

ICTS Summer School on Numerical Relativity 2013 – Mathematical formulation

Tutorial # 3

3.1 Consider a two-dimensional sphere embedded in a flat, three-dimensional, as in the example we discussed in class. Show that the geometries of the two spaces satisfy Ricci's equation.

3.2 Show that the Lie derivative of the projection operator γ^a_b along αn^a vanishes,

$$\mathcal{L}_{\alpha n} \gamma^a_b = 0, \quad (1)$$

and that this is not true in general for the Lie derivative along n^a alone. This result implies that spatial tensors that are Lie dragged along αn^a remain spatial.

3.3 Show that the spatial Christoffel symbols

$$\Gamma^a_{bc} = \frac{1}{2} \gamma^{ad} (\partial_b \gamma_{dc} + \partial_c \gamma_{db} - \partial_d \gamma_{bc}) \quad (2)$$

are spatial projections of the spacetime Christoffel symbols,

$$\Gamma^a_{bc} = \gamma^a_d \gamma^e_b \gamma^f_c {}^{(4)}\Gamma^d_{ef}. \quad (3)$$

3.4 Show that the determinant $g = \det(g_{ab})$ of the spacetime metric g_{ab} can be written as

$$\sqrt{-g} = \alpha \sqrt{\gamma}, \quad (4)$$

where $\gamma = \det(\gamma_{ij})$ is the determinant of the spatial metric γ_{ij} .

Hint: Recall that for any square matrix A_{ij} the inverse can be computed from $(A^{-1})_{ij} = \text{cofactor of } A_{jk} / \det(A_{ij})$.

3.5 Show that the determinant of the spatial metric γ satisfies

$$\partial_t \ln \gamma^{1/2} = -\alpha K + D_i \beta^i, \quad (5)$$

while the mean curvature $K \equiv \gamma^{ij} K_{ij}$ satisfies

$$\partial_t K = -D^2 \alpha + \alpha (K_{ij} K^{ij} + 4\pi(\rho + S)) + \beta^i D_i K. \quad (6)$$

3.6 (a) Revisit the Painlevé-Gullstrand metric of Problem 1.4 and identify the lapse α , the shift β^i and the spatial metric γ_{ij} .

(b) Compute the extrinsic curvature K_{ij} and the Ricci tensor R_{ij} .

(c) (Optional) Verify that this solution satisfies the constraint and evolution equations with $\partial_t K_{ij} = 0$.