}$
- $\mathrm{M}(\Delta)=2.31 \mathrm{GeV}$

CP-PACS Coll. S. Aoki et al., Phys. Rev. D65 (2002) 054505


BGL@KEK
[E: D67 (2003) 059901]

- Techniques
- Automatic Wick contraction code to handle 4 up- and 5 down-quarks
- Non-rela limit op is used to create 3 N state at source

$$
N^{s r c}=\epsilon_{a b c}\left(u_{a}^{T} C \gamma_{5} \frac{1+\gamma_{4}}{2} d_{b}\right) \frac{1+\gamma_{4}}{2} u_{c}
$$

$\rightarrow$ Factor of $2^{3}=8$ faster

## Results for wave functions



## Genuine Three Nucleon Force

Scalar/ I soscalar TNF



## I ndication of short-range repulsive TNF

## Note:

GS saturation may not be sufficient Constant energy shift not included
$\rightarrow O(10) M e V ~ s h i f t ~ p o s s i b l e ~$

$$
\mathrm{M}(\pi)=1.13 \mathrm{GeV}
$$

## Genuine Three Nucleon Force

Scalar/ I soscalar TNF



Comparison with improved Laplacian op.

】
Discretization error in Laplacian op. is small
$\begin{array}{ll}\nabla_{s t d}^{2} f(\vec{x})=\frac{1}{a^{2}} \sum_{i}\left[f\left(\vec{x}+a_{i}\right)+f\left(\vec{x}-a_{i}\right)-2 f(\vec{x})\right] & =\nabla^{2} f(\vec{x})+\mathcal{O}\left(a^{2}\right) \\ \nabla_{p r c t}^{2} f(\vec{x})=\frac{1}{-12 a^{2}} \sum_{i}\left[-\left(f\left(\vec{x}+2 a_{i}\right)+f\left(\vec{x}-2 a_{i}\right)\right)+16\left(f\left(\vec{x}+a_{i}\right)+f\left(\vec{x}-a_{i}\right)\right)-30 f(\vec{x})\right]=\nabla^{2} f(\vec{x})+\mathcal{O}\left(a^{4}\right)\end{array}$

## Summary/Outlook



- Potentials from Lattice QCD using NBS wave function
- Central and tensor potentials in parity-even channel
- Qualitative features of NN potentials are reproduced, Velocity expansion checked
- Significant step toward Nuclear Physics from QCD
- Lattice QCD can give useful predictions on unknown potentials
- YN, YY: Strangeness physics, hyperon matter in neutron star, SU(3) \& beyond SU(3)
- Meson-Baryon: N-K, N-ccbar (Kawanai-Sasaki), Q-Qbar: (I ida-I keda)
- The First calculation on Three Nucleon Force (TNF) from Lattice QCD
- 2 N subtraction is possible using only parity-even potentials
- Calculation of linear setup of $3 \mathrm{~N}\left({ }^{3} \mathrm{H}\right)$ system
- Indication of Repulsive TNF at short distance, further studies ongoing
- Various complementary approaches useful
- Operator-Product-Expansion (Aoki-Balog-Weisz)
- Lattice nuclei (Yamazaki-Kuramashi-Ukawa)
- Strong-coupling limit (de Forcrand-Fromm)
- Outlook
- Realistic potentials (and phase shifts) with physically light masses w/ large volume
- Parity-odd potentials, Higher derivative terms (LS-force and more) $\rightarrow$ More TNFs
- Understand the insight of nuclei: lattice nuclei vs. lattice potentials +ab initio calc.
- Resonances from potentials ? ( $\rho, \Delta, \mathrm{H}$-dibaryon, exotics...)
- TNF: other 3D-conf (triangle etc.) for spacial info, $\mathrm{I}=3 / 2, \mathrm{SU}(2)_{\mathrm{f}} \rightarrow \mathrm{SU}(3)_{\mathrm{f}}$ : Astro physics


