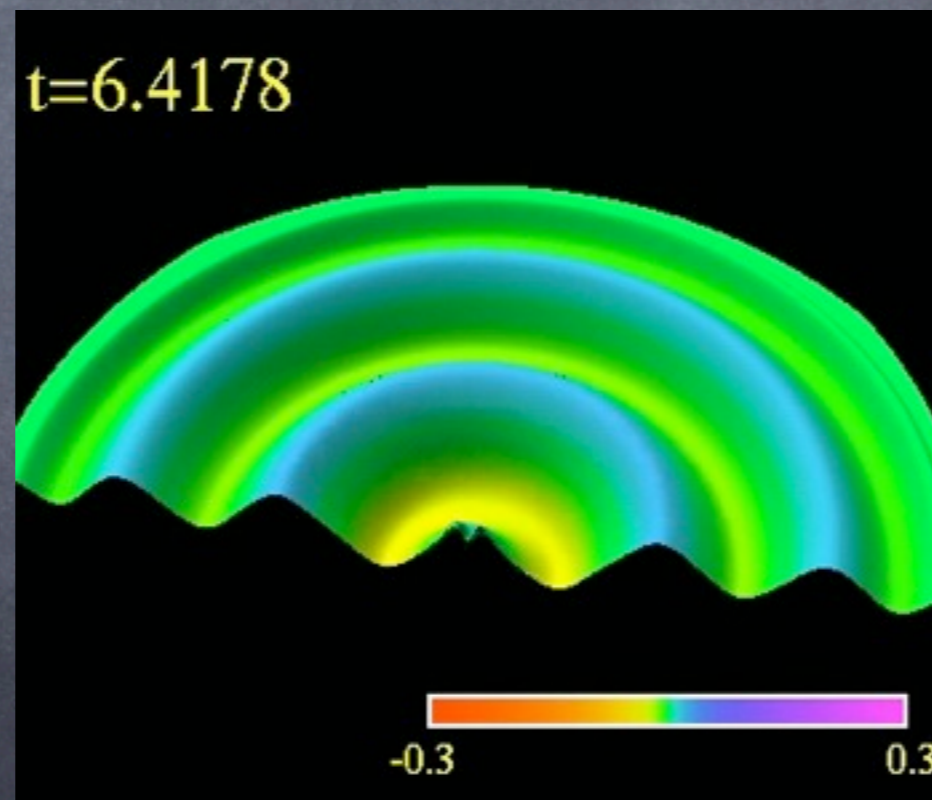


Introduction to Theory and Numerics of Partial Differential Equations VI: Wave equation and Einstein equations in spherical symmetry/1+1 dimensions



Sascha Husa

ICTS Summer School on Numerical Relativity

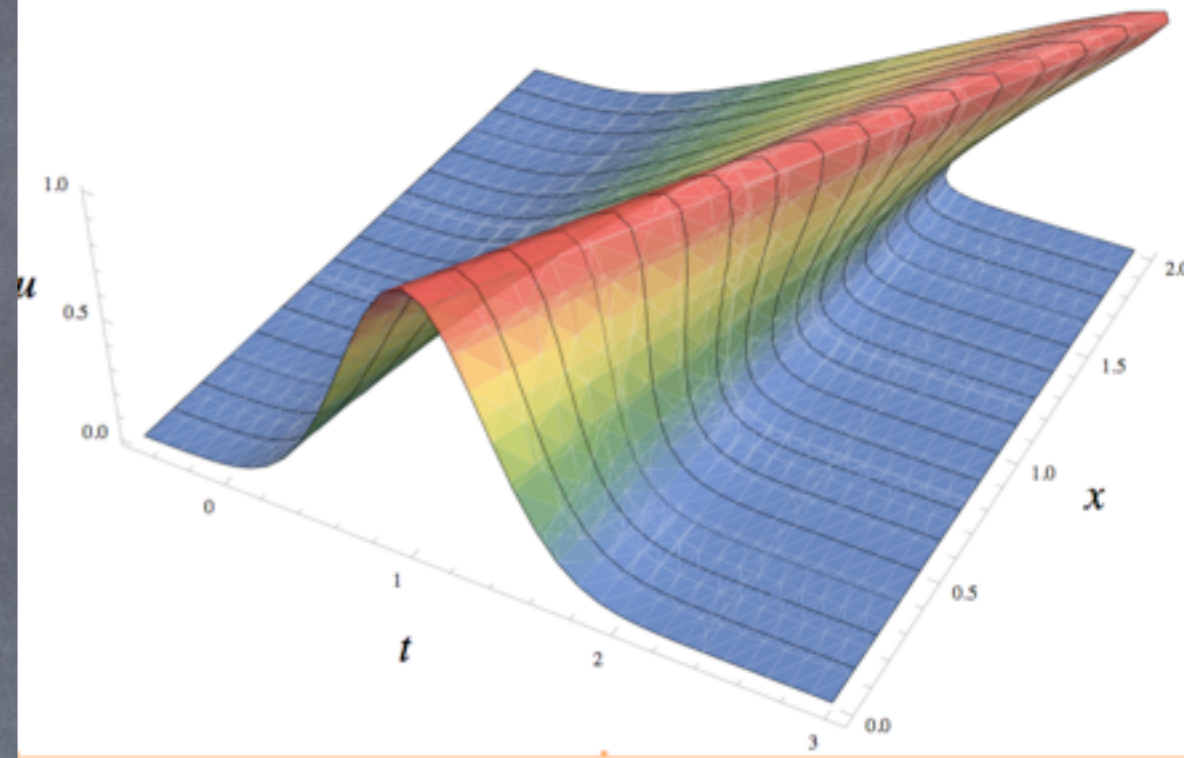
Bangalore, June 2013

Lab goals for today

- Look at energy, energy conservation and characteristic variables in your wave equation code.
- Upgrade your code to fourth order centered finite differences. What is the effect on accuracy/efficiency?
- Change boundary conditions to reflecting, incoming signal, and "outgoing".
- Implement the shifted wave equation.
- Start planning your coupled Einstein code (who can form a black hole tomorrow?)

smooth and distributional solutions

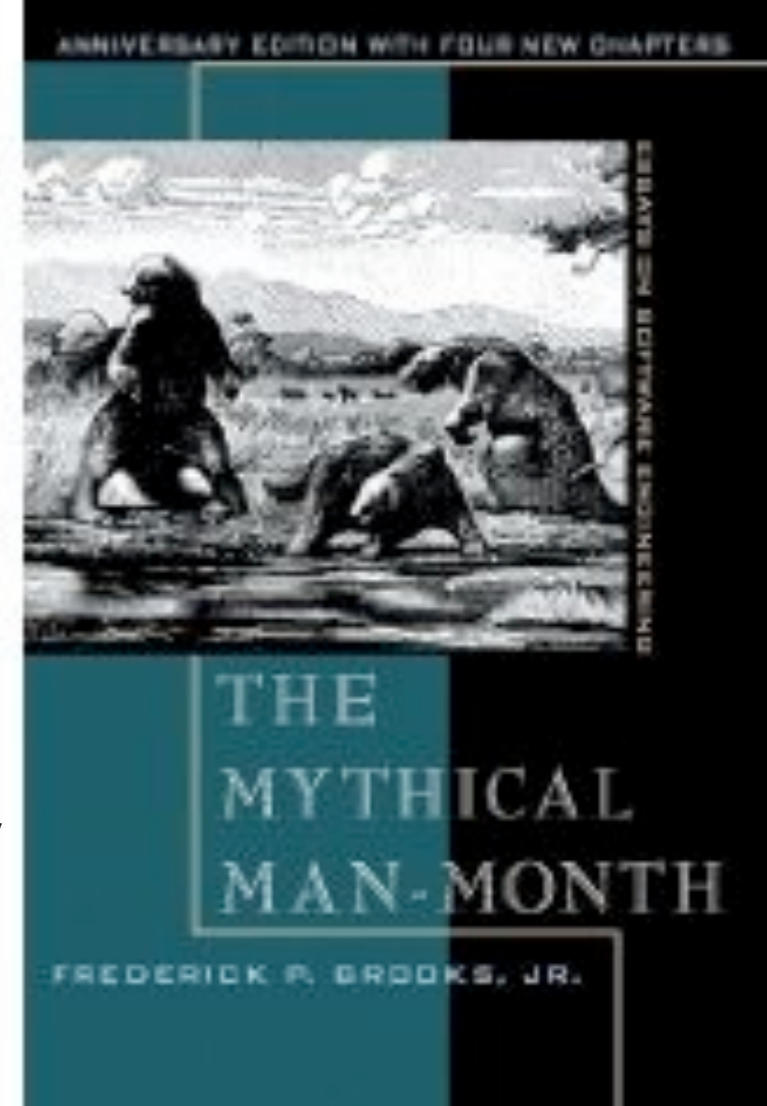
- Burger's equation: $u_t = u u_x$.
- Characteristic speeds depend on u , peak velocity overtakes rest of the wave after some time.



- More generally: characteristics can cross, typically signifies physical breakdown of underlying PDE, like in fluid dynamics.
- Unless a PDE is linearly degenerate (speeds independent of solution), shocks can form from smooth data in a finite time.
- Vacuum EE: can be written in linearly degenerate form, do not expect physical shocks, but shocks can form due to bad gauge conditions.
- Numerical methods for fluid dynamics are dominated by methods that deal with shocks – e.g. propagate shocks at correct physical speed.
- Solutions of vacuum GR are smooth except due to bad gauges or physical singularities, high order FD or spectral ideal!

Cost & error

- Want: approximate solution to a PDE with an error estimate at affordable cost.
 - affordable cost in running the code,
 - affordable cost in developing the code,
 - affordable cost modifying the code for new problems.
- We will only need a certain accuracy, but it may not be easy to understand what it should be, in particular if we want to study new phenomena.
- Don't waste too much time trying to interpret poor numerical data, use the time to produce better data. Don't become obsessed with machine accuracy (double precision $\sim 10^{-16}$, 64 bit). Usually this is more than enough, sometimes not!
- If you have a good idea and a working code, computer time will come to you.
- Often most of the human time is spent on debugging the code, and on trying to figure out “the physics” when numerical data are poor (inaccurate, noisy, ...)
- Defensive programming & good enough resolution & enough output!



Avoid problems early ...

Defensive programming is a form of **defensive design** intended to ensure the continuing function of a piece of **software** in spite of unforeseeable usage of said software. The idea can be viewed as reducing or eliminating the prospect of **Murphy's Law** having effect. Defensive programming techniques are used especially when a piece of software could be misused mischievously or inadvertently to catastrophic effect.

Defensive programming is an approach to improve software and source code, in terms of:

- General quality - Reducing the number of **software bugs** and problems.
 - Making the source code comprehensible - the source code should be readable and understandable so it is approved in a **code audit**.
 - Making the software behave in a predictable manner despite unexpected inputs or user actions.
-
- Keep it Simple!

Finite difference stencils in Fourier space

- Example: second order centered finite difference stencils.

$$\partial_x f \approx \frac{f_3 - f_1}{2h} \quad \partial_{xx} f \approx \frac{f_3 - 2f_2 + f_1}{h^2}$$

- apply them to a wave of frequency ω :

$$f(x) = e^{i\omega x}$$

- Apply finite difference operator to function:

$$\frac{e^{ih\omega} - e^{-ih\omega}}{2h}$$

- Simplify expression

$$\hat{D}_2 = \frac{i \sin(h\omega)}{h}$$

Numerical stability for first order hyperbolic systems

- P : linear constant coefficient differential operator

$$\partial_t u = P(\partial_x) u \quad \hat{P}(i\omega) : \quad \partial/\partial x_j \rightarrow i\omega_j = i \frac{\xi_j}{h} \quad (\text{i.e. } \hat{P} = i\omega_i A^i)$$

- WP is equivalent to $|e^{\hat{P}(i\omega)t}| \leq K e^{\alpha t} \rightarrow$ need \hat{P} diagonalizable

- discretize, e.g. 2nd order centered: $\partial_x \Rightarrow \frac{i}{h} \sin \xi$ (exercise!)

- n-th order Runge Kutta: $v^{n+1} = Q v^n = p(\Delta t P) v^n \quad p(x) = \sum_{l=0}^{l=n} \frac{x^l}{l!}$

- Fourier: $\hat{v}^{n+1}(\xi) = \hat{Q}(\xi) v^n(\xi) = p(\Delta t \hat{P}(\xi)) \hat{v}^n(\xi)$

- now we can solve: $\hat{v}^n(\xi) = \hat{Q}(\xi)^n v^0(\xi)$

- amplification matrix \hat{Q} diagonalizable if \hat{P} is!

- stability if eigenvalues satisfy: $|q_\mu| \leq 1, q_\mu = p(\Delta t p_\mu)$

- PDE does not explicitly depend on direction or dimension d

$$\lambda = \frac{\Delta t}{\Delta x} \leq \frac{\alpha_0}{\sigma(A) \sqrt{d}}, \quad \alpha_0 = 2(ICN), \sqrt{3}(RK3), \sqrt{8}(RK4)$$

nonlinear systems and dissipation

- Numerical schemes for quasi-linear hyperbolic PDEs: can use the same numerical methods, but need to dissipate high frequency modes to achieve numerical stability.
- Standard procedure: add Kreiss-Oliger dissipation for $2r-2$ accurate scheme, dissipation strength $\sigma > 0$:

$$\partial_t u \rightarrow \partial_t u + Qu, \quad Q_{2r} = \sigma \frac{(-\Delta x)^{2r-1}}{2^{2r}} (D_+)^r (D_-)^r$$

- does not degrade convergence order!
- Adding too much dissipation decreases time-step limit (makes equations behave more and more like heat equation).
- Artificial dissipation in fluid dynamics has traditionally been used to smear out shocks, superseded by "High resolution shock capturing" methods.

Second order in space systems: motivation

- Can we discuss well-posedness for second order in space systems like YADM and g-harmonic without first order reduction?
- Reduction to first order in time \rightarrow new evolution equations
- Reduction to first order in space \rightarrow new evolution & constraint equations.
 - enlarges solution space, new unphysical d.o.f. may give rise to instabilities (remember EM on curved background).
- General theory for WP of 2nd order in space only > 2004
 - How about accuracy of 1st vs. 2nd order in space?
- generalized wave equations: WP

example: mixed order wave equation

- Time domain: $h_{,t} = k, \quad k_{,t} = h_{,xx}$
- Frequency domain, $t \rightarrow \omega$: $\hat{h}_{,t} = \hat{k}, \quad \hat{k}_{,t} = -\omega^2 \hat{h}$
- Introduce new variable λ as the square root of $h_{,xx}$:

$$\hat{\lambda} := i\omega \hat{h} \quad \Rightarrow \quad \hat{\lambda}_{,t} = i\omega \hat{k}, \quad \hat{k}_{,t} = i\omega \hat{\lambda} \quad \hat{h}_{,t} = \hat{k}$$

$$\partial_t \begin{pmatrix} h \\ k \\ \lambda \end{pmatrix} = A \begin{pmatrix} h \\ k \\ \lambda \end{pmatrix}, \quad A = \begin{pmatrix} 0 & 0 & 0 \\ 0 & -i\omega & 0 \\ 0 & 0 & i\omega \end{pmatrix}$$

- Characteristic speeds are $-1, 1, 0$; problem is symmetric hyperbolic and well posed in the norm (L^2 does not always work!):

$$\|u\|^2 = \int (|h|^2 + |k|^2 + |\partial_x h|^2) dx$$

- In the Fourier domain this system could be treated in analogy with first order in space systems, using a "pseudo-differential reduction" - but variables play different roles depending on how often they are differentiated.
- In the discrete case, we will have to choose an appropriate discretization for the derivative in the norm!

second order in space hyperbolic systems

- normal form: P takes second derivatives of u , but not v .

$$\partial_t \begin{pmatrix} u \\ v \end{pmatrix} = P \begin{pmatrix} u \\ v \end{pmatrix}, \quad P = \begin{pmatrix} A^i \partial_i + B & C \\ D^{ij} \partial_i \partial_j + E^i \partial_i + F & G^i \partial_i + J \end{pmatrix}$$

- Second order principal symbol $\hat{P} = \begin{pmatrix} i\omega A^n & C \\ -\omega^2 D^{nn} & i\omega G^n \end{pmatrix}$

- Analyze WP & numerical stability by pseudo-differential reduction (first order reduction in Fourier space).

- WP reduces to diagonalizability of $\hat{P}_{\text{reduced}} = i\omega \begin{pmatrix} A^n & C \\ D^{nn} & G^n \end{pmatrix}$

- Discrete stability is **not** implied by WP + centered FD + small Δt

- $\partial_{xx} = \partial_x \partial_x$ does not carry over from continuum, e.g.

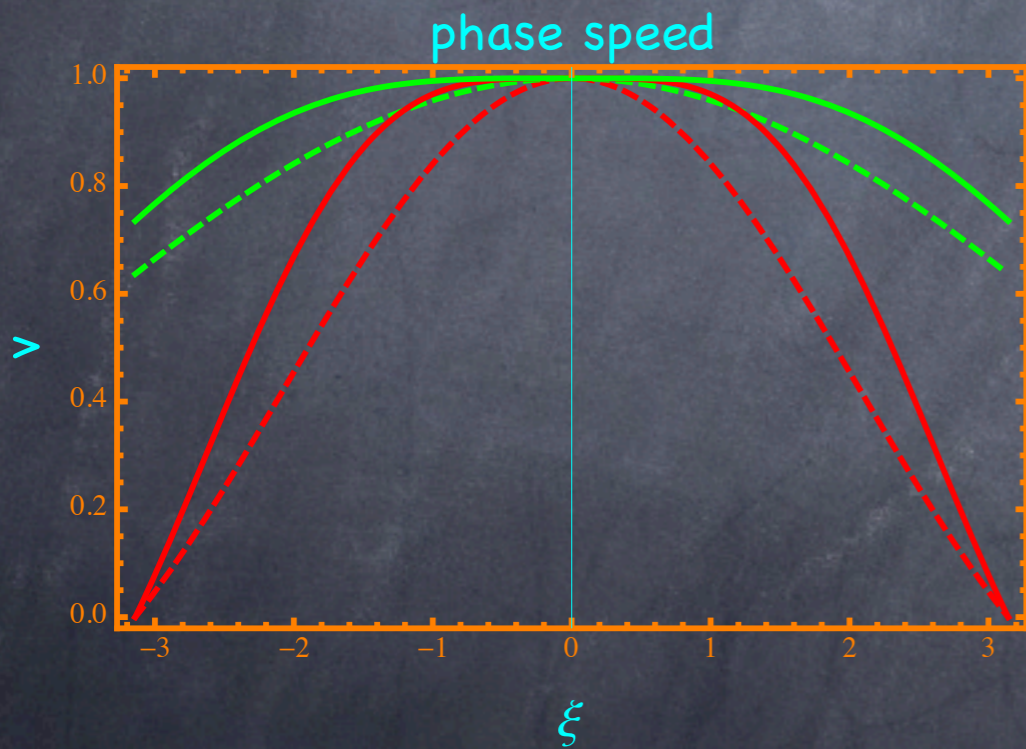
$$\hat{D}^{(2)} = -\frac{4}{\Delta x^2} \sin^2 \frac{\xi}{2} \neq \left(\frac{i}{\Delta x} \sin \xi \right)^2$$

- discrete norm: $\|u\|_h^2 + \|v\|_h^2 + \sum_{i=1}^d \|D_{+i} u\|_h^2, \quad D_{+} v_j = \frac{v_{j+1} - v_j}{\Delta x}$

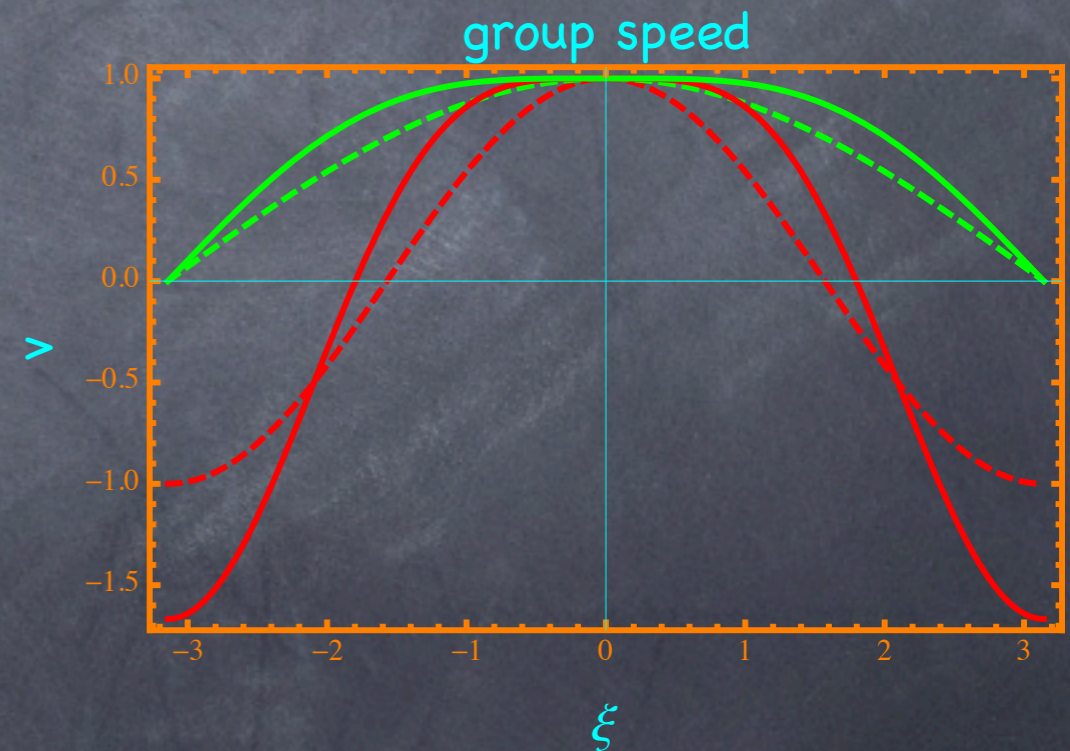
comparison 1st vs 2nd order in space

- $\lambda(\xi)$ eigenval. of $\hat{P}(\xi)$
- phase velocity $v_p = i \frac{\lambda}{\omega}$
- group vel. $v_g = i \frac{d\lambda}{d\omega}$

	2nd order accurate	
	advective	wave
v_p	$\frac{\sin \xi}{\xi} \approx 1 - \frac{\xi^2}{6} + O(\xi^4)$	$\frac{2}{\xi} \sin \frac{\xi}{2} \approx 1 - \frac{\xi^2}{24} + O(\xi^4)$
v_g	$\cos \xi \approx 1 - \frac{\xi^2}{2} + O(\xi^4)$	$\cos \frac{\xi}{2} \approx 1 - \frac{\xi^2}{8} + O(\xi^4)$
C.I.	α_0	$\alpha_0/2$
u.m.	$0, \pi$	0
f.u.m.	$\pm \frac{\pi}{2} \approx \pm 1.571$	π



advective eq.
 wave eq.
 - - 2nd order
 --- 4th order



- modes with speeds of the wrong sign will come out of BHs!
- second order in space systems have high frequency damping built in!

Some simple incarnations of the scalar wave equation

- Scalar WEQ defined with metric g_{ab} , may consider fixed metric, or couple scalar field to Einstein equations:

$$g^{ab}\nabla_a\nabla_b\phi = 0 \quad G_{ab}[g] = 8\pi GT_{ab}[\phi]$$

- e.g. WEQ on Minkowski space. 1+1 dimensional problems are obtained by considering plane waves

$$g^{\mu\nu} = \eta_{\mu\nu}, \quad \phi(x, t) \rightarrow \phi_{tt} = \phi_{xx}$$

- or spherically symmetric waves

$$g^{\mu\nu} = \eta_{\mu\nu}, \quad \phi(r, t) \rightarrow \phi_{tt} = \phi_{rr} + \frac{2}{r}\phi_{rr}$$

- Scaling of variables can do miracles:

$$\tilde{\phi}(r, t) := r\phi(r, t) \rightarrow \tilde{\phi}_{tt} = \tilde{\phi}_{rr}$$

Scalar field energy in flat space

- Energy density ρ gives rise to a conserved energy E :

$$E = \int_{R^3} \rho d^3x \quad \rho = \left(\frac{\partial \phi}{\partial t} \right)^2 + |\nabla \phi|^2$$

- For plane waves we get

$$\rho = \left(\frac{\partial \phi}{\partial t} \right)^2 + \left(\frac{\partial \phi}{\partial x} \right)^2$$

- Because of energy conservation, for plane waves the field strength can't decay.

- In spherical symmetry we expect decay with $1/r$

$$\phi(r, t) := \frac{\tilde{\phi}(r, t)}{r}$$

Boundary conditions

- Can consider 3 distinct cases:
 - finite grid without boundaries, use periodic boundary conditions = identify end points, DONE
 - finite grid with boundaries, need to impose boundary conditions (reflecting, incoming signal, outgoing=no incoming signal)
 - infinite grid. need to "pull in" infinity with a coordinate transformation, will lead to singular equations → investigate tomorrow

Spice up the wave equation with moving coordinates

- Restrict to plane waves in 1 space dimension:

$$ds^2 = -d\tilde{t}^2 + d\tilde{x}^2$$

- redefine x coordinate using shift (vector)

$$dt = d\tilde{t}, \quad dx = d\tilde{x} - \beta d\tilde{t}$$

- The metric becomes

$$g_{\mu\nu} = \begin{pmatrix} (-1 + \beta^2) & \beta \\ \beta & 1 \end{pmatrix} \quad g^{\mu\nu} = \begin{pmatrix} -1 & \beta \\ \beta & (1 - \beta^2) \end{pmatrix}$$

- Rewrite the WEQ using e.g.

$$\square\phi = \frac{1}{\sqrt{-g}} \partial_\mu [\sqrt{-g} g^{\mu\nu} \partial_\nu \phi] = 0$$

Shifted wave equation

$$\begin{aligned}\square\phi &= \frac{1}{\sqrt{-g}}\partial_\mu[\sqrt{-g}g^{\mu\nu}\partial_\nu\phi] \\ &= \partial_t[g^{t\nu}\partial_\nu\phi] + \partial_x[g^{x\nu}\partial_\nu\phi] \\ &= \partial_t[g^{tt}\partial_t\phi + g^{tx}\partial_x\phi] + \partial_x[g^{xt}\partial_t\phi + g^{xx}\partial_x\phi] \\ &= \partial_t[-\partial_t\phi + \beta\partial_x\phi] + \partial_x[\beta\partial_t\phi + (1 - \beta^2)\partial_x\phi] \\ &= 0\end{aligned}$$

- Suggests definition of new variables:

$$\psi := \partial_x\phi \quad \pi := \partial_t\phi - \beta\partial_x\phi$$

- Evolution equations: $\partial_t\phi = \pi + \beta\psi$

$$\partial_t\psi = \partial_x(\pi + \beta\psi) \quad \partial_t\pi = \partial_x(\psi + \beta\pi)$$

characteristic variables

- matrix formulation:

$$\mathbf{u} = (\pi, \psi)^T$$

$$\partial_t \mathbf{u} + \mathbf{A} \partial_x \mathbf{u} = \mathbf{u} \partial_x \beta \quad \mathbf{A} = - \begin{pmatrix} \beta & 1 \\ 1 & \beta \end{pmatrix}$$

- \mathbf{A} is diagonalizable with eigenvalues = characteristic speeds $\lambda_1 = -\beta + \alpha$, $\lambda_2 = -\beta - \alpha$, and eigenvectors

$$v_1 = (1, -1)^T, \quad v_2 = (1, 1)^T$$

- characteristic variables, propagating with characteristic speeds:

$$u_1 = u_R = \frac{1}{2} (\pi - \partial_x \phi) \quad u_2 = u_L = \frac{1}{2} (\pi + \partial_x \phi)$$

- Plot the characteristic variables in your code!
Observe that these quantities propagate as expected.

Boundary conditions

- Putting boundary conditions on outgoing characteristic fields is not logically consistent - initial boundary value problem will not be well-posed.
- Can only put boundary conditions on incoming characteristic fields!

- Examples:

- reflecting boundary conditions

$$\phi = \partial_t \phi = \pi = \partial_t \pi = \partial_x \psi = 0$$

- outgoing boundary conditions: incoming signal set to zero, e.g. at left boundary:

$$u_R = 0 = \pi - \psi \quad \Rightarrow \quad \pi = \psi$$

Scalar field coupled to gravity

- Simple form of the metric in spherical symmetry with zero shift

$$ds^2 = -\alpha(r, t)^2 dt^2 + a(r, t)^2 dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\varphi^2)$$

- Definitions:

$$\psi = \partial_r \phi, \quad \pi = \frac{a}{\alpha} \partial_t \phi$$

- Einstein equations:

$$\frac{\partial_r \alpha}{\alpha} = \frac{\partial_r a}{a} \frac{a^2 - 1}{r} \quad \frac{\partial_r a}{a} = \frac{1 - a^2}{2r} + \frac{r}{4} (\psi^2 + \pi^2) \quad \partial_t a = \frac{1}{2} r a \alpha \phi \pi$$

- Scalar field equations:

$$\partial_t \pi = \frac{1}{r^2} \left(\frac{r^2 \alpha \psi}{a} \right) \quad \partial_t \psi = \partial_r \left(\frac{\alpha \pi}{a} \right)$$